



June 12, 2025

# SBND Detector Commissioning and Early Performance

**Lauren Yates (Fermilab) on behalf of the SBND Collaboration**

International Workshop on Weak Interactions and Neutrinos (WIN 2025) — University of Sussex, Brighton, UK

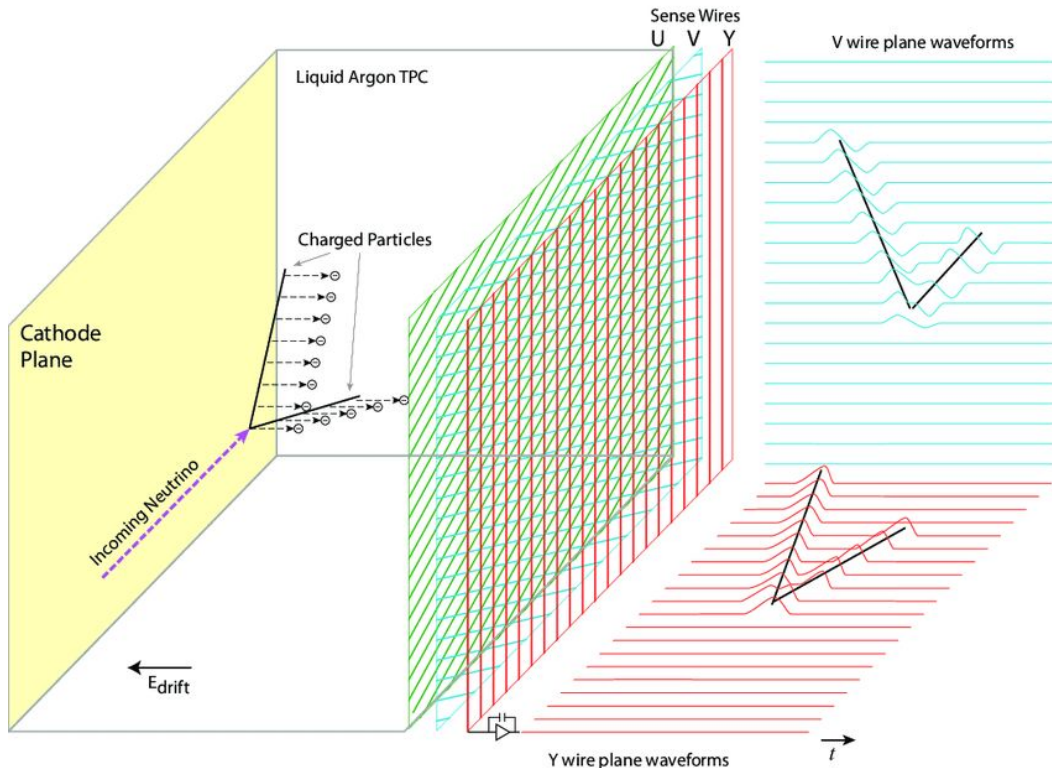


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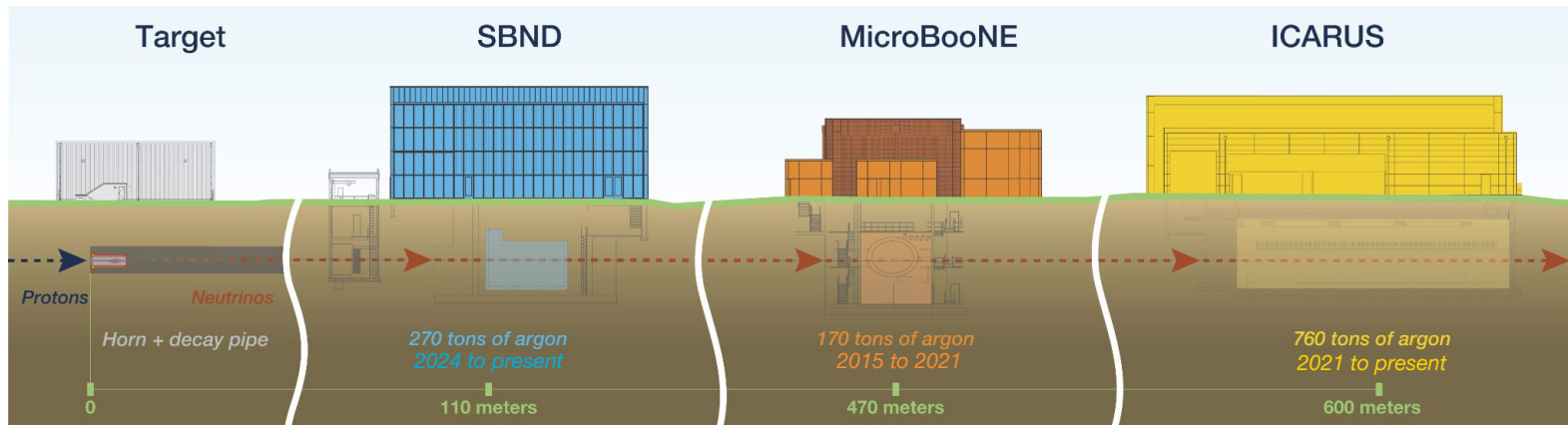


# Detecting Neutrino Interactions with LArTPCs



- LArTPC detectors are highly-capable, fully-active tracking calorimeters with mm-scale spatial resolution
- Enables disentangling complex final states with low thresholds and excellent particle identification
- Precise timing information also available via scintillation light

# Short Baseline Neutrino Program at Fermilab



- Consists of liquid argon time projection chambers (LArTPCs) located at different baselines in Fermilab's Booster Neutrino Beam (BNB), aiming to conclusively address short-baseline neutrino oscillations at the eV-scale
  - SBND has a critical role in enabling this as the program's near detector
- SBND also has a rich single-detector physics program including neutrino–argon cross section measurements and new and rare physics searches
- We have been stably collecting physics-quality neutrino data since December 2024

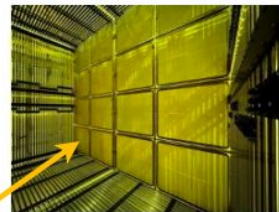
# The SBND Detector: TPC

## LArTPC

Active mass is 112 t  
Active vol. is  $4 \times 4 \times 5$   
 $\text{m}^3$



**Cold Electronics (in LAr)**  
pre-amplify and digitize  
TPC wire signals

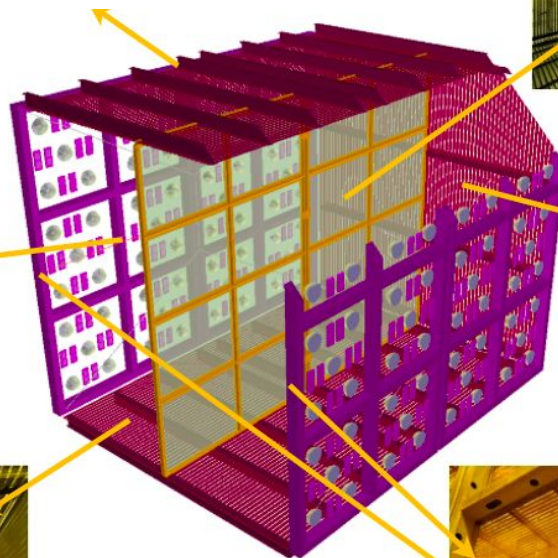


**Cathode Plane at  $-100\text{kV}$**   
divides the detector into  
two drift volumes

Drift distance is 2m,  
max. drift time is  $\sim 1.28\text{ms}$



**Field Cage** wraps around  
the two TPCs to step down  
the voltage and ensure a uniform  
electric field of  $500 \text{ V/cm}$

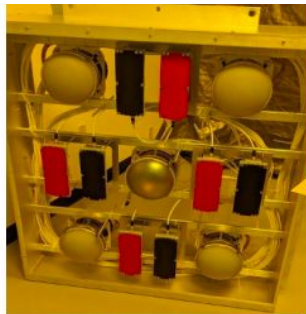


**Anode Plane** on either side,  
each with three wire planes  
with 3 mm wire spacing and  
different orientation per plan

Total of 11,260 wires

# The SBND Detector: PDS, CRT, & Trigger

## Photon Detection System



**24 PDS Boxes**

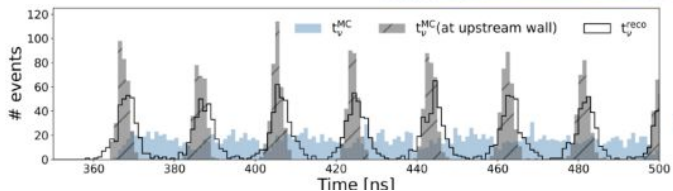
Behind the anode wire planes

5 × 24 = **120 8" PMTs**  
80% TPB-coated,  
20% uncoated

8 × 24 = **192 X-ARAPUCAs**  
Half with wavelength shifting

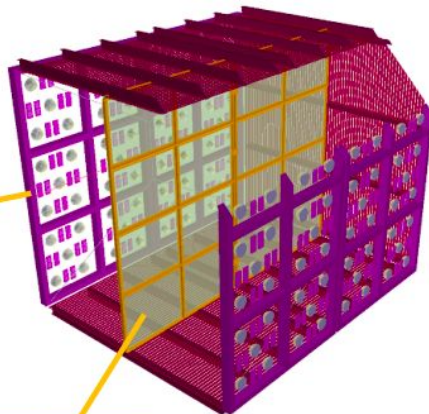


**Cathode Plane** with  
TPB-coated reflective foils  
mounted behind mesh panels



Sophisticated PDS  
reconstruction techniques  
developed on simulation  
demonstrate ns-scale timing

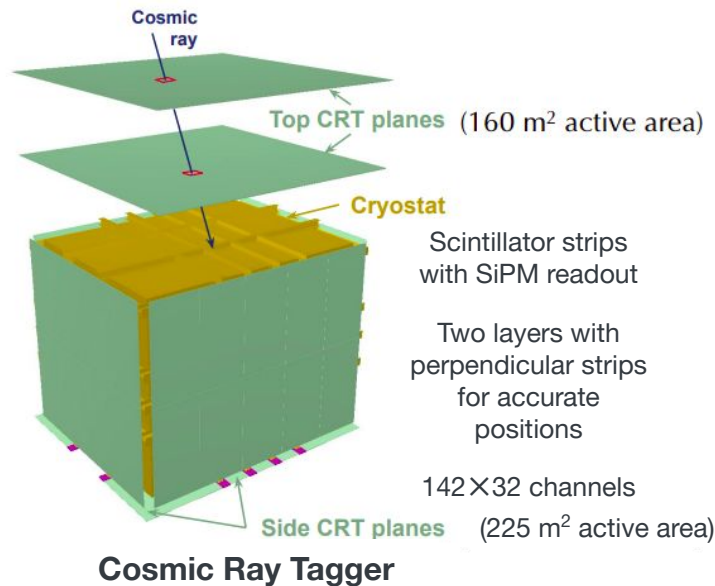
[EPJ C 84, 1046 \(2024\)](#)



Hardware trigger  
system capable of  
incorporating inputs  
from PMTs and CRT,  
plus beam signals



**Trigger System**



# Detector Commissioning

Detector commissioning was the process of activating all subsystems, and verifying their operational parameters were set to be sufficient for SBND's physics running

I will give brief descriptions of a few selected commissioning topics:

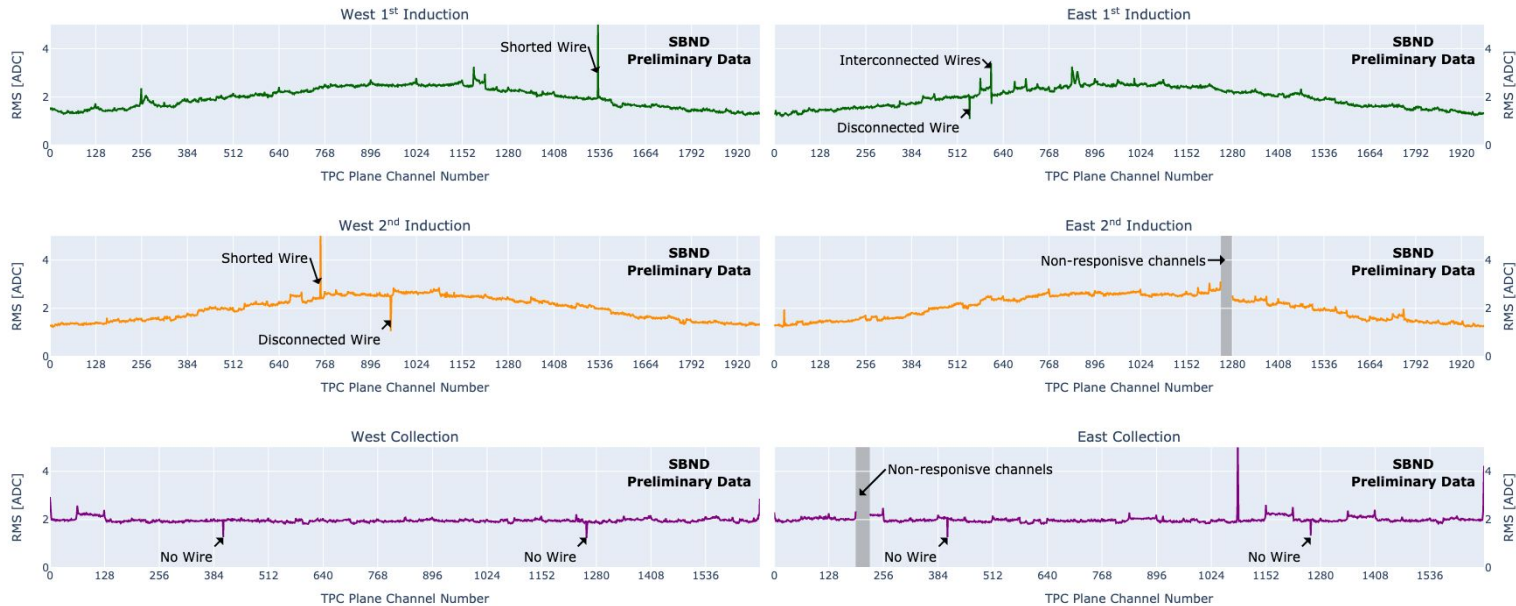
- TPC electronics noise
- PMT gain equalization
- CRT system activation and timing
- DAQ, trigger, and timing systems verification

There are many other aspects, and you're welcome to ask me questions if there's something else you'd like to hear about!

# TPC Electronics Noise

SBND's TPC electronics noise is close to the intrinsic floor, resulting in excellent signal-to-noise ratio on all planes — comparable to others after their noise filtering

TPC Noise per Channel



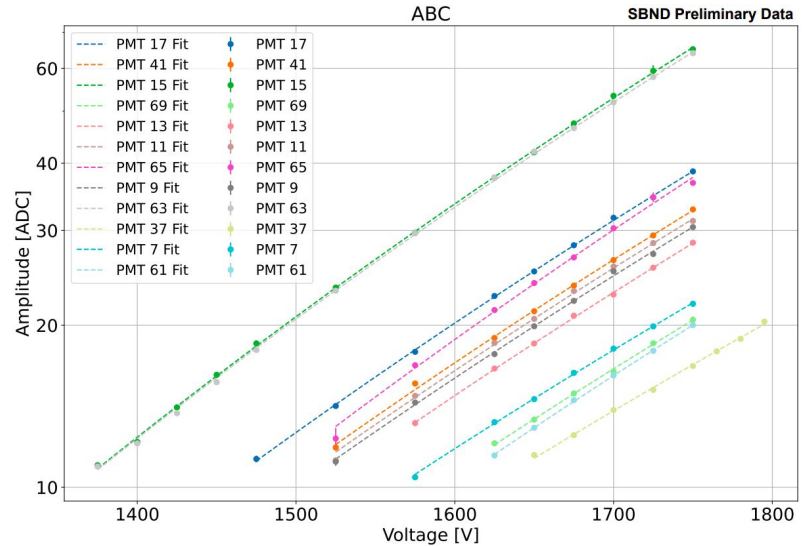
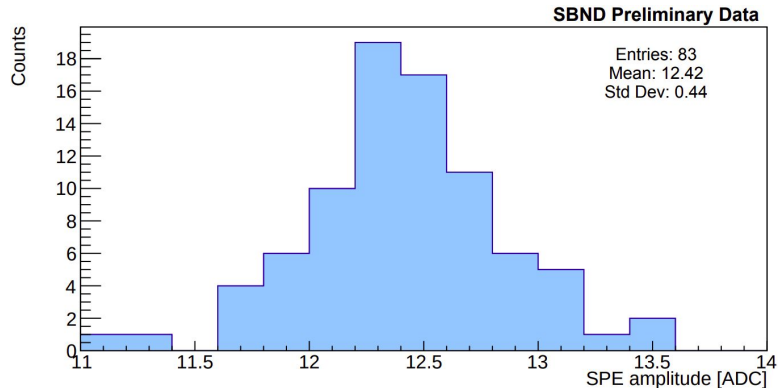
# PMT Gain Equalization

PMT gain equalization is an essential prerequisite to a light-based trigger

Performed a voltage scan and then set each channel to the bias voltage that provides a gain of about 12.5 ADC/PE

Balancing gain with PMT longevity

Resulting distribution of PMT gains for is shown below, equalized to ~5% level



# CRT Commissioning

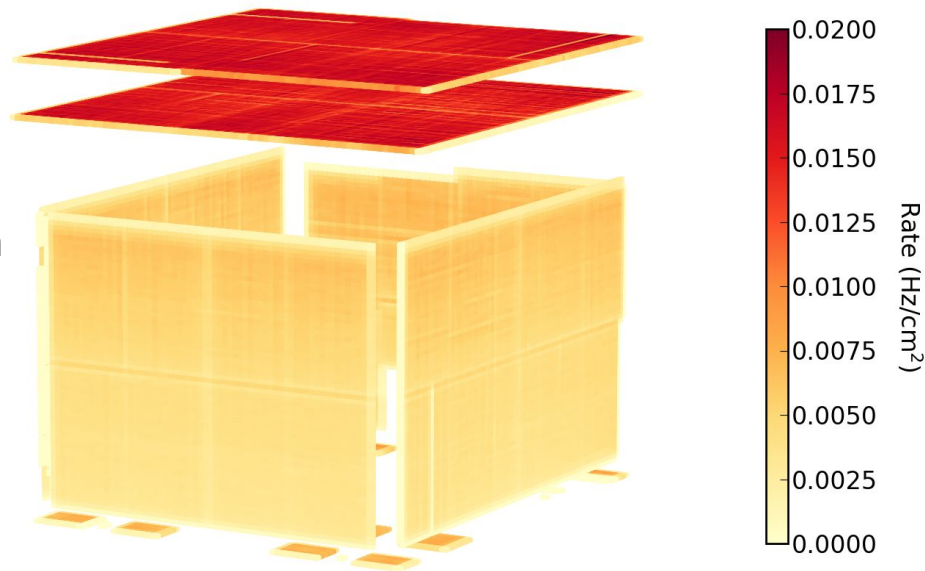
Rejecting cosmic backgrounds is important for many analyses

Tagging cosmics can be useful for calibration analyses, as the CRT can provide a known time and position for through-going muons

Can also measure the cosmic ray rate around the detector



SBND Preliminary  
CRT Off-Beam Data

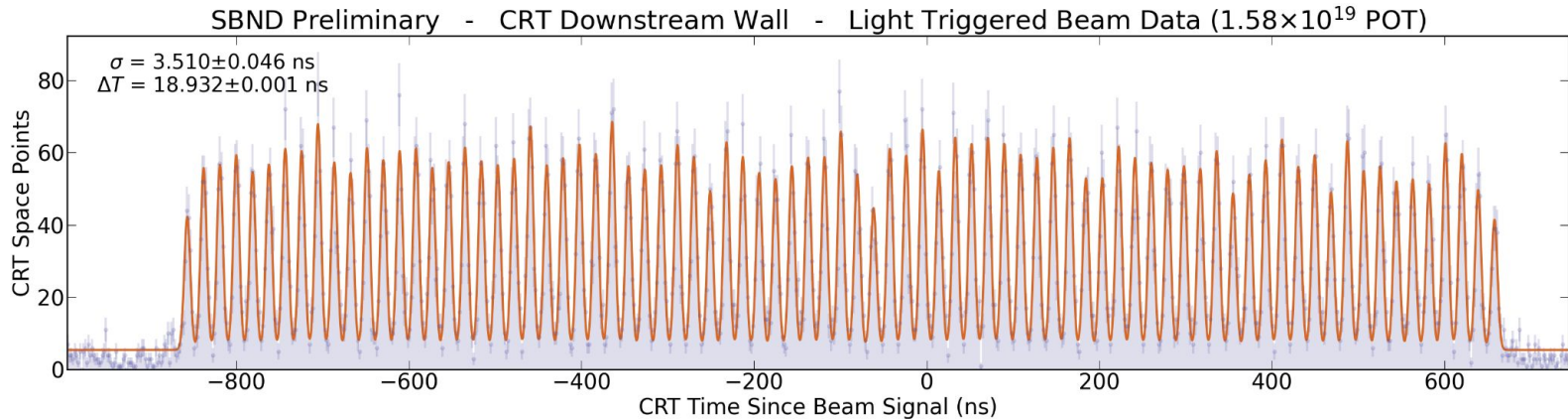


# CRT Timing

CRT system is designed to have timing resolution of  $O(1 \text{ ns})$ , and we have already been able to demonstrate that in SBND data

We can observe this bucket structure in the CRT activity, and buckets have an measured width of about 3.5ns

Precise timing can help distinguish entering cosmics vs. exiting neutrinos

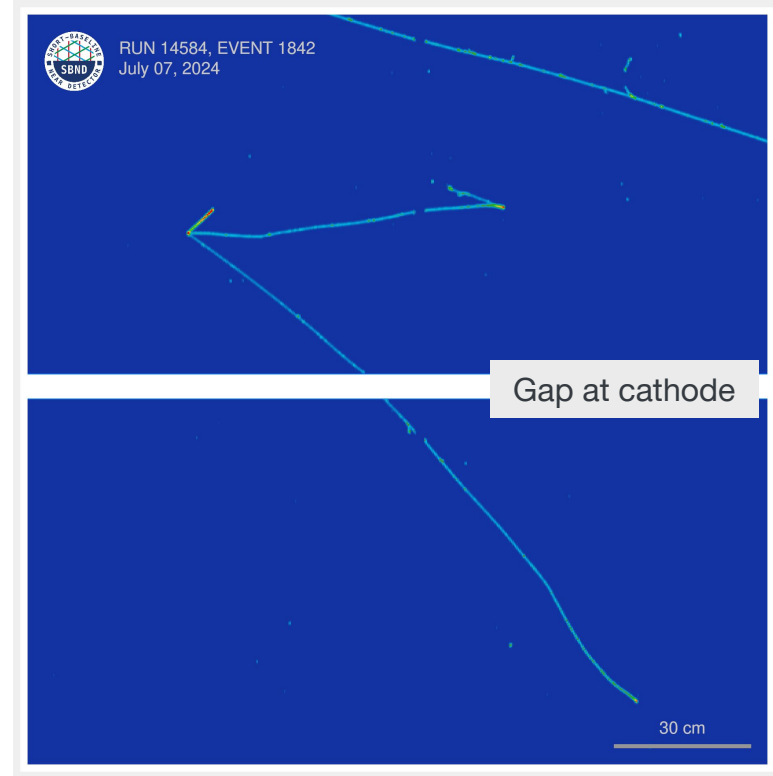
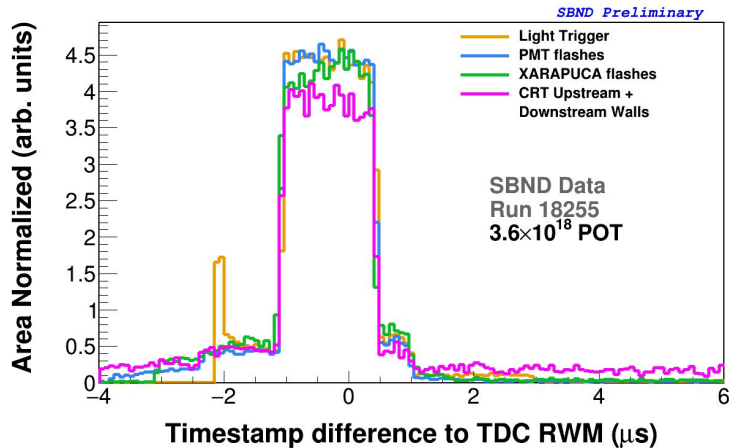




# Data Acquisition & System Synchronization

For the PDS, CRT, and trigger, system synchronization can be verified by seeing the beam activity appears at the same time

For the TPC, we can look at event displays, and check that we see neutrino topologies consistent with in-time interactions

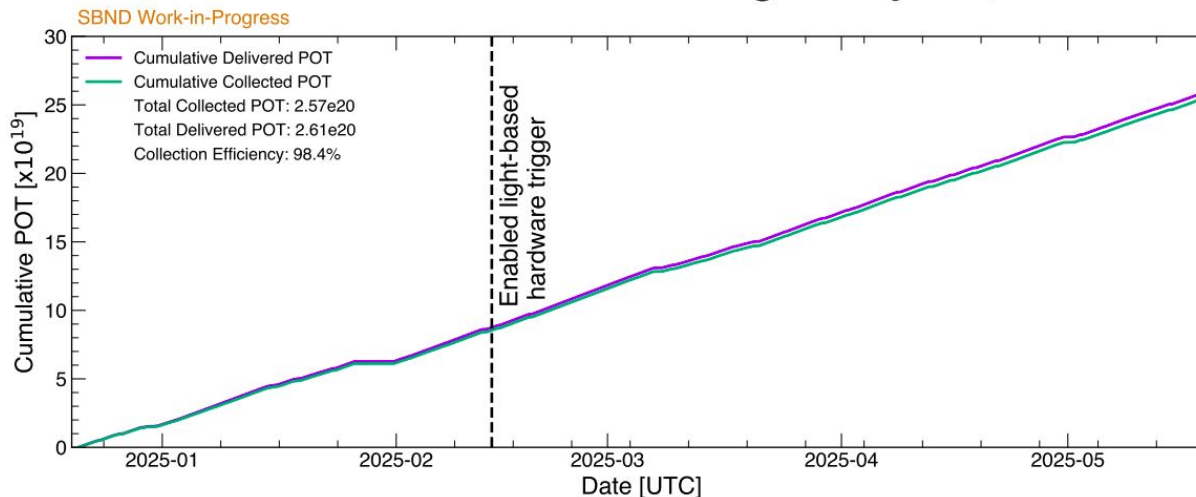


# SBND Operations to Date

SBND has been smoothly collecting physics-quality neutrino data since December, and implemented our light-based hardware trigger in February (>98% uptime)

To date, have collected more than  $2.5e20$  POT, which corresponds to an expectation of more than 2 million neutrino interactions in the detector

## SBND Cumulative POT through May 20, 2025



# Summary & Outlook

SBND detector commissioning is complete and we are stably collecting physics-quality data, having already accumulated world-leading statistics for neutrino–argon interactions

Our early detector performance is outstanding, with all systems consistently meeting or exceeding their design requirements

We are making great progress in understanding and calibrating our detector, which is an important step towards high-precision physics measurements

See [Lynn Tung](#)'s talk

We have an exciting physics program that will keep us busy for many years to come!  
Stay tuned for future results to come from SBND and the SBN Program

See [Josie Paton](#), [Amy Filkins](#), and [Vincent Basque](#)'s talks

 Thank you!

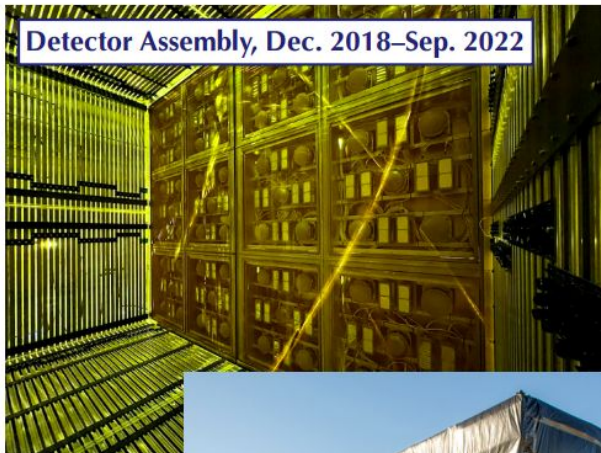


A grayscale photograph of the interior of a particle accelerator tunnel, showing the complex machinery and the circular structure of the tunnel. The image is used as a background for the slide.

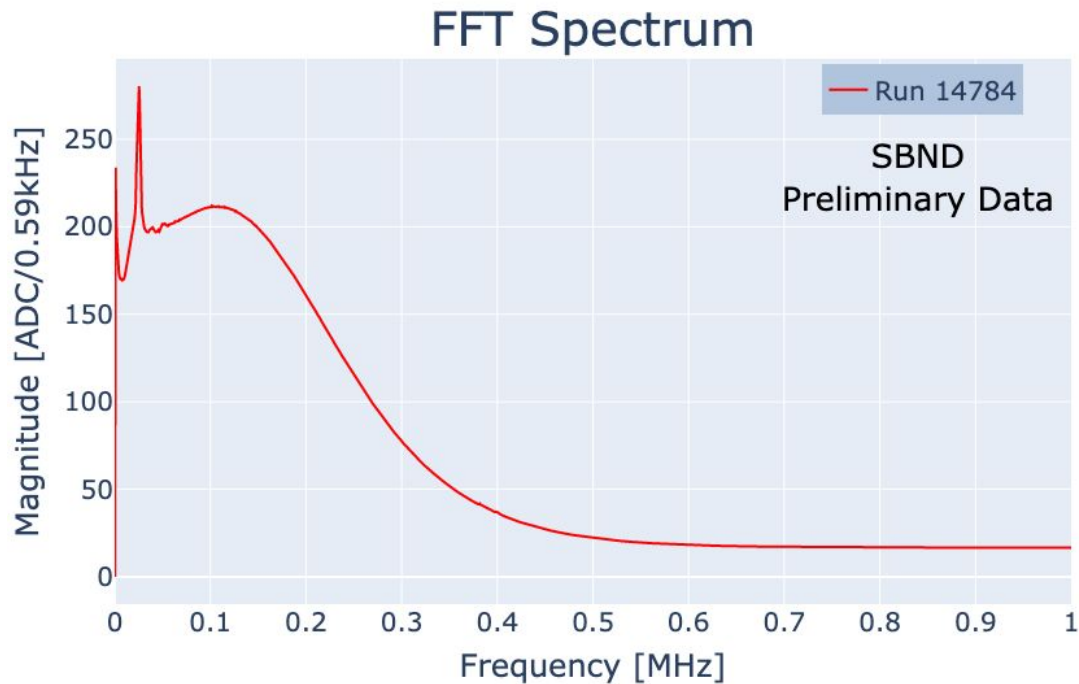
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# Additional Slides

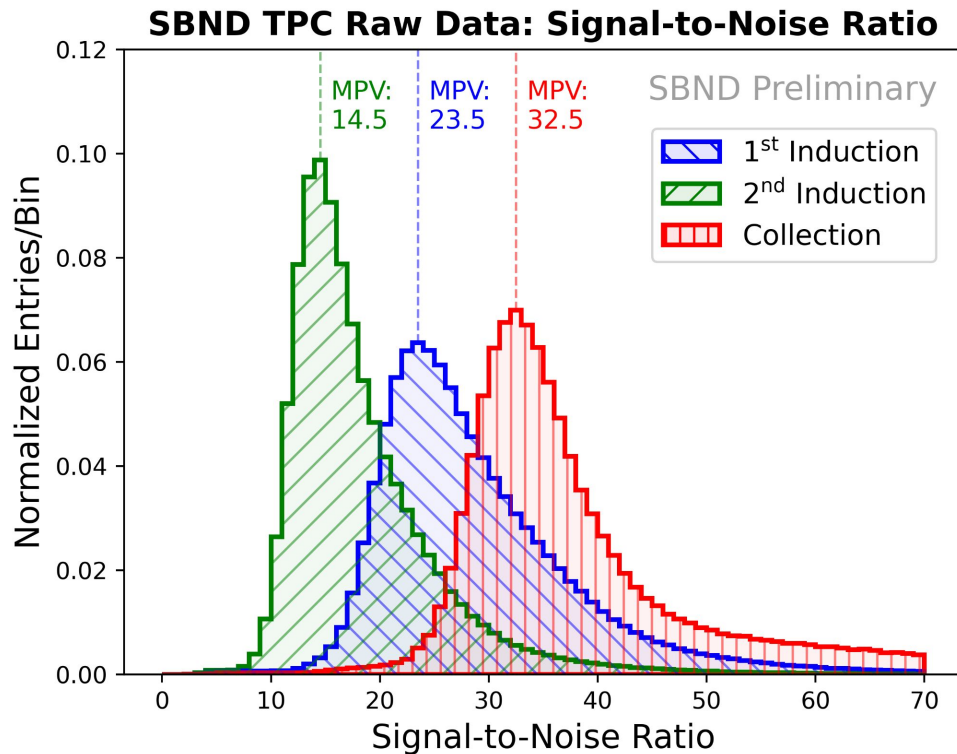
# SBND Timeline



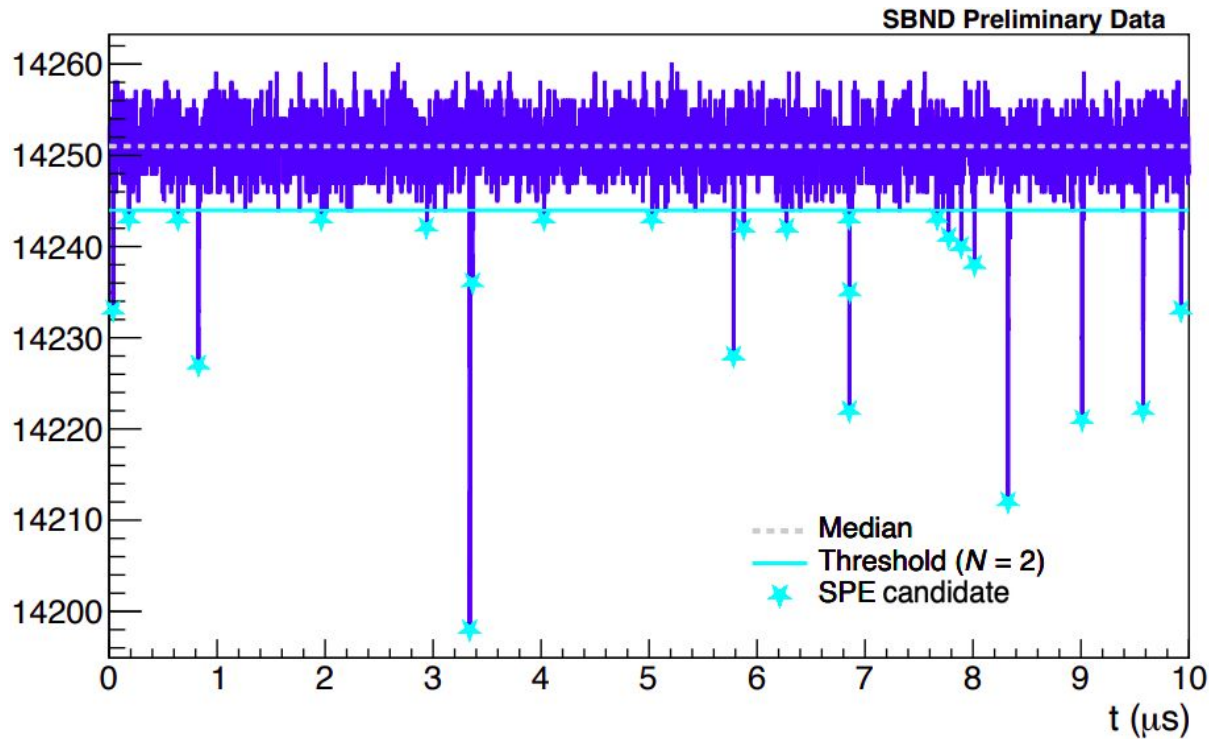
# TPC Electronics Noise Spectrum



# TPC Electronics Signal-to-Noise



# PMT Gain Equalization: SPE-Finding



# PMT Gain Equalization: SPE Fitting

