

PTOLEMY

**PonTecorvo Observatory for Light
Early-universe Massive-neutrino Yield**

Neutrino mass and Cosmic Neutrino Background with PTOLEMY experiment

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Università degli studi dell'Aquila

WIN25 10/06/2025



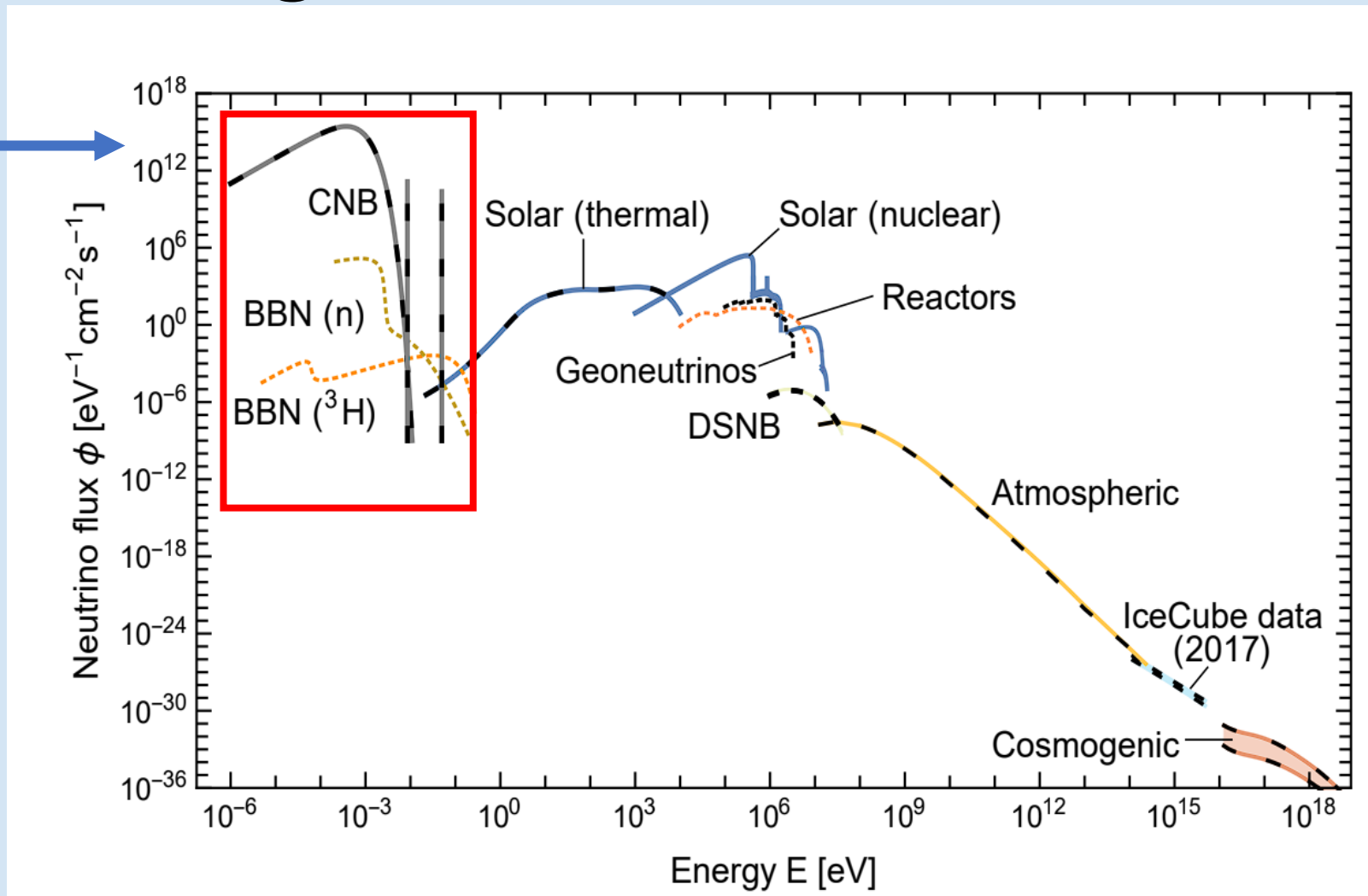
Cosmic Neutrino Background direct detection

- Most abundant neutrino source in the Universe!
- But extremely low Energy

The only neutrino source at low energy

$$K_\nu \approx 10^{-4} \text{ eV}$$

- Neutrinos do not provide energy for interaction
- Very difficult to detect:



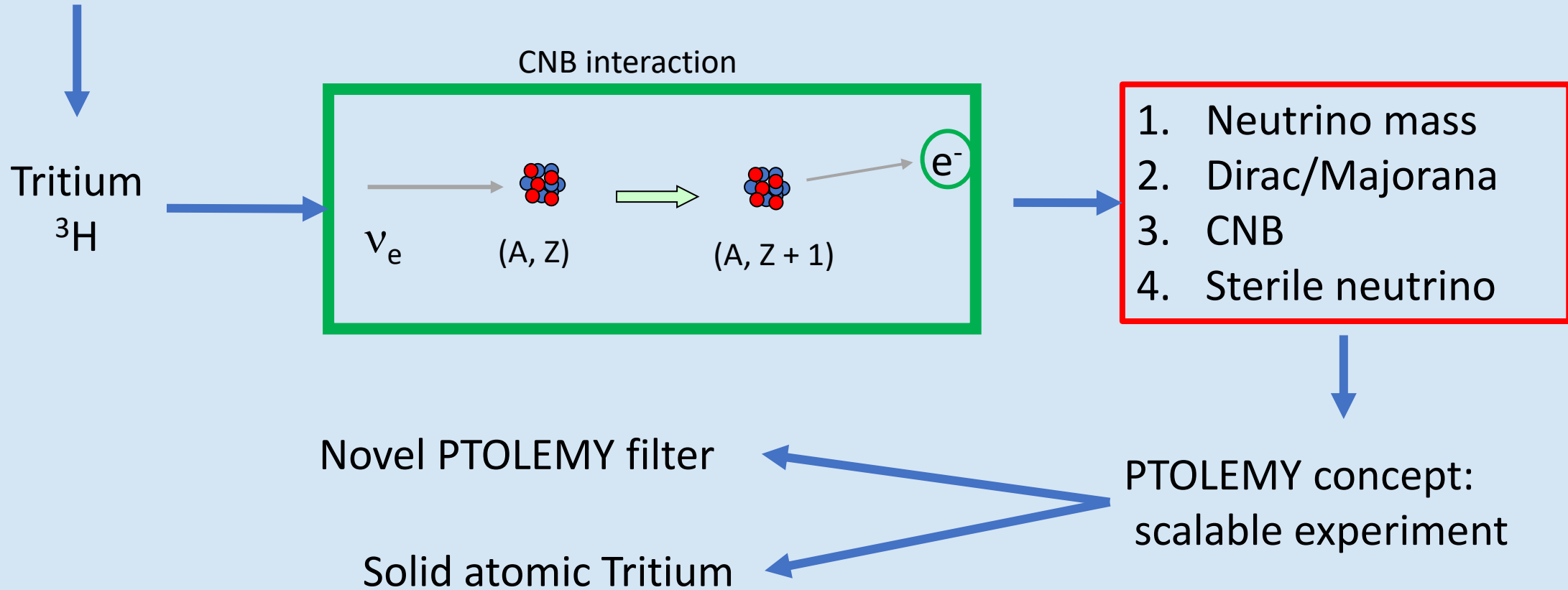
How?

PTOLEMY Project

Cosmic Neutrino Background direct detection: PTOLEMY

- Neutrino capture on beta decaying Nuclei reaction
- Need beta unstable element
- Want maximize interaction with CNB

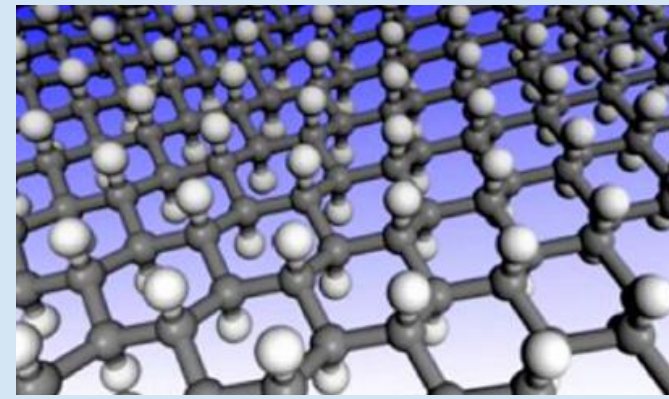
No energy needed!



JCAP 0706:015,2007

JCAP 07 (2019) 047

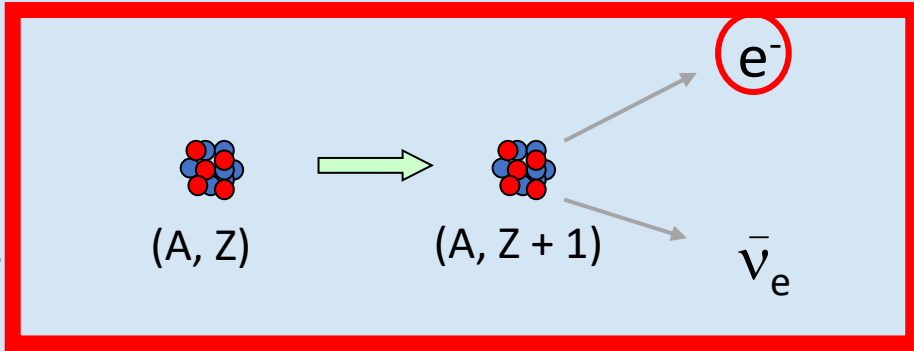
Solid Tritium target



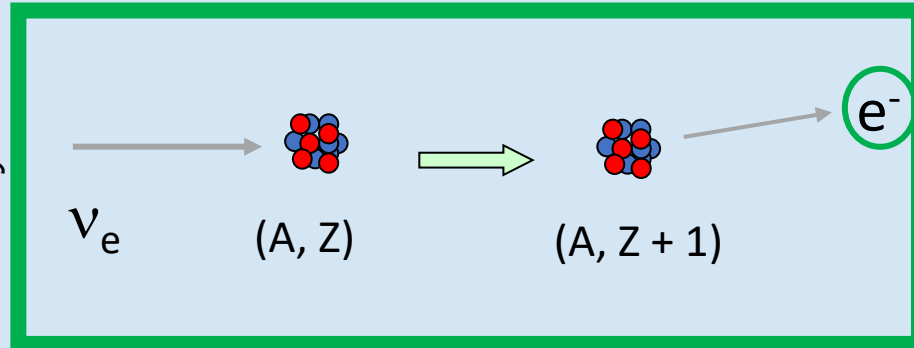
- Atomic Tritium on graphene



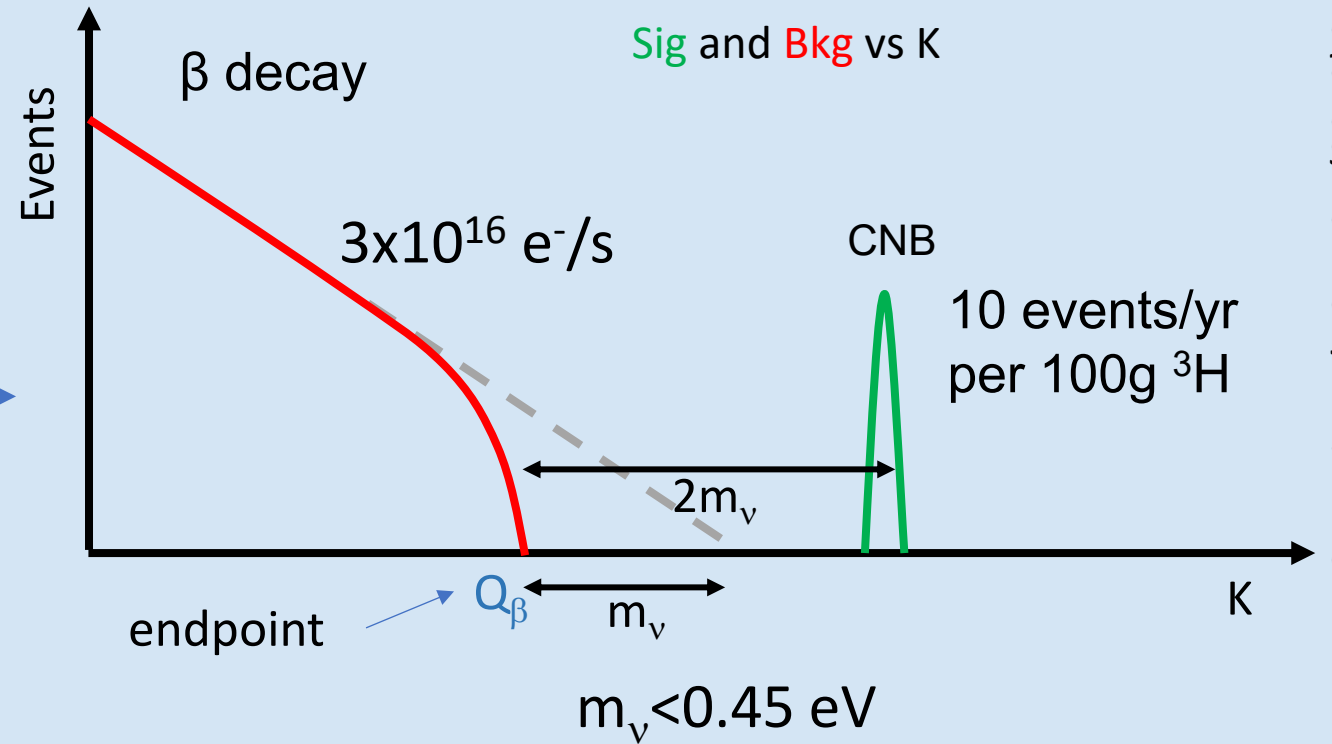
β decay (Background)



CNB interaction



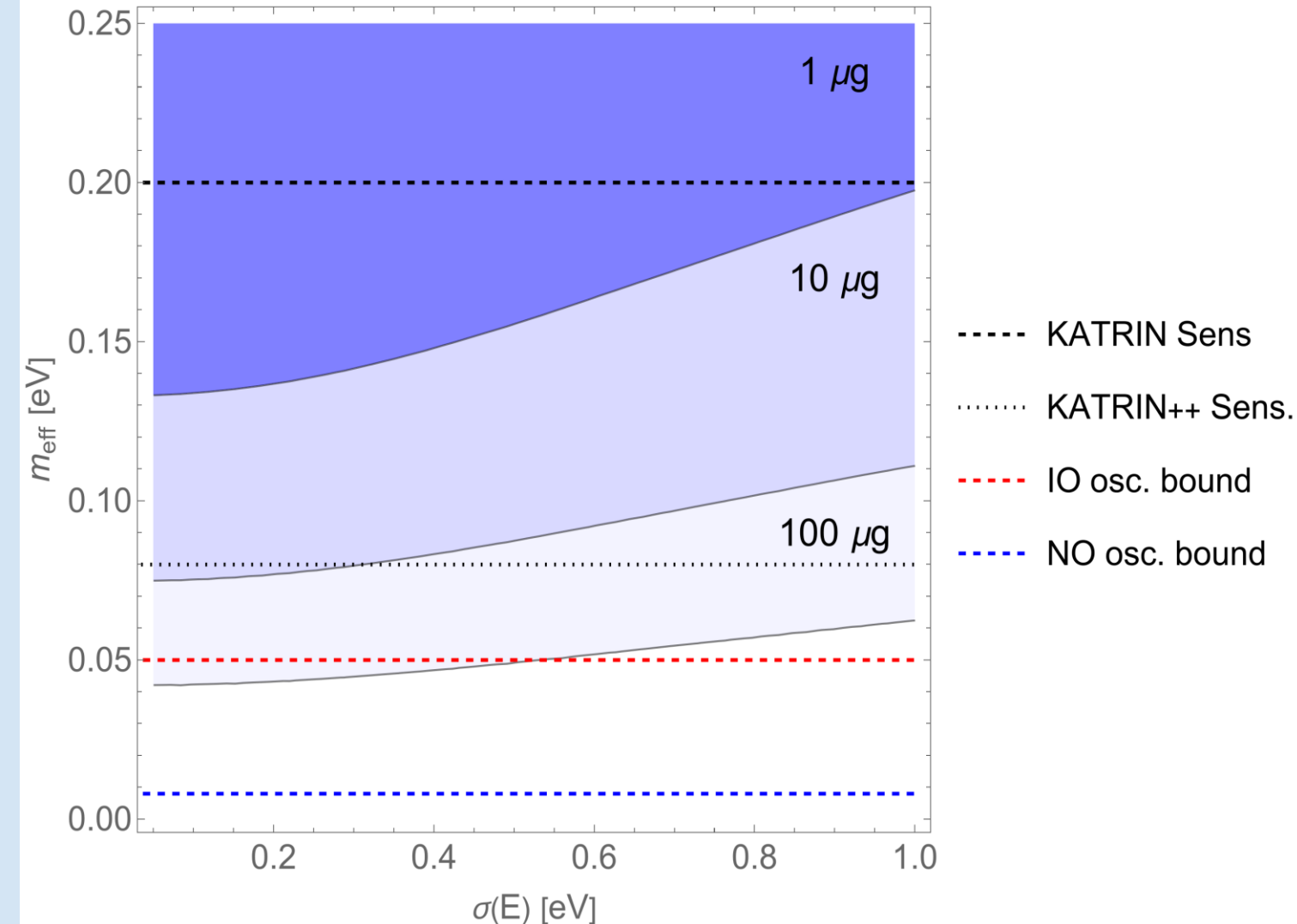
- High rate
- Need good energy resolution



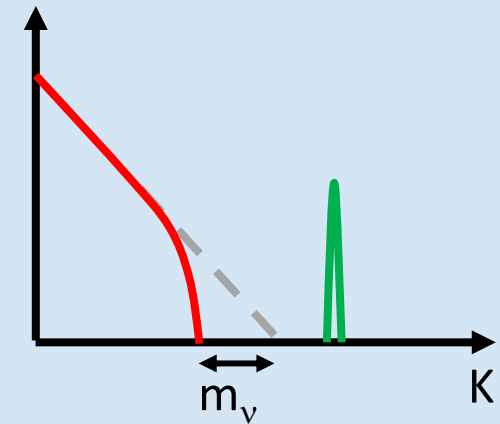
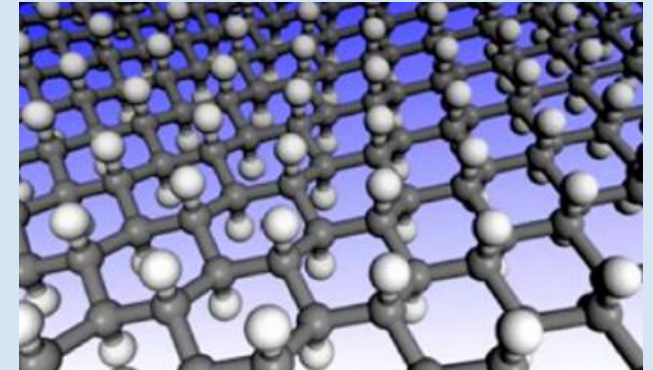
PTOLEMY neutrino mass sensitivity

1 μg = 7x7cm² of Tritiated graphene

PTOLEMY Sensitivity (3y - 50% eff)



Tritium target



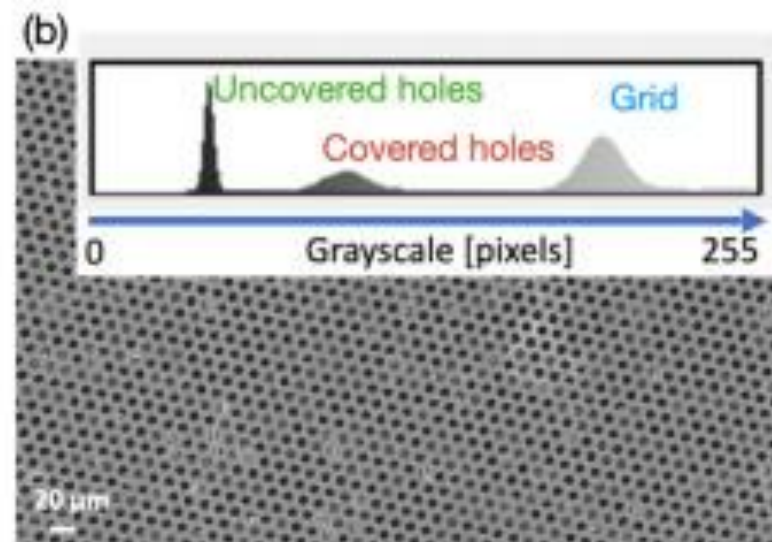
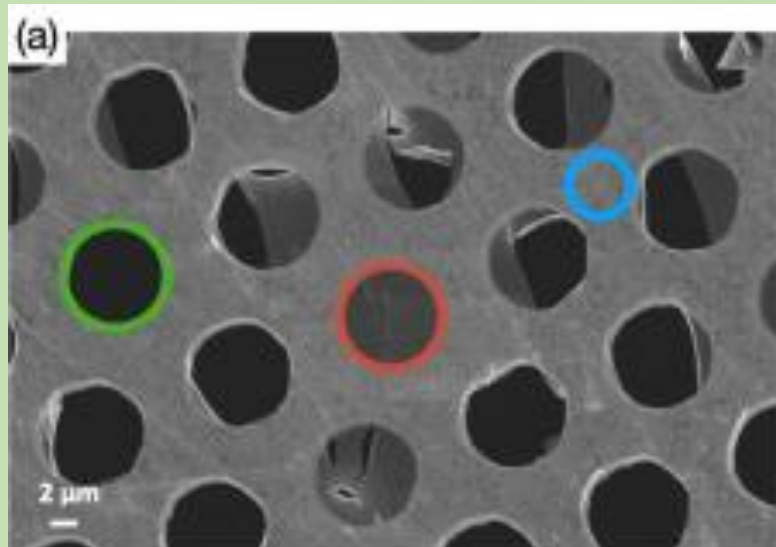
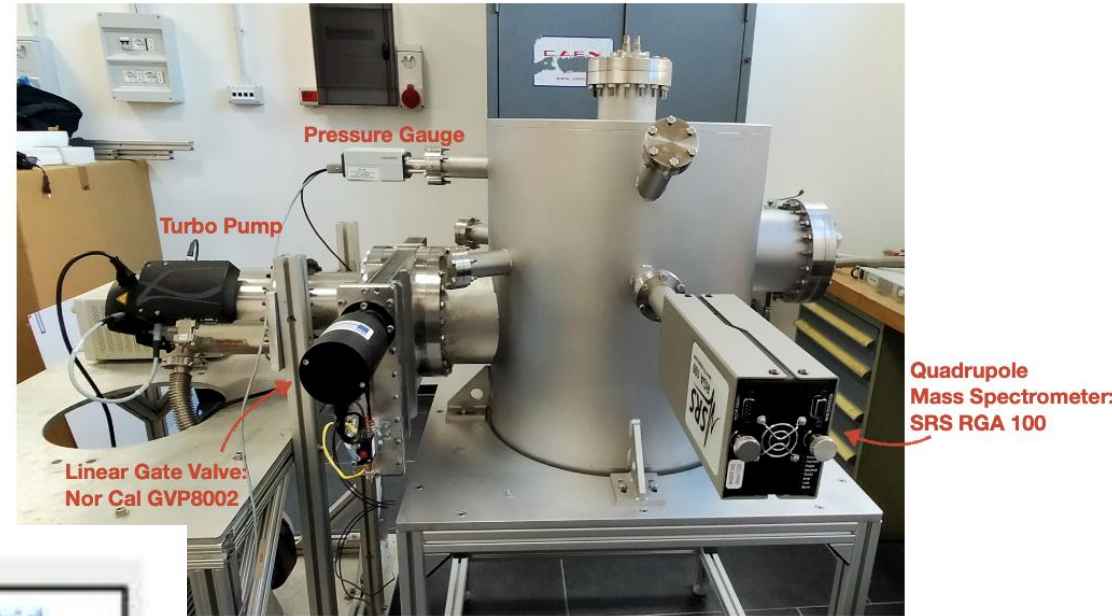
How could be possible to measure this?

Tritium: state of the art



- Tritium @ UKAEA: feasibility study and design requirements of new T loading chamber (quote received)
- Novel Tritium deposition technique **Done**
- High graphene loading **Done**
- Tritium Vacuum stability **Done**

T-chamber R side view



<https://arxiv.org/pdf/2504.10238>

NanoLett. 2022,22,2971–2977

<https://arxiv.org/abs/2504.11853>

RF region: filter trigger



Requirements:

- Measurement K, θ for each electron in the ROI
- $O(10 \text{ eV})$ energy resolution
- In $O(100 \mu\text{s})$
- Non destructive measurement
- Uniform 1T field

How?

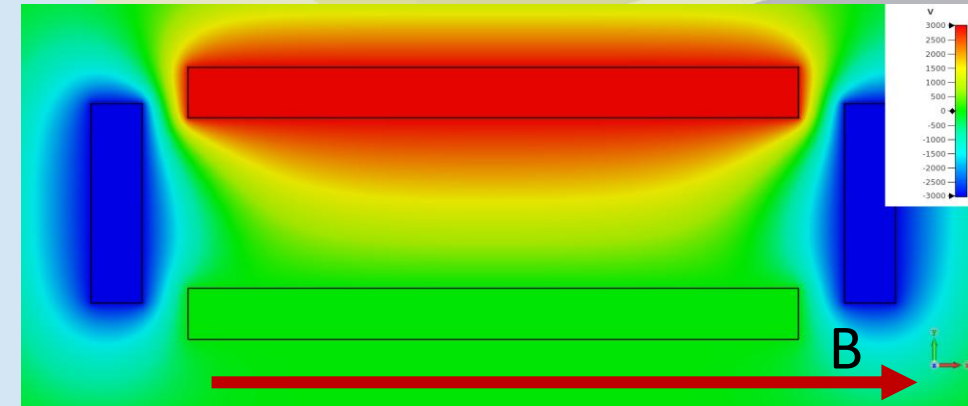
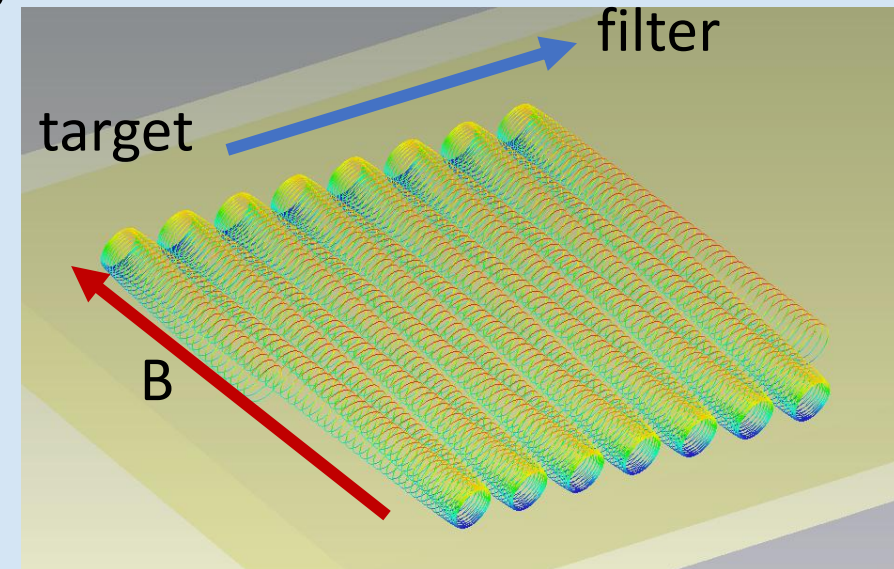
Electron spectroscopy with cyclotron radiation (CRES)

- Cyclotron radiation
- In uniform 1T
- Bouncing motion

PROJECT 8
Thank you!

$$f_c = 27\text{GHz}$$

$$P=1\text{fW}$$

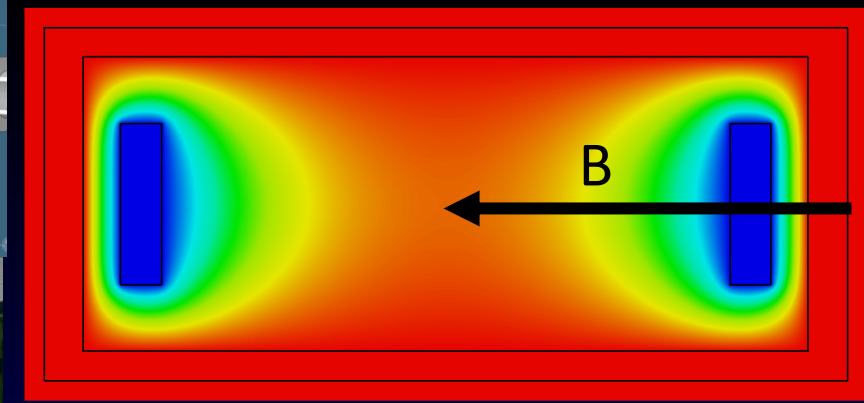
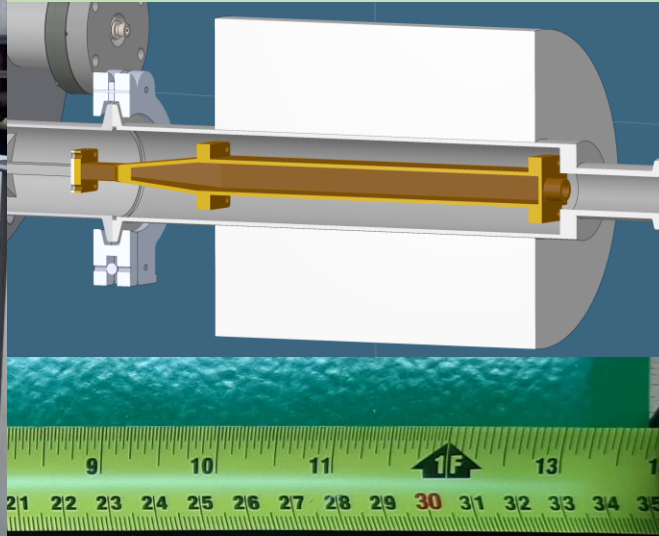
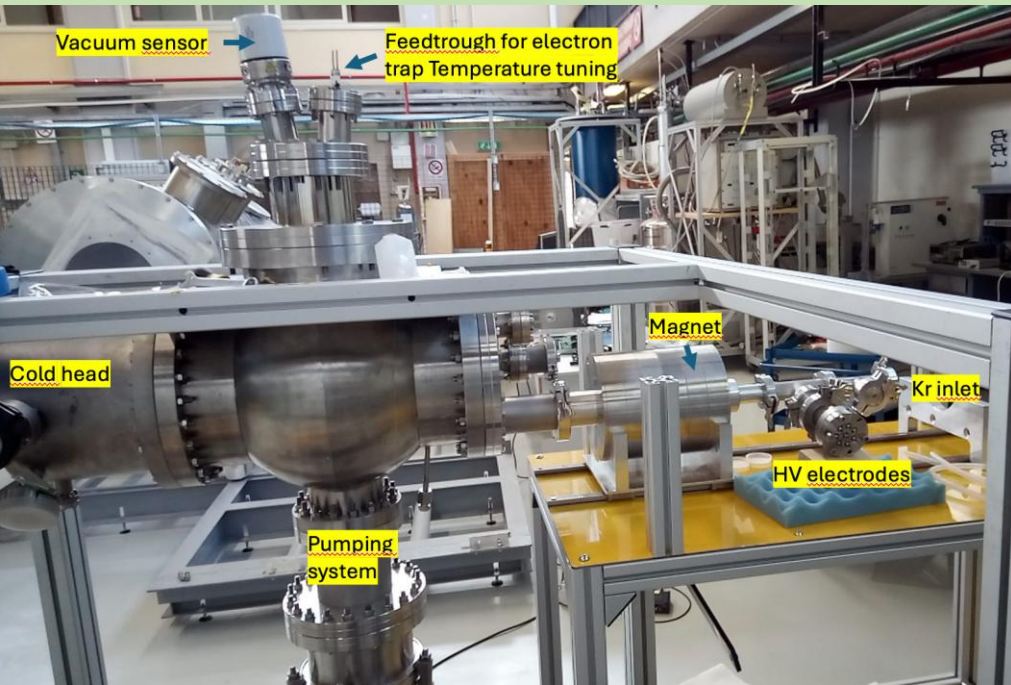
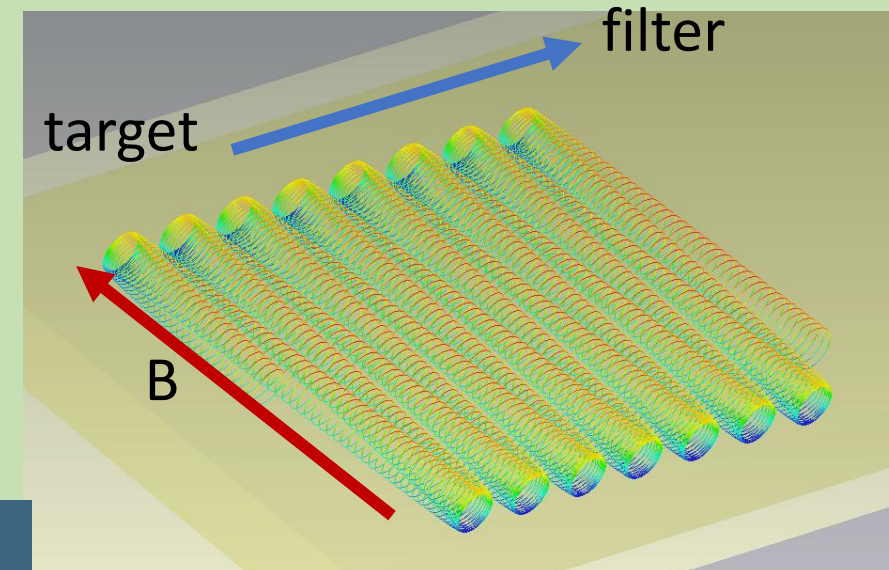


$$f_c = \frac{1}{2\pi} \frac{|q|B}{m} \frac{1}{K/m + 1}$$

Electron Trap: test setup for RF Region

Ongoing measurements @ LNGS

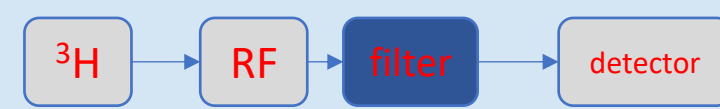
- ^{83m}Kr source
- 0.98T dipole permanent magnet
- Detection of RF emitted by electron in bouncing motion



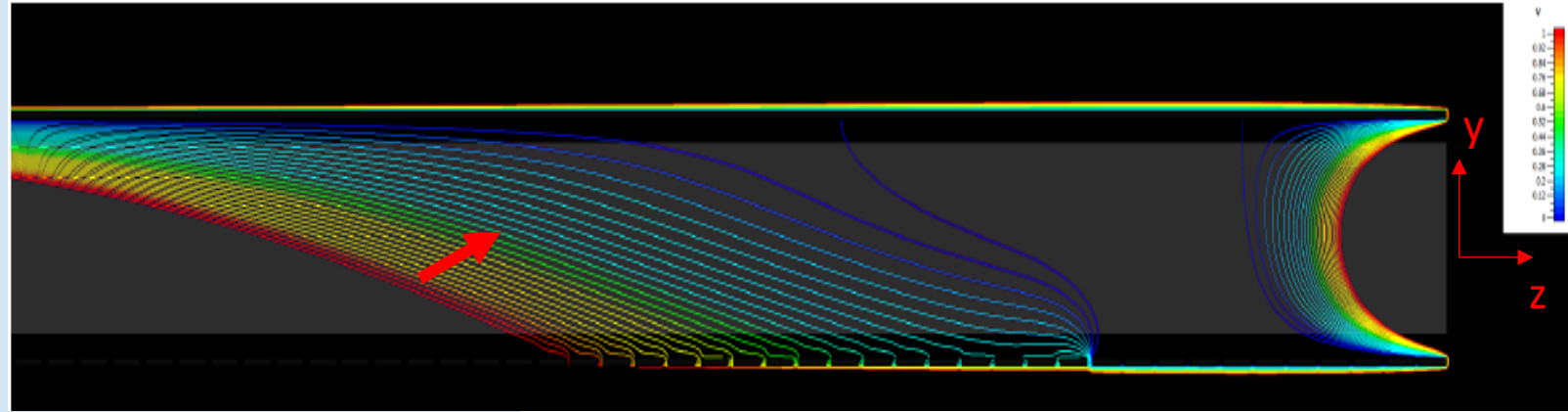
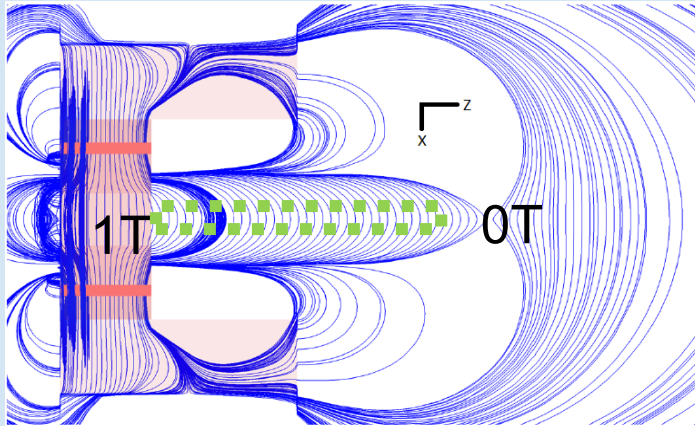
RF readout



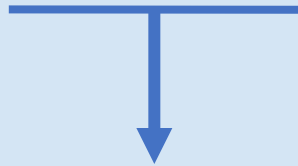
PTOLEMY filter



- Exponential decaying magnetic field
- Exponential decaying electric field



$$\begin{aligned}
 B_x &= B_0 \cos\left(\frac{x}{\lambda}\right) e^{-z/\lambda}, \\
 B_y &= 0, \\
 B_z &= -B_0 \sin\left(\frac{x}{\lambda}\right) e^{-z/\lambda}.
 \end{aligned}$$



$$\begin{aligned}
 E_x &= 0, \\
 E_y &= E_0 \cos\left(\frac{y}{\lambda}\right) e^{-z/\lambda}, \\
 E_z &= -E_0 \sin\left(\frac{y}{\lambda}\right) e^{-z/\lambda}.
 \end{aligned}$$

$$V_{E \times B}^y(z)|_{x,y=0} = \frac{\mathbf{E} \times \mathbf{B}}{B_x^2} = \frac{E_z B_x \hat{y}}{B_x^2} = \frac{E_z}{B_x} \hat{y}$$

$$V_{\nabla B-C} = \frac{1}{2} m (v_{\perp}^2 + 2v_{\parallel}^2) \frac{\mathbf{B} \times \nabla_{\perp} B}{qB^3} = (T_{\perp} + 2T_{\parallel}) \frac{\mathbf{B} \times \nabla_{\perp} B}{qB^3}$$



Dynamic EM Filter



Pitch angle Θ = angle between e^- velocity and B

- Electron deceleration:

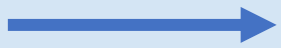
Drain K, Θ

K up to O(100 eV) on TES

Energy Drain ✓

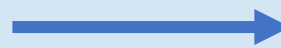
- If electron has:

$$K < Q_\beta$$



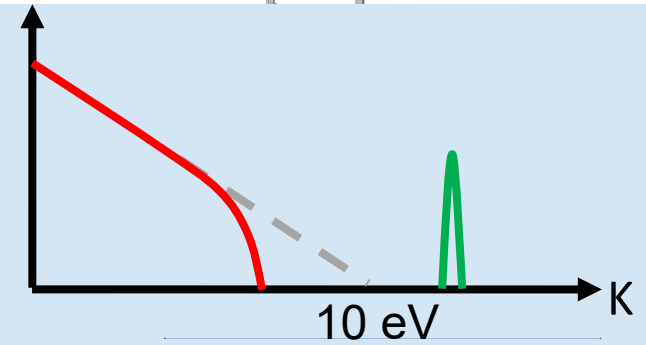
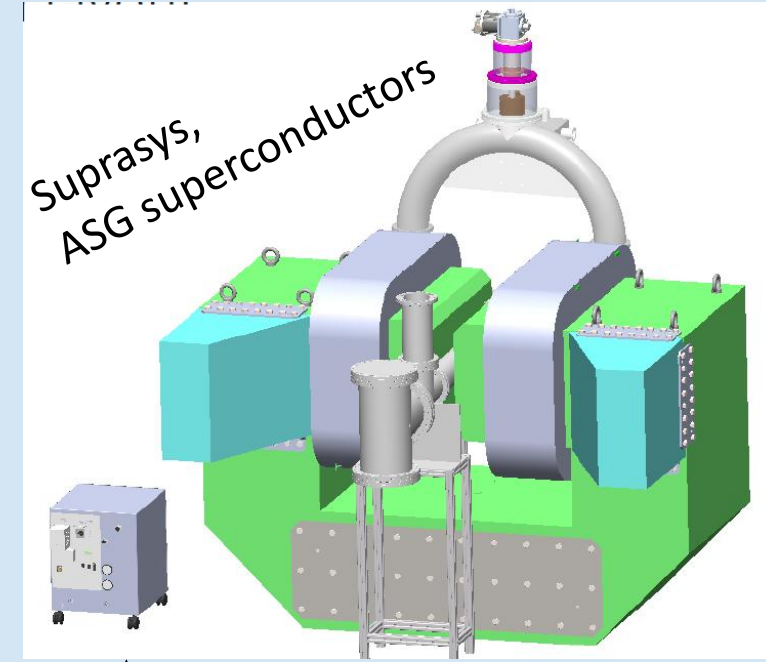
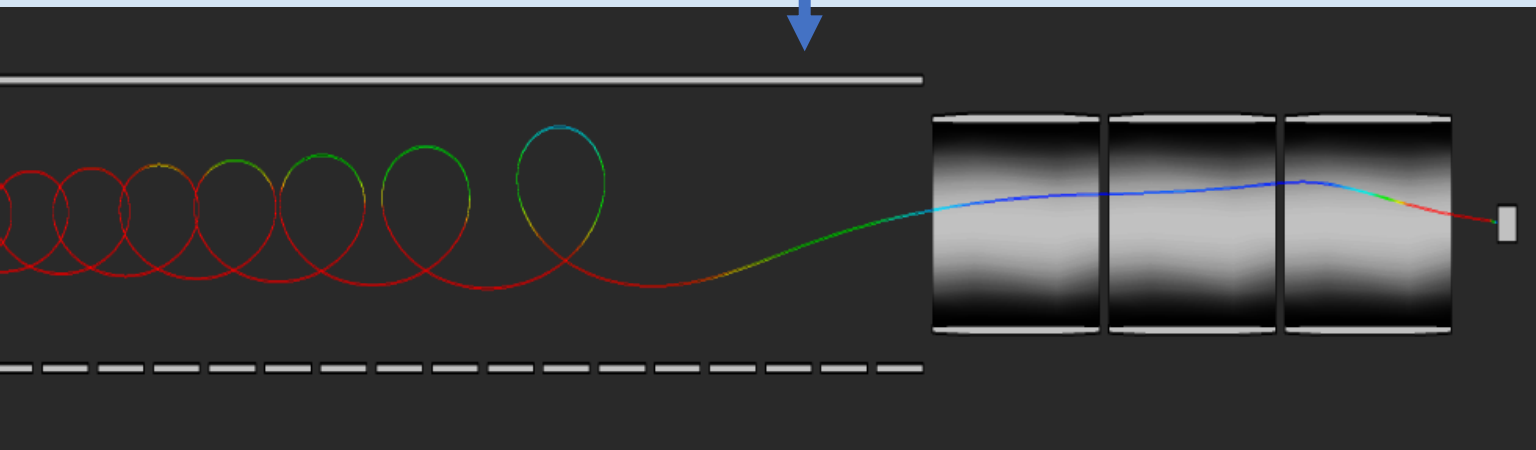
Discarded

$$K \sim Q_\beta$$



TES

Background rejection ✓



Event by event basis
If K, Θ known

RF Region

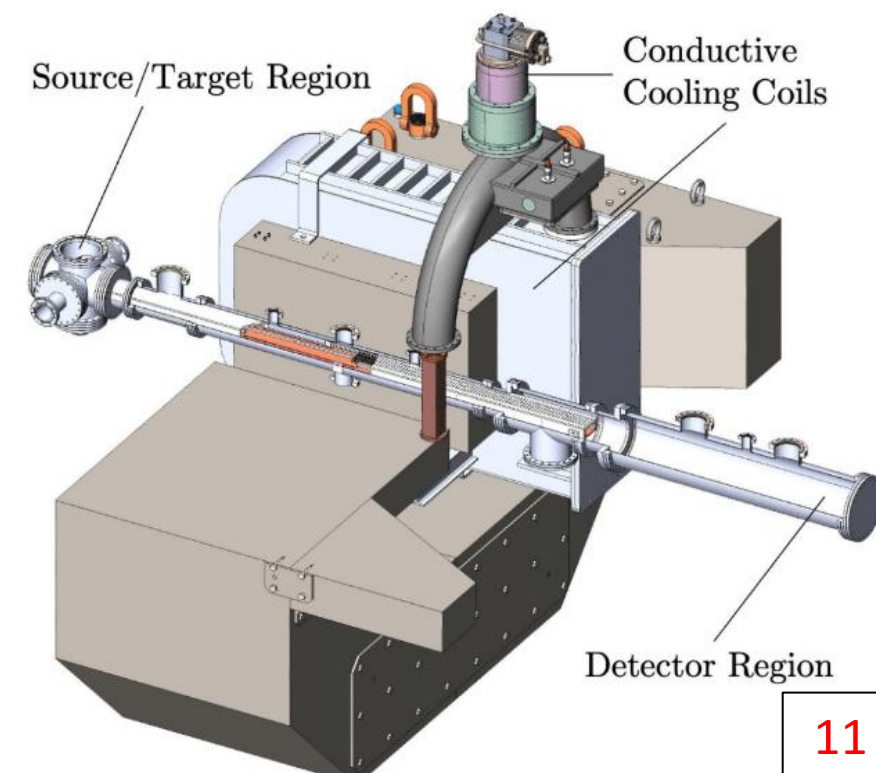
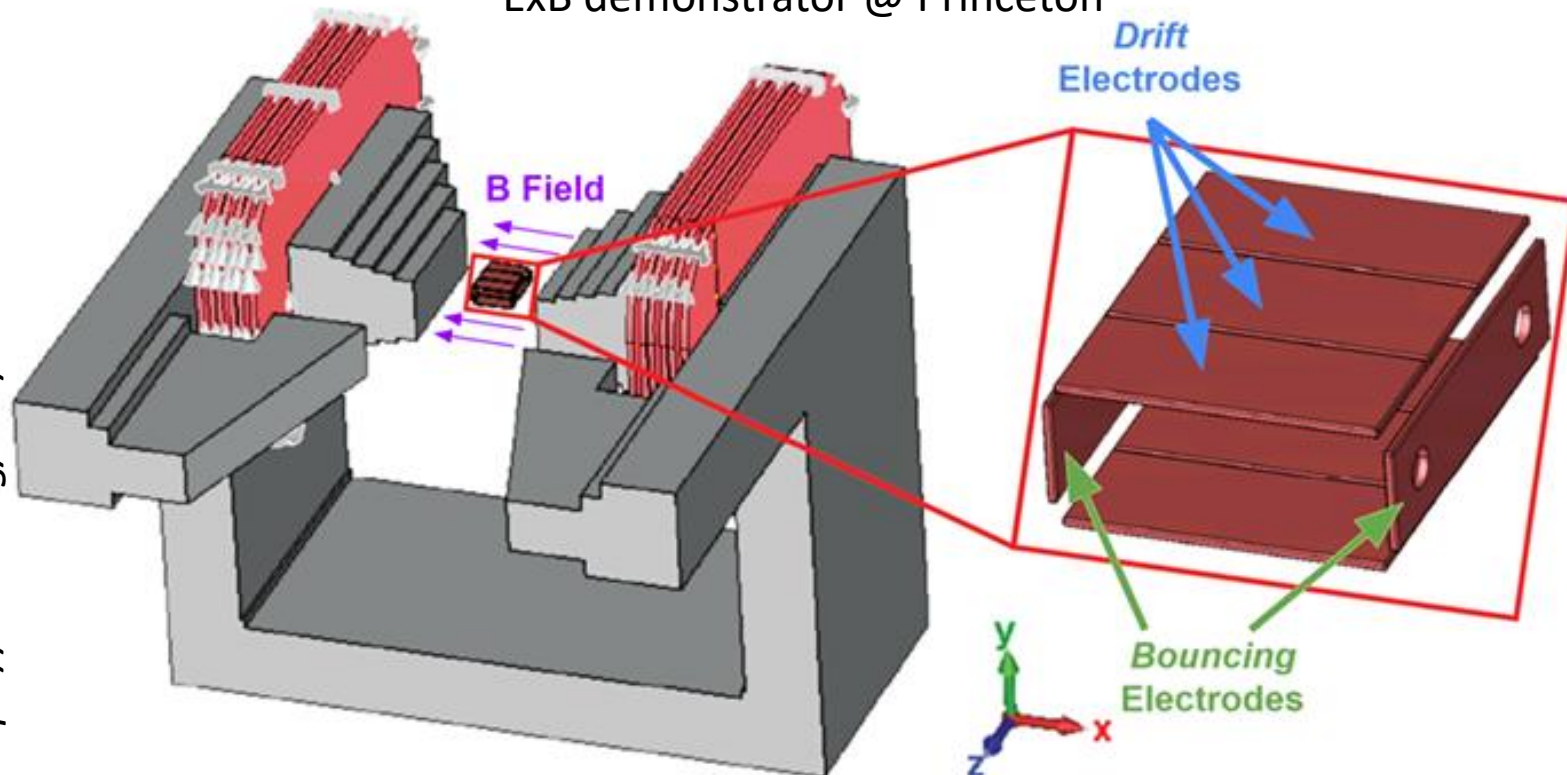
PTOLEMY filter: state of the art



- ExB demonstrator at Princeton: slow electron drift
- PTOLEMY magnet under construction: shipment to LNGS nov '25

Done

ExB demonstrator @ Princeton



<https://arxiv.org/abs/2503.10025>

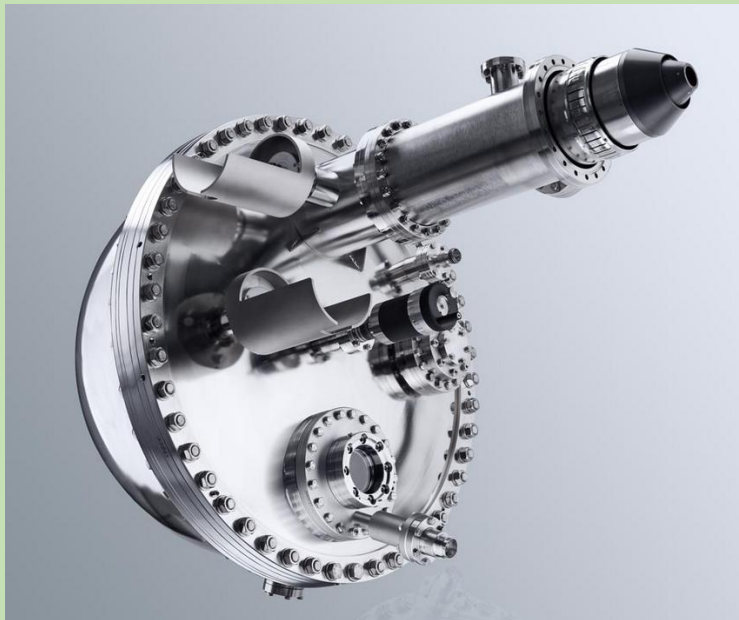
Obtaining desired energy resolution



Option A

Hemispherical electron analyser

- Typically used in spectroscopy
- Capability to reach $\sigma_K=50\text{meV}$
- Room temperature



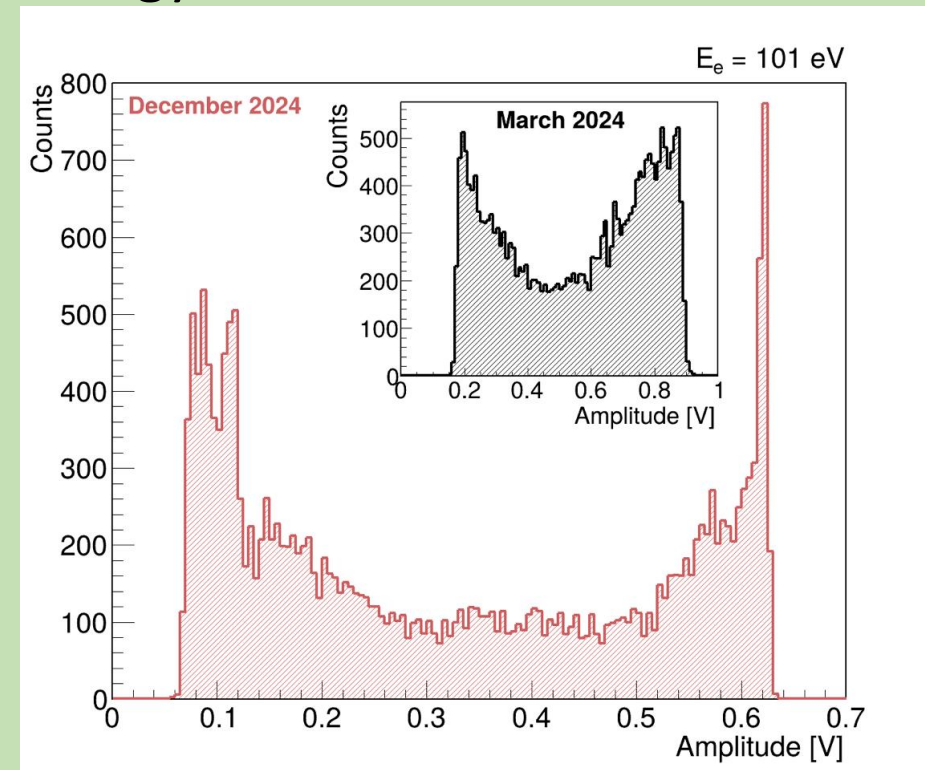
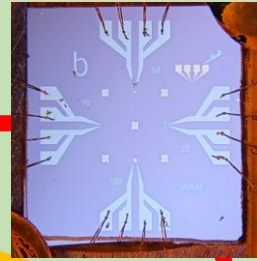
Commercial availability

Option B

TES for electrons

- Typically used for photons
- First time used for electrons in our energy ROI

Done



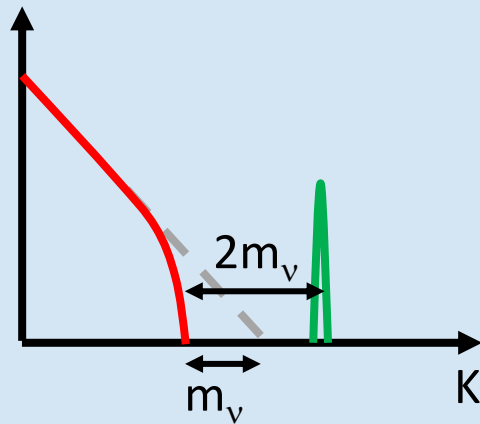
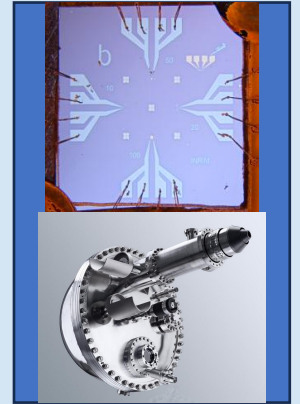
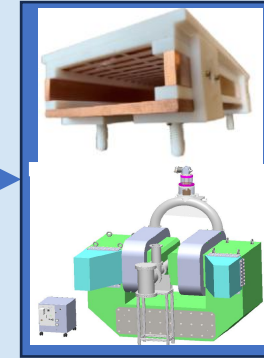
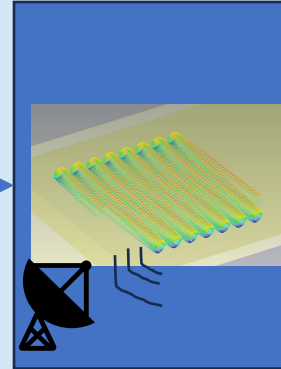
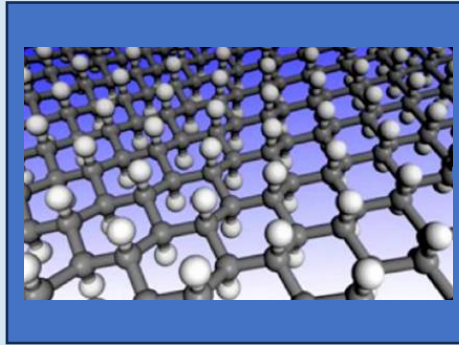
PTOLEMY Concept

Tritium target

RF region

Filter

High energy resolution measurement



Filter trigger
K and Θ , fast and rough
measurement

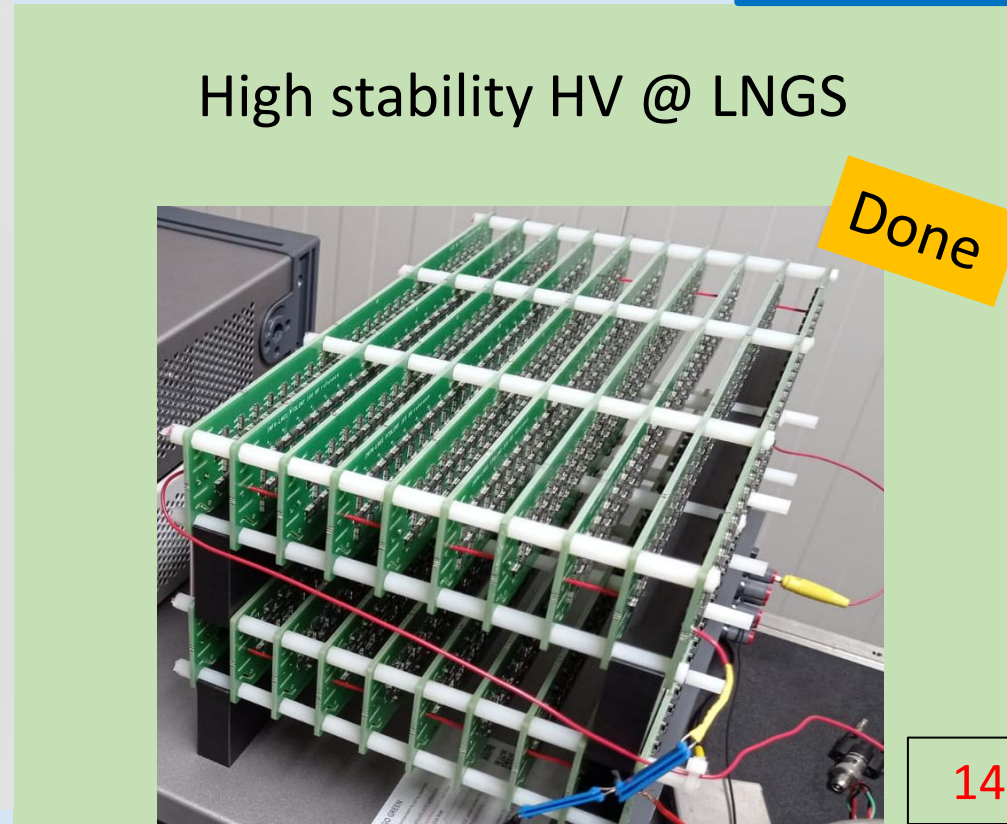
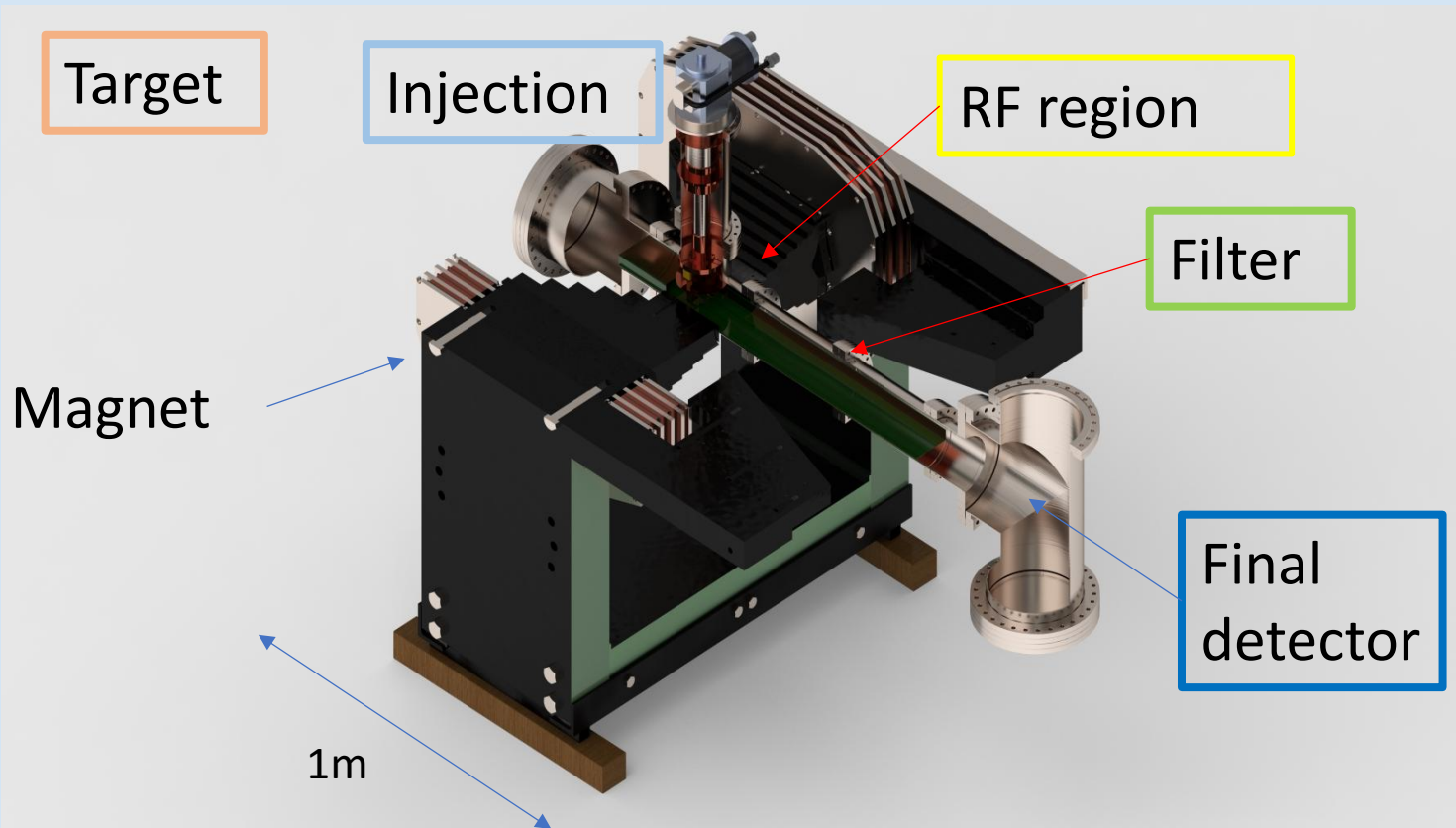
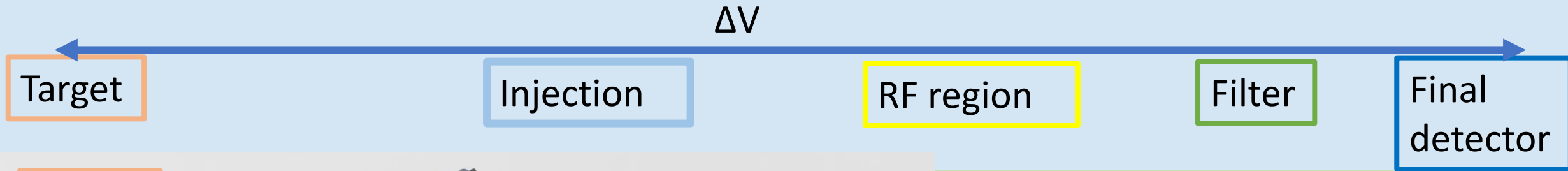
Energy Drain
+filter

$\sigma_K = 50$ meV
TOF information
combined with RF

PTOLEMY Concept

How to reconstruct target's electron kinetic energy K_i ?

$$\longrightarrow K_i = q\Delta V + E_{\text{RF,corr}} + E_{\text{final}}$$



High stability HV @ LNGS

The PTOLEMY Collaboration



Kavli Institute
for Cosmological Physics
at The University of Chicago



Radboud University



UNIVERSITÀ DI PISA



UNIVERSITÀ DEGLI STUDI DI NAPOLI
FEDERICO II



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Conclusion and outlook

PTOLEMY experimental method



Atomic, solid Tritium target

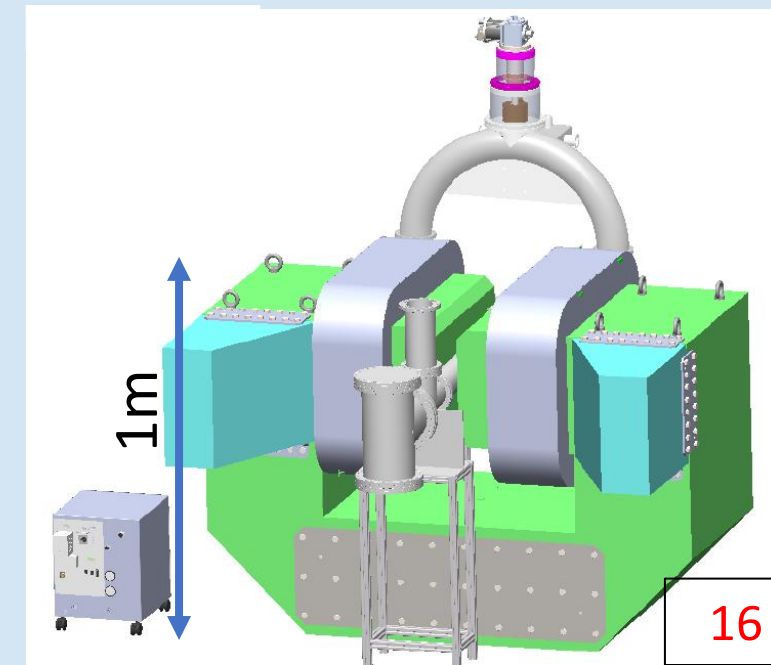
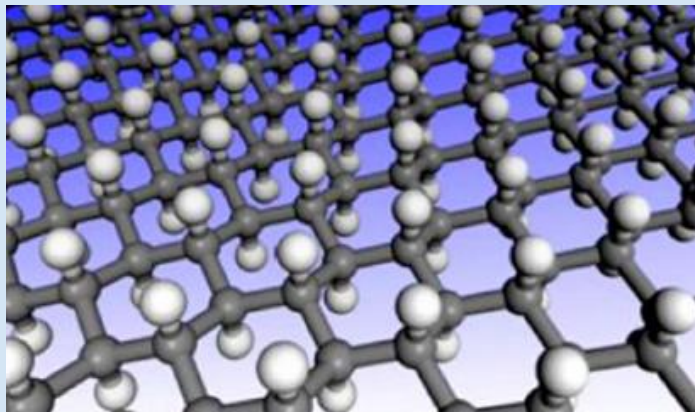
Background rejection by realtime filter
+ small size detector

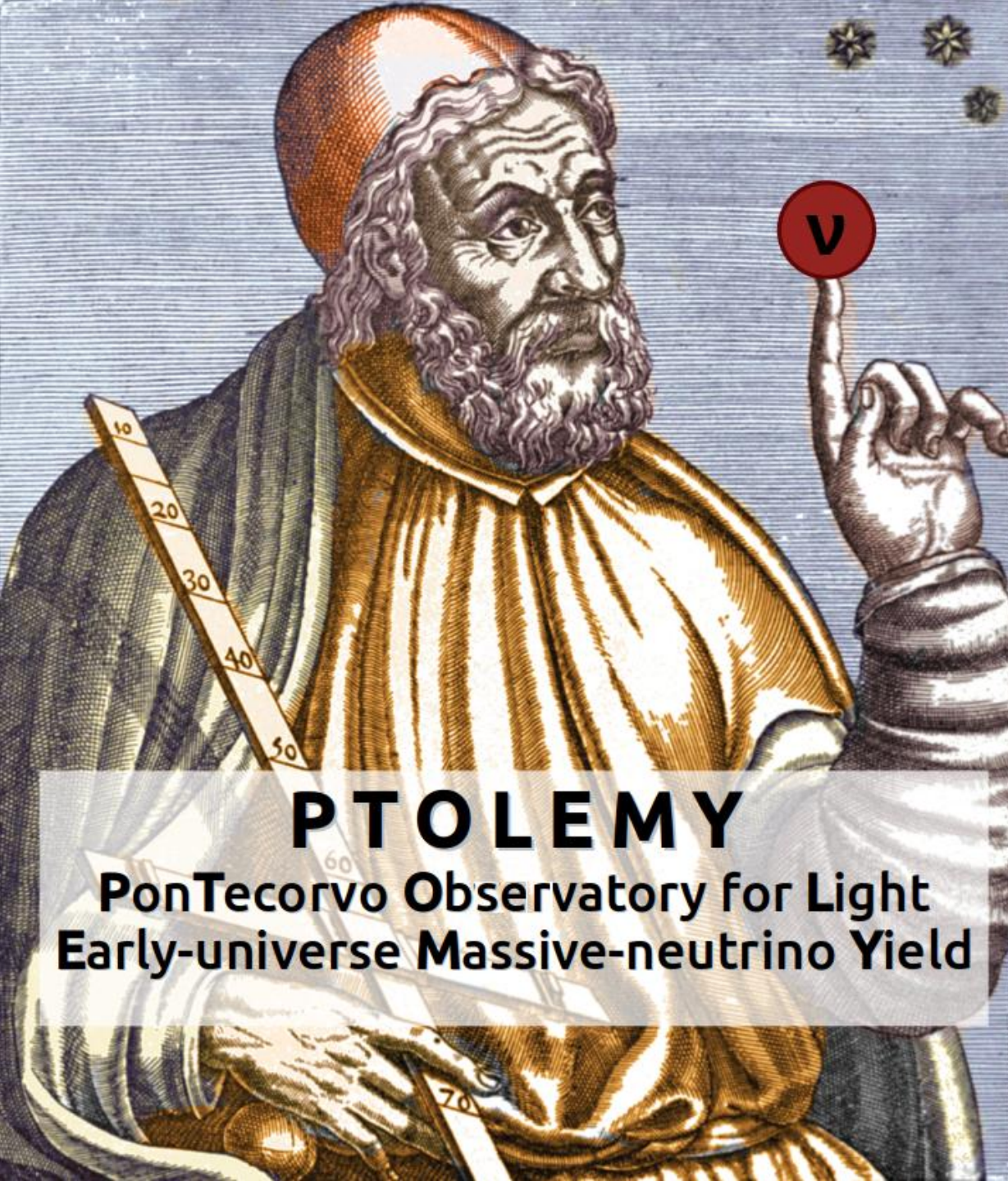
High precision energy measurement



Keys for future
high Tritium mass
experiment

- Phase 0) Demonstrator @ LNGS
- Phase 1) Measurement m_ν
- Phase 2) Mid Scale demonstrator
- Phase 3) Measurement CNB





PTOLEMY

PonTecorvo Observatory for Light
Early-universe Massive-neutrino Yield

Backup slides

b