

# Inhomogeneous phases of dense nuclear matter

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Theo F. Motta (IFT UNESP)

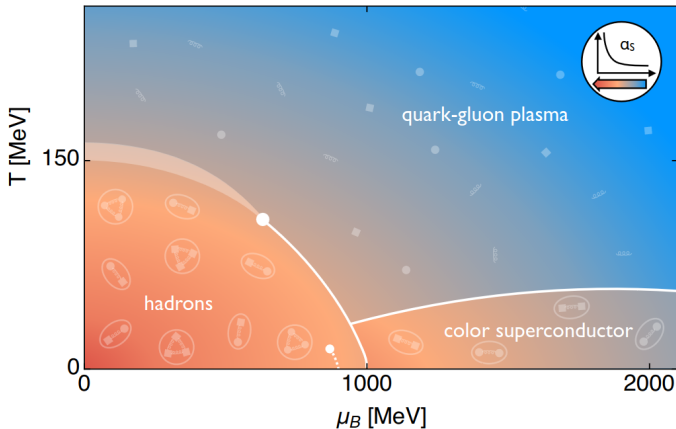
18<sup>th</sup> August, 2026 SEWM Helsinki

In collaboration with Gastão Krein & Randall Pradinett

[PhysRevD-L071503] & [PhysRevD-113-9-094028]

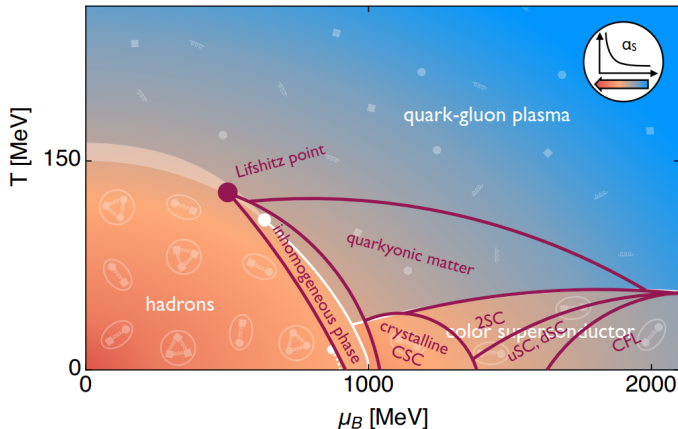
# QCD Phase Structure

$$\mathcal{L} = \frac{1}{4} F_{\mu\nu}^a F_{\mu\nu}^a + \bar{\psi}(\not{D} + m_q)\psi$$

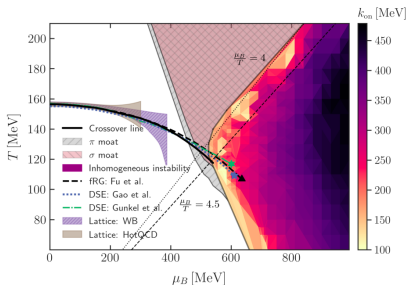
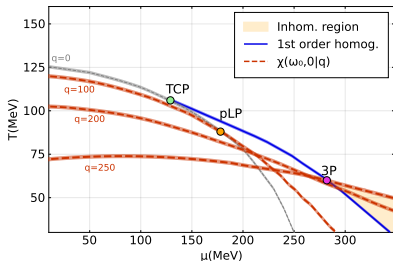


# QCD Phase Structure

$$\mathcal{L} = \frac{1}{4} F_{\mu\nu}^a F_{\mu\nu}^a + \bar{\psi}(\not{D} + m_q)\psi$$



# Some QCD results



- DSE based calculation

[TFM, J. Bernhardt, M. Buballa, C. Fischer, Phys.Rev.D 111 (2025) 7, 074030]

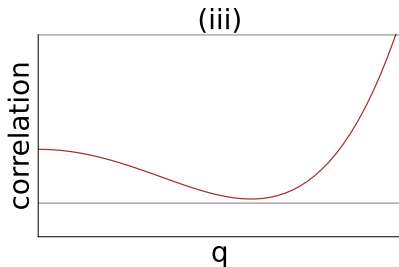
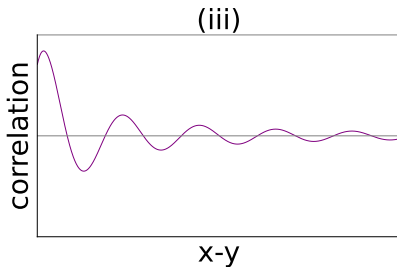
- FRG based calculation

[J. Pawłowski, F. Rennecke, F. Sattler, arXiv: 2512.20510]

# Correlation Functions

$$\langle \psi(\textcolor{red}{x})^\dagger \hat{O} \psi(\textcolor{red}{x}) \psi(\textcolor{blue}{y})^\dagger \hat{O} \psi(\textcolor{blue}{y}) \rangle = f(\textcolor{red}{x} - \textcolor{blue}{y})$$

- There is something “in between”, however, the moat regime
- The phase is homogeneous, but it's correlation function...



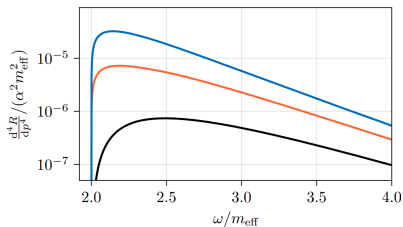
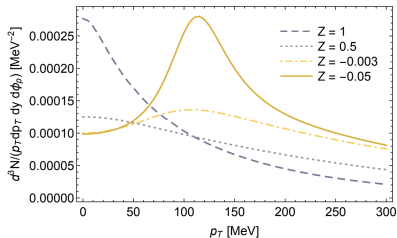
# Moat signatures

- Particles produced at finite momentum

[Phys.Rev.Lett. 127 (2021) 15, 152302]

- Increased dilepton production

[Phys.Rev.Lett. 135 (2025) 10, 101904]



# What about dense nuclear matter?

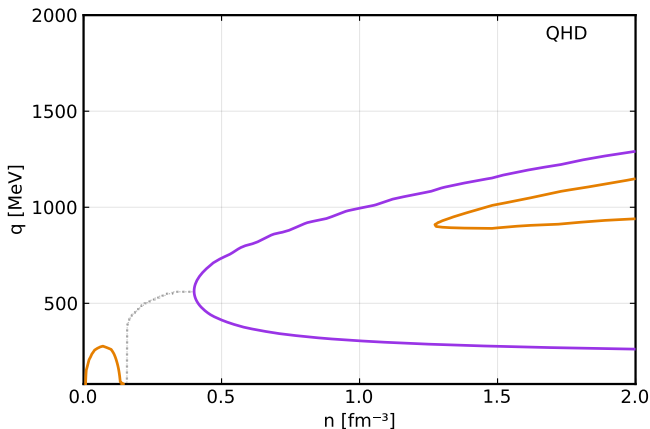
- Migdal's work in the 70's about  $\pi$  condensation
- "The Possibility of a liquid crystal phase in nuclear matter"

PLB by Migdal and P. Watson

- From the classic 1997 Review by Serot & Walecka:
  2. In nuclear matter in the RRP, poles appear in the polarization propagator at zero frequency  $q_0 = 0$  and finite wavenumber  $|\mathbf{q}| \neq 0$ ; the value of this wavenumber is a few times the nucleon mass in QHD-I.<sup>57,84-86,91</sup> Such poles imply an instability of the system against density fluctuations of the corresponding wavelength. There are several possible interpretations of these results:
    - (a) The RRP for the propagators is inadequate.
    - (b) Vertex modifications are important. (Phenomenological form factors affect the numerical results significantly.<sup>84</sup>)
    - (c) The composite structure of the baryon must be taken into account before one reaches distance scales where these poles develop.
    - (d) The instability is real.

# Zeros of the SDF in QHD

$$\epsilon(n_b, \vec{q}) = \det\left(\begin{array}{c} \text{---} \bullet \text{---} \\ \text{---} \end{array}^{-1}\right) = \epsilon_L(n_b, \vec{q}) \epsilon_T(n_b, \vec{q})^2$$



[Reproducing K. Lim, C. Horowitz Nucl.Phys.A 501 (1989) 729-750]

# Quark-Meson-Coupling model (QMC)

$$(i\not{\partial} - m_q + g_\sigma^q \sigma - g_\omega^q \gamma^0 \omega_0) \psi = 0$$

$$i\gamma \cdot \hat{r} \psi(R) = -\psi(R)$$



- Quarks are confined in spherical bags
- Quarks interact with quarks via meson exchange

$$\psi_m(\vec{r}) = \mathcal{N} \left( \begin{array}{c} j_0(xr/R) \\ i\beta_q \vec{\sigma} \cdot \hat{r} j_1(xr/R) \end{array} \right) \frac{\chi_m}{\sqrt{4\pi}}$$

- Use this to calculate the energy of a stationary bag, the baryon mass  $M^*(n_s)$

# Quark-Meson-Coupling model (QMC)

- So if I take this result for  $M^*$  and put it into the baryon propagator

$$G_F(k) = \frac{\not{k} + M^*}{k^2 - M^{*2} + i\epsilon}$$

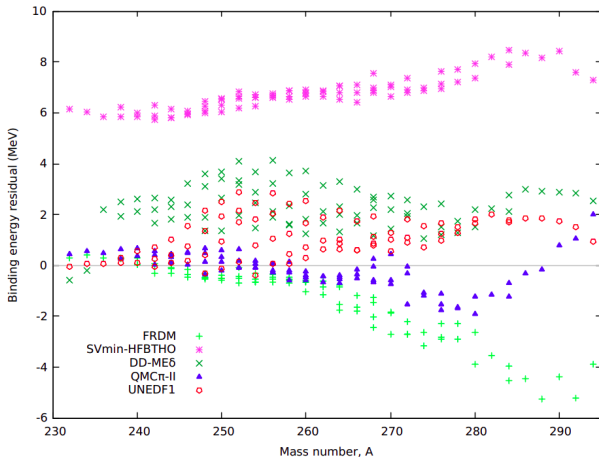
$$G_D(k) = 2\pi i(\not{k} + M^*)\delta(k^2 - M^{*2})\eta(k_0)$$

I can do nuclear many-body physics

- I can also obtain the effective coupling in MF by

$$\Gamma(n_s) \propto \frac{\partial M^*}{\partial n_s}$$

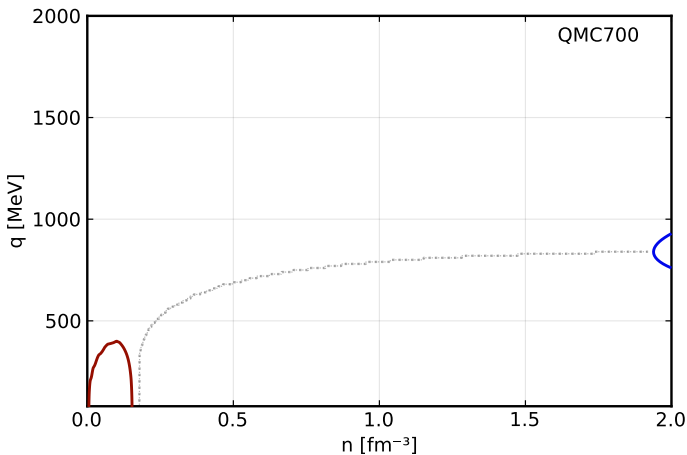
# Quark-Meson-Coupling model (QMC)



- Quark-Meson-Coupling (QMC) model for **finite nuclei**, nuclear matter and beyond [P.A.M. Guichon, J. Stone, A.W. Thomas, Prog.Part.Nucl.Phys. 100 (2018) 262-297]
- The role of baryon structure in **neutron stars** [T.F. Motta, A.W. Thomas, Mod.Phys.Lett.A 37 (2022)]

# Quark-Meson-Coupling model (QMC)

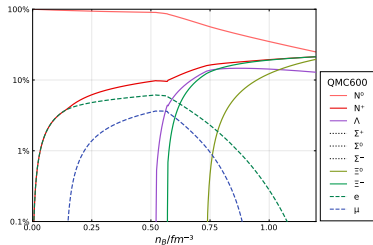
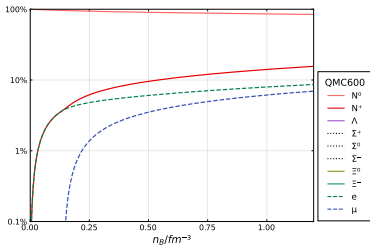
$$\epsilon(n_b, \vec{q}) = \det\left(\widetilde{\text{---}} \bullet \text{---}^{-1}\right) = \epsilon_L(n_b, \vec{q}) \epsilon_T(n_b, \vec{q})^2$$



# Quark-Meson-Coupling model (QMC)

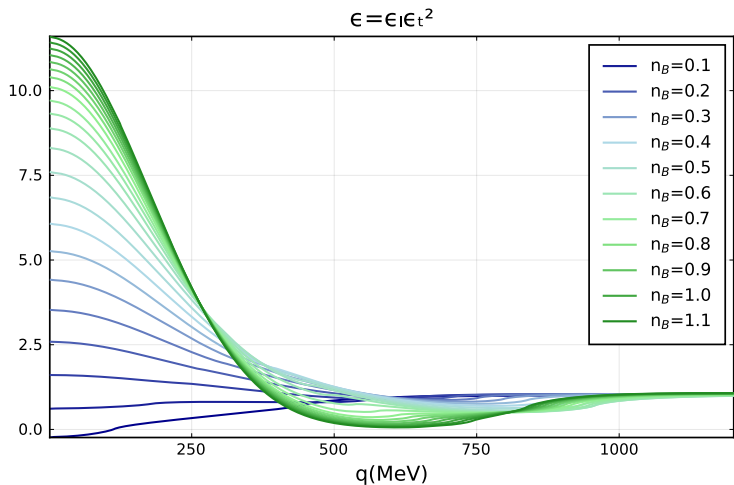
- We can write the energy density of INM as

$$\mathcal{E}(\{n_b\}) = \int \frac{d^3k}{(2\pi)^3} \sqrt{k^2 + M_b^{*2}} + \text{meson exchange diagrams},$$



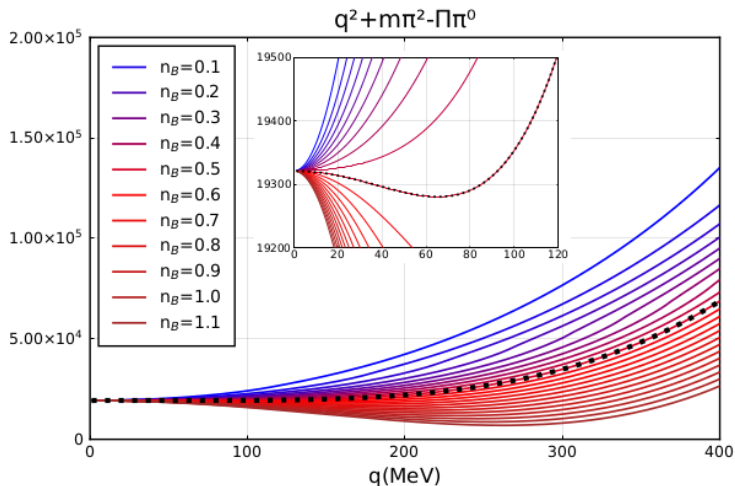
# Nucleons only $\beta$ -equilibrium

- The stability analysis of the **isoscalar** channels



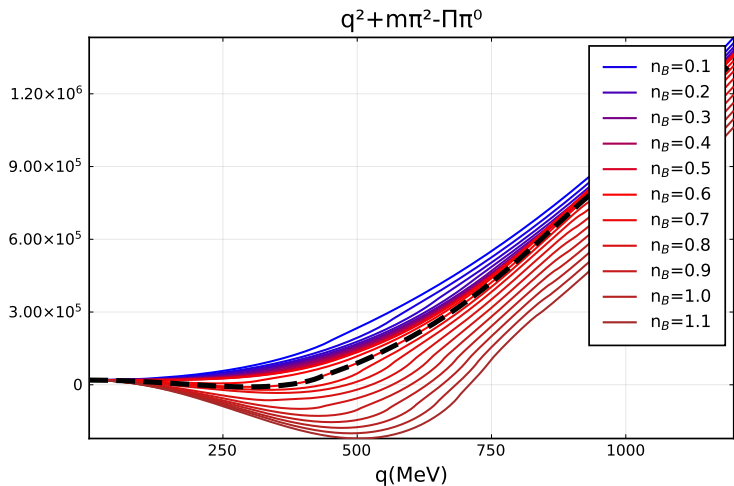
# Nucleons only $\beta$ -equilibrium

- Looking at the pions specifically



# Full baryon octet $\beta$ -equilibrium

- Again... Looking at the pions specifically



- Moat regimes are most likely present in dense nuclear matter
- No instability is present with nucleons only
- Instabilities show up for high density hyperon rich matter

backup

# Nuclear Matter (RMF)

- Take a model, any model

$$\begin{aligned}\mathcal{L} = & \bar{\psi}_b (i \not{\partial} - M_b) \psi_b + g_\sigma^b \bar{\psi}_b \sigma \psi_b - g_\omega^b \bar{\psi}_b \not{\omega} \psi_b \\ & - g_\rho^b \bar{\psi}_b \not{\rho}_a \tau_a \psi_b - \frac{g_A}{2f_\pi} \bar{\psi}_b \gamma_5 \not{\partial} \pi_a \tau_a \psi_b + \mathcal{L}_{\sigma\omega\rho\pi},\end{aligned}$$

- Essentially what we want to do is

$$\Omega[n_s(\vec{x}), n_{ps}(\vec{x}), n_b(\vec{x}), \dots] \quad \text{around} \quad n(\vec{x}) \rightarrow \bar{n} + \delta n(\vec{x})$$

Which we can write as

$$\Omega[n(x)] \approx \Omega[\bar{n}] + \delta\Omega[\bar{n}, \delta n(\vec{x})].$$

- So the key piece of the analysis is

$$\delta\Omega[\bar{n}, \delta n(\vec{x})] = \int_{\vec{q}} |\delta n(\vec{q})|^2 D_{\phi}^{-1}(q_0 = 0, \vec{q})$$

- So the key piece of the key piece of the analysis really is  $D_{\phi}^{-1}$

$$\text{wavy line with a dot}^{-1} = \text{wavy line}^{-1} - \text{wavy line with a bubble}^{-1}$$

- Whith the added caveat that

$$\text{wavy line with a bubble} = \begin{pmatrix} \Pi_s & \Pi_{\mu} \\ \Pi_{\nu} & \Pi_{\mu\nu} \end{pmatrix}.$$

# Quark-Meson-Coupling model (QMC)

