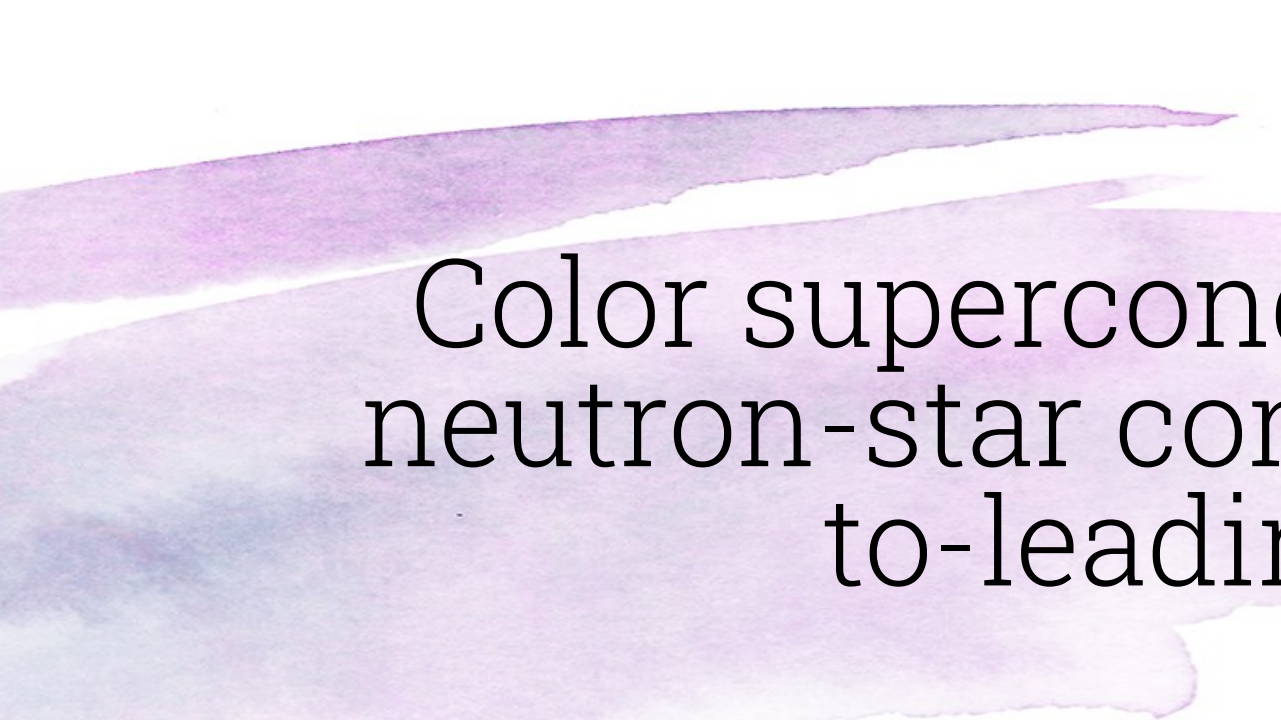




Color superconductivity under neutron-star conditions at next- to-leading order

A. Geißel, TG, J. Braun, PRL 135, 211901 (2025) [[2504.03834](#)]

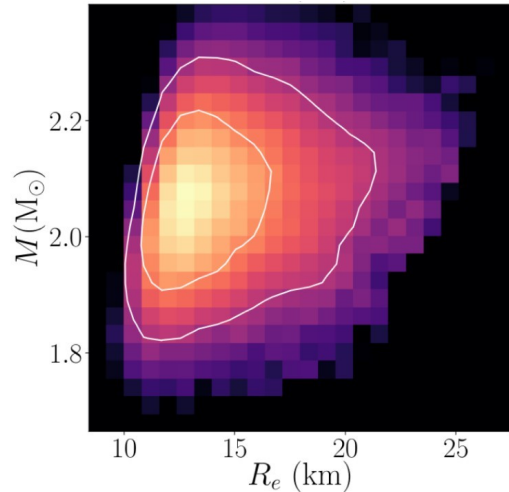
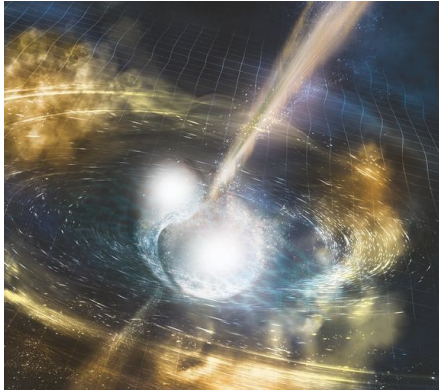
Tyler Gorda
The Ohio State University

SEWM - Helsinki

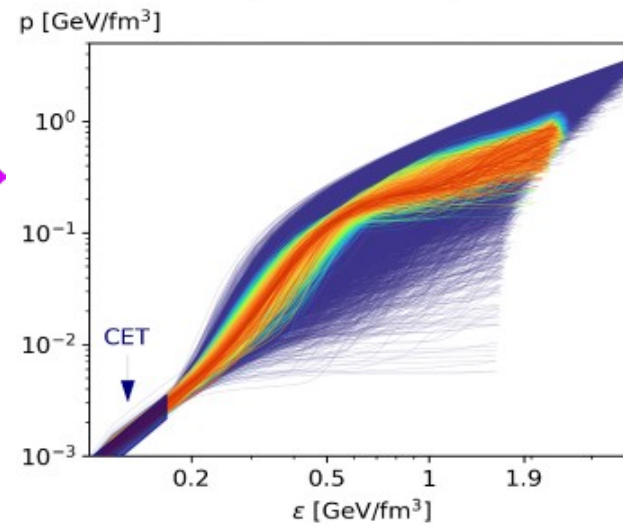
18 August 2026

Interdisciplinary approach to learn about dense matter

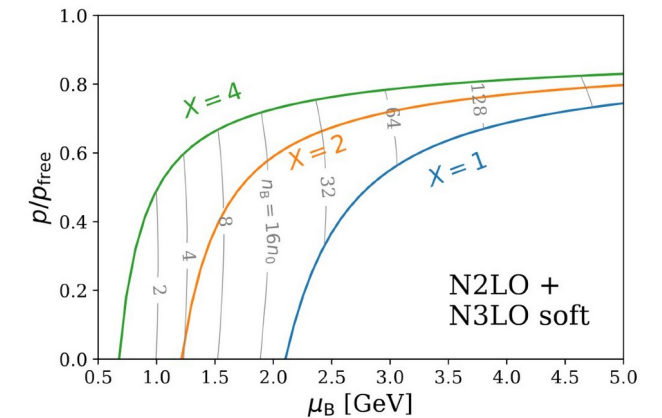
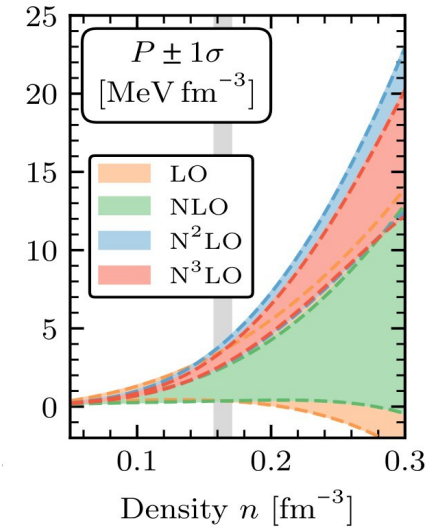
Astronomy + Astrophysics



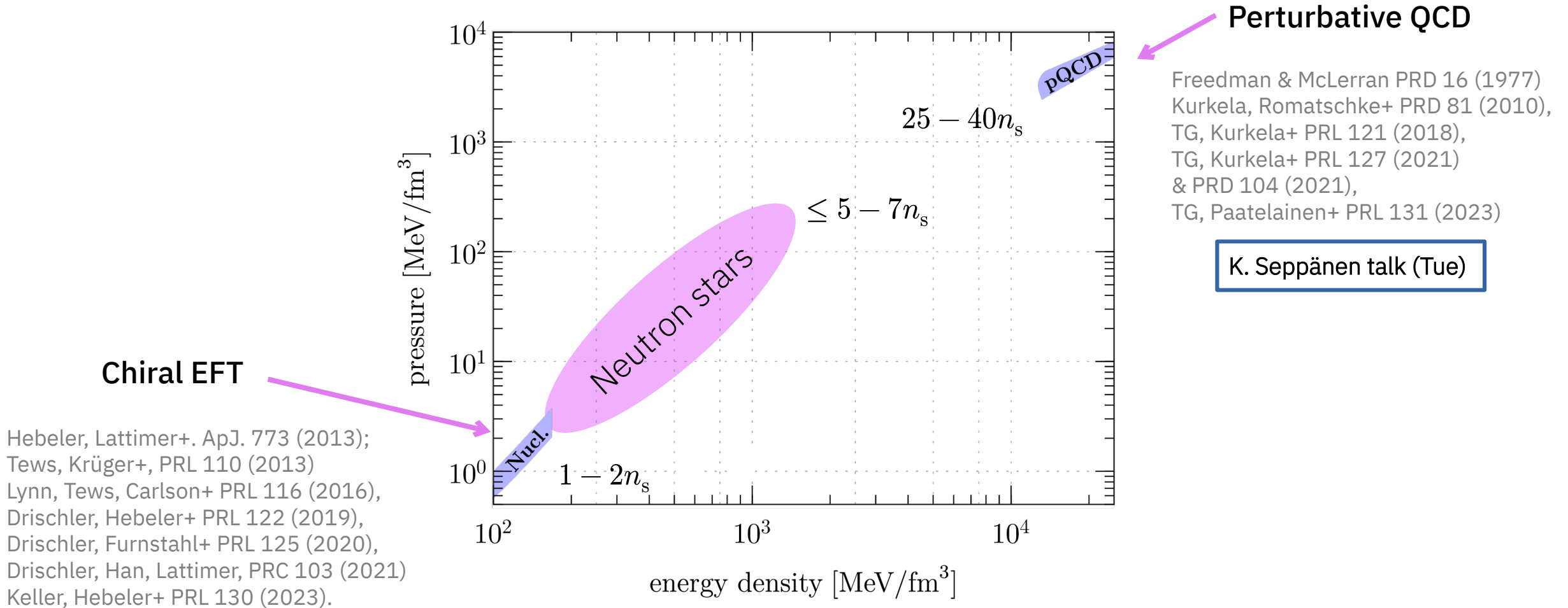
Equation-of-state (EoS) Inference & Properties of dense matter



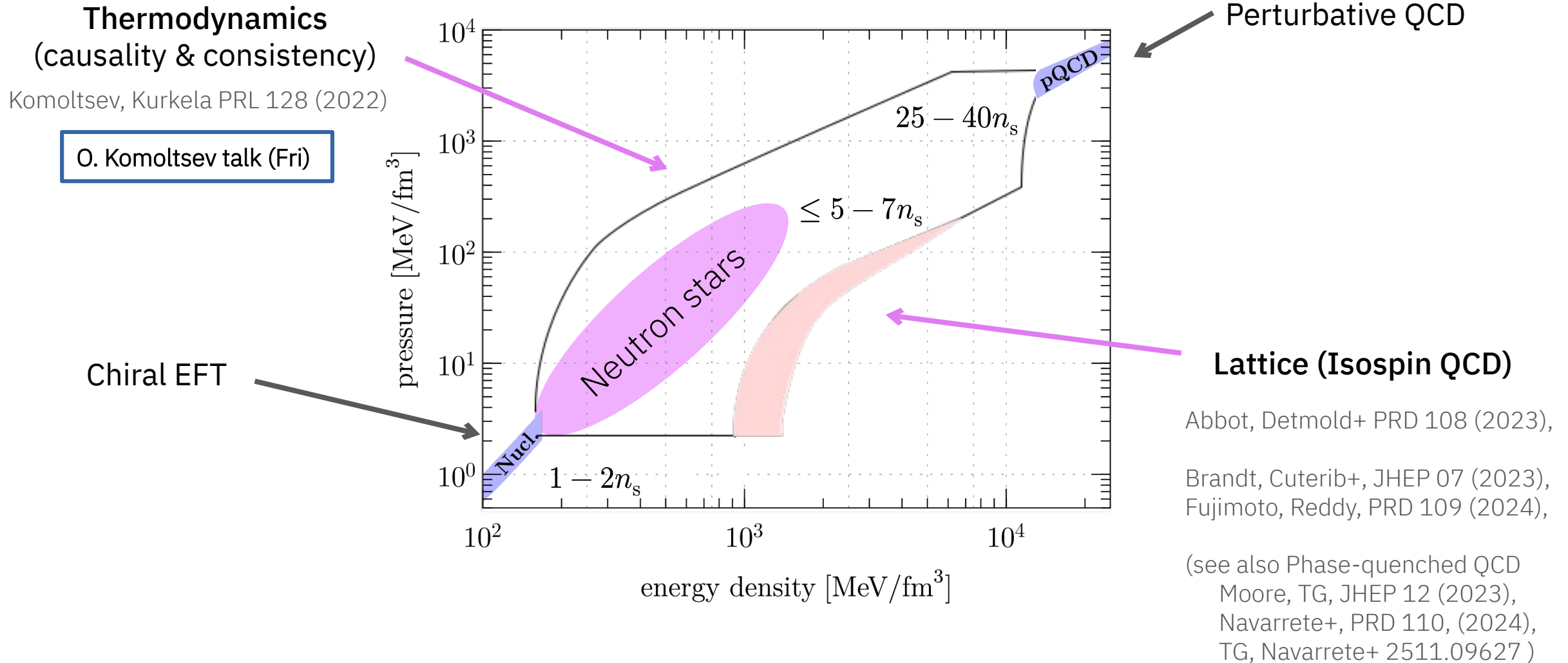
Nuclear + Particle theory



Different inputs constrain different parts of the EoS



Different inputs constrain different parts of the EoS



State-of-the-art unpaired pQCD EoS

- Describes **(nearly) massless quarks, gluons** interacting. Quarks are weakly interacting with $O(20\%)$ corrections when expanding results in the strong coupling α_s

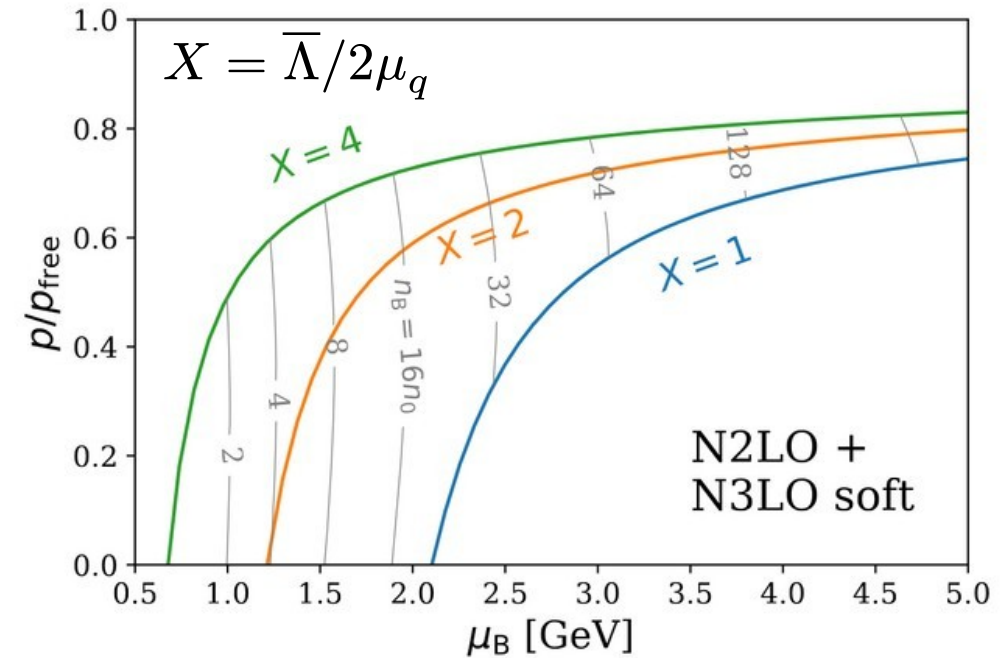
$$\frac{p}{p_{\text{free}}} = 1 + a_1 \alpha_s(\bar{\Lambda}) + a_2 \alpha_s^2(\bar{\Lambda}) + a_3 \alpha_s^3(\bar{\Lambda}) + \dots$$

free quark gas*

Freedman, McLerran, PRD 16 (1977)

TG, Kurkela+ PRL 121 (2018) – ongoing

- Calibrated by **collider data**
- State-of-the-art results converge well above $25\text{--}40n_s$



TG, Kurkela+ PRL 127 (2021), PRD 104 (2021)
TG, Paatelainen, Säppi, Seppänen, PRL 131 (2023)

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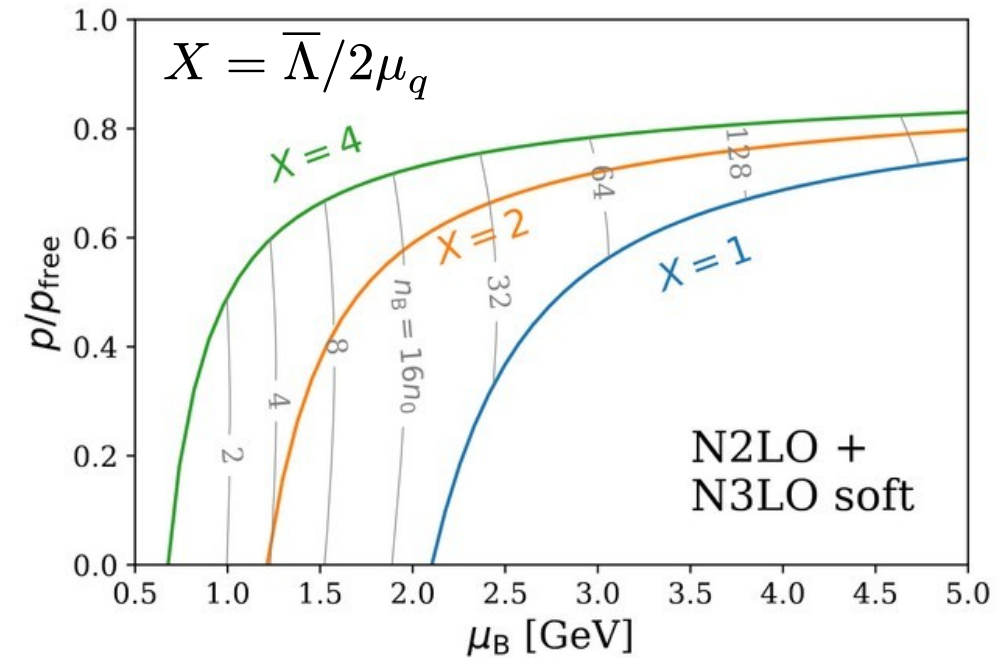
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**but this is not the whole story...*

Ground state of high-density quark matter is *paired*

- Gluonic attraction between quarks with opposite 3-momenta leads to **instability of the quark Fermi surface** and creates a nonzero diquark expectation value (condensate):

$$\Delta_{ab}^{ij} \sim \bar{\psi}_a^i \gamma^5 \mathcal{C} \bar{\psi}_b^j T \quad (\text{f = flavor, c = color, rest is Dirac})$$

Barrois, Nucl. Phys. B129 (1977),
Bailin, Love, Phys. Rep. 107 (1984),
Alford, Rajagopal, Wilczek, PLB 422 (1998),
Rapp, Schäfer, Shuryak+, PRL 81 (1998),
Alford+, Rev. Mod. Phys. 80, (2008)

- This leads to an **energy gap** in the quark excitations on Fermi surface
- Leads to additional pressure $\sim \Delta^2 \mu_B^2$; typically $\Delta \sim 100$ MeV
- **Complication**: flavor breaking from strange quark $\sim -m_s^2 \mu_B^2$, with $m_s \sim 100$ MeV (as long as m_s small enough, pairing still happens)

- **Goal**: Compute **NLO corrections** to QCD in the **paired vacuum**, assuming the coupling and the gap are both small (NLO correction to γ_1):

$$p = p_{\text{free}} \left[\gamma_0(\alpha_s, \bar{m}_s^2) + \gamma_1(\alpha_s, \bar{m}_s^2) \bar{\Delta}_{\text{CFL}}^2 + \dots \right], \quad \text{with } \bar{x} := x/(\mu_B/3)$$

Color-flavor-locked (CFL) matter under NS conditions

- CFL matter pattern **couples color and flavor** $\langle \psi_a^i T \gamma^5 \mathcal{C} \psi_b^i \rangle \sim \Delta_{\text{CFL}} \varepsilon_{abA} \varepsilon^{ijA}$
- NS matter neutral under
 - i. color
 - ii. flavor
 - iii. charge**locked & broken by explicit m_s**

- Means that general chemical potential is of the form

Alford & Rajagopal, JHEP 06, 031 (2002)

$$\mu_{f,c} = \mu_B/3 - \mu_e Q_{f,c} + \mu_3 T_{f,c}^3 + \mu_8 T_{f,c}^8$$

- **Important:** μ_e, μ_3, μ_8 parametrically small, so we can expand **in them** (they **vanish** when $m_s = 0$)

$$Q_{f,c} := \text{diag}_f \left(\frac{2}{3}, -\frac{1}{3}, -\frac{1}{3} \right) \otimes 1_c$$

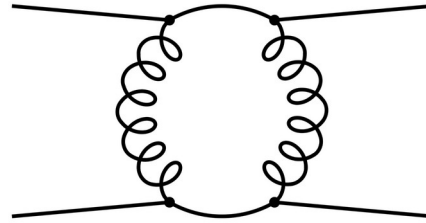
$$T_{f,c}^8 := 1_f \otimes \text{diag} \left(\frac{1}{3}, \frac{1}{3}, -\frac{2}{3} \right)$$

$$T_{f,c}^3 := 1_f \otimes \text{diag} \left(\frac{1}{2}, -\frac{1}{2}, 0 \right)$$

Deriving Effective Action in the gapped phase

Geißel, TG, Braun, PRD 110, 014034 (2024)

- Motivated by Renormalization-group (RG) approach, we consider **infinitesimal RG step** down in energy scale from the Euclidean QCD action S ,
- This generates an infinite set of new interactions, **including 4-quark interaction** from



- Leads to nonzero expectation value of the diquark operator $\Delta_{ab}^{ij} \sim \bar{\psi}_a^i \gamma^5 \mathcal{C} \bar{\psi}_b^j{}^T$
- Make diquark explicit by using exact **Hubbard-Stratonovich transformation** to exchange 4-quark interaction for quark-quark-diquark interaction + mass term $m^2 \Delta^2$
- In principle, m must be determined self-consistently. Here **simplify by taking Δ a constant background field**, whose form is fixed to results from the literature. i.e., $\partial\Gamma/\partial\Delta|_{\Delta=\Delta_{\text{gap}}} = 0$ fixes the mass m (Γ is the effective action we compute)

Computing the correction (1)

Geißel, TG, Braun, PRL 135, 211901 (2025)

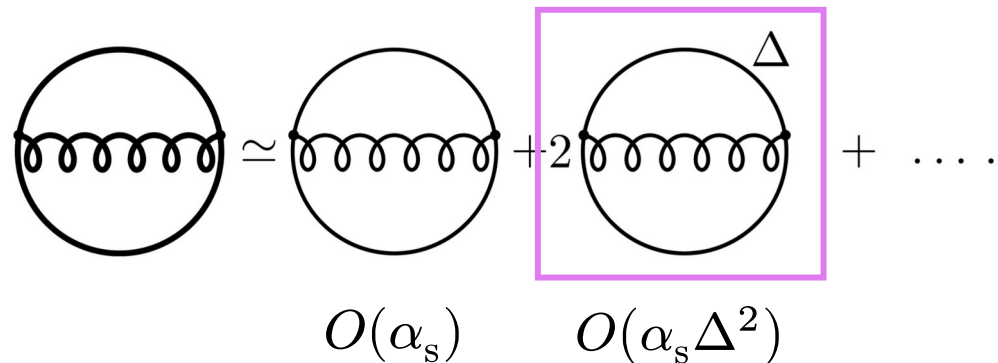
- To compute the term in the pressure, we only need the following modified action:

$$S_{\text{eff}} := \int_x \left\{ \bar{\psi} (\mathbf{i}\not{D} - \mathbf{i}\hat{\mu}_{f,c}\gamma_0 - \hat{M}_{f,c}) \psi + m^2 \Delta^2 + \frac{1}{2} (\psi^T \mathcal{C} \gamma_5 \Delta \epsilon_a^f \epsilon_a^c \psi + \text{h.c.}) \right\} + S_{\text{gf}} + S_{\text{gh}}$$

- Using S_{eff} , compute 2x2 propagators for (ψ, ψ^T) and the gauge field
- Expand each propagator in two parts to make gap part explicit, **power count in the gap**

$$\mathcal{P} = \mathcal{P}^0 + (\mathcal{P} - \mathcal{P}^0) := \mathcal{P}^0 + \mathcal{P}^\Delta$$

- Must compute corrections to following diagram(s)! Unfortunately must be done numerically...



Computing the correction (1)

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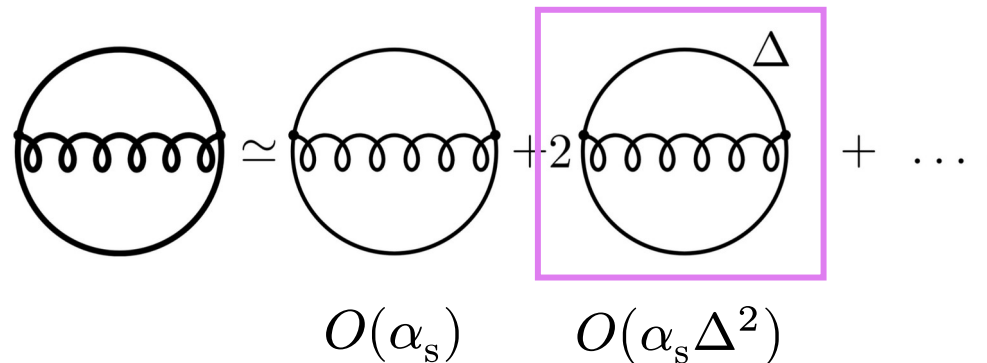
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- note this is not a **full** NLO result: double Δ insertions also have a NLO contribution at $O(\alpha_s \Delta^2)$

Computing the correction (2)

Geißel, TG, Braun, PRL 135, 211901 (2025)

- Numerically evaluate this diagram with a high cutoff Λ . **Vary large Λ and small Δ and use fitting function** to evaluate the effective action Γ for these different values

$$\frac{\Gamma_{\text{quark}}^{2\text{-loop}}}{V_4} = 128\pi\alpha_s\Delta^2\mu^2 \left[-2.44 + 0.0078 \ln \left(\frac{\Delta^2}{\mu^2} \right) \right]$$

- Minimize the effective action wrt Δ to fix m (introduce counterterm to remove divergences) and convert the diquark field Δ into the known gaps from the literature

$$\Delta_{\text{gap}} \propto \frac{\mu}{\alpha_s^{5/2}} \exp(-A/\sqrt{\alpha_s})$$

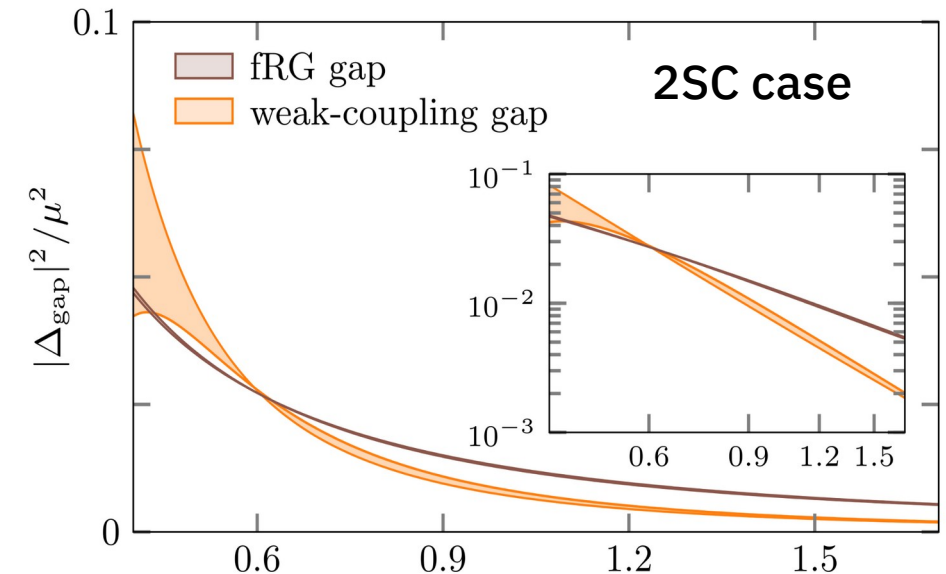
weak coupling
Son PRD 59 (1999)

$$\Delta_{\text{gap}} \sim \Lambda_{\text{QCD}} \exp[-B/(\alpha_s^2\mu^2)]$$

fRG
Braun, Schallmo, PRD
105 (2022)

Both approximately power laws

Geißel, TG, Braun, PRD 110, 014034 (2024)



Results under NS conditions (1)

Geißel, TG, Braun, PRL 135, 211901 (2025)

- NS conditions give $\bar{\mu}_e = \bar{\mu}_3 = 0, \bar{\mu}_8 = -\bar{m}_s^2/2$, which gives the pressure

Alford & Rajagopal, JHEP 06, 031 (2002)

$$p_{\text{CFL}}^{\text{NS}} = p_{\text{NQM}}^{\text{NS}} + p_{\text{free}} \left[\gamma_1(\alpha_s, \bar{m}_s^2) \bar{\Delta}_{\text{CFL}}^2 - \frac{\bar{m}_s^4}{4} \right]$$

- CFL favored when 2nd term is positive: $\bar{\Delta}_{\text{CFL}} > \frac{\bar{m}_s^2}{4} \underbrace{\left[1 - 5.11\alpha_s + \frac{\bar{m}_s^2}{6} \right]}_{< 1 \text{ when converged}}$
 $\Rightarrow$ CFL made **more stable by loop corrections**

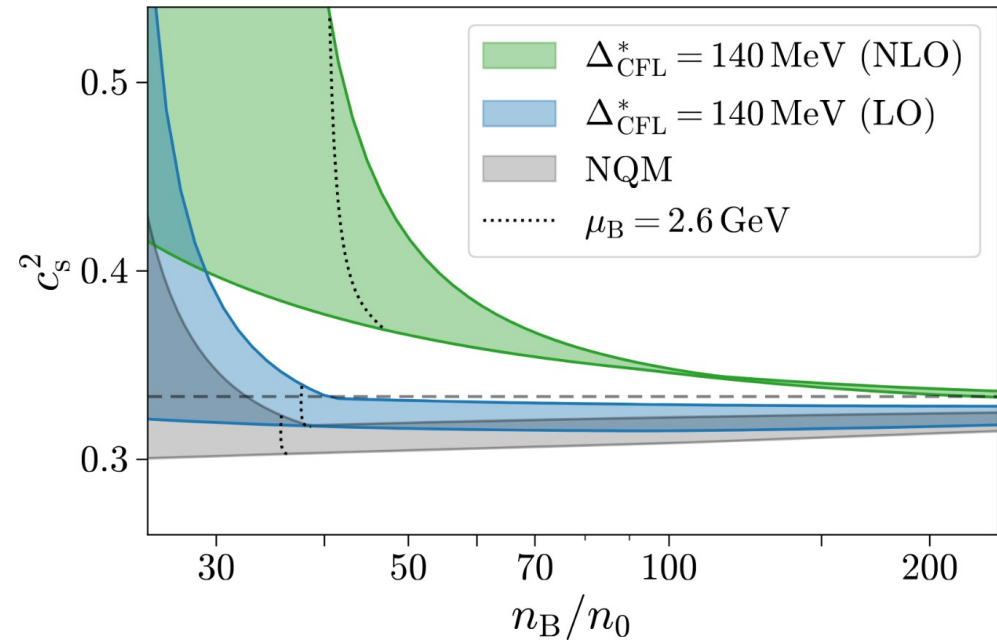
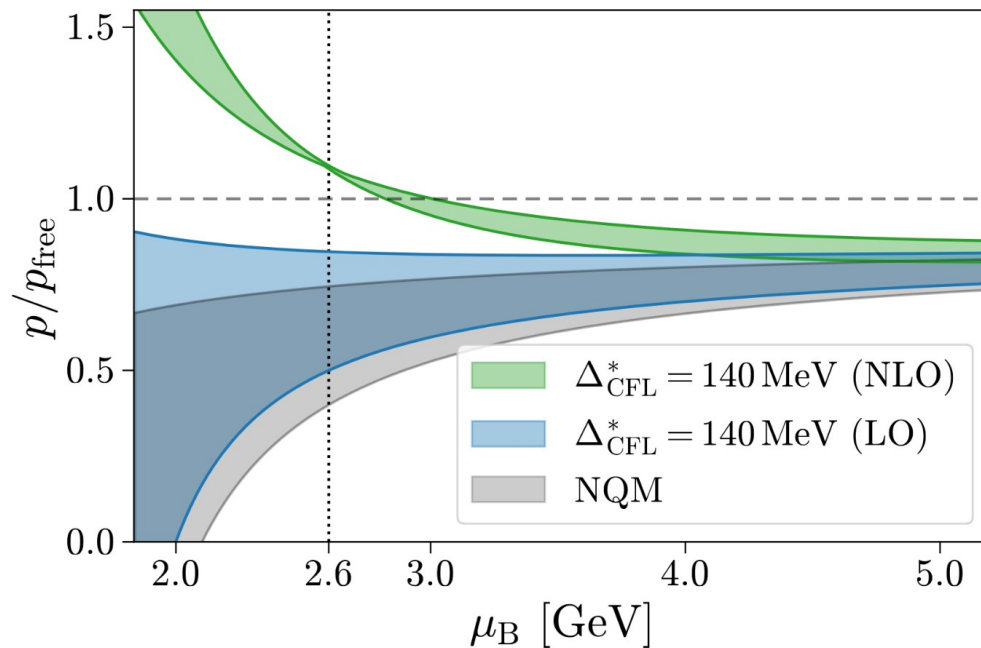
- Also find that **CFL more favored over 2SC phase at NLO**

Results under NS conditions (2)

Geißel, TG, Braun, PRL 135, 211901 (2025)

$$\gamma_1(\alpha_s, \overline{m}_s^2) = 4(1 - \overline{m}_s^2/3 + 10.2\alpha_s)$$

- Term shifts the pressure and the speed of sound **towards larger values**



- NLO shift is similar size as LO shift. **Worrisome???**

Constraining the gap with NS data

Geißel, TG, Braun, PRL 135, 211901 (2025)

- Shown previously that CEFT + Astro place upper bound on the gap at high densities

Reason:

Komoltsev, Kurkela PRL 128 (2022),

Kurkela, Rajagopal, Steinhorst, PRL 132, 262701 (2024)

$$p_H - p_L \leq \frac{n_H}{2} \left(\mu_H - \frac{\mu_L^2}{\mu_H} \right) \& \text{ EoS}_L \text{ constrained} \Rightarrow \text{upper bound on } \Delta_{\text{gap}}$$

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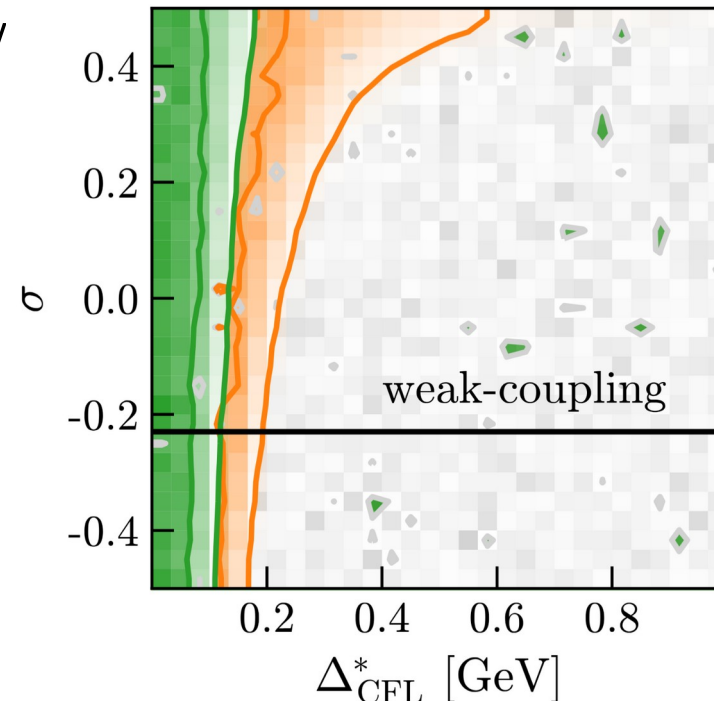
- Assuming power law scaling, with $\mu_B^* = 2.6 \text{ GeV}$

$$\Delta_{\text{gap}} = \Delta_{\text{CFL}}^* (\mu_B / \mu_B^*)^\sigma$$

$$\sigma \approx -0.23 \text{ (weak coupling)}$$

$$\sigma \approx +0.45 \text{ (fRG)}$$

- At 95% credence, $\Delta_{\text{CFL}}^* \lesssim 140 \text{ MeV}$ for the “symmetric” ensemble, after marginalizing over the unknown exponent (green)



Orange:

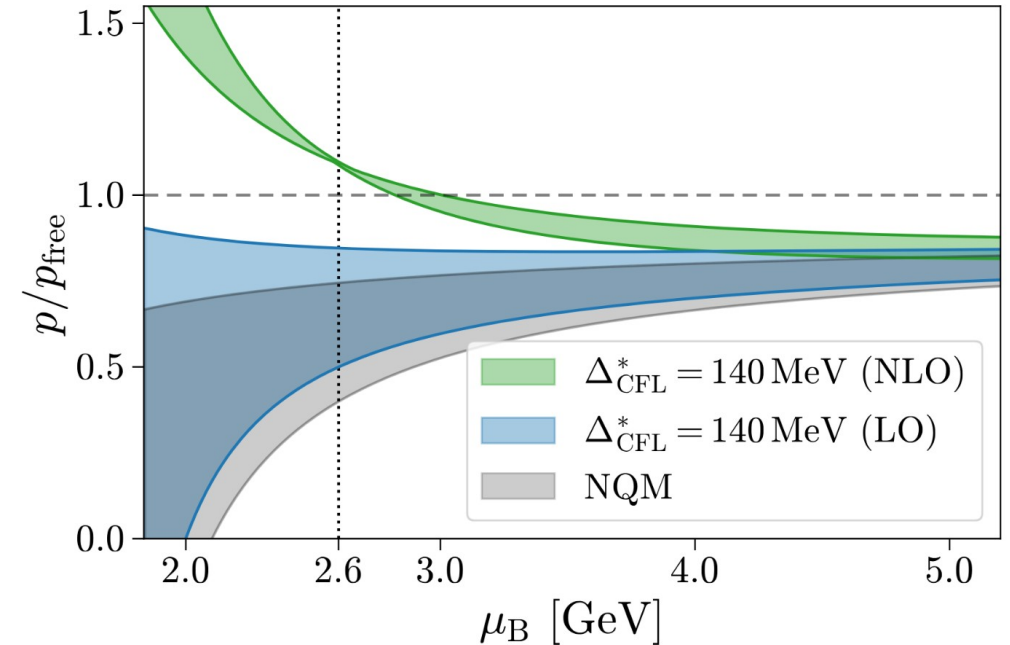
1. existence of $2M_\odot$,
2. causal connection to pQCD

Green:

1. NS prior to M_{TOV} ,
2. $c_s^2 < 2/3$ for extensions

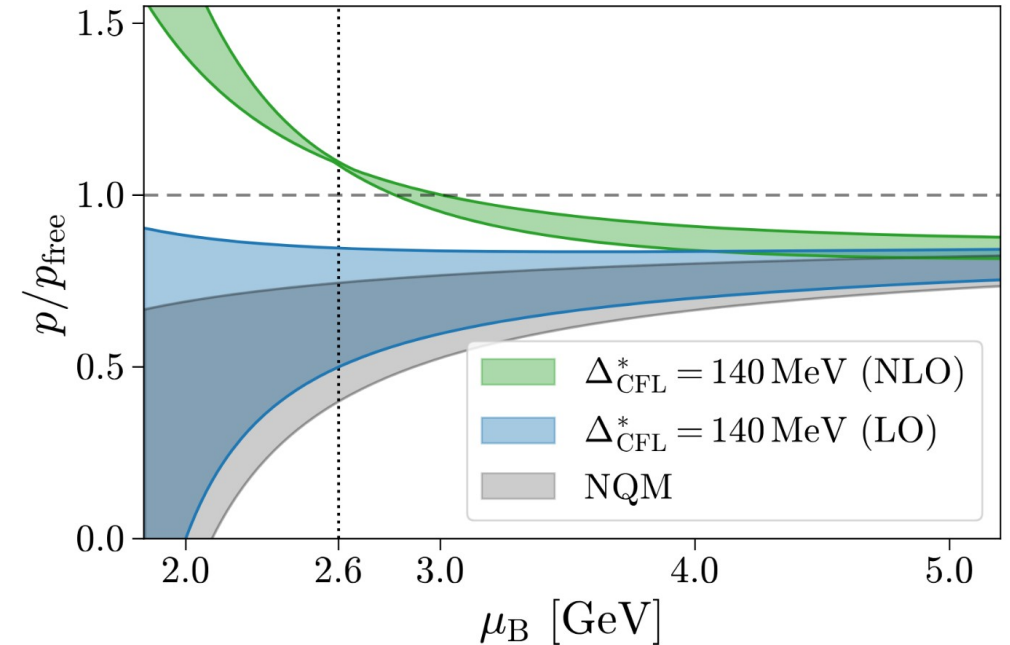
Summary and Outlook

- ✓ Computed NLO correction to pressure, speed of sound of paired quark matter under NS conditions at $O(\alpha_s \Delta^2)$
 - Assumed small, constant diquark background field
 - Derived effective action S_{eff} to capture pairing effects at this order
 - Result is numerical and small compared to free pressure, but large correction to LO correction
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Thanks for your attention!