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Revisiting the Rhoades-Ruffini Bound

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We revisit the derivation of the Rhoades-Ruffini bound [1] of $3.2 M_{\odot}$ for the upper limit of the maximum mass of neutron stars. We find that the assumption made there for the onset density of an ultimately stiff phase of high-density matter is not stringent. Relaxing this assumption and allowing for an onset of stiff non-nucleonic constant-speed-of-sound matter under neutron star constraints at the saturation density or below boosts the upper limit of the theoretically possible maximum mass to $4 M_{\odot}$ or higher. We provide a fit formula for the dependence of the maximum mass on the speed of sound and the onset density of the phase transition [2]. Our findings are relevant for discussing the nature of the so-called mass gap objects in the mass range between 2.5 and $5 M_{\odot}$.

[1] C.E. Rhoades and R. Ruffini, Maximum Mass of a Neutron Star, Phys. Rev. Lett. 32, 324 (1974) [2] D. Blaschke and A. Wójcik, Revisiting the Rhoades-Ruffini Bound, in preparation (2025)

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