

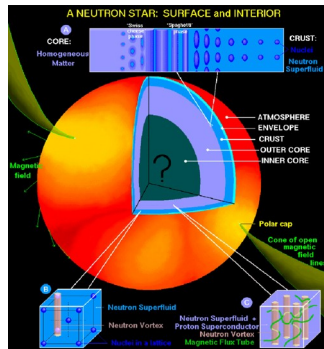
# Phenomenology and *ab initio* for heavy and light nuclei

Alex Gezerlis



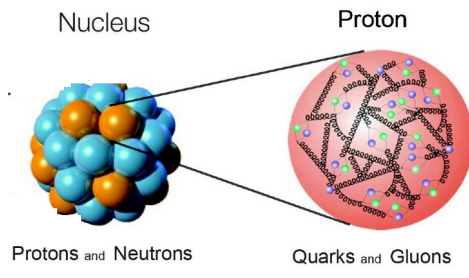
Theory Canada 17  
University of Regina  
June 7, 2025

# Outline

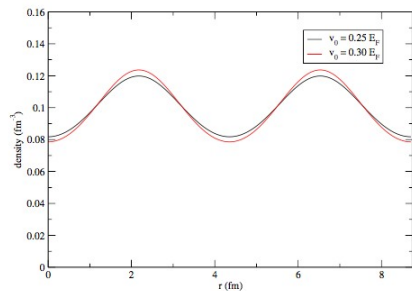


Credit: Dany Page

## Motivation

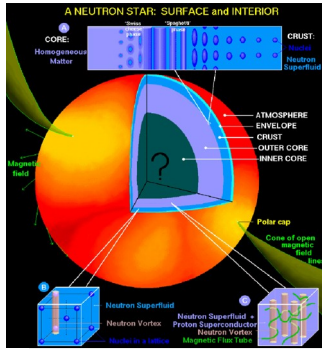


## Nuclear methods



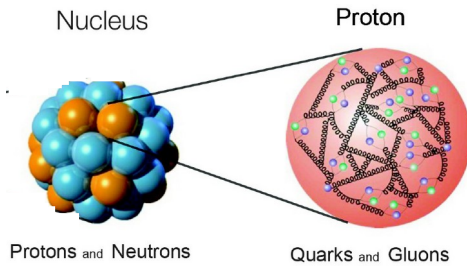
## Recent results

# Outline

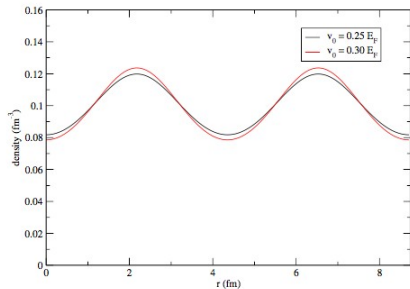


Credit: Dany Page

## Motivation



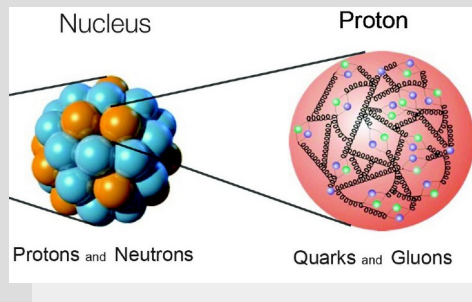
## Nuclear methods



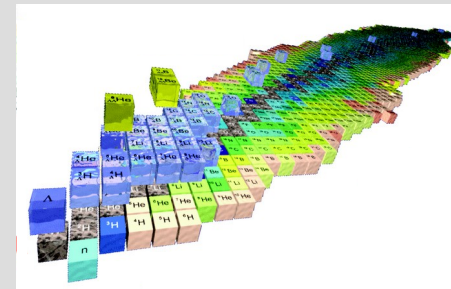
## Recent results

# Physical systems studied

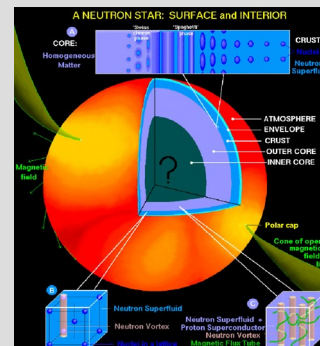
## Nuclear forces



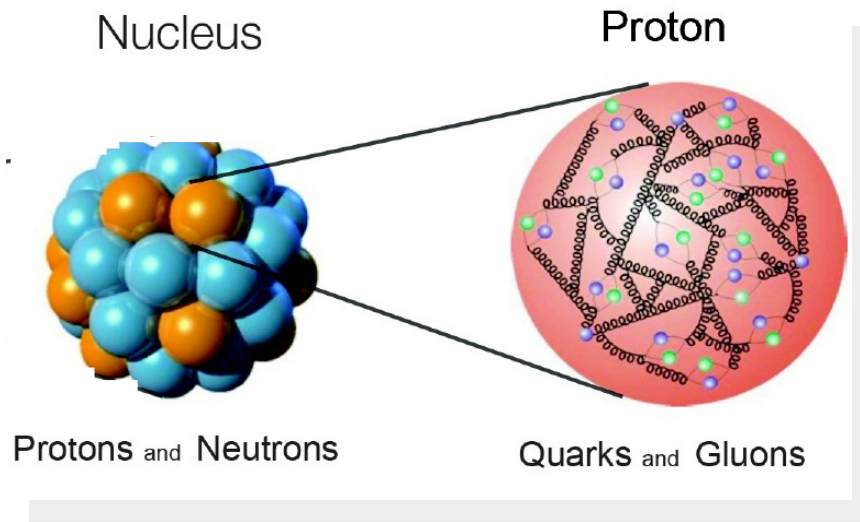
## Nuclear structure



## Nuclear astrophysics

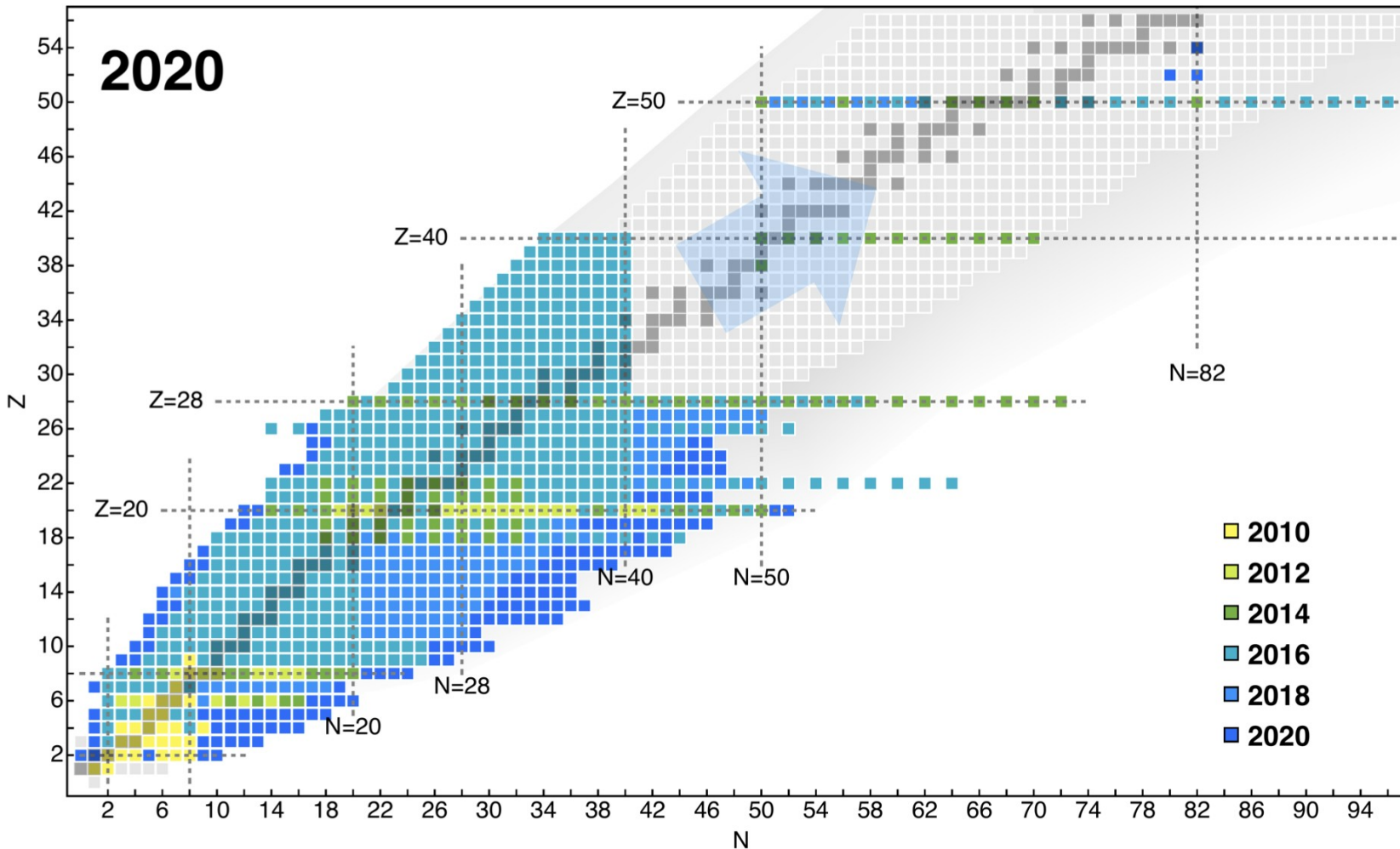


# Key system: few nucleons



- No unique nuclear potential
- Preferable to use combination of phenomenological (high-quality) and more modern (conceptually clean) approach
- Desirable to make contact with underlying level
- New era, where practitioners design interactions themselves

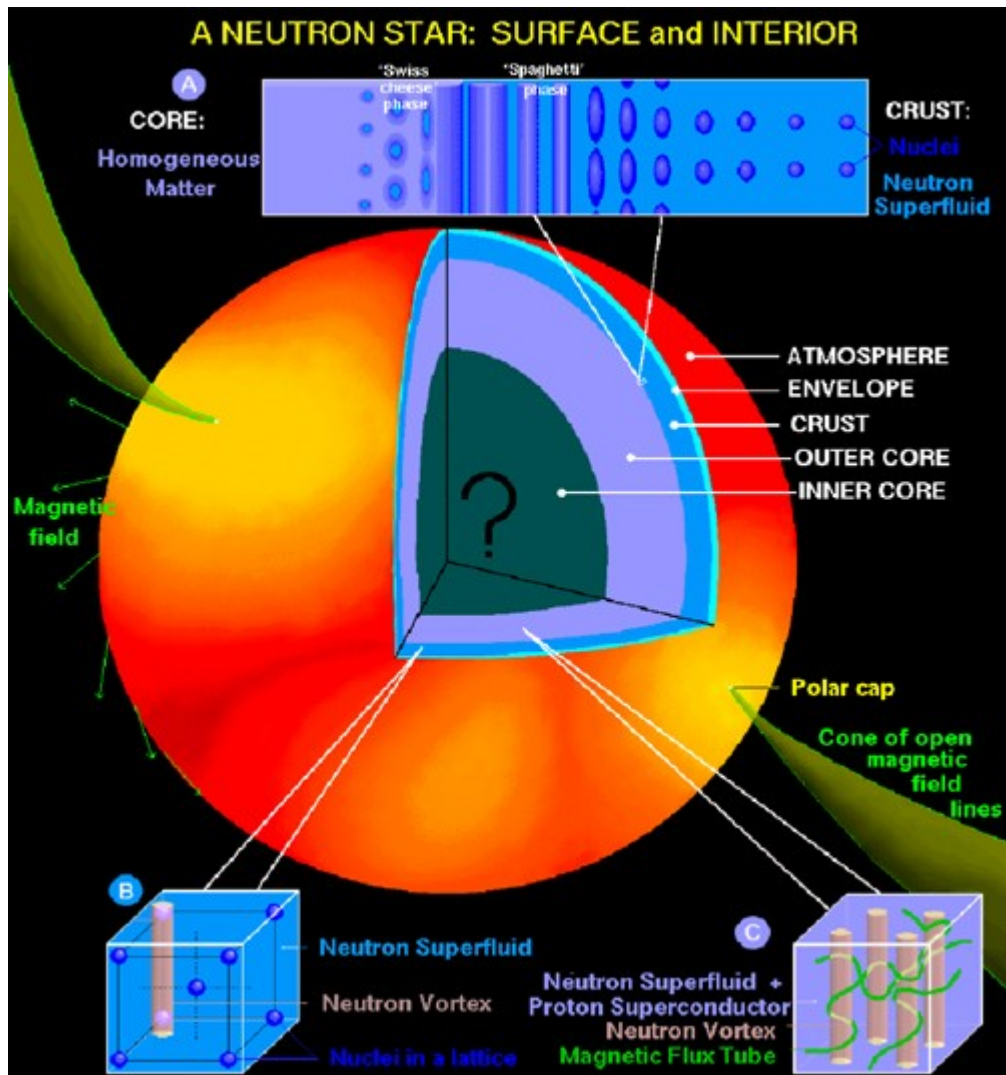
# Key system: nuclei



Credit: Heiko Hergert

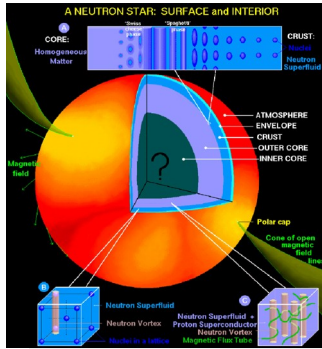
- Lots of recent progress
- Open-shell nuclei are the current frontier
- Goal is to study nuclei *from first principles* (when possible)

# Key system: neutron stars



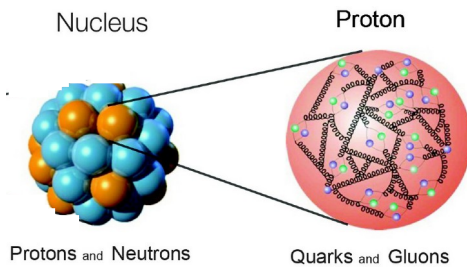
- Ultra-dense: 1.4 solar masses (or more) within a radius of 10 kilometres
- Terrestrial-like (outer layers) down to exotic (core) behaviour
- Observationally probed, i.e., not experimentally accessible
- Goal is to study neutron stars *from first principles* (when possible)

# Outline

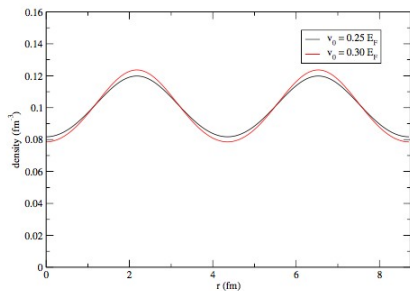


Credit: Dany Page

## Motivation

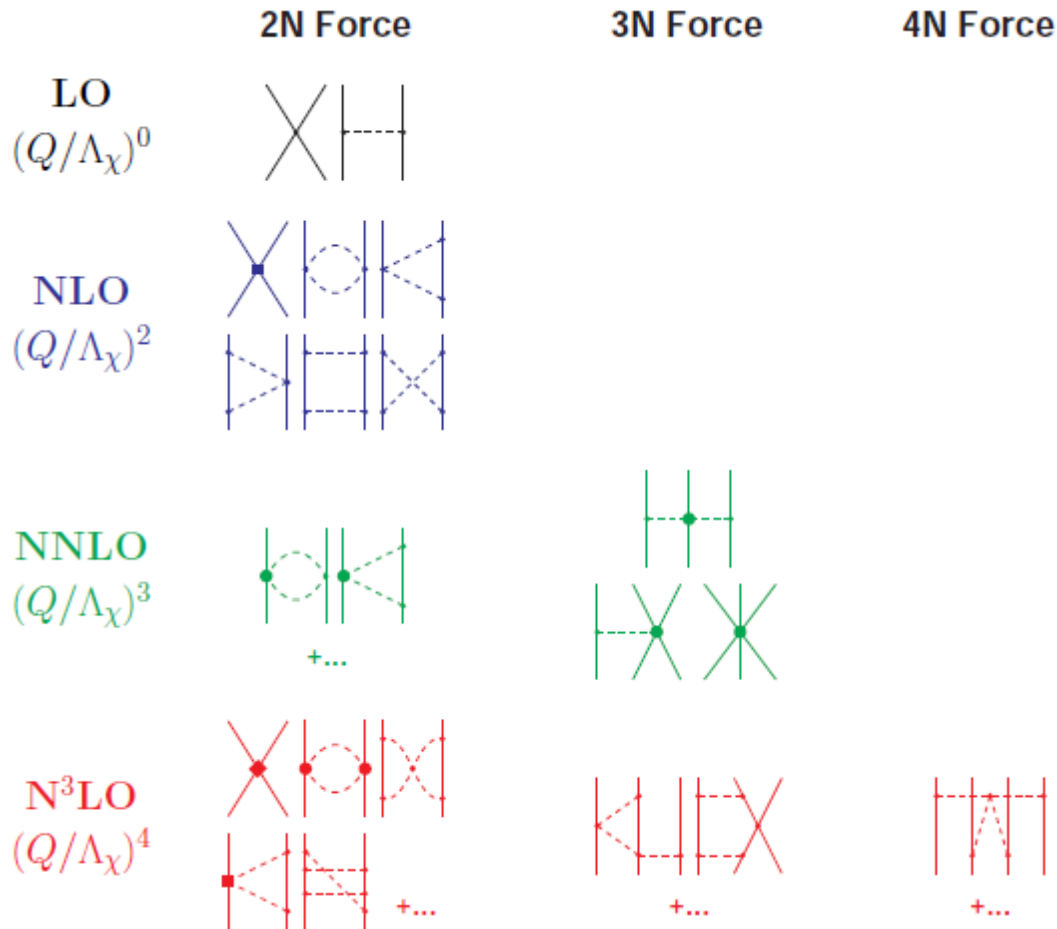


## Nuclear methods



## Recent results

# Nuclear interactions



- Attempts to connect with underlying theory (QCD)
- Low-momentum expansion
- Naturally emerging many-body forces
- Low-energy constants from experiment or lattice QCD
- Now available in non-local, local, or semi-local varieties
- Power counting's relation to renormalization actively investigated

**But even with the interaction in place,  
how do you solve the many-body problem?**

# Nuclear many-body problem

$$H\Psi = E\Psi$$

where

$$H = \sum_i K_i + \sum_{i<j} V_{ij} + \sum_{i<j<k} V_{ijk}$$

Wave function depends on coordinates, spin projections, and isospin projections, so we are faced with a large number of complex coupled second-order differential equations

# Nuclear many-body methods

- Phenomenological (fit to  $A$ -body experiment)
- Ab initio (fit to few-body experiment)

# Two complementary approaches

## Phenomenological (fit to A-body experiment)

- *Shell model*  
mainstay of nuclear physics, still very important
- *Hartree-Fock/Hartree-Fock-Bogoliubov (HF/HFB)*  
mean-field theory, a priori inapplicable, unreasonably effective
- *Energy-density functionals (EDF)*  
like mean-field but with wider applicability

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# Two complementary approaches

## Ab initio (fit to few-body experiment)

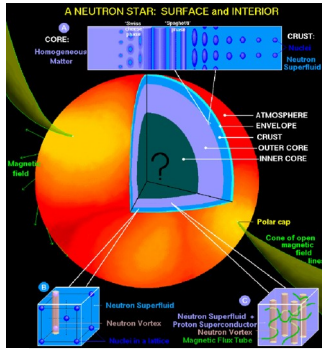
- *Quantum Monte Carlo (QMC)*  
stochastically solve the many-body problem “exactly”
- *Perturbative Theories (PT)*  
first few orders only
- *Resummation schemes (e.g. SCGF)*  
selected class of diagrams up to infinite order
- *Coupled cluster (CC)*  
generate  $n_p$ - $n_h$  excitations of a reference state
- *No-core shell model (NCSM)*  
fully ab initio, in contradistinction to traditional SM

# Two complementary approaches

## Ab initio (fit to few-body experiment)

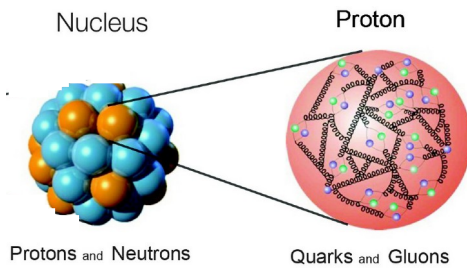
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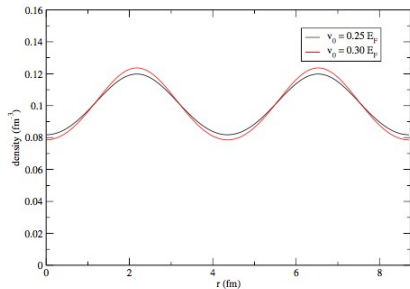


Credit: Dany Page

## Motivation



## Nuclear methods



## Recent results

## **Recent results**

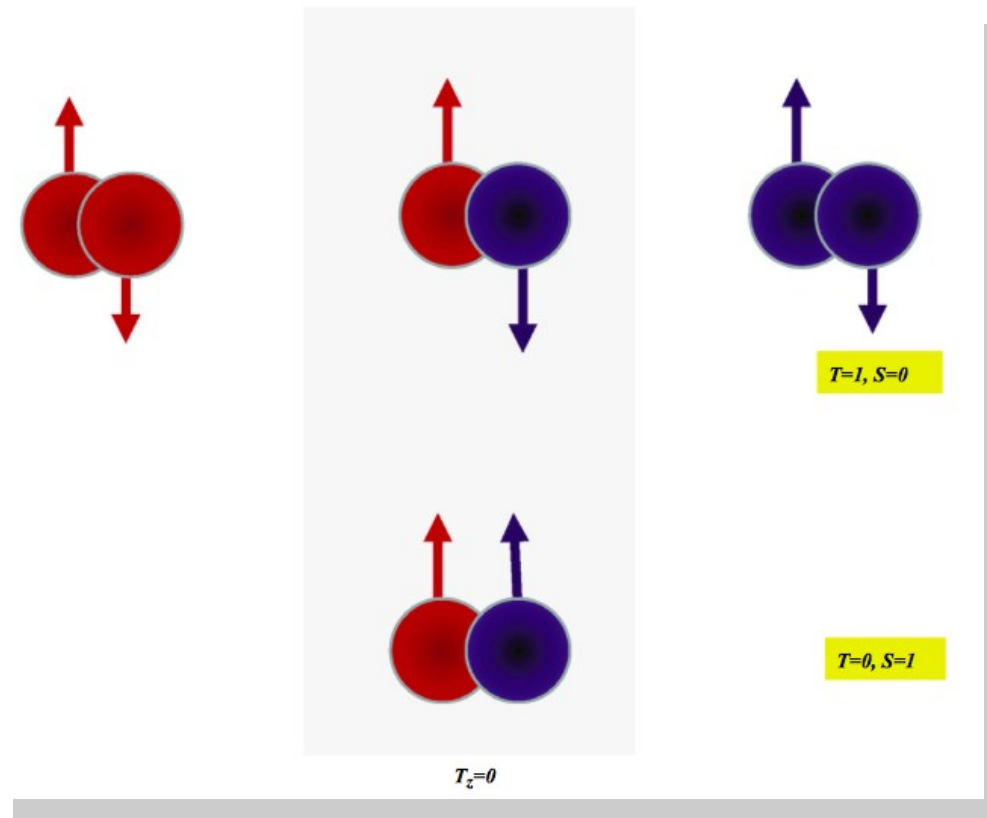
- **Mixed-spin pairing in heavy nuclei**
- **Neural-network wave functions for light nuclei**

## Mixed-spin pairing in heavy nuclei

G. Palkanoglou, M. Stuck, and A. Gezerlis, Phys. Rev. Lett. **134**, 032501 (2025)

G. Palkanoglou and A. Gezerlis, arXiv:2505.08879

# Types of pairs in nuclei



S. Frauendorf and A. O. Macchiavelli, Prog. Nucl. Part. Phys. **78**, 24 (2014)

# Deuteron-like pairing

**central  
“paradox”**

- $np$  interaction stronger than  $nn$  and  $pp$  interaction  
(*there is no bound dineutron in vacuum*)
- However, known nuclei exhibit  $nn$  and  $pp$  pairing.

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(*there is no bound dineutron in vacuum*)
  - However, known nuclei exhibit  $nn$  and  $pp$  pairing.

## Possible answer I:

- Isospin polarization discourages spin-triplet pairing:  
look at  $N=Z$  nuclei

A. L. Goodman, Phys. Rev. C **58**, R3051 (1998)

A. O. Macchiavelli *et al.*, Phys. Rev. C **61**, 041303(R) (2000)

R. Chasman, Phys. Lett. B **524**, 81 (2002)

## Possible answer II:

- Spin-orbit field interferes with spin-triplet pairing more:  
look at heavy nuclei

A. Poves and G. Martinez-Pinedo, Phys. Lett. B **430**, 203 (1998)

G. F. Bertsch and Y. L. Luo, Phys. Rev. C **81**, 064320 (2010)

S. Baroni, A. O. Macchiavelli, A. Schwenk, Phys. Rev. C **81**, 064308 (2010)

# Hamiltonian

$$\hat{H} = \sum_i \langle i | H_{sp} | j \rangle a_i^\dagger a_j + \sum_{i>j, k>l} \langle ij | v | kl \rangle a_i^\dagger a_j^\dagger a_l a_k$$

- $H_{sp}$  : kinetic + potential well + spin-orbit
- $\langle ij | v | kl \rangle$  contact pairing interaction in 6 channels

$$\langle ij | v | kl \rangle = \frac{1}{4} \langle ij | (3v_t + v_s + (v_t - v_s) \vec{\sigma} \cdot \vec{\sigma}') \delta^{(3)}(\vec{r} - \vec{r}') P_{L=0} | kl \rangle$$

where  $v_s$  and  $v_t$  are fit (and varied)

# Hartree-Fock-Bogoliubov theory

- Applying the Bogoliubov  $U$  and  $V$  we go to the quasiparticle representation

- The ordinary and anomalous densities are:

$$\rho = V^* V^T \quad \text{and} \quad \kappa = V^* U^T$$

- Hartree-Fock-Bogoliubov equations:

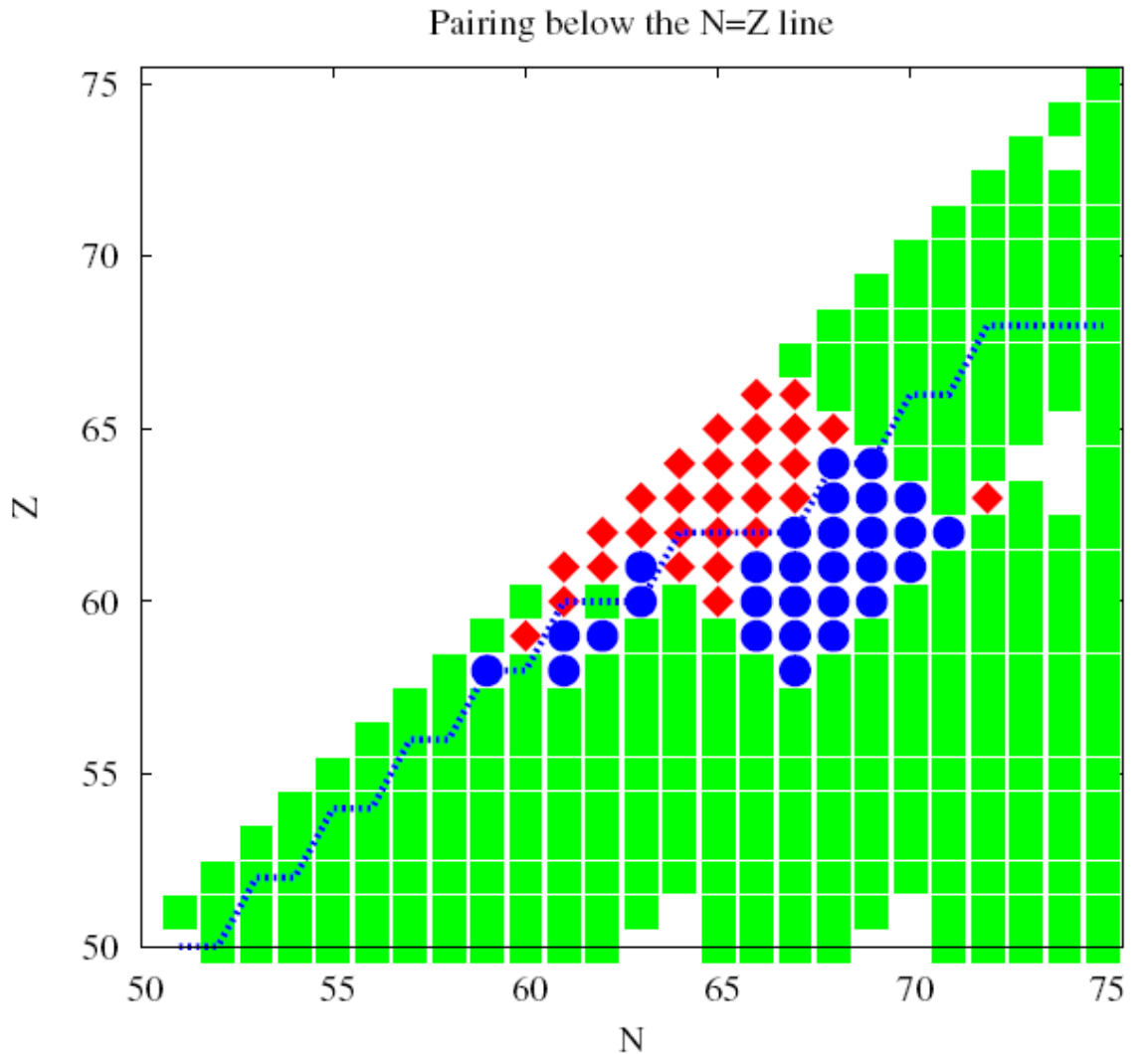
$$\begin{bmatrix} h' & \Delta \\ -\Delta^* & -h'^* \end{bmatrix} \begin{pmatrix} U_k \\ V_k \end{pmatrix} = \begin{pmatrix} U_k \\ V_k \end{pmatrix} E_k$$

where  $h = \varepsilon + \Gamma$  and the interaction is buried inside

$$\Gamma_{ij} = \sum_{kl} \bar{v}_{iljk} \rho_{kl}$$

$$\Delta_{ij} = \frac{1}{2} \sum_{kl} \bar{v}_{ijkl} \kappa_{kl}$$

# Pairing in heavy nuclei



## Correlation energies

- Blue line: proton drip
- Green: spin-singlet
- Red: spin-triplet
- Blue: mixed-spin
- Spin-triplet pairing persists off N=Z line
- Mixed-spin pairing appears to be energetically stable (note: no deformation)

# What about deformation?

To go beyond, take a deformed Woods-Saxon well:

$$\psi_{nl_z}(\rho, \phi, z) = \frac{e^{l_z \phi}}{\sqrt{2\pi}} \frac{u(\rho, z)}{\sqrt{\rho}}$$
$$-\frac{\hbar^2}{2m} \left( \frac{\partial^2}{\partial \rho^2} + \frac{\partial^2}{\partial z^2} \right) u + \left[ V_{\text{WS}}(\rho, z) + \frac{\hbar^2}{2m} \frac{l_z}{\rho^2} - \frac{\hbar^2}{2m} \frac{1}{4\rho^2} \right] u = E_{nl_z} u$$

# What about deformation?

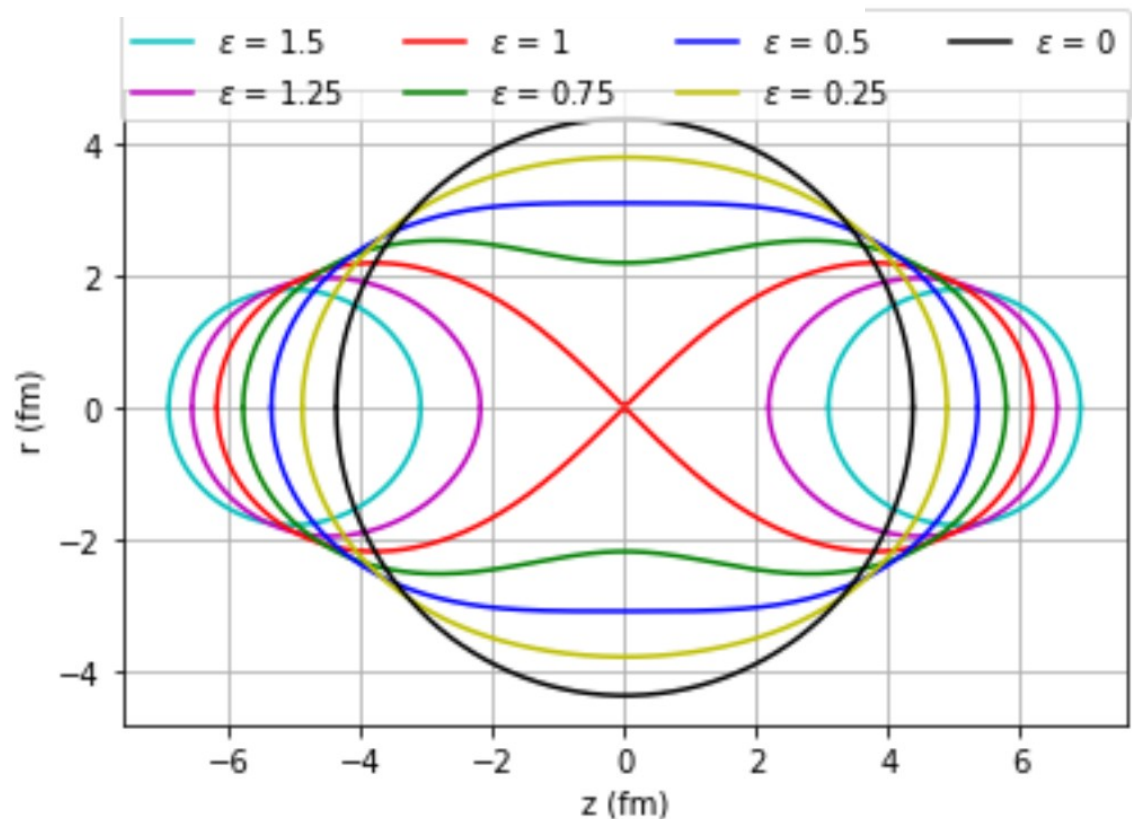
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$$V_{\text{WS}}(\rho, z) = \frac{V_0}{1 + e^{d(\rho, z)/a}}$$

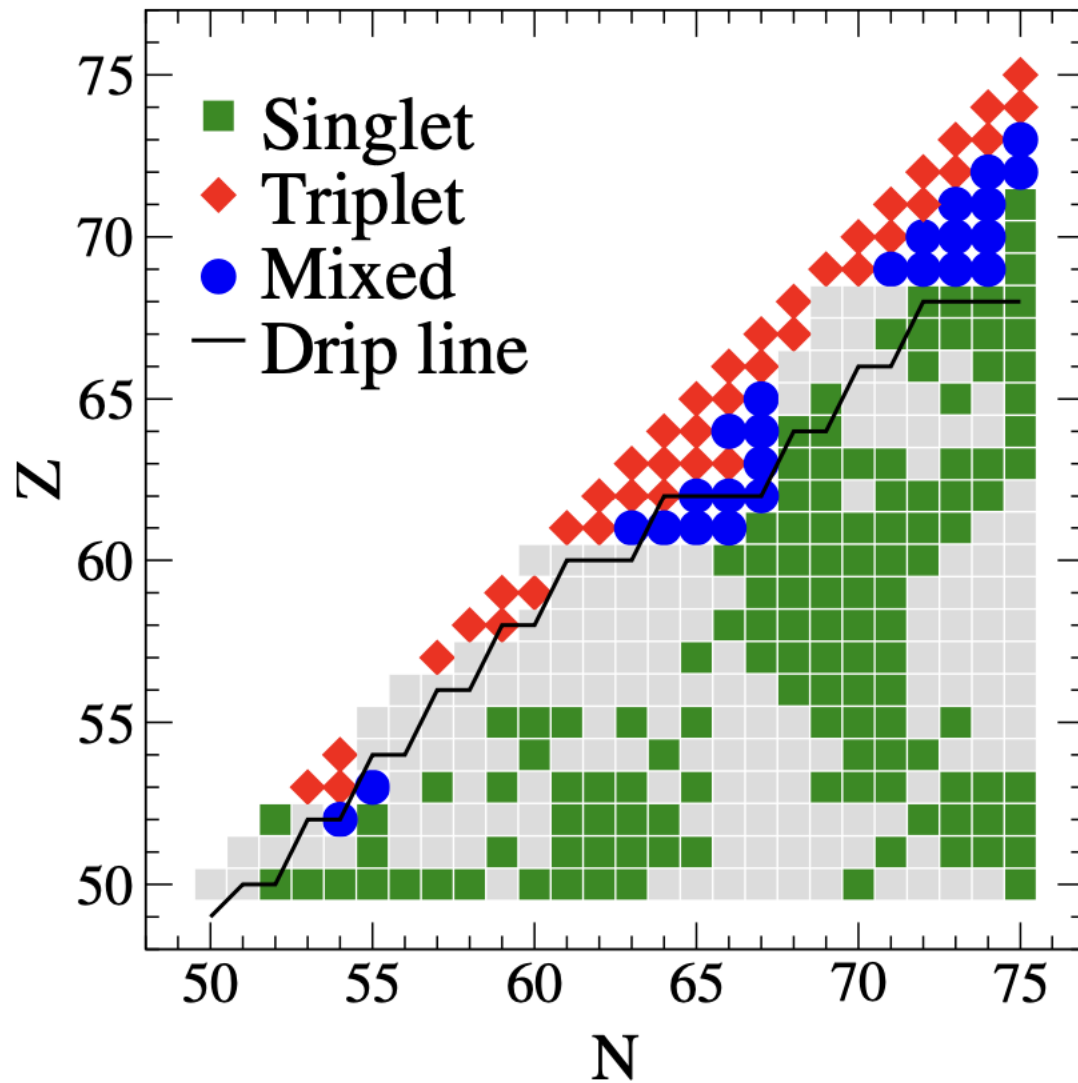
Surface is a  
Cassini oval:



# Deformation and exotic pairing

Deformation impacts  
but does not kill  
mixed-spin pairing:

- Quadrupole deformation is main factor
- For  $N=Z$  deformation actually *enhances* spin-triplet pairing

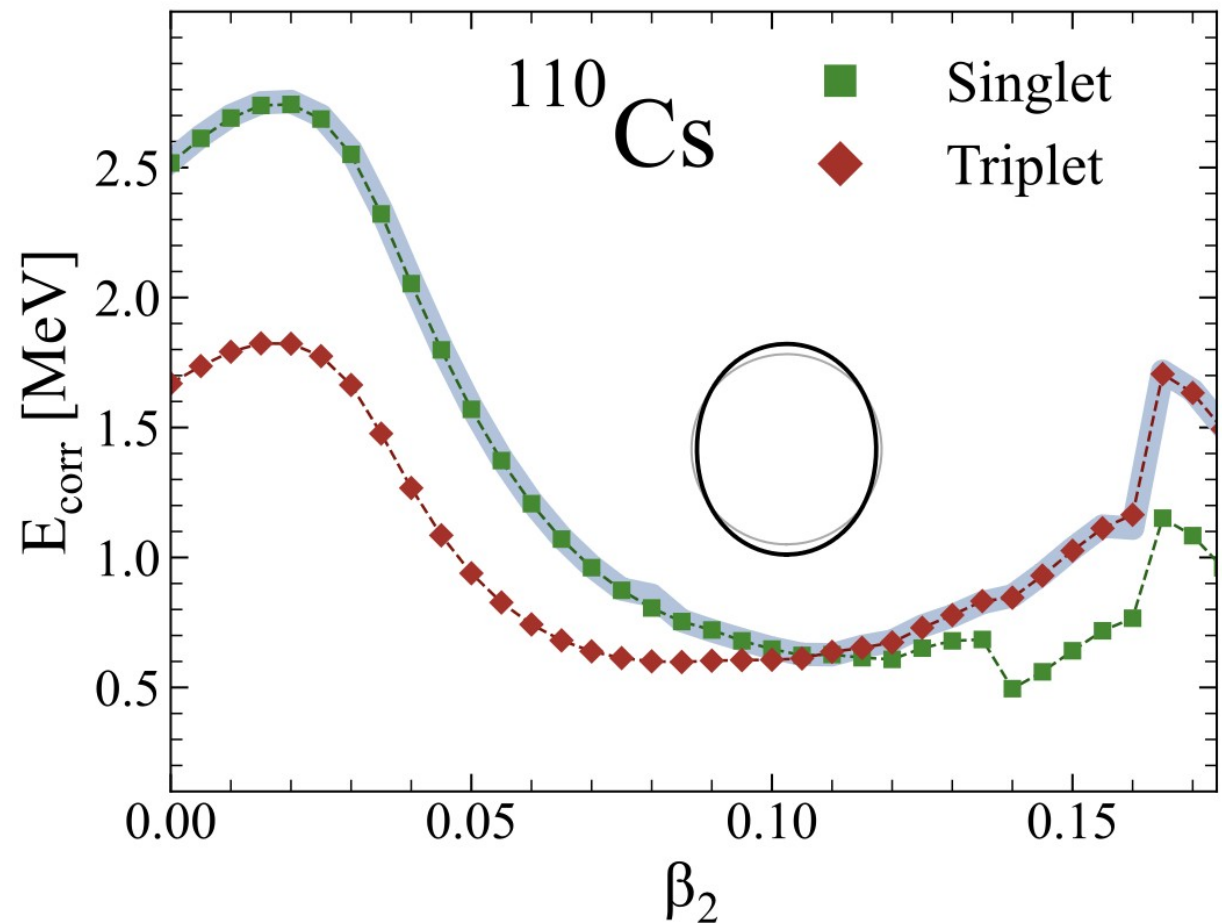


# Deformation and exotic pairing

Vary quadrupole deformation and compare constrained with unconstrained runs.

Deformation impacts the surface and therefore the effect of spin-orbit.

But spin-triplet pairs move to the interior of the nucleus and are unimpacted.

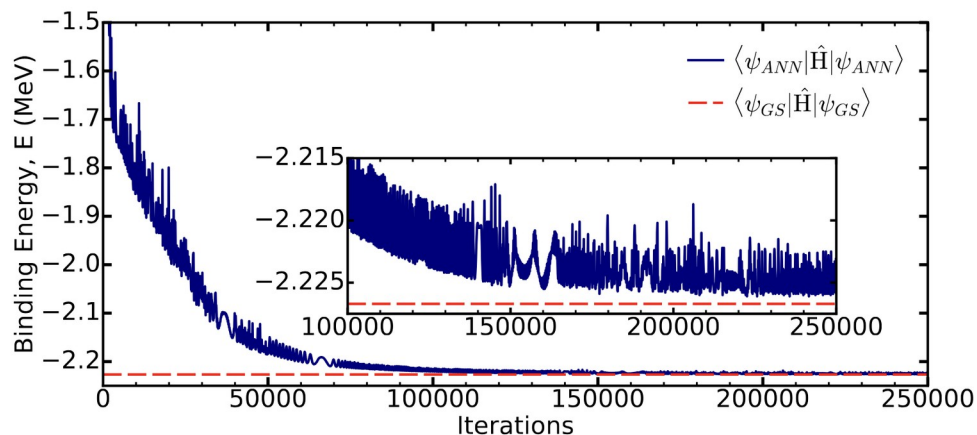


# Neural-network wave functions for light nuclei

P. Weng, A. Gezerlis, and J. Holt, arXiv:2505.11442

# Earlier work

## Deuteron



J. W. T. Keeble and A. Rios,  
Phys. Lett. B **809**, 135743 (2020)

## Light nuclei

	$\Lambda$	VMC-ANN	VMC-JS	GFMC	GFMC <sub>c</sub>
${}^2\text{H}$	$4 \text{ fm}^{-1}$	-2.224(1)	-2.223(1)	-2.224(1)	-
	$6 \text{ fm}^{-1}$	-2.224(4)	-2.220(1)	-2.225(1)	-
${}^3\text{H}$	$4 \text{ fm}^{-1}$	-8.26(1)	-7.80(1)	-8.38(2)	-7.82(1)
	$6 \text{ fm}^{-1}$	-8.27(1)	-7.74(1)	-8.38(2)	-7.81(1)
${}^4\text{He}$	$4 \text{ fm}^{-1}$	-23.30(2)	-22.54(1)	-23.62(3)	-22.77(2)
	$6 \text{ fm}^{-1}$	-24.47(3)	-23.44(2)	-25.06(3)	-24.10(2)

C. Adams, G. Carleo, A. Lovato, N. Rocco,  
Phys. Rev. Lett. **127**, 022502 (2021)

N.B. Limited to pionless Hamiltonian

# Neural networks for light nuclei

Spin-isospin correlations

$$|\psi\rangle = \mathcal{S} \prod_{i < j} \left( 1 + \sum_{\mathbf{x}} u_{ij}^{(\mathbf{x})} \hat{O}_{ij}^{(\mathbf{x})} \right) f_{ij}^{(c)} |\Phi\rangle$$

$$|\psi\rangle \rightarrow \left( 1 + \sum_{i < j < k} \sum_{\text{cyc}} \sum_{\mathbf{x}} \epsilon^{(\mathbf{x})} \hat{V}_{ijk}^{(\mathbf{x})} \right) |\psi\rangle$$

for N2LO chiral Hamiltonian

P. Weng, A. Gezerlis, and J. Holt, arXiv:2505.11442

# Neural networks for light nuclei

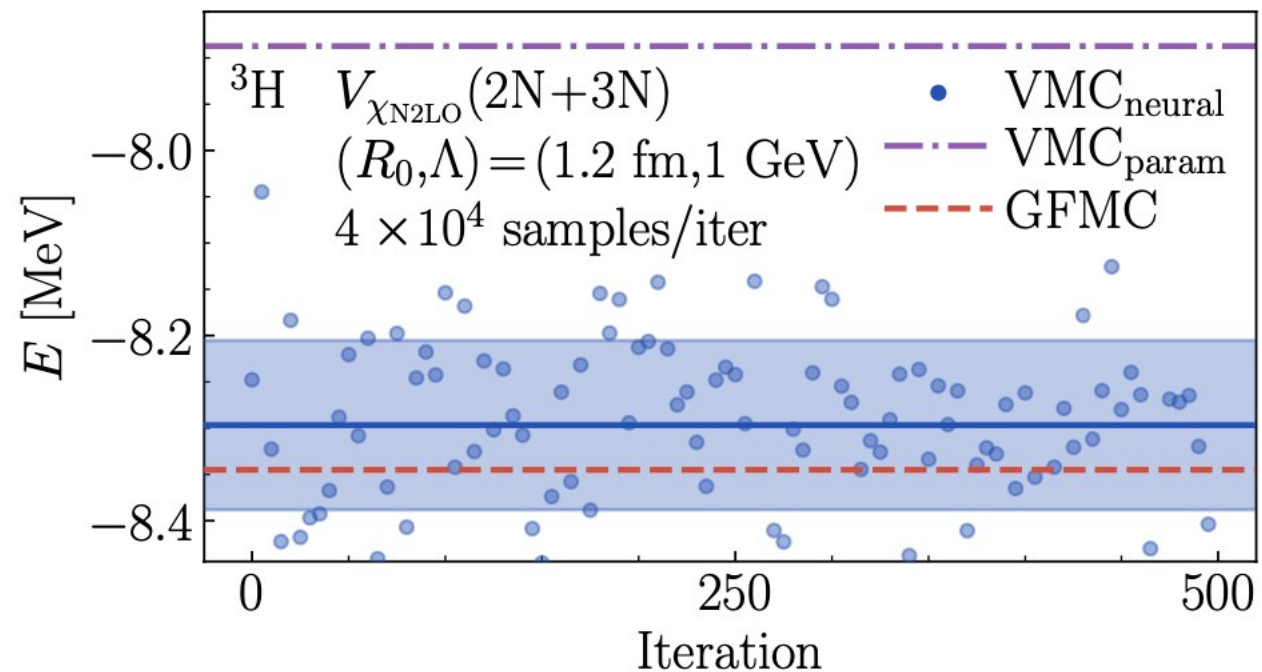
Nearly reproduces GFMC results  
already at the VMC level

$E = E_k + V_{\chi\text{N}^2\text{LO}}(2\text{N})$				
	$R_0$ [fm]	$E_{\text{neural}}$ [MeV]	$E_{\text{GFMC}}$ [MeV]	$ \Delta E / E_{\text{GFMC}} $
${}^3\text{H}$	1.0	$-7.338 \pm 0.008$	$-7.554 \pm 0.007$	2.9%
	1.1	$-7.500 \pm 0.006$	$-7.625 \pm 0.005$	1.6%
	1.2	$-7.678 \pm 0.005$	$-7.740 \pm 0.005$	0.8%
${}^2\text{H}$	1.0	$-2.217 \pm 0.005$	$-2.21 \pm 0.02$	0.3%
	1.2	$-2.212 \pm 0.004$	$-2.20 \pm 0.03$	0.5%

P. Weng, A. Gezerlis, and J. Holt, arXiv:2505.11442

# Neural networks for light nuclei

Dramatic improvement over  
standard/parametric VMC  
employed before, e.g., AFDMC



# Conclusions

- Rich connections between physics of nuclei, cold atoms, and compact stars
- Exciting time in terms of interplay between nuclear interactions and many-body approaches
- Ab initio and phenomenology are mutually beneficial
- Mixed-spin pairing is present even after you introduce deformation

# Acknowledgments

## Collaborators

### Guelph

- Michael Stuck

### Texas A&M

- Jeremy Holt
- Pengsheng Weng

### TRIUMF

- Georgios Palkanoglou

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