

# Infinite Heat Order in $3+1$ dimensions

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New directions in the large-charge expansion - Giulia Muco, 22.06.26

# In this talk:

- Some background
- Symmetry non-restoration in QFT: a (3+1)-d example?
- The finite-N story

The question we are dealing with:

## Does spontaneous symmetry breaking survive at arbitrarily high temperatures?

- Thermodynamics naively suggests **no**:

At finite  $T$  we minimise

$$F = E - ST$$

= > we want to maximise  $S$ , usually achieved with disordered (symmetric) phases

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- More concretely, a “cluster decomposition” property can be proven:

$$\exists \beta_* : \langle A(x)B(y) \rangle_\beta \xrightarrow{|x-y| \gg \beta} \langle A(x) \rangle_\beta \langle B(y) \rangle_\beta, \quad \beta < \beta_*$$

## ... But:

There are examples (experimentally observed!) in which the systems acquire order when heated (for a short range of T):

- Rochelle salt becomes a little more ordered than at lower temperatures
- Liquid  $\text{He}_3$  solidifies when heated

# How?

... going back to:

$$\exists \beta_* : \langle A(x)B(y) \rangle_\beta \xrightarrow{|x-y| \gg \beta} \langle A(x) \rangle_\beta \langle B(y) \rangle_\beta, \quad \beta < \beta_*$$

This statement is based on assumptions!

e.g.:

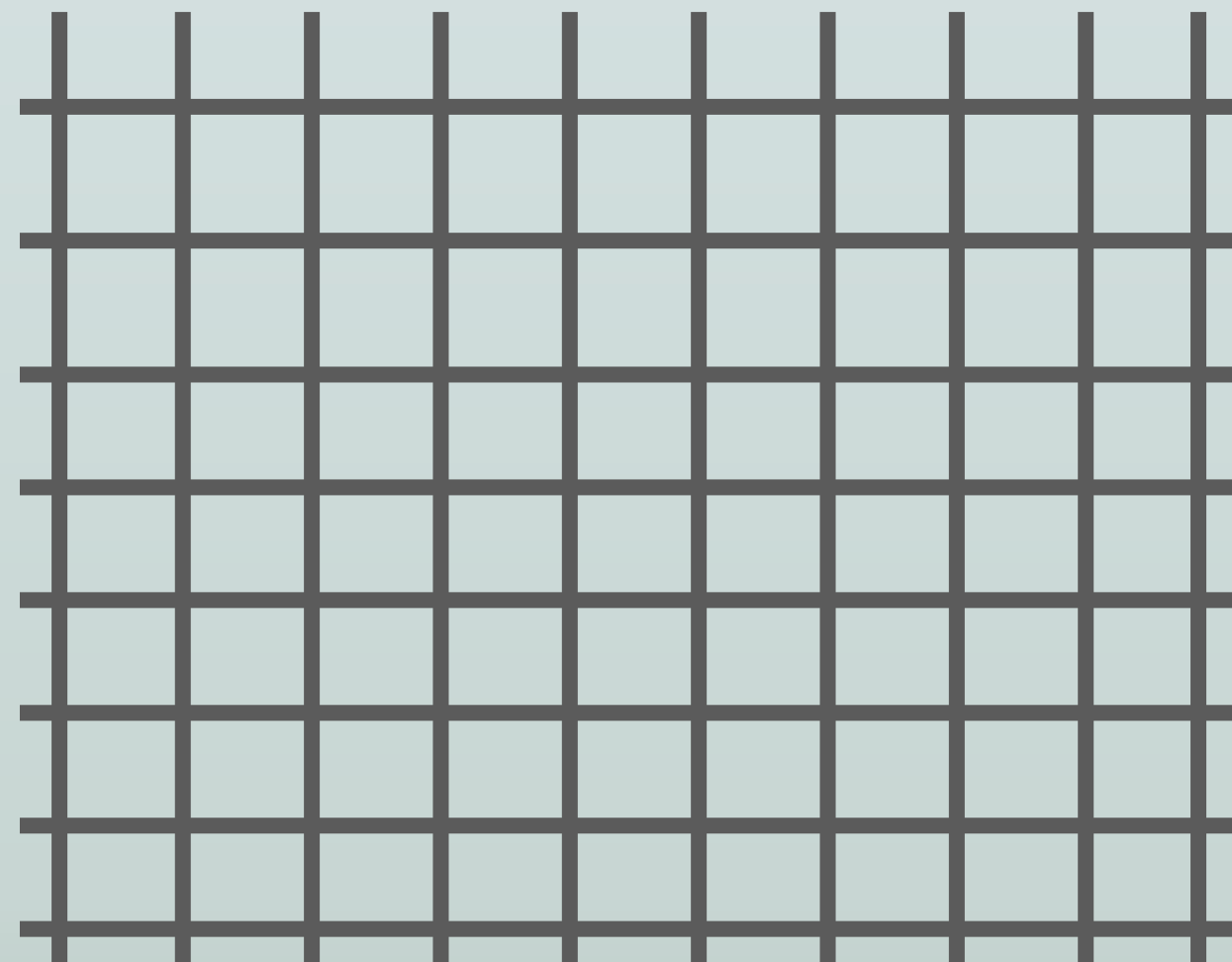
1.  $\exists$  a uniform probability distribution over the space of events  $\Omega$
2. The space of events is a direct product:  $\Omega = \Omega_1 \times \dots \times \Omega_N$

It is possible to escape these in QFT!

# A 2-d example on a lattice

Consider a system on an  $L \times L$  square lattice with Hamiltonian

$$H = U \sum_{u \sim v} n_u^2 n_v^2 + \sum_v n_v$$

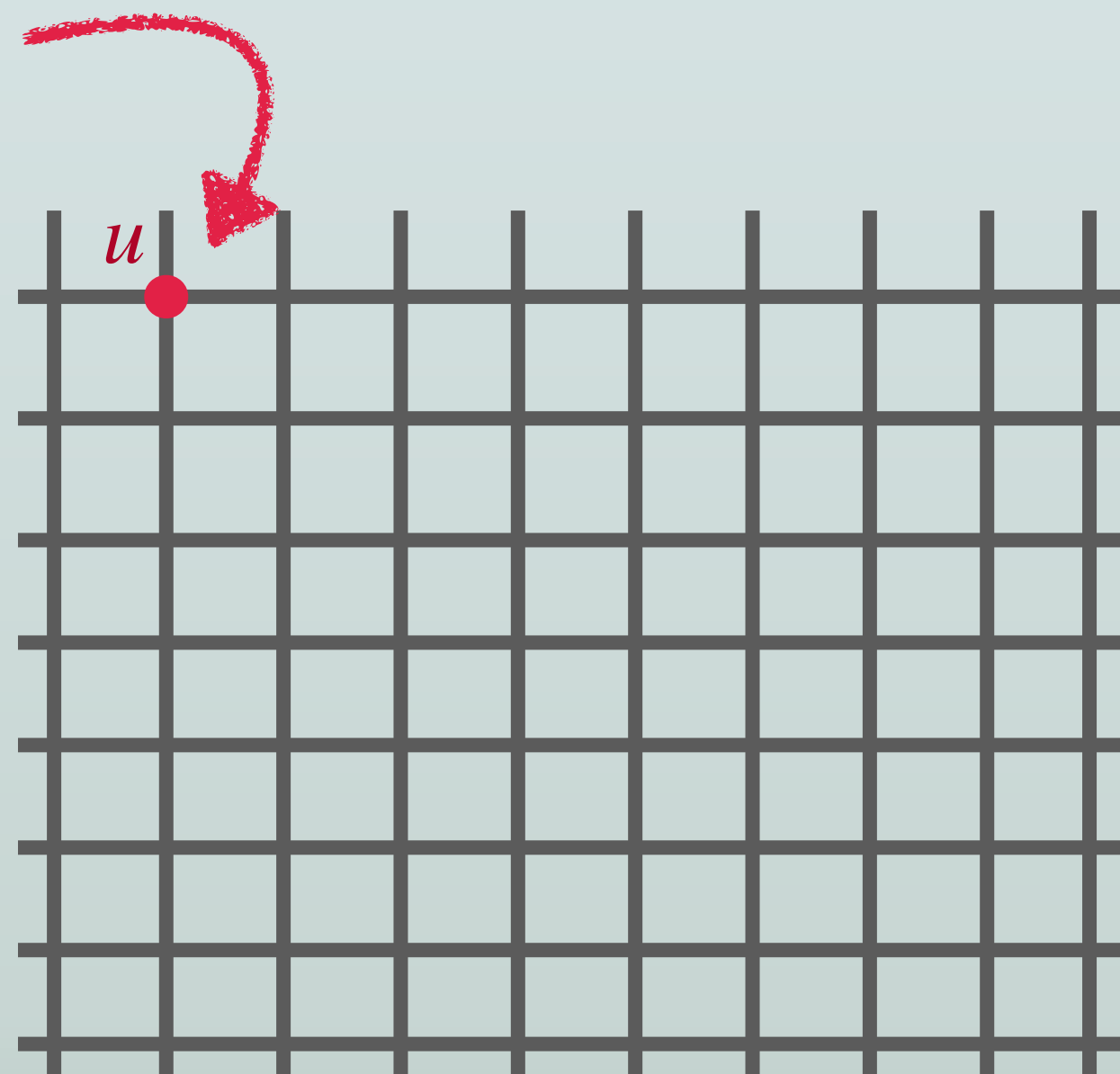


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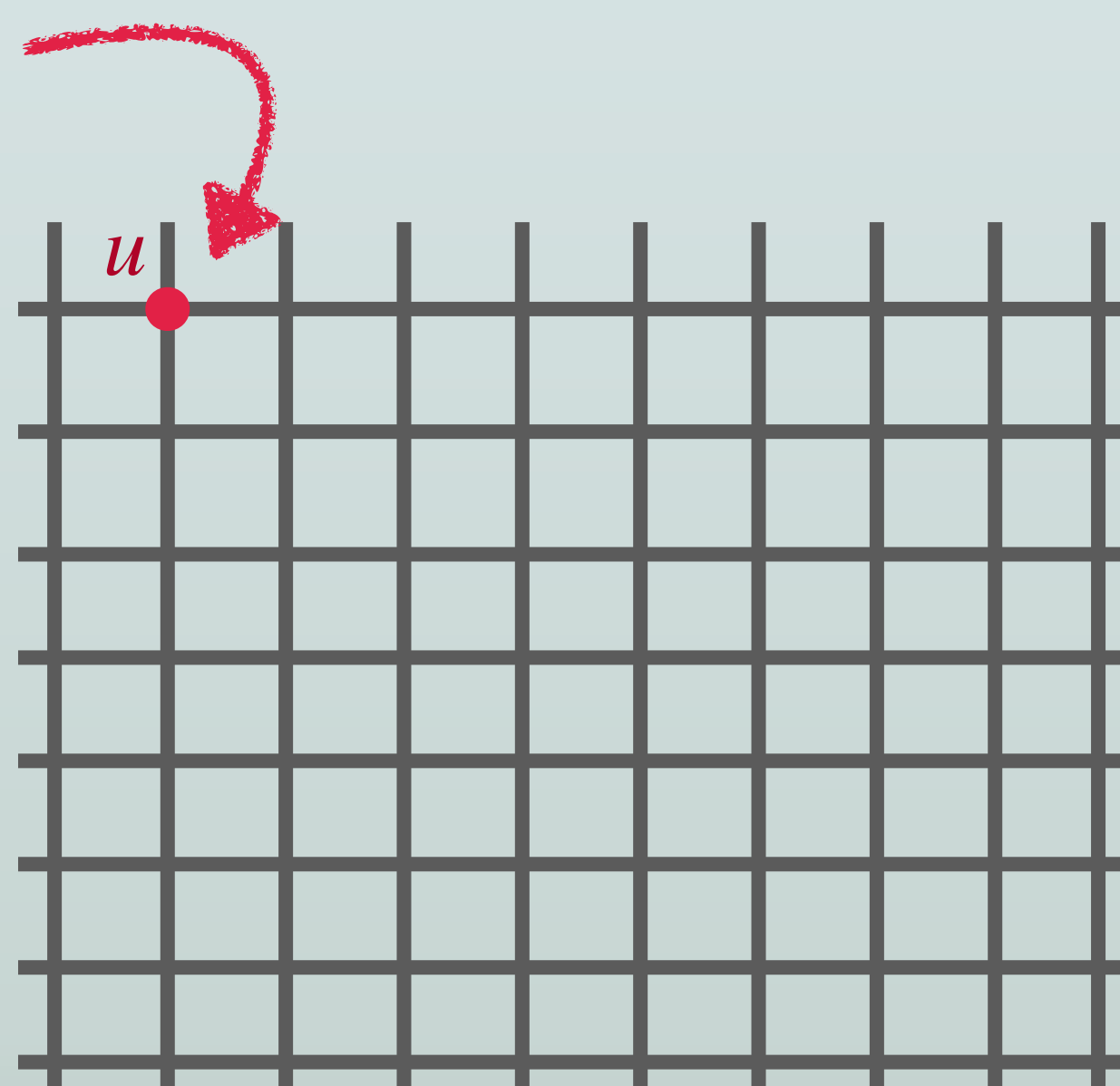
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Infinite number of possible microstates!



Potential to evade the theorem!

## Claim:

For sufficiently small  $\beta < \beta_c$ , the model is an entropic ordered solid.

$\hbar$ QTC

SDU 

## Claim:

hQTC

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SDU 

## Why?

The free energy  $F$  of the solid is smaller than the one of the gas.

$$Z(\beta) = \sum_{\mathbf{n}} e^{-\beta H(\mathbf{n})} = e^{-\beta F}$$

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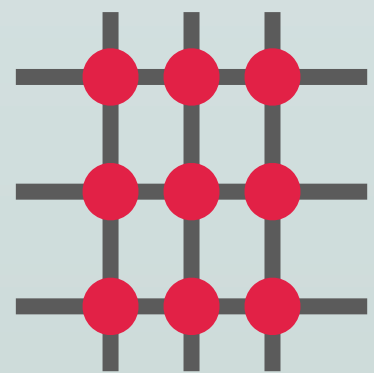
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## Gas:

All sites are occupied



$$\beta \bar{n}_g^4 \sim 1 \quad \Rightarrow \quad \bar{n}_g \sim T^{1/4}$$

$$Z_g \sim \bar{n}_g^{L^2} \sim T^{L^2/4}$$

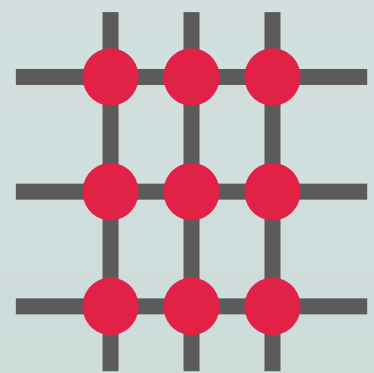
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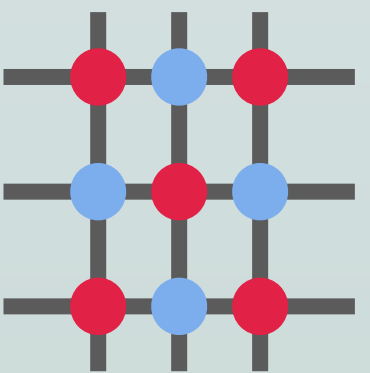
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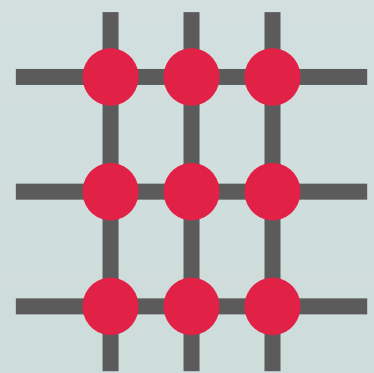
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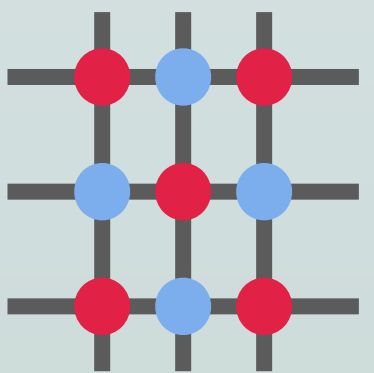
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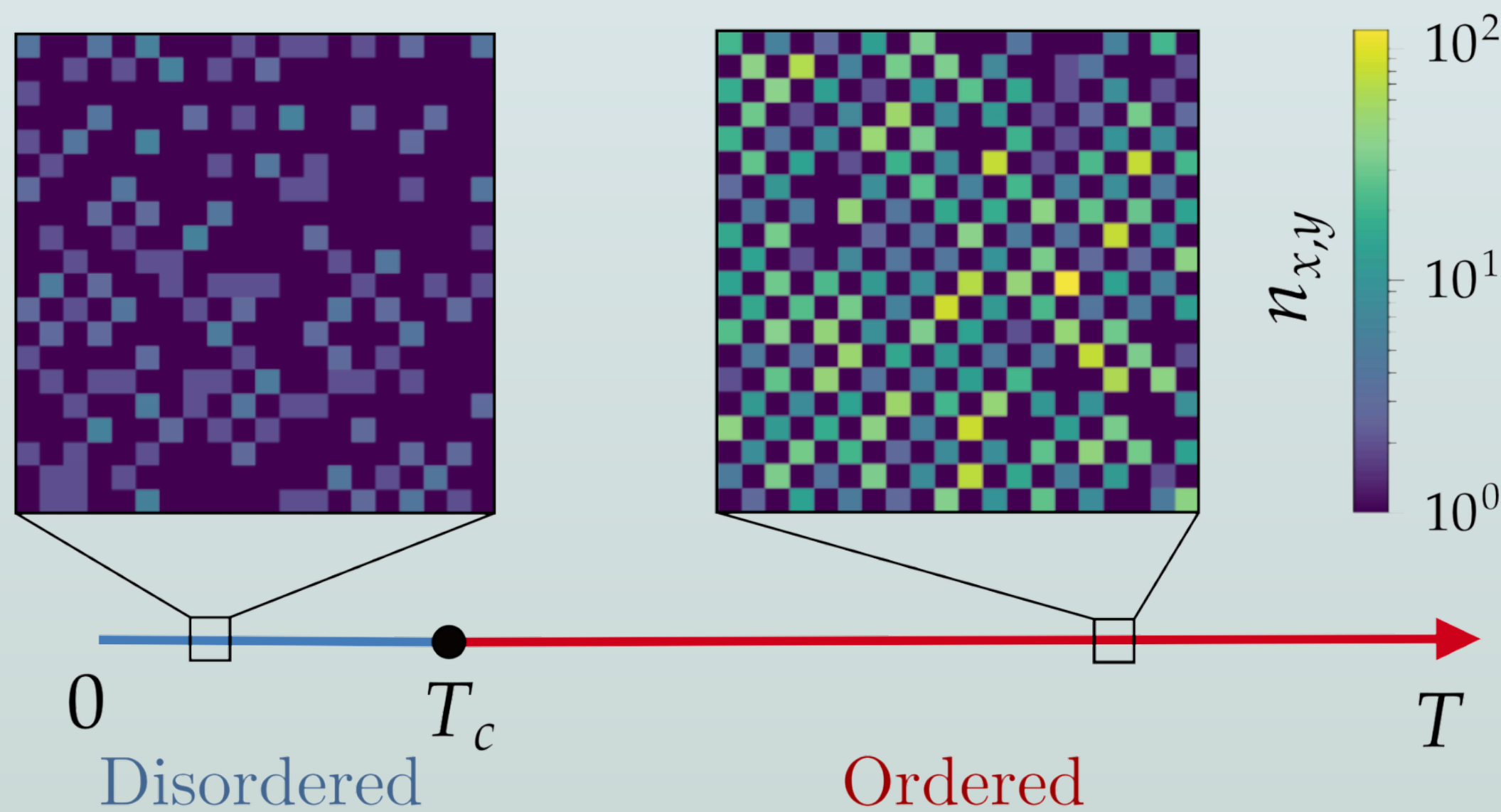
$$\beta \bar{n}_s \sim 1 \quad \Rightarrow \quad \bar{n}_s \sim T$$

$$Z_s \sim \bar{n}_s^{L^2/2} \sim T^{L^2/2}$$

The dominant contribution to  $Z(\beta)$  comes from the checkerboard pattern for sufficiently high  $T$ !

... in fact:

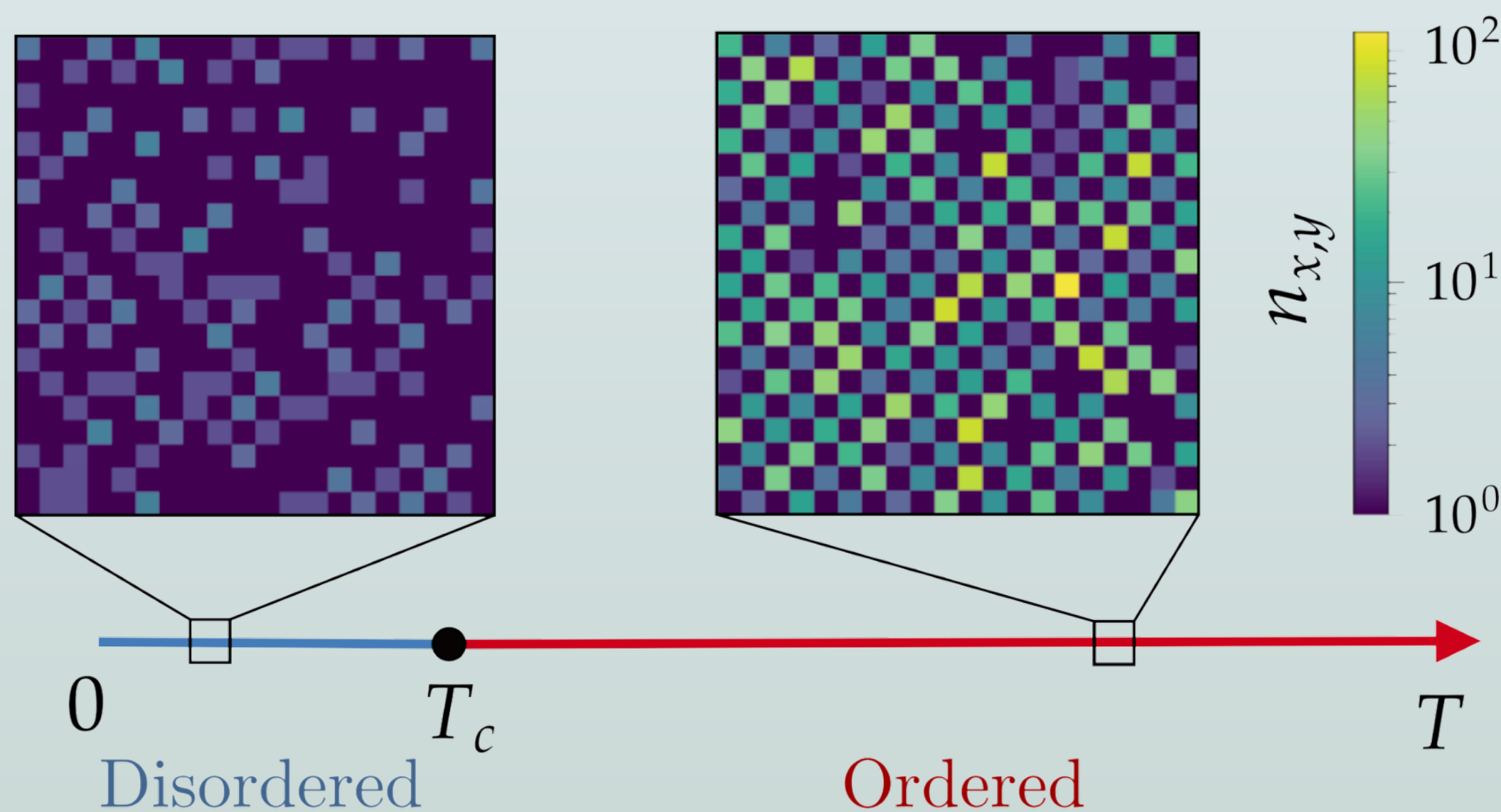
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The **underlying mechanism** that makes the solid phase overall entropically favourable is that sacrificing some entropy to have order opens up a large phase space of configurations through the fluctuations permitted at each site.

In other words: order in one degree of freedom enables other degrees of freedom to strongly fluctuate. The same idea applies in **QFT**.

# **Symmetry non-restoration in $(3+1)$ -d QFTs**

- How do we investigate symmetry non-restoration in QFT?
  1. Put our models at finite  $T$
  2. Compute the resulting **thermal masses** (from  $\Delta V_T$ )
  3. If the mass matrix of the theory has negative eigenvalues, there is condensation

- This topic in the realm of QFT was first studied by Weinberg (1974)

$$V = \frac{\lambda_1}{4}\phi_1^4 + \frac{\lambda_2}{4}\phi_2^4 - \frac{\lambda}{2}\phi_1^2\phi_2^2 \quad \lambda_{1,2} > 0 \quad , \quad \lambda^2 < \lambda_1\lambda_2$$

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... at high T:

$$\Delta V_T = \frac{T^2}{24} \left( (3\lambda_1 - \lambda)\phi_1^2 + (3\lambda_2 - \lambda)\phi_2^2 \right)$$

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for  $\lambda > 3\lambda_2$  :  $\langle \phi_2 \rangle \neq 0$ ,  $\mathbb{Z}_2 \times \mathbb{Z}_2 \rightarrow \mathbb{Z}_2$

- But: this theory is **not UV complete!** The scalar couplings increase with the energy
- What we can actually say: assuming a physical cutoff, for temperatures below it, one observes the phenomenon of symmetry non-restoration
- Our task is then to find a **UV complete** theory in **(3+1)-d** in which symmetries remain broken at arbitrarily high temperatures

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Bajc, Lugo, Sannino,  
arXiv:2012.08428

# Strategy:

- Look for **complete asymptotically free** theories (these require the presence of gauge fields!)  $\Rightarrow$  ansatz for the RGEs and the couplings  $\lambda_I$
- Compute  $\Delta V_T(\lambda_I)$

  $m_T^2(\lambda_I) < 0 \quad ?$

# What kind of theories are considered?

## Complete asymptotically free




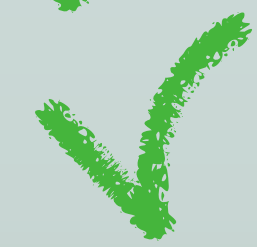
- Gauge singlet scalars
- Gauged scalars
  1.  $SU(N)$  gauge group
  2.  $SU(N_{c_1}) \times SU(N_{c_2})$ 
    - A.  $\phi_i \sim \square$
    - B.  $\phi_i \sim \mathbf{adj}$

## Asymptotically safe



- Litim-Sannino (arXiv:1406.2337)
- A gauged scalar variant of the LS model

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# The case of gauge singlet scalar fields

Consider a  $SU(N_c)$  gauge theory with  $N_f = N_{f_1} + N_{f_2}$  Dirac fermions ( $\square$ ) coupled to the  $\phi_{1,2}$  scalars:

$$\mathcal{L}_Y = \phi_1 \sum_{i=1}^{N_{f_1}} y_{1i} \bar{\psi}_{1i} \psi_{1i} + \phi_2 \sum_{i=1}^{N_{f_2}} y_{2i} \bar{\psi}_{2i} \psi_{2i} \quad \mathbb{Z}_2 \times \mathbb{Z}_2 \quad : \quad \begin{cases} \phi_i \rightarrow -\phi_i \\ \psi \rightarrow i\gamma_5 \psi \end{cases}$$

asymptotic freedom  $\Rightarrow \alpha_g \propto 1/t$  for large  $t = \log(\mu/\mu_0)$

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**complete** asymptotic freedom  $\Rightarrow \alpha_g \propto 1/t$  for large  $t = \log(\mu/\mu_0)$

 **all** couplings  $\lambda_I$  must vanish in the UV at least as fast as  $\alpha_g$ .

With the requirement of  $m_T^2 < 0$  it is sufficient to study the fixed flow solution in which  $\lambda_I \propto 1/t$ .

- This transforms the problem of solving the RGEs (coupled partial differential eq.s) to the problem of solving a system of *polynomial* equations!
- In practice, we define

$$\alpha_g = \frac{g^2}{(4\pi)^2}, \quad \alpha_{y_i} = \frac{y_i^2}{(4\pi)^2}, \quad \alpha_{\lambda_i} = \frac{\lambda_i}{(4\pi)^2}, \quad \alpha_\lambda = \frac{\lambda}{(4\pi)^2}$$

And look for solutions of asymptotic form  $\alpha_a = \frac{\tilde{\alpha}_a}{t}$ ,  $a = g, y_i, \lambda_i, \lambda$  ( $\tilde{\alpha}_a$  constant)

- ... turns out: it is **not** possible to find a solution of the RGEs  $\tilde{\alpha}_a$  such that  $m_T^2 < 0$
- We turn our attention to **gauged scalars**.

# $SU(N_{c_1}) \times SU(N_{c_2})$ with adjoint scalars

## Infinite-N limit

$$V = \frac{\lambda'_1}{4} \text{Tr}(\Sigma_1^4) + \frac{\lambda'_2}{4} \text{Tr}(\Sigma_2^4) + \frac{\lambda_1}{4} [\text{Tr}(\Sigma_1^2)]^2 + \frac{\lambda_2}{4} [\text{Tr}(\Sigma_2^2)]^2 - \frac{\lambda}{2} \text{Tr}(\Sigma_1^2) \text{Tr}(\Sigma_2^2).$$

Field	Spin	$SU(N_{c_1})$	$SU(N_{c_2})$	Multiplicity
$A_\mu^{(1)}$	1	adj	<b>1</b>	—
$A_\mu^{(2)}$	1	<b>1</b>	adj	—
$\Sigma_1$	0	adj	<b>1</b>	1
$\Sigma_2$	0	<b>1</b>	adj	1
$\Psi_1$	$\frac{1}{2}$	$\square$	<b>1</b>	$N_{f1}$
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$$\Delta V_T = \frac{(4\pi)^2 T^2}{48 \log T} \left( \left( \tilde{\lambda}_1 + 2\tilde{\lambda}'_1 - \frac{N_{c_2}}{N_{c_1}} \tilde{\lambda} + 12\tilde{\alpha}_1 \right) \text{Tr} \Sigma_1^2 + \left( \tilde{\lambda}_2 + 2\tilde{\lambda}'_2 - \frac{N_{c_1}}{N_{c_2}} \tilde{\lambda} + 12\tilde{\alpha}_2 \right) \text{Tr} \Sigma_2^2 \right)$$

A solution for bounded potential with negative thermal mass square exists 

# $SU(N_{c_1}) \times SU(N_{c_2})$ with fundamental scalars

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$$V = \frac{\lambda_1}{2} \left( \vec{\varphi}_1^* \cdot \vec{\varphi}_1 \right)^2 + \frac{\lambda_2}{2} \left( \vec{\varphi}_2^* \cdot \vec{\varphi}_2 \right)^2 - \lambda \left( \vec{\varphi}_1^* \cdot \vec{\varphi}_1 \right) \left( \vec{\varphi}_2^* \cdot \vec{\varphi}_2 \right)$$

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we define:

$$g_i^2 = \frac{(4\pi)^2 \tilde{\alpha}_i}{N_{c_i} t} \quad , \quad \lambda_i = \frac{(4\pi)^2 \tilde{\lambda}_i}{N_{c_i} t} \quad ,$$

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$\mu_1^2$

$\mu_2^2$

... defining  $\tilde{\lambda}_{\pm} = \frac{1}{2}(\tilde{\lambda}_1 \pm \tilde{\lambda}_2)$  ,  $\tilde{\alpha}_{\pm} = \frac{1}{2}(\tilde{\alpha}_1 \pm \tilde{\alpha}_2)$

the RGEs become

$$-\tilde{\lambda}_1 = 2\tilde{\lambda}_1^2 + 2\tilde{\lambda}^2 - 6\tilde{\alpha}_1\tilde{\lambda}_1 + \frac{3}{2}\tilde{\alpha}_1^2$$

$$-\tilde{\lambda}_2 = 2\tilde{\lambda}_2^2 + 2\tilde{\lambda}^2 - 6\tilde{\alpha}_2\tilde{\lambda}_2 + \frac{3}{2}\tilde{\alpha}_2^2$$

$$-\tilde{\lambda} = 2(\tilde{\lambda}_1 + \tilde{\lambda}_2)\tilde{\lambda} - 3(\tilde{\alpha}_1 + \tilde{\alpha}_2)\tilde{\lambda}$$

$$\dots \text{defining } \tilde{\lambda}_{\pm} = \frac{1}{2}(\tilde{\lambda}_1 \pm \tilde{\lambda}_2) \quad , \quad \tilde{\alpha}_{\pm} = \frac{1}{2}(\tilde{\alpha}_1 \pm \tilde{\alpha}_2)$$

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$$-\tilde{\lambda} = 2(\tilde{\lambda}_1 + \tilde{\lambda}_2)\tilde{\lambda} - 3(\tilde{\alpha}_1 + \tilde{\alpha}_2)\tilde{\lambda}$$

A solution is:  $\tilde{\alpha}_- = 0 \quad , \quad \tilde{\lambda}_+ = \frac{6\tilde{\alpha}_+ - 1}{4} \quad , \quad \tilde{\lambda}_-^2 + \tilde{\lambda}^2 = \frac{1}{16}(24\tilde{\alpha}_+^2 - 12\tilde{\alpha}_+ + 1) \quad ,$

$$\tilde{\alpha}_+ \geq (3 + \sqrt{3})/12 \quad .$$

This solution supports symmetry non-restoration at arbitrary high temperatures 

... But, is this a valid solution? One must check:

1. Our solution to the RGEs leads to a potential which is bounded from below

$$\lambda_1 \lambda_2 - \lambda^2 > 0 \quad \Leftrightarrow \quad \tilde{\lambda}_+^2 - \tilde{\lambda}_-^2 - \tilde{\lambda}^2 > 0$$





# The finite number degrees of freedom story

- We focus on the  $SU(N_{c_1}) \times SU(N_{c_2})$  gauge theory with fundamental scalars
- At finite-N:

$$-\tilde{\lambda}_1 = \left(2 + \frac{8}{xN_{c_2}}\right) \tilde{\lambda}_1^2 + 2\tilde{\lambda}^2 - 6 \left(1 - \frac{1}{x^2 N_{c_2}^2}\right) \tilde{\alpha}_1 \tilde{\lambda}_1$$

$$+ \frac{3}{2} \tilde{\alpha}_1^2 \left(1 + \frac{1}{xN_{c_2}} - \frac{4}{x^2 N_{c_2}^2} + \frac{2}{x^3 N_{c_2}^3}\right)$$

$$-\tilde{\lambda}_2 = \left(2 + \frac{8}{N_{c_2}}\right) \tilde{\lambda}_2^2 + 2\tilde{\lambda}^2 - 6 \left(1 - \frac{1}{N_{c_2}^2}\right) \tilde{\alpha}_2 \tilde{\lambda}_2$$

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$$\left. - 3 \left( \tilde{\alpha}_1 \left(1 - \frac{1}{x^2 N_{c_2}^2}\right) + \tilde{\alpha}_2 \left(1 - \frac{1}{N_{c_2}^2}\right) \right) \right]$$

$$\mu_i^2 = 3\tilde{\alpha}_i \left(1 - \frac{1}{N_{c_i}^2}\right) + 2\tilde{\lambda}_i \left(\frac{1}{N_{c_i}} + 1\right) - 2\sqrt{\frac{N_{c_j}}{N_{c_i}}} \tilde{\lambda} \quad ,$$

$(i, j) = (1, 2), (2, 1)$

- We focus on the  $SU(N_{c_1}) \times SU(N_{c_2})$  gauge theory with fundamental scalars
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$$\tilde{\lambda}_{\pm} = \tilde{\lambda}_{\pm}^{(0)} + \frac{1}{N_{c_2}} \tilde{\lambda}_{\pm}^{(1)},$$

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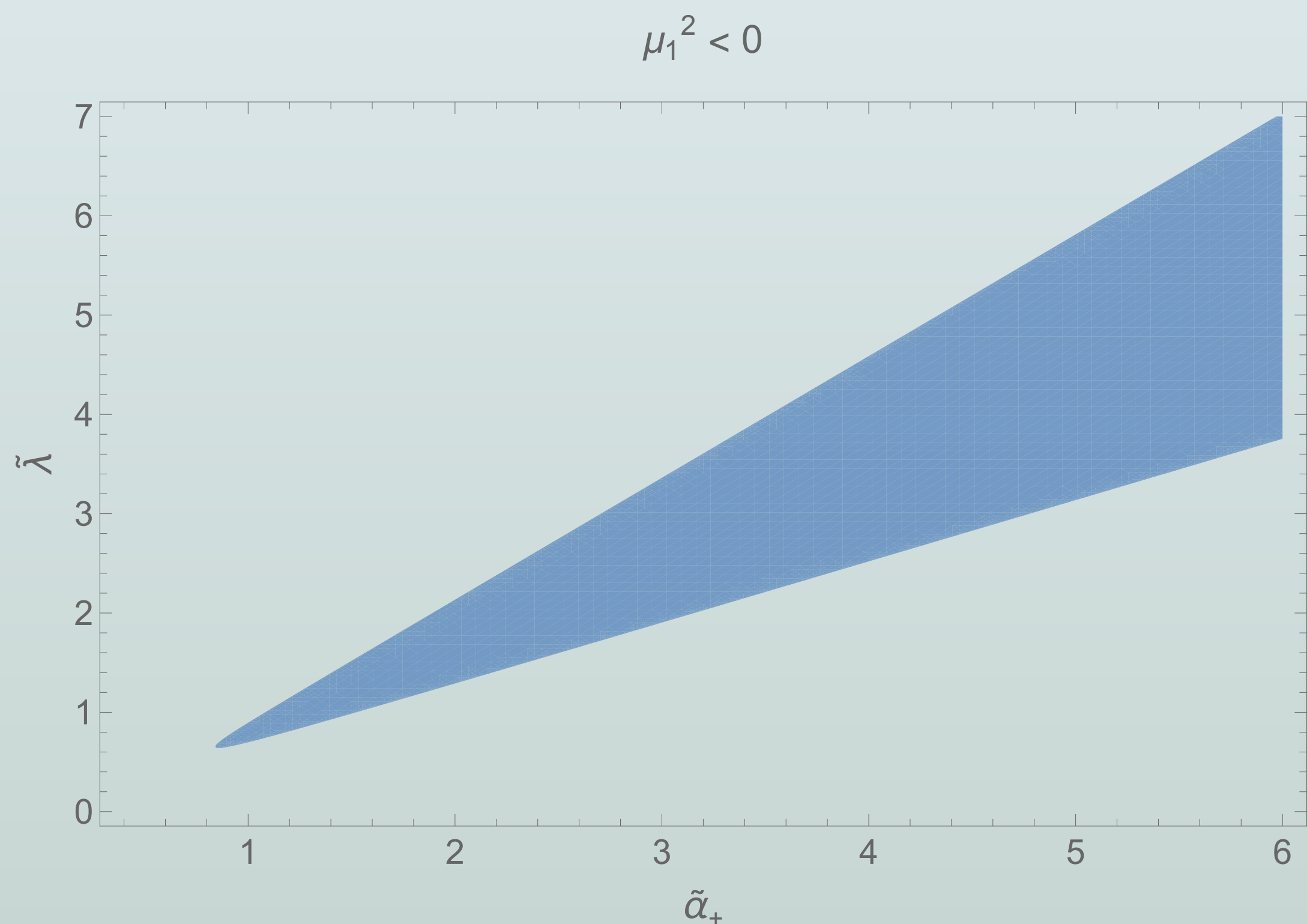
$$\tilde{\lambda} = \tilde{\lambda}^{(0)} + \frac{1}{N_{c_2}} \tilde{\lambda}^{(1)},$$

$$\tilde{\alpha}_\pm = \tilde{\alpha}_\pm^{(0)} + \frac{1}{N_{c_2}} \tilde{\alpha}_\pm^{(1)}.$$

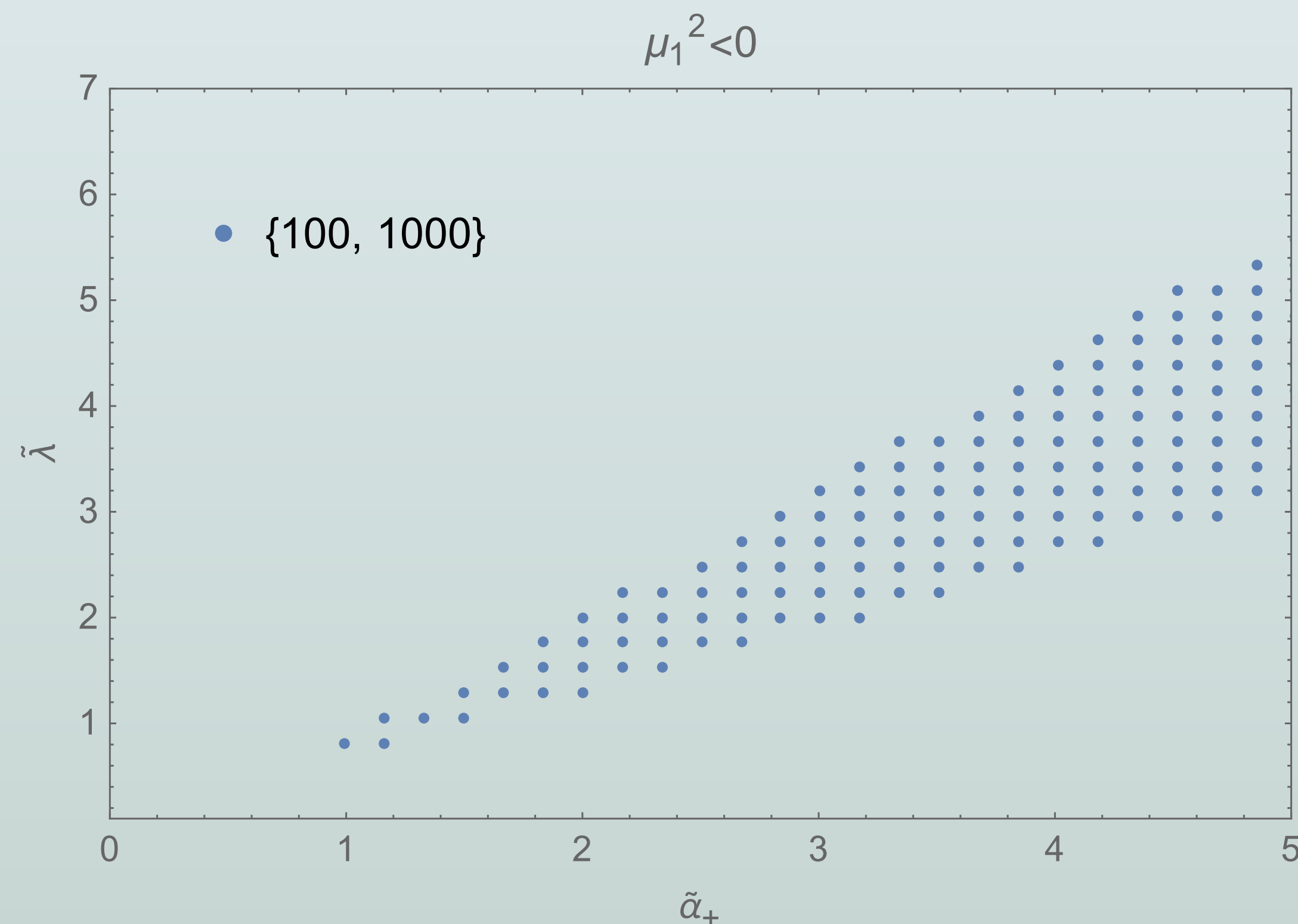
NLO: we now solve for them

LO: already solved!

**1/N result:**



**Numerical evaluation  
(no 1/N truncation)**

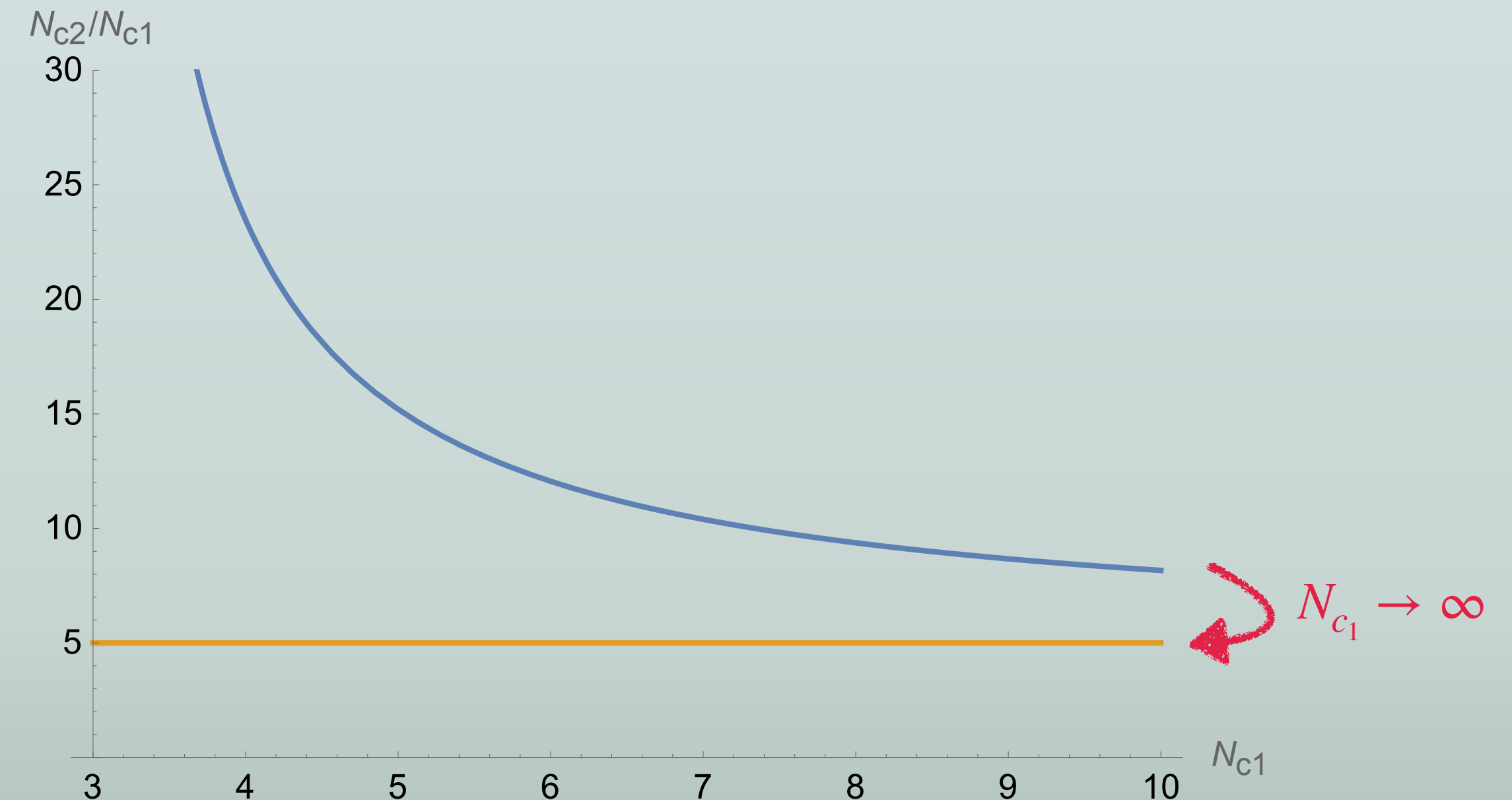
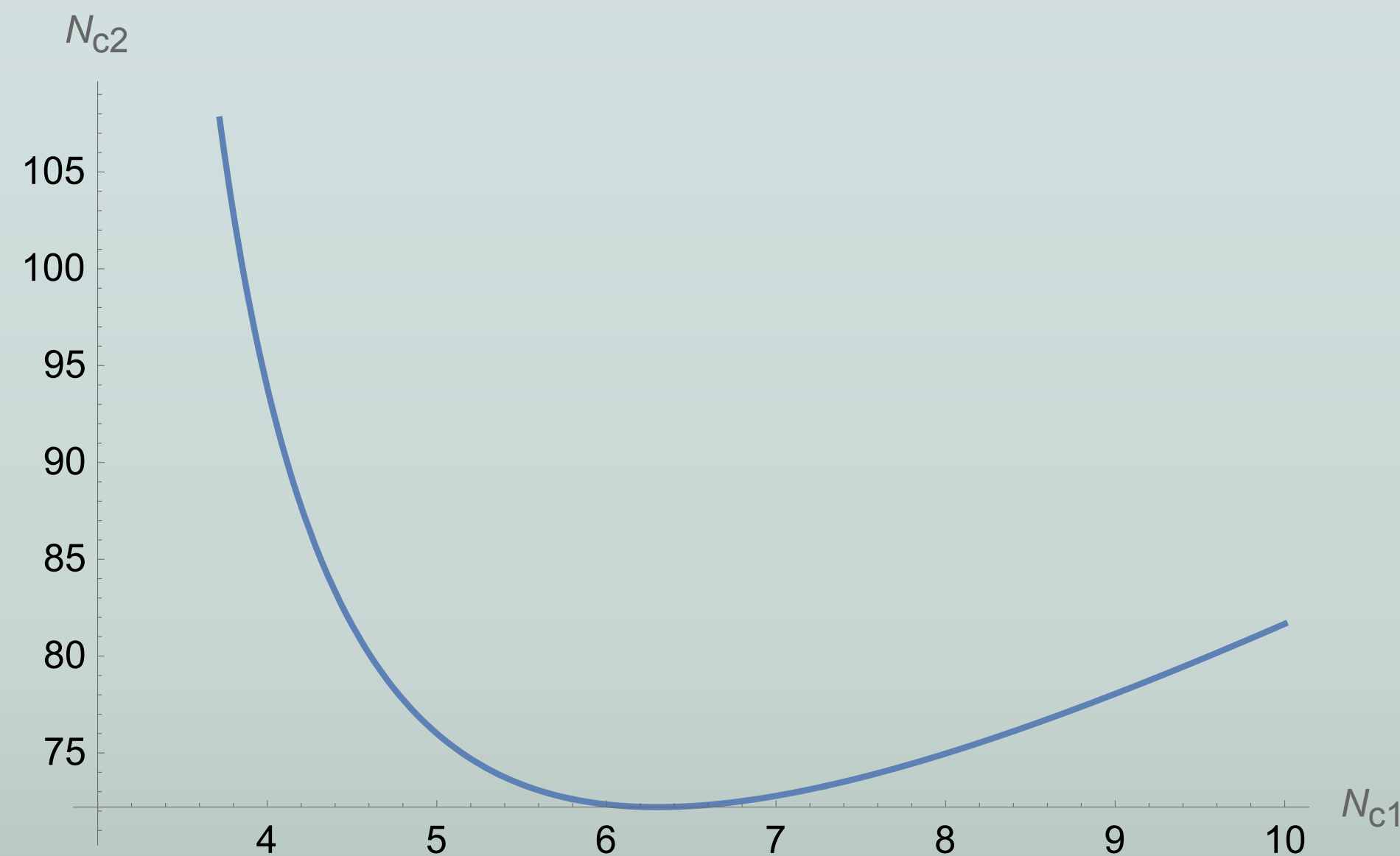


**Bottom line: symmetry non-restoration survives with a finite number of degrees of freedom at arbitrarily high T!**

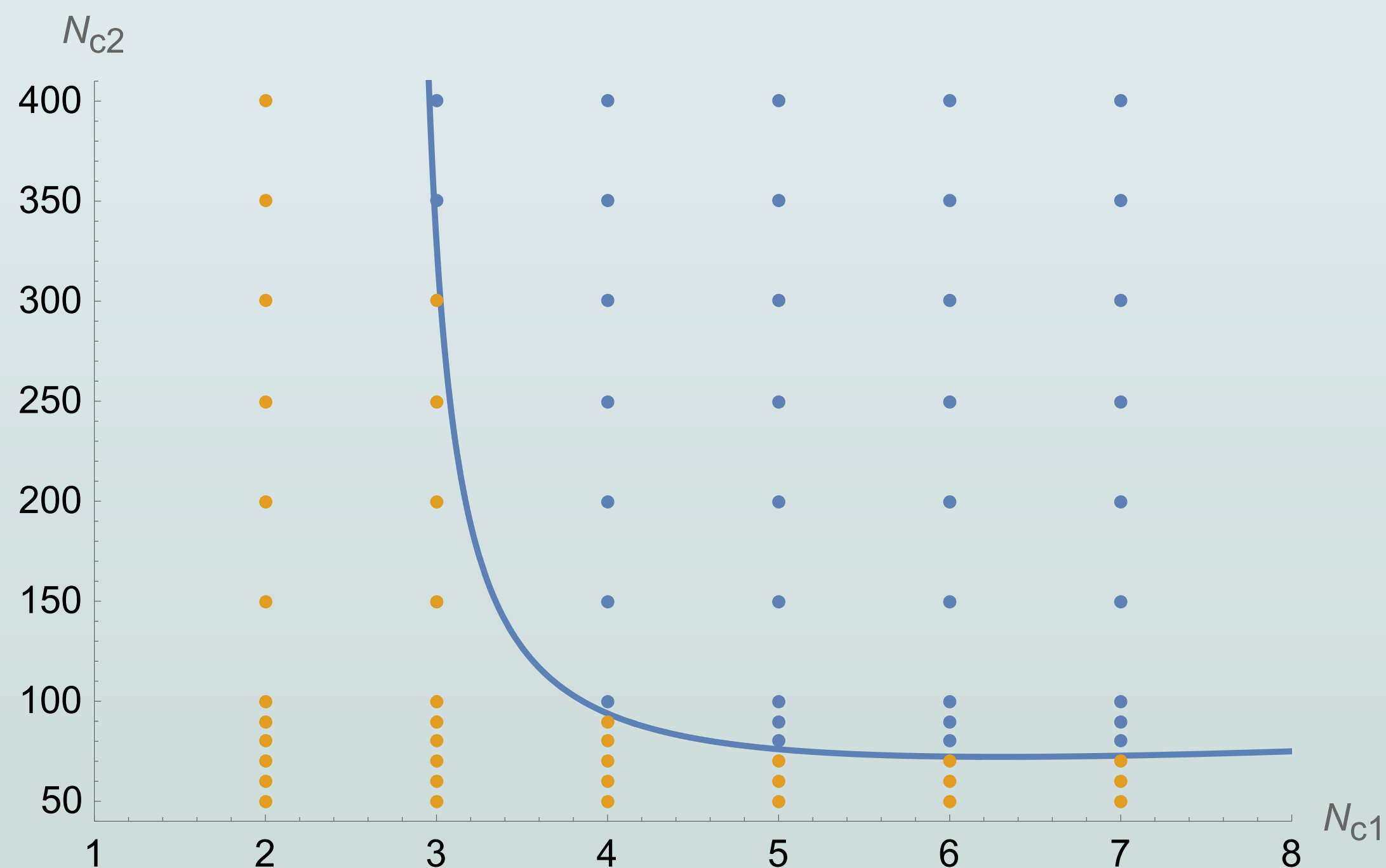
# On the lower bound on the colour degrees of freedom

- What's the lowest possible bound on  $N_{c_i}$  ( $i = 1, 2$ ) below which **symmetry must restore at high temperature?**
- Within perturbation theory we find:

$$N_{c2} \geq \frac{10N_{c1}^4 + 31N_{c1}^3 + 12N_{c1}^2 - 29N_{c1} - 20}{2N_{c1}^3 - 3N_{c1}^2 - 9N_{c1} + 5} .$$

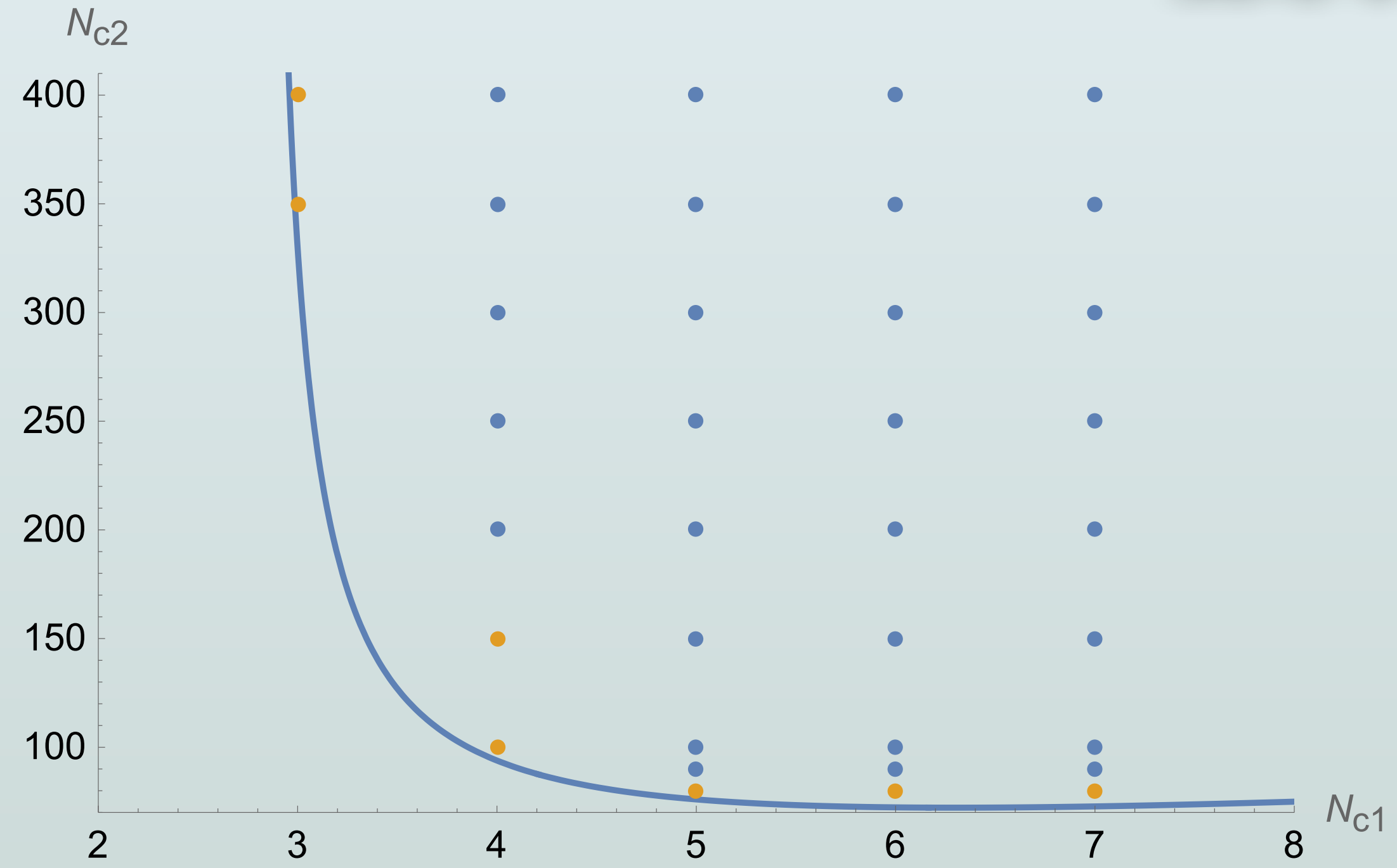
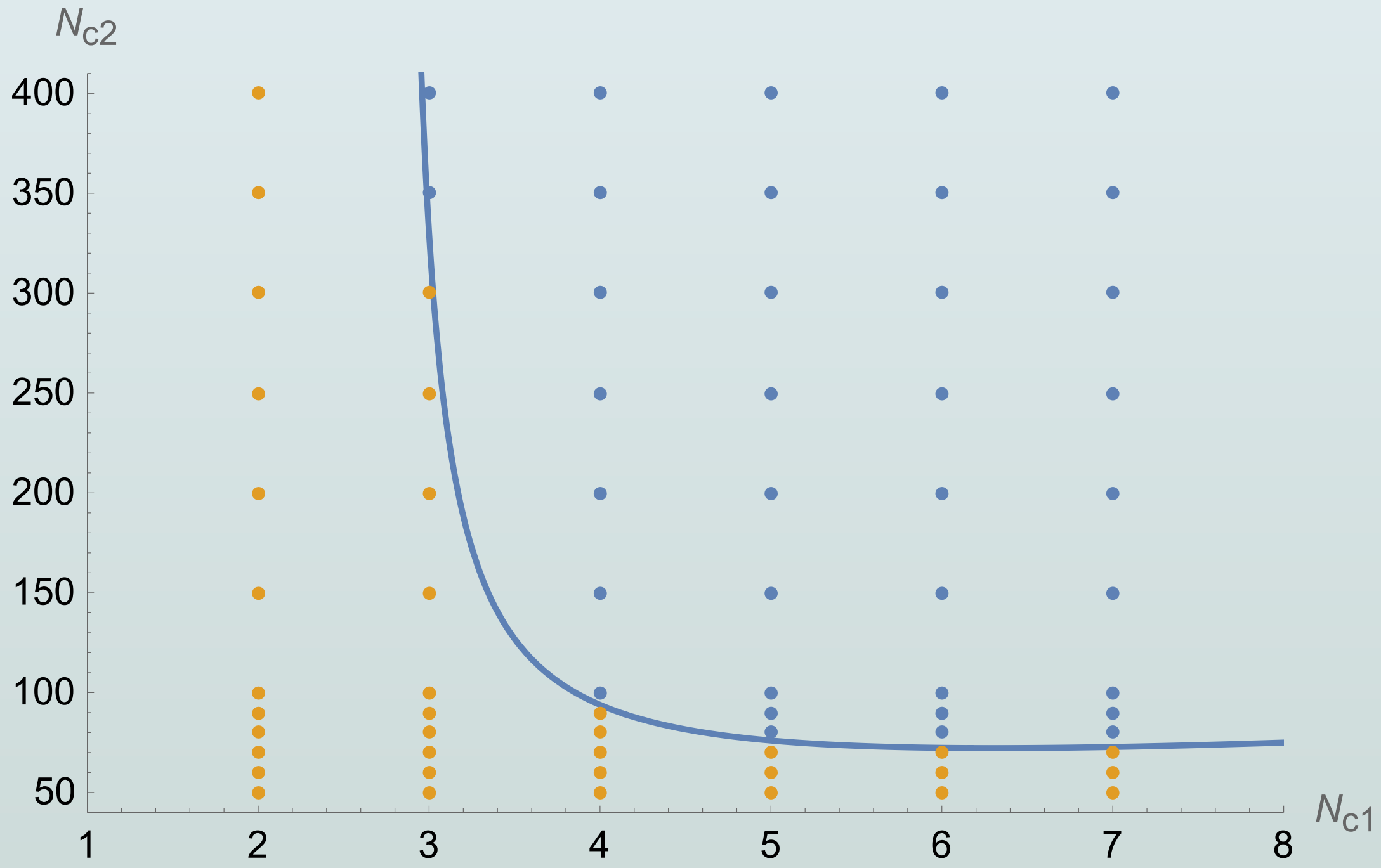


## Numerical results (no $1/N$ truncation)



Regions in the  $N_{c_1} - N_{c_2}$  plane with (blue) and without (yellow) a solution with symmetry non-restoration. The boundary between them is well described by the perturbative constraint!

# Numerical results (no $1/N$ truncation)



Regions in the  $N_{c_1} - N_{c_2}$  plane with (blue) and without (yellow) a solution with symmetry non-restoration. The boundary between them is well described by the perturbative constraint!



We lose some solutions when taking into account integer  $N_{f_{1,2}}$

# Conclusions & Summary

- We now know that there exists (at least) one example of a (3+1)-d theory with a finite number of degrees of freedom that realises symmetry non-restoration at arbitrarily high T
- Symmetry non-restoration might be the answer to some longstanding phenomenological issues, e.g. the problems of domain walls and magnetic monopoles simply wouldn't arise in cosmology.

**Thank you!**