

JHEP 08 (2024) 080 (2405.17056 [hep-th])

JHEP 01 (2025) 111 (2410.22988 [hep-th])

Phys.Rev.D 111 (2025) 12, L121702 (2501.07372 [hep-th])

Compactly packaging PM dynamics: a fresh look at the eikonal

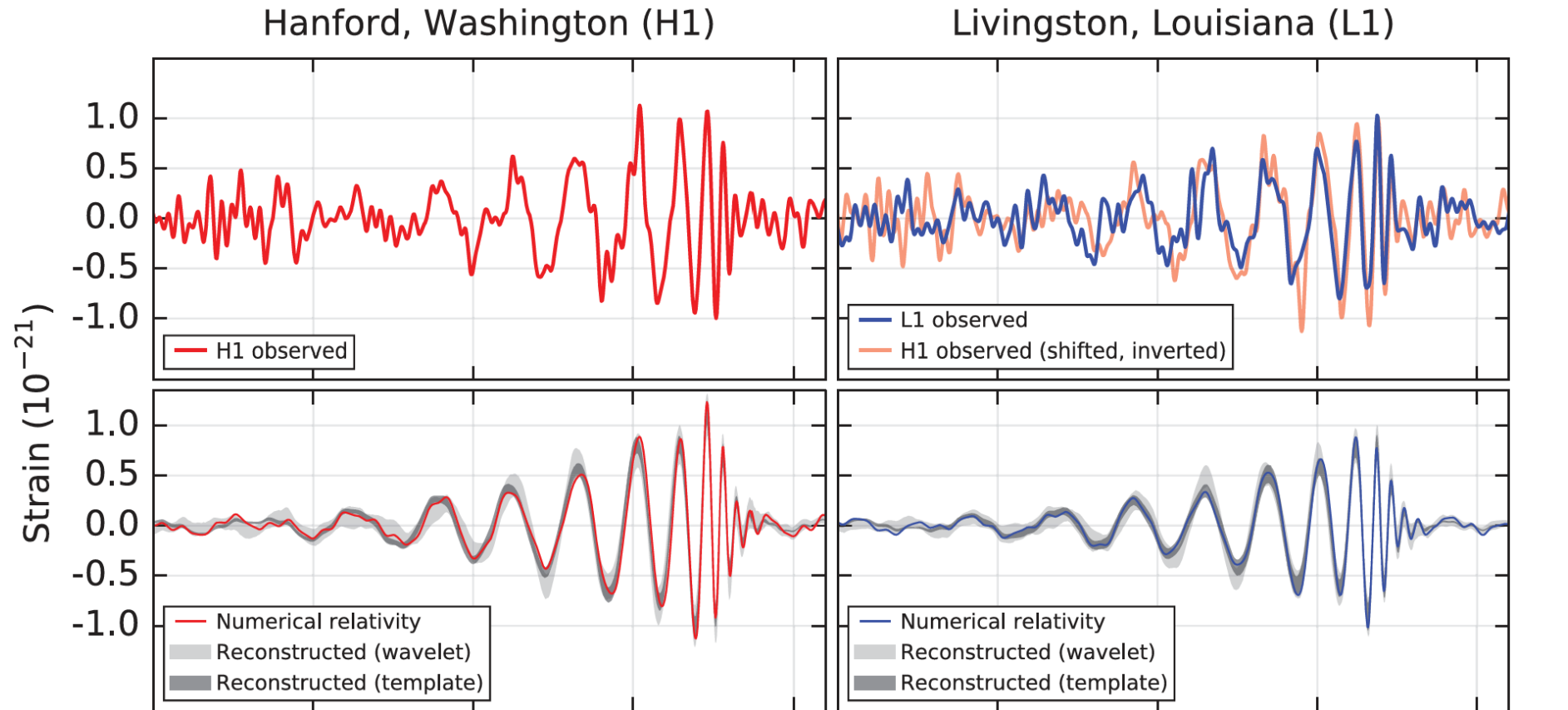
Jung-Wook Kim [MPI for Gravitational Physics (AEI), Potsdam]

Works in collaboration with : Joon-Hwi Kim, Sangmin Lee [2405.17056]

Joon-Hwi Kim, Sungsoo Kim, Sangmin Lee [2410.22988]

PM dynamics: Why did we get here?

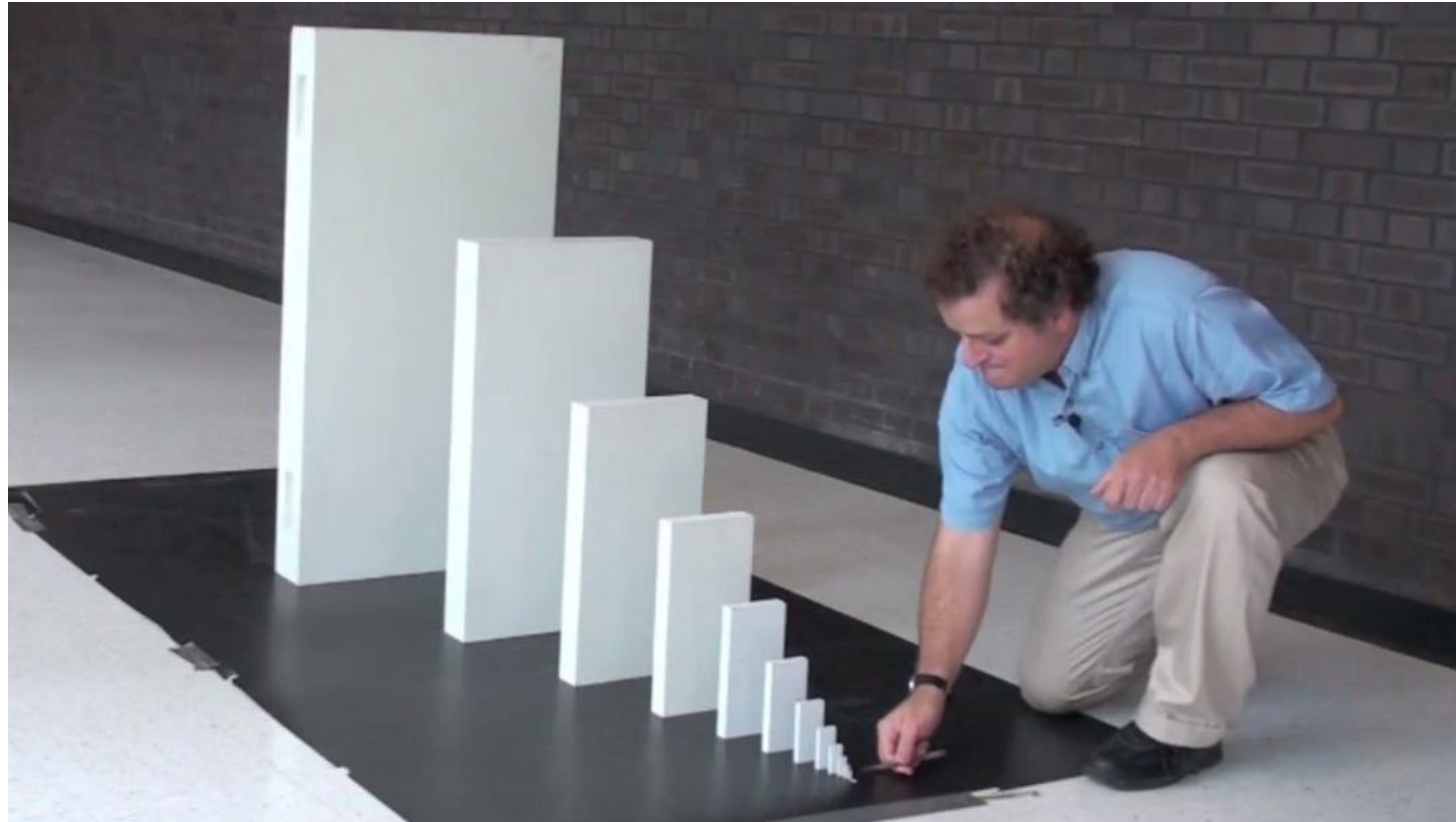
- 10 years since GW150914



LIGO Scientific & Virgo [1602.03837]

PM dynamics: Why did we get here?

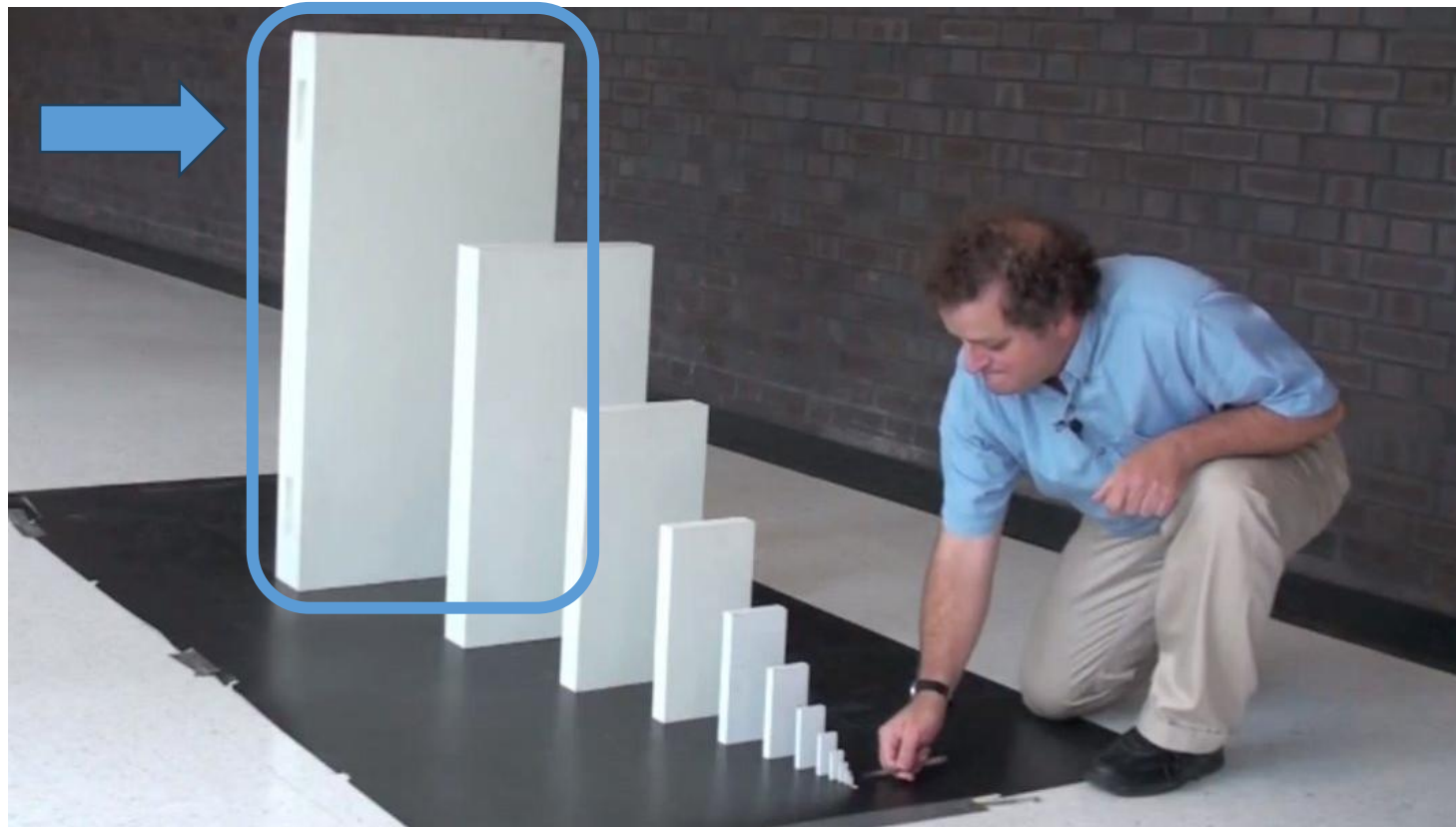
- 10 years since GW150914: huge literature on PM dynamics



PM dynamics: Why did we get here?

- 10 years since GW150914: huge literature on PM dynamics

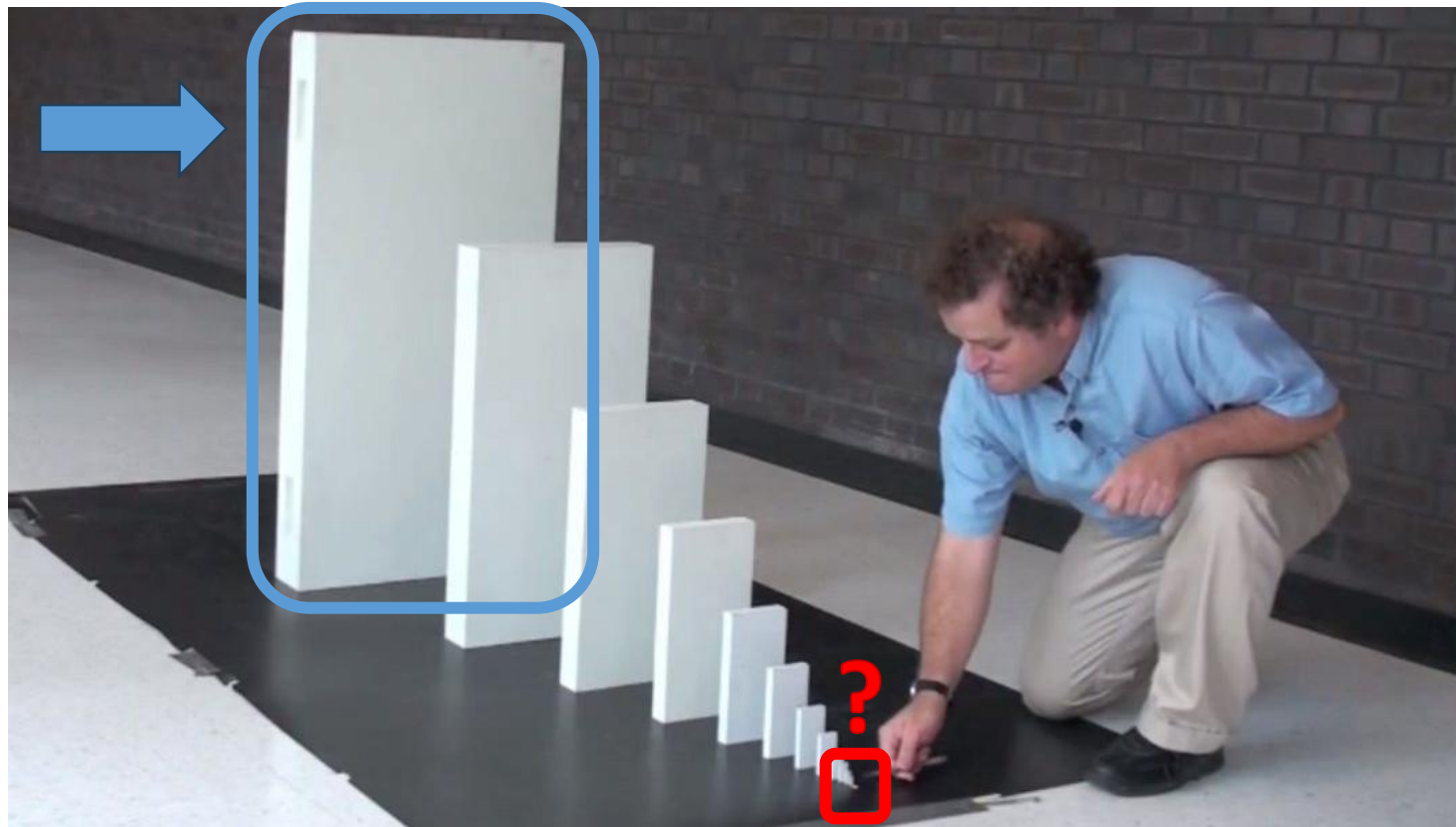
massive
PM
literature
(pun intended)



PM dynamics: Why did we get here?

- 10 years since GW150914: huge literature on PM dynamics

massive
PM
literature
(pun intended)



PM dynamics: Why did we get here?

- 10 years since GW150914: huge literature on PM dynamics

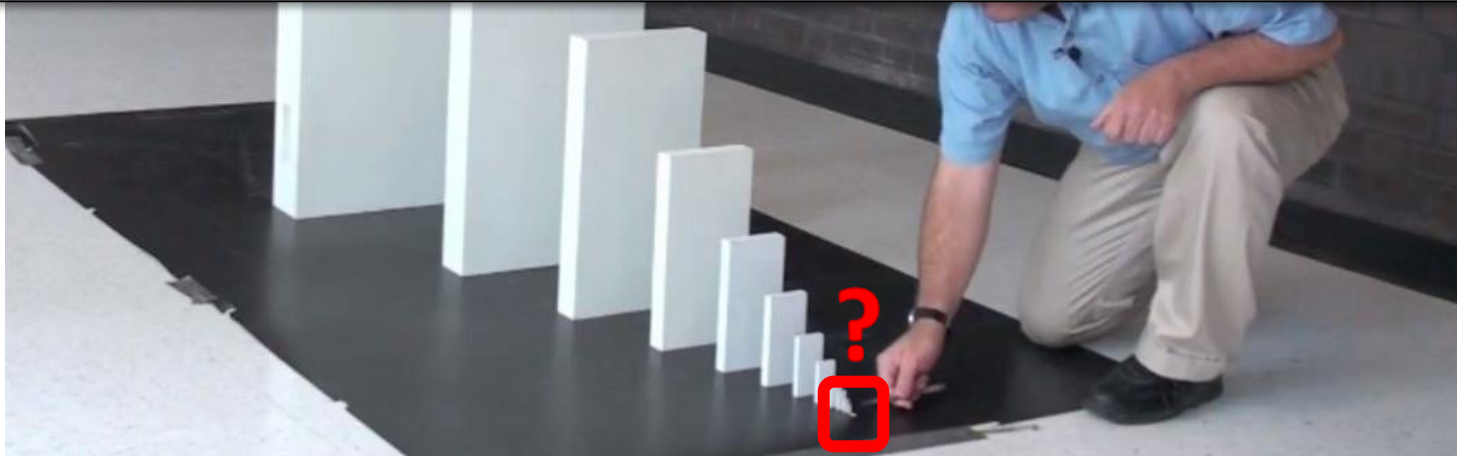
High-energy gravitational scattering and the general relativistic two-body problem

Thibault Damour*

Institut des Hautes Etudes Scientifiques, 35 route de Chartres, 91440 Bures-sur-Yvette, France



(Received 29 October 2017; published 26 February 2018)



PM dynamics: Why did we get here?

- 10 years since GW150914: huge literature on PM dynamics

High-energy gravitational scattering and the general relativistic two-body problem

Thibault Damour*

We derive several consequences of our second post-Minkowskian Hamiltonian: (i) the need to use special phase-space gauges to get a tame high-energy limit; and (ii) predictions about a (rest-mass independent) linear Regge trajectory behavior of high-angular-momenta, high-energy circular orbits. Ways of testing these predictions by dedicated numerical simulations are indicated. We finally indicate a way to connect our classical results to the quantum gravitational scattering amplitude of two particles, and we urge amplitude experts to use their novel techniques to compute the two-loop scattering amplitude of scalar masses, from which one could deduce the third post-Minkowskian effective one-body Hamiltonian.

PM dynamics: Why did we get here?

- 10 years since GW150914: huge literature on PM dynamics

High-energy gravitational scattering and the general relativistic two-body problem

Thibault Damour*

We derive several consequences of our second post-Minkowskian Hamiltonian: (i) the need to use special phase-space gauges to get a tame high-energy limit; and (ii) predictions about a (rest-mass independent) linear Regge trajectory behavior of high-angular-momenta, high-energy circular orbits. Ways of testing these predictions by dedicated numerical simulations are indicated. We finally indicate a way to connect our classical results to the quantum gravitational scattering amplitude of two particles, and we urge amplitude experts to use their novel techniques to compute the two-loop scattering amplitude of scalar masses, from which one could deduce the third post-Minkowskian effective one-body Hamiltonian.

PM dynamics: Why did we get here?

- 10 years since GW150914: huge literature on PM dynamics

High-energy gravitational scattering and the general relativistic two-body problem

Scattering Amplitudes and the Conservative Hamiltonian for Binary Systems at Third Post-Minkowskian Order

Zvi Bern,¹ Clifford Cheung,² Radu Roiban,³ Chia-Hsien Shen,¹ Mikhail P. Solon,² and Mao Zeng⁴

¹*Mani L. Bhaumik Institute for Theoretical Physics, University of California at Los Angeles, Los Angeles, California 90095, USA*

²*Walter Burke Institute for Theoretical Physics, California Institute of Technology, Pasadena, California 91125*

³*Institute for Gravitation and the Cosmos, Pennsylvania State University, University Park, Pennsylvania 16802, USA*

⁴*Institute for Theoretical Physics, ETH Zürich, 8093 Zürich, Switzerland*



(Received 28 January 2019; published 24 May 2019)

amplitude experts to use their novel techniques to compute the two-loop scattering amplitude of scalar masses, from which one could deduce the third post-Minkowskian effective one-body Hamiltonian.

PM dynamics: Why did we get here?

- 10 years since GW150914: huge literature on PM dynamics
- Keywords from Damour's request
 - Effective-one-body (EOB): weak-field \rightarrow analytic strong-field [Buonanno, Damour]
 - Resum by leveraging background-probe limit
 - Motivation: GW loudest in strong-field regime
 - Post-Minkowskian (PM): (gauge-inv.) scattering dynamics
 - Singularities *physical* \rightarrow leverage for resummation (e.g. \mathcal{L} -resum [Damour, Rettegno])
- What else can we leverage from PM? How should we do it?
 - Spin resummation / Through the eikonal (= scattering generator)

PM dynamics: Why did we get here?

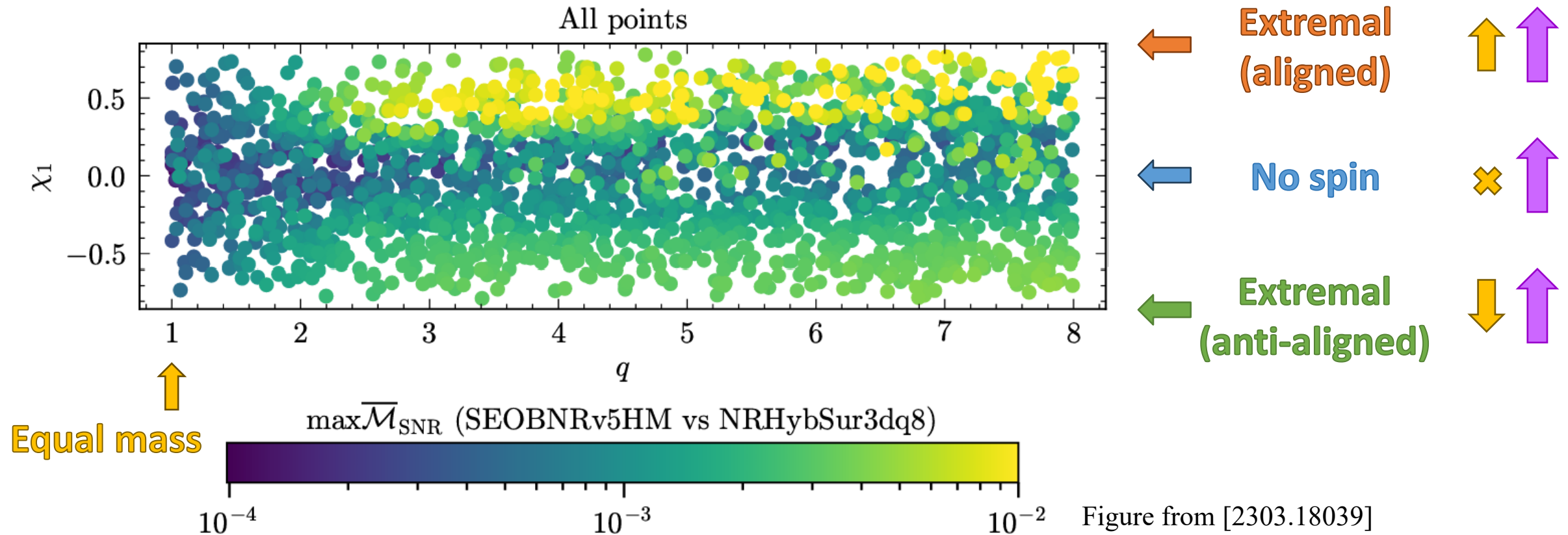
- 10 years since GW150914: huge literature on PM dynamics
- Keywords from Damour's request
 - Effective-one-body (EOB): weak-field \rightarrow analytic strong-field [Buonanno, Damour]
 - Resum by leveraging background-probe limit
 - Motivation: GW loudest in strong-field regime
 - Post-Minkowskian (PM): (gauge-inv.) scattering dynamics
 - Singularities *physical* \rightarrow leverage for resummation (e.g. \mathcal{L} -resum [Damour, Rettegno])
- What else can we leverage from PM? How should we do it?
 - Spin resummation / Through the eikonal (= scattering generator)

PM dynamics: Why did we get here?

- 10 years since GW150914: huge literature on PM dynamics
- Keywords from Damour's request
 - Effective-one-body (EOB): weak-field \rightarrow analytic strong-field [Buonanno, Damour]
 - Resum by leveraging background-probe limit
 - Motivation: GW loudest in strong-field regime
 - Post-Minkowskian (PM): (gauge-inv.) scattering dynamics
 - Singularities *physical* \rightarrow leverage for resummation (e.g. \mathcal{L} -resum [Damour, Rettegno])
- What else can we leverage from PM? How should we do it?
 - Spin resummation / Through the eikonal (= scattering generator)

Why spin resummation?

- Waveform models perform worse for high spins/mass ratios
 - Comparison with numerical relativity (NR) simulations



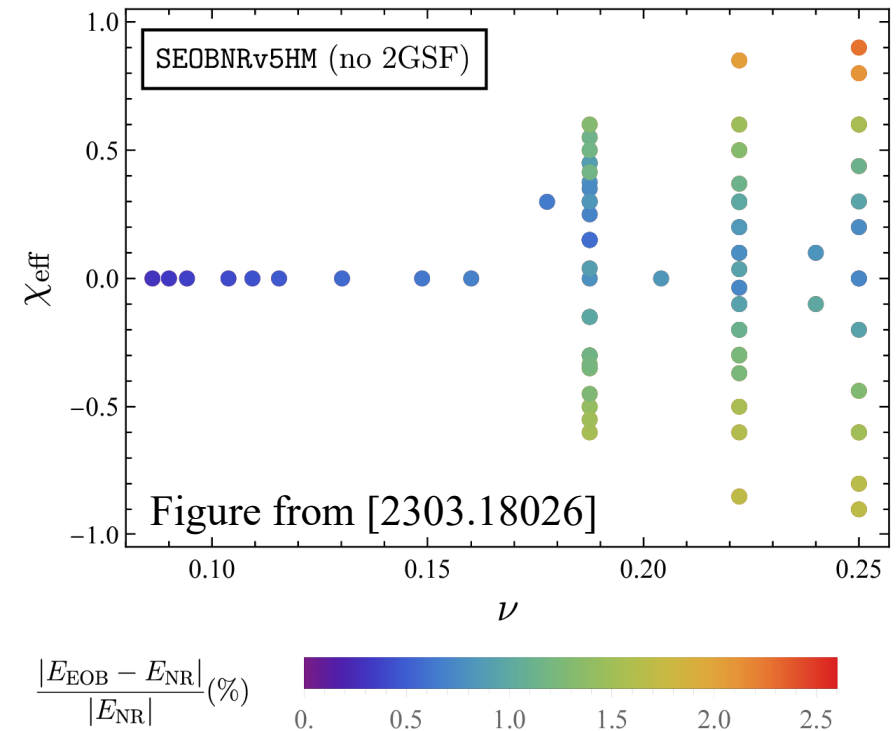
Why spin resummation?

- Waveform models perform worse for high spins/mass ratios
 - Comparison with numerical relativity (NR) simulations
- Limiting factor for (next-gen) GW physics
 - “we ascertain that current waveforms can accurately recover the distribution of masses in the LVK astrophysical population, **but not spins**”

[Dhani, Völkel, Buonanno, Estelles, Gair, Pfeiffer, Pompili, Toubiana]

Parameter	Notation	Astrophysically relevant range
Total mass in the detector frame	M	$10^5 - 10^7 M_\odot$
Mass ratio (> 1)	q	$1 - 10$
Dimensionless spin	$\max \chi_i $	$0 - 0.998$
Eccentricity entering LISA band	e_{init}	$0 - 0.99$
Eccentricity at last stable orbit	e_{merge}	< 0.1
Signal to noise ratio	SNR	$10 - 10^4$

Table from [2311.01300]



Why spin resummation?

- Why expect spin to be resumable?

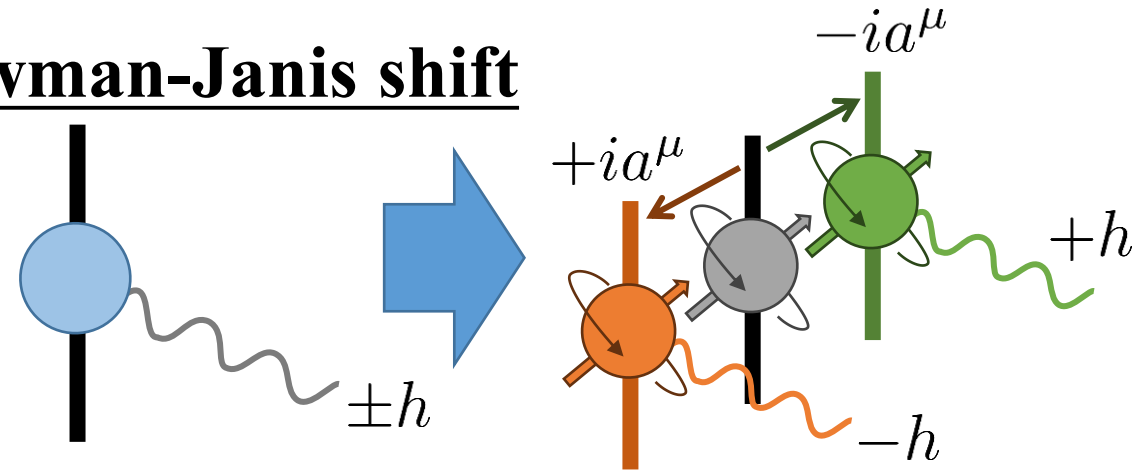
- Black hole spin moments given by Newman-Janis shift

- Kerr couples minimally to gravitons

[Guevara, Ochirov, Vines; Chung, Huang, **JWK**, Lee]

- Minimal coupling \approx on-shell NJ shift

[Arkani-Hamed, Huang, O'Connell]



- Expectation: *singularities* of binary dynamics governed by NJ shift

- Resummations leverage singularities \Rightarrow *singularities* $>$ *pert. exactness*

- Several known models implement NJ shift dynamically

- Worldsheet [Guevara, Maybee, Ochirov, O'Connell, Vines], Twistor worldline [JH Kim, **JWK**, Lee],

- Bootstrap [Bjerrum-Bohr, Chen, Skowronek], Higher-spin gauge [Cangemi, Chiodaroli, Johansson,

Ochirov, Pichini, Skvortsov]

Why spin resummation?

- Why expect spin to be resumable?

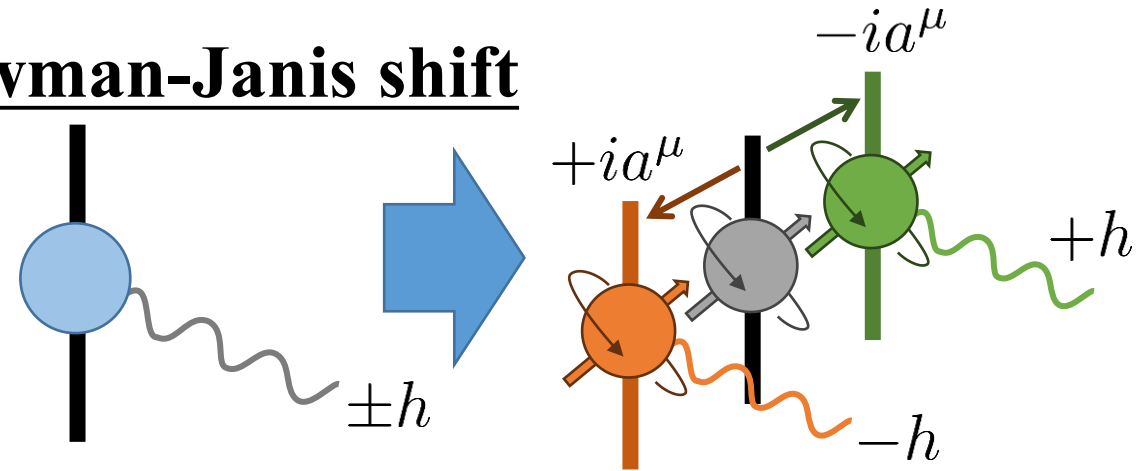
- Black hole spin moments given by Newman-Janis shift

- Kerr couples minimally to gravitons

[Guevara, Ochirov, Vines; Chung, Huang, **JWK**, Lee]

- Minimal coupling \approx on-shell NJ shift

[Arkani-Hamed, Huang, O'Connell]



- Expectation: *singularities* of binary dynamics governed by NJ shift

- Resummations leverage singularities \Rightarrow *singularities* $>$ *pert. exactness*

- Several known models implement NJ shift dynamically

- Worldsheet [Guevara, Maybee, Ochirov, O'Connell, Vines], Twistor worldline [JH Kim, **JWK**, Lee],

- Bootstrap [Bjerrum-Bohr, Chen, Skowronek], Higher-spin gauge [Cangemi, Chiodaroli, Johansson,

Ochirov, Pichini, Skvortsov]

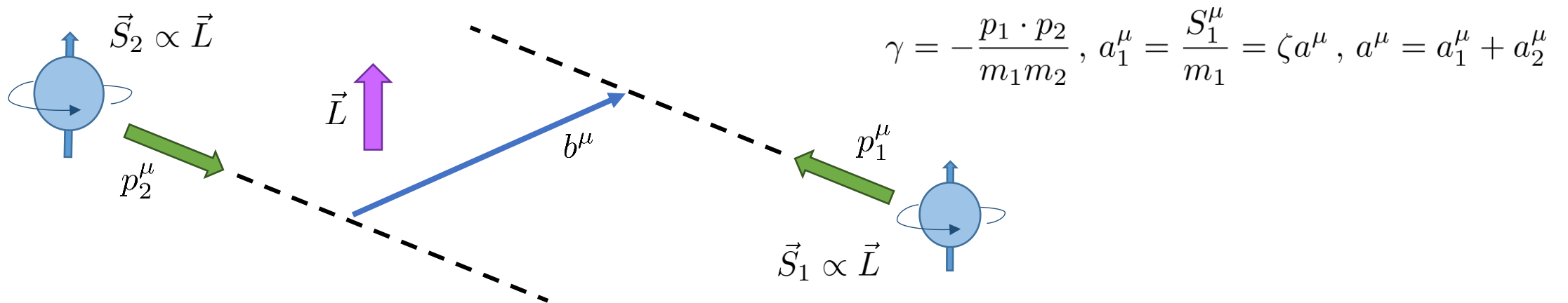
Spin resummation @ NLO

[JH Kim, **JWK**, Lee]

- Spin-resummed 2PL twistor worldline eikonal
 - Eikonal is a double infinite sum of 2F1, but simplifies in special kinematics
 - Aligned-spin configuration: *square roots*

[Aoude, Haddad, Helset; Damgaard, Hoogeveen, Luna, Vines; Chen, **JWK**, Wang; Bohnenblust, Cangemi, Johansson, Pichini]

$$\chi(2,\text{aligned}) = \frac{(q_1 q_2)^2 \left(b^2 + \frac{(\zeta-2)\gamma}{(\gamma^2-1)} \epsilon[b, v_1, v_2, a] + \frac{\gamma^2(1-\zeta)+\zeta}{\gamma^2-1} a^2 \right)}{32\pi m_1 \sqrt{\gamma^2 - 1} (b^2 - a^2)^{3/2}} + (1 \leftrightarrow 2)$$

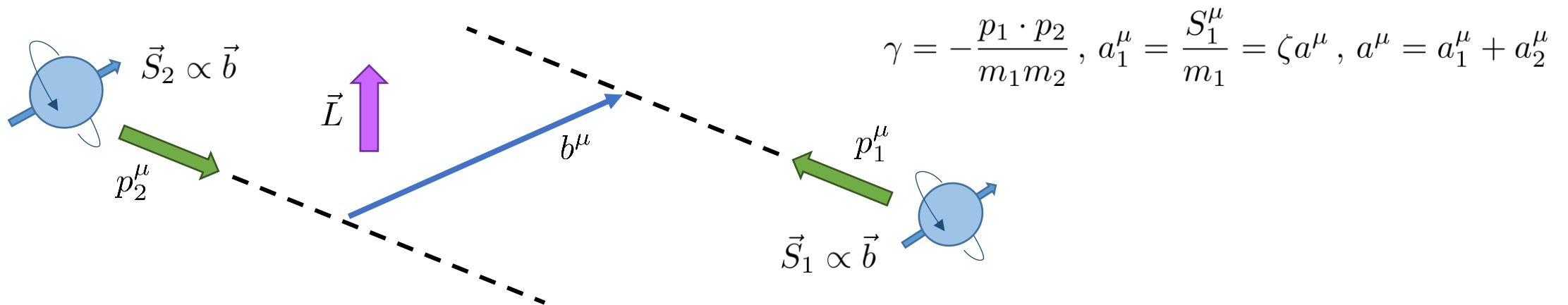


Spin resummation @ NLO

[JH Kim, JWK, Lee]

- Spin-resummed 2PL twistor worldline eikonal
 - Eikonal is a double infinite sum of 2F1, but simplifies in special kinematics
 - Axial scattering configuration: *elliptic integrals*

$$\chi_{(2,\text{axial})} = \frac{(q_1 q_2)^2 \sqrt{b^2}}{16\pi^2 m_1 (\gamma^2 - 1)^{3/2}} \left[\frac{\gamma^2 (\zeta - 1) - \zeta}{b^2} K\left(-\frac{a^2}{b^2}\right) - \frac{\gamma^2 (\zeta - 2) - (\zeta - 1)}{b^2 + a^2} E\left(-\frac{a^2}{b^2}\right) \right] + (1 \leftrightarrow 2)$$



What is the eikonal?

- Eikonal = scattering generator (See also Riccardo's talk) [JH Kim, JWK, Lee; Gonzo, Shi]

$$\mathcal{O}_{\text{out}} = e^{\{\chi, \bullet\}}[\mathcal{O}_{\text{in}}] = \mathcal{O}_{\text{in}} + \{\chi, \mathcal{O}_{\text{in}}\} + \frac{1}{2!} \{\chi, \{\chi, \mathcal{O}_{\text{in}}\}\} + \frac{1}{3!} \{\chi, \{\chi, \{\chi, \mathcal{O}_{\text{in}}\}\}\} + \dots$$

- (KMOC) + ($S = e^{\frac{i}{\hbar} N}$) + ($\hbar \rightarrow 0$) [Kosower, Maybee, O'Connell; Lehmann, Symanzik, Zimmermann; Damgaard, Plante, Vanhove; Hansen]

$$\mathcal{O}_{\text{out}} = S^\dagger \mathcal{O}_{\text{in}} S = e^{-\frac{i}{\hbar} N} \mathcal{O}_{\text{in}} e^{+\frac{i}{\hbar} N} = \sum_{n=0}^{\infty} \frac{1}{n!} \left(-\frac{i}{\hbar}\right)^n \underbrace{[N, [N, [N, \dots, [N, \mathcal{O}_{\text{in}}]] \dots]}_{n \text{ times}}$$

- *Magnus series*: power series expansion of N -matrix [JH Kim, JWK, S Kim, Lee]

- Commutator expansion \Rightarrow causal propagators (tree graphs)
 - Especially relevant for (classical) WQFT [Mogull, Plefka, Steinhoff; Jakobsen, Sauer]

- Real-valued & IR-finite 3PM eikonal from WQFT [JH Kim, JWK, S Kim, Lee]

What is the eikonal?

- Eikonal = scattering generator (See also Riccardo's talk) [JH Kim, JWK, Lee; Gonzo, Shi]

$$\mathcal{O}_{\text{out}} = e^{\{\chi, \bullet\}}[\mathcal{O}_{\text{in}}] = \mathcal{O}_{\text{in}} + \{\chi, \mathcal{O}_{\text{in}}\} + \frac{1}{2!} \{\chi, \{\chi, \mathcal{O}_{\text{in}}\}\} + \frac{1}{3!} \{\chi, \{\chi, \{\chi, \mathcal{O}_{\text{in}}\}\}\} + \dots$$

- (KMOC) + ($S = e^{\frac{i}{\hbar}N}$) + ($\hbar \rightarrow 0$) [Kosower, Maybee, O'Connell; Lehmann, Symanzik, Zimmermann; Damgaard, Plante, Vanhove; Hansen]

$$\mathcal{O}_{\text{out}} = S^\dagger \mathcal{O}_{\text{in}} S = e^{-\frac{i}{\hbar}N} \mathcal{O}_{\text{in}} e^{+\frac{i}{\hbar}N} = \sum_{n=0}^{\infty} \frac{1}{n!} \left(-\frac{i}{\hbar}\right)^n \underbrace{[N, [N, [N, \dots, [N, \mathcal{O}_{\text{in}}]] \dots]}_{n \text{ times}}$$

- *Magnus series*: power series expansion of N -matrix [JH Kim, JWK, S Kim, Lee]

- Commutator expansion \Rightarrow causal propagators (tree graphs)
 - Especially relevant for (classical) WQFT [Mogull, Plefka, Steinhoff; Jakobsen, Sauer]
- Real-valued & IR-finite 3PM eikonal from WQFT [JH Kim, JWK, S Kim, Lee]

The role of nested brackets

- Impact parameter space is “noncommutative” (due to $(b \cdot p_i) = 0$)

$$\{b^\mu, b^\nu\} \doteq \frac{b^\mu w_1^\nu - b^\nu w_1^\mu}{m_1} - \frac{b^\mu w_2^\nu - b^\nu w_2^\mu}{m_2}$$

$$\begin{aligned} \gamma &= -(v_1 \cdot v_2) \\ p_i^\mu &= m_i v_i^\mu \\ w_i \cdot v_j &= -\delta_{ij} \end{aligned}$$

- Poisson brackets = *causality cuts* [JH Kim, JWK, S Kim, Lee]

- Special case of Peierls brackets (\approx covariantised Poisson brackets)
- Join vertices by retarded-minus-advanced Green’s ftn.

$$-i\{V_1, V_2\} = \sum_{\phi} \left[(\Delta V_1^+) \xrightarrow{\phi} (\Delta V_2^+) - (\Delta V_1^+) \xleftarrow{\phi} (\Delta V_2^+) \right]$$

- Poisson brackets can be computed *diagrammatically*
 - $(\chi \text{ from Magnus series}) + (e^{\{\chi, \bullet\}} [p_1^\mu] \text{ as c. cuts}) = (\text{impulse diagrams} = \text{rooted trees})$

[Jakobsen, Mogull, Plefka, Sauer]

The role of nested brackets

- Impact parameter space is “noncommutative” (due to $(b \cdot p_i) = 0$)

$$\{b^\mu, b^\nu\} = \frac{b^\mu w_1^\nu - b^\nu w_1^\mu}{m_1} - \frac{b^\mu w_2^\nu - b^\nu w_2^\mu}{m_2}$$

$$\begin{aligned} \gamma &= -(v_1 \cdot v_2) \\ p_i^\mu &= m_i v_i^\mu \\ w_i \cdot v_j &= -\delta_{ij} \end{aligned}$$

- Poisson brackets = *causality cuts* [JH Kim, JWK, S Kim, Lee]

- Special case of Dirac brackets (commutated Dirac brackets)

- Join vertices

Adopt manifestly orthogonal definition of IP!

$$b^\mu = \Delta x^\mu + (\Delta x \cdot v_1) w_1^\mu + (\Delta x \cdot v_2) w_2^\mu$$

$$\Delta x^\mu := x_1^\mu - x_2^\mu, \quad w_1^\mu := \frac{\gamma v_2^\mu - v_1^\mu}{\gamma^2 - 1}, \quad w_2^\mu := \frac{\gamma v_1^\mu - v_2^\mu}{\gamma^2 - 1}$$

- Poisson brackets

- (χ from Magnus series) + ($e^{\{\chi, \bullet\}} [p_1^\mu]$ as c. cuts) = (impulse diagrams = rooted trees)

[Jakobsen, Mogull, Plefka, Sauer]

The role of nested brackets

- Impact parameter space is “noncommutative” (due to $(b \cdot p_i) = 0$)

$$\{b^\mu, b^\nu\} = \frac{b^\mu w_1^\nu - b^\nu w_1^\mu}{m_1} - \frac{b^\mu w_2^\nu - b^\nu w_2^\mu}{m_2}$$

$$\begin{aligned} \gamma &= -(v_1 \cdot v_2) \\ p_i^\mu &= m_i v_i^\mu \\ w_i \cdot v_j &= -\delta_{ij} \end{aligned}$$

- Poisson brackets = *causality cuts* [JH Kim, JWK, S Kim, Lee]

- Special case of Peierls brackets (\approx covariantised Poisson brackets)
- Join vertices by retarded-minus-advanced Green’s ftn.

$$-i\{V_1, V_2\} = \sum_{\phi} \left[(\Delta V_1^+) \xrightarrow{\phi} (\Delta V_2^+) - (\Delta V_1^+) \xleftarrow{\phi} (\Delta V_2^+) \right]$$

- Poisson brackets can be computed *diagrammatically*
 - $(\chi \text{ from Magnus series}) + (e^{\{\chi, \bullet\}} [p_1^\mu] \text{ as c. cuts}) = (\text{impulse diagrams} = \text{rooted trees})$

[Jakobsen, Mogull, Plefka, Sauer]

The role of nested brackets

- Impact parameter space is “noncommutative” (due to $(b \cdot p_i) = 0$)

$$\{b^\mu, b^\nu\} = \frac{b^\mu w_1^\nu - b^\nu w_1^\mu}{m_1} - \frac{b^\mu w_2^\nu - b^\nu w_2^\mu}{m_2}$$

$$\gamma = -(v_1 \cdot v_2)$$

$$p_i^\mu = m_i v_i^\mu$$

$$w_i \cdot v_j = -\delta_{ij}$$

- Poisson brackets = *causality cuts* [JH Kim, JWK, S Kim, Lee]

- Special case of Peierls brackets (\approx covariantised Poisson brackets)
- Join vertices by retarded-minus-advanced Green’s ftn.

$$-i\{V_1, V_2\} = \sum_{\phi} \left[(\Delta V_1^+) \xrightarrow{\phi} (\Delta V_2^+) - (\Delta V_1^+) \xleftarrow{\phi} (\Delta V_2^+) \right]$$

- Poisson brackets can be computed *diagrammatically*
 - $(\chi$ from Magnus series) + $(e^{\{\chi, \bullet\}} [p_1^\mu]$ as c. cuts) = (impulse diagrams = rooted trees)
[Jakobsen, Mogull, Plefka, Sauer]

Nested brackets in twistor WL

[JH Kim, JWK, Lee; JH Kim, JWK, S Kim, Lee]

$$\Delta_{(1)} p_1^\mu = \dots \begin{array}{c} \blacksquare \\ \vdots \\ \blacksquare \end{array} \dots = \{\mathcal{X}_{(1)}, p_1^\mu\}, \quad \Delta_{(2)} p_1^\mu = \dots \begin{array}{c} \blacksquare \\ \vdots \\ \blacksquare \end{array} \begin{array}{c} \rightarrow \\ \vdots \\ \rightarrow \end{array} \begin{array}{c} \blacksquare \\ \vdots \\ \blacksquare \end{array} \dots + \dots \begin{array}{c} \blacksquare \\ \vdots \\ \blacksquare \end{array} \begin{array}{c} \vdots \\ \vdots \\ \vdots \end{array} \begin{array}{c} \blacksquare \\ \vdots \\ \blacksquare \end{array} \dots$$

$$i\mathcal{X}_{(1)} = \dots \begin{array}{c} \blacksquare \\ \vdots \\ \blacksquare \end{array} \dots, \quad i\mathcal{X}_{(2)} = \dots \begin{array}{c} \blacksquare \\ \vdots \\ \blacksquare \end{array} \begin{array}{c} \text{---} \\ \vdots \\ \text{---} \end{array} \begin{array}{c} \blacksquare \\ \vdots \\ \blacksquare \end{array} \dots + (1 \leftrightarrow 2).$$

Nested brackets in twistor WL

[JH Kim, JWK, Lee; JH Kim, JWK, S Kim, Lee]

$$\begin{aligned}
 \Delta_{(1)} p_1^\mu &= \text{[diagram: vertical wavy line with squares]} = \{\mathcal{X}_{(1)}, p_1^\mu\}, & \Delta_{(2)} p_1^\mu &= \text{[diagram: two vertical wavy lines with a triangle pointing right]} + \text{[diagram: two vertical wavy lines with a triangle pointing left]} \dots \\
 & & & \downarrow \\
 & & & \text{[diagram: two vertical wavy lines with a triangle pointing right]} = \text{[diagram: two vertical wavy lines]} + \frac{1}{2} \left(\text{[diagram: two vertical wavy lines with a crossed triangle]} \right) \\
 i\mathcal{X}_{(1)} &= \text{[diagram: vertical wavy line with squares]}, & i\mathcal{X}_{(2)} &= \text{[diagram: two vertical wavy lines with squares]} + (1 \leftrightarrow 2) \dots \\
 & & & \text{(ret.)} = \text{(sym.)} + \frac{1}{2} (\text{ret.} - \text{adv.}) \\
 & & & = \text{(sym.)} + \frac{1}{2} (\text{causality cut})
 \end{aligned}$$

Nested brackets in twistor WL

[JH Kim, JWK, Lee; JH Kim, JWK, S Kim, Lee]

$$\Delta_{(1)} p_1^\mu = \dots \begin{array}{c} \text{---} \\ | \\ \text{---} \end{array} \dots = \{\mathcal{X}_{(1)}, p_1^\mu\}, \quad \Delta_{(2)} p_1^\mu = \dots \begin{array}{c} \text{---} \blacktriangleright \text{---} \\ | \quad | \\ \text{---} \quad \text{---} \end{array} \dots + \dots \begin{array}{c} \text{---} \quad \text{---} \\ | \quad | \\ \text{---} \quad \text{---} \end{array} \dots \blacktriangleright \dots$$

$$\dots \begin{array}{c} \text{---} \text{---} \\ | \quad | \\ \text{---} \quad \text{---} \end{array} \dots + \dots \begin{array}{c} \text{---} \quad \text{---} \\ | \quad | \\ \text{---} \quad \text{---} \end{array} \dots = \{\mathcal{X}_{(2)}, p_1^\mu\}$$

(sym.) = (eikonal)

$$\dots \begin{array}{c} \text{---} \times \text{---} \\ | \quad | \\ \text{---} \quad \text{---} \end{array} \dots + \dots \begin{array}{c} \text{---} \quad \text{---} \\ | \quad | \\ \text{---} \quad \text{---} \end{array} \dots \times \dots = -i \left\{ \dots \begin{array}{c} \text{---} \quad \text{---} \\ | \quad | \\ \text{---} \quad \text{---} \end{array} \dots, \dots \begin{array}{c} \text{---} \text{---} \\ | \quad | \\ \text{---} \quad \text{---} \end{array} \dots \right\} = \{\mathcal{X}_{(1)}, \{\mathcal{X}_{(1)}, p_1^\mu\}\}$$

(c. cut) = (nested P.B.)

Nested brackets in twistor WL

[JH Kim, JWK, Lee; JH Kim, JWK, S Kim, Lee]

$$\Delta_{(1)} p_1^\mu = \begin{array}{c} \dots \\ \text{---} \blacksquare \text{---} \\ | \\ \text{wavy line} \\ | \\ \blacksquare \text{---} \\ \dots \end{array} = \{\chi_{(1)}, p_1^\mu\}$$

(1PL eikonal)

$$\begin{array}{c} \dots \\ \text{---} \blacksquare \text{---} \blacksquare \text{---} \\ | \quad | \\ \text{wavy line} \quad \text{wavy line} \\ | \quad | \\ \blacksquare \text{---} \blacksquare \text{---} \\ \dots \end{array} + \begin{array}{c} \dots \\ \blacksquare \text{---} \dots \blacksquare \text{---} \\ | \quad | \\ \text{wavy line} \quad \text{wavy line} \\ | \quad | \\ \blacksquare \text{---} \blacksquare \text{---} \\ \dots \end{array} = \{\chi_{(2)}, p_1^\mu\}$$

(2PL eikonal)

$$\mathcal{O}_{\text{out}} = e^{\{\chi, \bullet\}}[\mathcal{O}_{\text{in}}] = \mathcal{O}_{\text{in}} + \underbrace{\{\chi, \mathcal{O}_{\text{in}}\}} + \frac{1}{2!} \underbrace{\{\chi, \{\chi, \mathcal{O}_{\text{in}}\}\}} + \dots$$

(c. cut) = (nested P.B.)

$$\begin{array}{c} \dots \\ \text{---} \blacksquare \text{---} \times \text{---} \blacksquare \text{---} \\ | \quad | \\ \text{wavy line} \quad \text{wavy line} \\ | \quad | \\ \blacksquare \text{---} \blacksquare \text{---} \\ \dots \end{array} + \begin{array}{c} \dots \\ \blacksquare \text{---} \dots \blacksquare \text{---} \\ | \quad | \\ \text{wavy line} \quad \text{wavy line} \\ | \quad | \\ \blacksquare \text{---} \blacksquare \text{---} \\ \dots \end{array} = -i \left\{ \begin{array}{c} \dots \\ \text{---} \blacksquare \text{---} \\ | \\ \text{wavy line} \\ | \\ \blacksquare \text{---} \\ \dots \end{array}, \begin{array}{c} \dots \\ \text{---} \blacksquare \text{---} \\ | \\ \text{wavy line} \\ | \\ \blacksquare \text{---} \\ \dots \end{array} \right\} = \{\chi_{(1)}, \{\chi_{(1)}, p_1^\mu\}\}$$

Remarks on nested brackets

- The reorganisation extends to spinning observables
 - Impulse & spin kick @ $\mathcal{O}(G^3 S^4)$ (See also Mao's talk) [Akpinar, Cordero, Kraus, Smirnov, Zeng]
 - Dirac brackets = Poisson brackets (for variables manifestly satisfying constraints)

- Nested brackets implement rotations [JWK]

$$\{\chi, \bullet\} = \sum_{i=1,2} \{\chi, p_i^\mu\} \frac{\partial}{\partial p_i^\mu} + \sum_{i=1,2} \{\chi, S_i^{\mu\nu}\} \frac{\partial}{\partial S_i^{\mu\nu}} + \{\chi, b^\mu\} \frac{\partial}{\partial b^\mu}$$

- Conventional understanding: saddle-point shift [Di Vecchia, Heissenberg, Russo, Veneziano; Luna, Moynihan, O'Connell, Ross]
- Nested brackets are (causality) cuts
 - “KMOC cuts” as LO rotation contributions [Caron-Huot, Giroux, Hannesdottir, Mizera; Georgoudis, Heissenberg, Russo; Bini, Damour, De Angelis, Geralico, Herderschee, Roiban, Teng; Alessio, Gonzo, Shi]

Remarks on nested brackets

- The reorganisation extends to spinning observables
 - Impulse & spin kick @ $\mathcal{O}(G^3 S^4)$ (See also Mao's talk) [Akpinar, Cordero, Kraus, Smirnov, Zeng]
 - Dirac brackets = Poisson brackets (for variables manifestly satisfying constraints)

- Nested brackets implement rotations [JWK]

$$\{\chi, \bullet\} = \sum_{i=1,2} \{\chi, p_i^\mu\} \frac{\partial}{\partial p_i^\mu} + \sum_{i=1,2} \{\chi, S_i^{\mu\nu}\} \frac{\partial}{\partial S_i^{\mu\nu}} + \{\chi, b^\mu\} \frac{\partial}{\partial b^\mu}$$

- Conventional understanding: saddle-point shift [Di Vecchia, Heissenberg, Russo, Veneziano; Luna, Moynihan, O'Connell, Ross]
- Nested brackets are (causality) cuts
 - “KMOC cuts” as LO rotation contributions [Caron-Huot, Giroux, Hannesdottir, Mizera; Georgoudis, Heissenberg, Russo; Bini, Damour, De Angelis, Geralico, Herderschee, Roiban, Teng; Alessio, Gonzo, Shi]

Remarks on nested brackets

- The reorganisation extends to spinning observables
 - Impulse & spin kick @ $\mathcal{O}(G^3 S^4)$ (See also Mao's talk) [Akpinar, Cordero, Kraus, Smirnov, Zeng]
 - Dirac brackets = Poisson brackets (for variables manifestly satisfying constraints)

- Nested brackets implement rotations [JWK]

$$\{\chi, \bullet\} = \sum_{i=1,2} \{\chi, p_i^\mu\} \frac{\partial}{\partial p_i^\mu} + \sum_{i=1,2} \{\chi, S_i^{\mu\nu}\} \frac{\partial}{\partial S_i^{\mu\nu}} + \{\chi, b^\mu\} \frac{\partial}{\partial b^\mu}$$

- Conventional understanding: saddle-point shift [Di Vecchia, Heissenberg, Russo, Veneziano; Luna, Moynihan, O'Connell, Ross]
- Nested brackets are (causality) cuts
 - “KMOC cuts” as LO rotation contributions [Caron-Huot, Giroux, Hannesdottir, Mizera; Georgoudis, Heissenberg, Russo; Bini, Damour, De Angelis, Geralico, Herderschee, Roiban, Teng; Alessio, Gonzo, Shi]

Scattering generator approach to 3PM impulse

- The 3PM eikonal [Damgaard, Plante, Vanhove; Brandhuber, Chen, Travaglini, Wen; JH Kim, **JWK**, S Kim, Lee]

$$\chi_{(1)} = -Gm_1m_2 \frac{2\gamma^2 - 1}{\sqrt{\gamma^2 - 1}} \log |b^2| ,$$

$$\chi_{(2)} = +\frac{3\pi}{4} G^2 m_1 m_2 (m_1 + m_2) \frac{5\gamma^2 - 1}{\sqrt{\gamma^2 - 1}} \frac{1}{|b^2|^{1/2}} ,$$

$$\chi_{(3)} = \frac{G^3 m_1 m_2}{3\sqrt{\gamma^2 - 1} |b^2|} \left[(m_1^2 + m_2^2) X_{(3,0)} + m_1 m_2 (X_{(3,1c)} + X_{(3,1r)}) \right] ,$$

$$X_{(3,0)} = \frac{(64\gamma^6 - 120\gamma^4 + 60\gamma^2 - 5)}{(\gamma^2 - 1)^2} ,$$

$$X_{(3,1c)} = \frac{2\gamma(36\gamma^6 - 114\gamma^4 + 132\gamma^2 - 55)}{(\gamma^2 - 1)^2} - \frac{12(4\gamma^4 - 12\gamma^2 - 3)}{\sqrt{\gamma^2 - 1}} \operatorname{arccosh}(\gamma) ,$$

$$X_{(3,1r)} = (2\gamma^2 - 1)^2 \left[-2 \frac{(5\gamma^2 - 8)}{(\gamma^2 - 1)^{3/2}} + \frac{6\gamma(2\gamma^2 - 3)}{(\gamma^2 - 1)^2} \operatorname{arccosh}(\gamma) \right] .$$

Scattering generator approach to 3PM impulse

- Reference available [Herrmann, Parra-Martinez, Ruf, Zeng; Jakobsen, Mogull, Plefka, Sauer; Kälin, Liu, Neef, Porto]

- Misses (radiation loss) = (momentum lost by radiated graviton)

$$\Delta_{(3)} p_1^\mu = \Delta_{(3,c)} p_1^\mu + \Delta_{(3,rr)} p_1^\mu + \Delta_{(3,rl)} p_1^\mu$$

$$e^{\{\chi, \bullet\}} p_1^\mu \Big|_{\mathcal{O}(G^3)} = \Delta_{(3,c)} p_1^\mu + \Delta_{(3,rr)} p_1^\mu$$

$$\Delta_{(3,rl)} p_1^\mu = (v_2 \cdot P_{\text{rad}}) w_2^\mu, \quad P_{\text{rad}}^\mu := -\Delta_{(3)} p_1^\mu - \Delta_{(3)} p_2^\mu$$

[2207.00569]

(Humboldt3PM // ListToPlus) - impulse3PMsimp (* exactly radiation loss term *)

$$\frac{\gamma (33 - 112 \gamma^2 + 165 \gamma^4 - 70 \gamma^6) G^3 m_1^2 m_2^2 \pi \text{ArcCosh}[\gamma] (u_\perp)_1^i}{16 (-1 + \gamma^2)^2 (b_i b^i)^{3/2}} + \frac{(-5 + \gamma (76 + 5 \gamma (-30 + \gamma (12 + 7 \gamma)))) G^3 m_1^2 m_2^2 \pi \text{Log}\left[\frac{1+\gamma}{2}\right] (u_\perp)_1^i}{8 \sqrt{-1 + \gamma^2} (b_i b^i)^{3/2}} + \frac{(-1151 + \gamma (3336 + \gamma (-3148 + 3 \gamma (304 + \gamma (-113 + 2 (92 - 35 \gamma) \gamma)))) G^3 m_1^2 m_2^2 \pi (u_\perp)_1^i}{48 ((-1 + \gamma^2) (b_i b^i))^{3/2}}$$

Scattering generator approach to 3PM impulse

- Reference available [Herrmann, Parra-Martinez, Ruf, Zeng; Jakobsen, Mogull, Plefka, Sauer; Kälin, Liu, Neef, Porto]

- Misses (radiation loss) = (momentum lost by radiated graviton)

$$\Delta_{(3)} p_1^\mu = \Delta_{(3,c)} p_1^\mu + \Delta_{(3,rr)} p_1^\mu + \Delta_{(3,rl)} p_1^\mu$$

$$e^{\{\chi, \bullet\}} p_1^\mu \Big|_{\mathcal{O}(G^3)} = \Delta_{(3,c)} p_1^\mu + \Delta_{(3,rr)} p_1^\mu$$

[2207.00569]

$$\Delta_{(3,rl)} p_1^\mu = (v_2 \cdot P_{\text{rad}}) w_2^\mu, \quad P_{\text{rad}}^\mu := -\Delta_{(3)} p_1^\mu - \Delta_{(3)} p_2^\mu$$

`(Humboldt3PM // ListToPlus)` - `impulse3PMsimp` (* exactly radiation loss term *)

$$\frac{\gamma (33 - 112 \gamma^2 + 165 \gamma^4 - 70 \gamma^6) G^3 m_1^2 m_2^2 \pi \text{ArcCosh}[\gamma] (u_\perp)_1^i (-5 + \gamma (76 + 5 \gamma (-30 + \gamma (12 + 7 \gamma)))) G^3 m_1^2 m_2^2 \pi \text{Log}\left[\frac{1+\gamma}{2}\right] (u_\perp)_1^i}{16 (-1 + \gamma^2)^2 (b_i b^i)^{3/2}} + \frac{8 \sqrt{-1 + \gamma^2} (b_i b^i)^{3/2}}{(-1151 + \gamma (3336 + \gamma (-3148 + 3 \gamma (304 + \gamma (-113 + 2 (92 - 35 \gamma) \gamma)))) G^3 m_1^2 m_2^2 \pi (u_\perp)_1^i}{48 ((-1 + \gamma^2) (b_i b^i))^{3/2}}$$

Scattering generator approach to 3PM impulse

- Reference available [Herrmann, Parra-Martinez, Ruf, Zeng; Jakobsen, Mogull, Plefka, Sauer; Kälin, Liu, Neef, Porto]

- Misses (radiation loss) = (momentum lost by radiated graviton)

$$\Delta_{(3)} p_1^\mu = \Delta_{(3,c)} p_1^\mu + \Delta_{(3,rr)} p_1^\mu + \Delta_{(3,rl)} p_1^\mu$$

$$e^{\{\chi, \bullet\}} p_1^\mu \Big|_{\mathcal{O}(G^3)} = \Delta_{(3,c)} p_1^\mu + \Delta_{(3,rr)} p_1^\mu$$

$$\Delta_{(3,rl)} p_1^\mu = (v_2 \cdot P_{\text{rad}}) w_2^\mu, \quad P_{\text{rad}}^\mu := -\Delta_{(3)} p_1^\mu - \Delta_{(3)} p_2^\mu$$

- Which contributions did we miss?

- (Graviton phase space contributions) = (diagrams from graviton c. cuts)

- Computed from *radiation eikonal* as $\Delta_{(3,rl)} p_1^\mu = \frac{1}{2} \{ \chi_{(1.5)}^{\text{rad}}, \{ \chi_{(1.5)}^{\text{rad}}, p_1^\mu \}_\circ \}_\circ$ [JWK]

The radiation eikonal

[JWK]

- Missing contributions from graviton phase space / causality cuts
 - No vertex rules where graviton legs can be attached \Rightarrow add new vertices!
 - Formal implementation: turn on background $H_{\mu\nu}$
 - Practical implementation: add diagrams with external graviton legs
- Expanded set of eikonal-derived PM observables
 - Impulse radiation loss @ 3PM
 - Radiative observables: scattering waveform & radiated energy
 - New perspective on “KMOC cuts” in scattering WFs [Caron-Huot, Giroux, Hannesdottir, Mizera; Georgoudis, Heissenberg, Russo; Bini, Damour, De Angelis, Geralico, Herderschee, Roiban, Teng; [Alessio, Gonzo, Shi](#)]
 - Initial disagreement btw. PM & PN [Brandhuber, Brown, Chen, De Angelis, Gowdy, Travaglini; Herderschee, Roiban, Teng; Elkhidir, O’Connell, Sergola, Vazquez-Holm; Georgoudis, Heissenberg; Bini, Damour, Geralico]

The radiation eikonal

[JWK]

- Missing contributions from graviton phase space / causality cuts
 - No vertex rules where graviton legs can be attached \Rightarrow add new vertices!
 - Formal implementation: turn on background $H_{\mu\nu}$
 - Practical implementation: add diagrams with external graviton legs
- Expanded set of eikonal-derived PM observables
 - Impulse radiation loss @ 3PM
 - Radiative observables: scattering waveform & radiated energy
 - New perspective on “KMOC cuts” in scattering WFs [Caron-Huot, Giroux, Hannesdottir, Mizera; Georgoudis, Heissenberg, Russo; Bini, Damour, De Angelis, Geralico, Herderschee, Roiban, Teng; [Alessio, Gonzo, Shi](#)]
 - Initial disagreement btw. PM & PN [Brandhuber, Brown, Chen, De Angelis, Gowdy, Travaglini; Herderschee, Roiban, Teng; Elkhidir, O’Connell, Sergola, Vazquez-Holm; Georgoudis, Heissenberg; Bini, Damour, Geralico]

The radiation eikonal

[JWK]

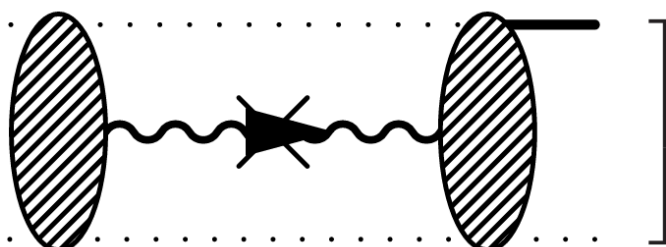
- Missing contributions from graviton phase space / causality cuts
 - No vertex rules where graviton legs can be attached \Rightarrow add new vertices!
 - Formal implementation: turn on background $H_{\mu\nu}$
 - Practical implementation: add diagrams with external graviton legs
- Expanded set of eikonal-derived PM observables
 - Impulse radiation loss @ 3PM
 - Radiative observables: scattering waveform & radiated energy
 - New perspective on “KMOC cuts” in scattering WFs [Caron-Huot, Giroux, Hannesdottir, Mizera; Georgoudis, Heissenberg, Russo; Bini, Damour, De Angelis, Geralico, Herderschee, Roiban, Teng; [Alessio, Gonzo, Shi](#)]
 - Initial disagreement btw. PM & PN [Brandhuber, Brown, Chen, De Angelis, Gowdy, Travaglini; Herderschee, Roiban, Teng; Elkhidir, O’Connell, Sergola, Vazquez-Holm; Georgoudis, Heissenberg; Bini, Damour, Geralico]

Impulse radiation loss @ 3PM

[JWK]

- Computed as $\Delta_{(3,rl)} p_1^\mu = \frac{1}{2} \{ \chi_{(1.5)}^{\text{rad}}, \{ \chi_{(1.5)}^{\text{rad}}, p_1^\mu \}_\circ \}_\circ$, *exact match!*

$$\Delta_{(3,rl)} P_1^\mu = \frac{1}{2} \int \frac{d^D k}{(2\pi)^D} \frac{d^D q}{(2\pi)^D} \left(\frac{\partial \chi_{(1.5)}^{H^1}}{\partial H_{\alpha\beta}(k)} \right) \times \{ H_{\alpha\beta}(k), H_{\lambda\sigma}(q) \} \times \left(\frac{\partial}{\partial H_{\lambda\sigma}(q)} \left[-\frac{\partial \chi_{(1.5)}^{H^1}}{\partial X_{1\mu}} \right] \right)$$

$$= \frac{1}{2} \left[\text{Diagram} \right]$$


$$= \frac{G^3 m_1^2 m_2^2}{|b^2|^{3/2}} \left[\frac{\pi}{16} \frac{\gamma(33 - 112\gamma^2 + 165\gamma^4 - 70\gamma^6)}{(\gamma^2 - 1)^2} \text{arccosh}(\gamma) \right. \\ \left. + \frac{\pi}{8} \frac{(-5 + 76\gamma - 150\gamma^2 + 60\gamma^3 + 35\gamma^4)}{(\gamma^2 - 1)^{1/2}} \log \left(\frac{1 + \gamma}{2} \right) \right. \\ \left. + \frac{\pi}{48} \frac{(-1151 + 3336\gamma - 3148\gamma^2 + 912\gamma^3 - 339\gamma^4 + 552\gamma^5 - 210\gamma^6)}{(\gamma^2 - 1)^{3/2}} \right] w_2^\mu,$$

NLO scattering waveform

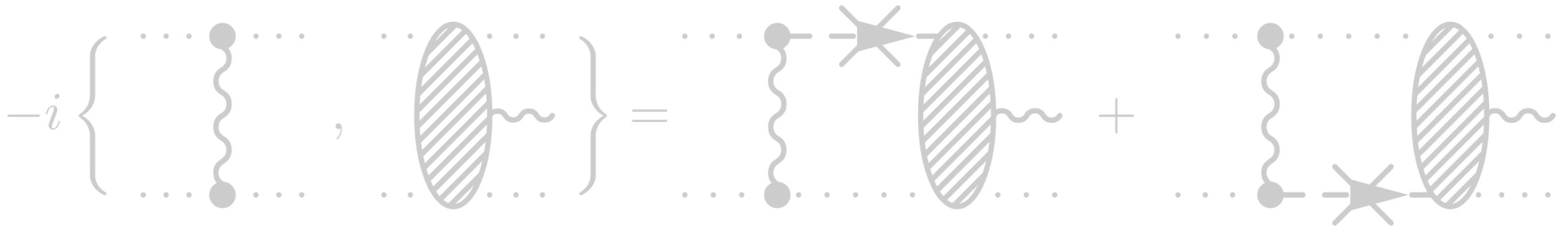
[JWK]

- Nested brackets (= cut contributions) compute rotated LO WF

$$S_{\mu\nu}^{\text{LO}} = \{\chi_{(1.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\}$$

$$S_{\mu\nu}^{\text{NLO}} = \{\chi_{(2.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\} + \frac{1}{2}\{\chi_{(1)}, \{\chi_{(1.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\}\} = \{\chi_{(2.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\} + \frac{1}{2}\{\chi_{(1)}, S_{\mu\nu}^{\text{LO}}\}$$

$$\{\chi_{(1)}, \bullet\} = \{\chi_{(1)}, b^\mu\} \frac{\partial}{\partial b^\mu} + \sum_{i=1,2} \{\chi_{(1)}, p_i^\mu\} \frac{\partial}{\partial p_i^\mu} + \sum_{i=1,2} \{\chi_{(1)}, S_i^{\mu\nu}\} \frac{\partial}{\partial S_i^{\mu\nu}}$$



NLO scattering waveform

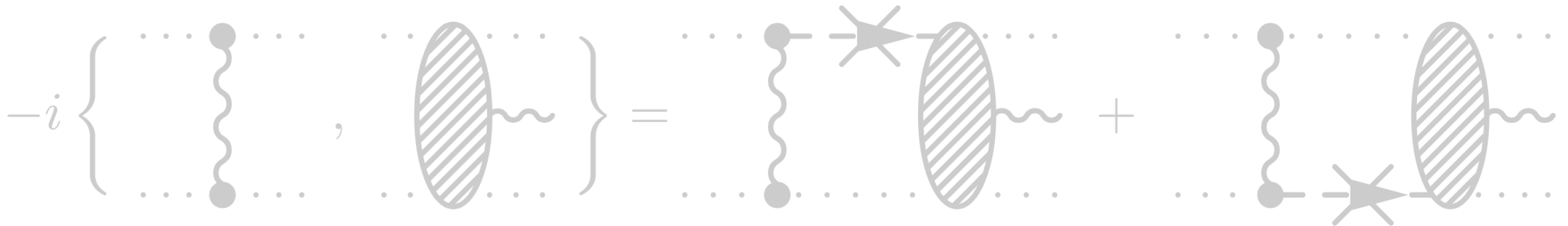
[JWK]

- Nested brackets (= cut contributions) compute rotated LO WF

$$S_{\mu\nu}^{\text{LO}} = \{\chi_{(1.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\}$$

$$S_{\mu\nu}^{\text{NLO}} = \{\chi_{(2.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\} + \frac{1}{2} \{\chi_{(1)}, \{\chi_{(1.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\}\} = \{\chi_{(2.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\} + \frac{1}{2} \{\chi_{(1)}, S_{\mu\nu}^{\text{LO}}\}$$

$$\{\chi_{(1)}, \bullet\} = \{\chi_{(1)}, b^\mu\} \frac{\partial}{\partial b^\mu} + \sum_{i=1,2} \{\chi_{(1)}, p_i^\mu\} \frac{\partial}{\partial p_i^\mu} + \sum_{i=1,2} \{\chi_{(1)}, S_i^{\mu\nu}\} \frac{\partial}{\partial S_i^{\mu\nu}}$$



NLO scattering waveform

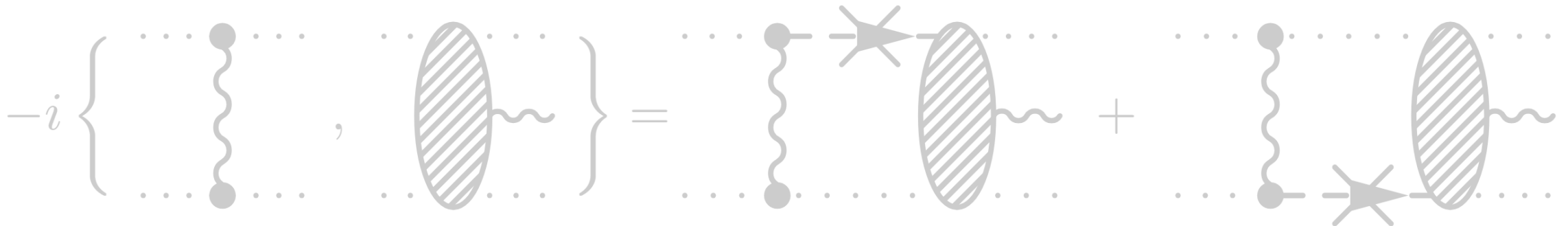
[JWK]

- Nested brackets (= cut contributions) compute rotated LO WF

$$S_{\mu\nu}^{\text{LO}} = \{\chi_{(1.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\}$$

$$S_{\mu\nu}^{\text{NLO}} = \{\chi_{(2.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\} + \frac{1}{2} \{\chi_{(1)}, \{\chi_{(1.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\}\} = \{\chi_{(2.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\} + \frac{1}{2} \{\chi_{(1)}, S_{\mu\nu}^{\text{LO}}\}$$

$$\{\chi_{(1)}, \bullet\} = \underbrace{\{\chi_{(1)}, b^\mu\}}_{\text{IP rotation by LO SA}} \frac{\partial}{\partial b^\mu} + \sum_{i=1,2} \underbrace{\{\chi_{(1)}, p_i^\mu\}}_{\text{LO impulse}} \frac{\partial}{\partial p_i^\mu} + \sum_{i=1,2} \underbrace{\{\chi_{(1)}, S_i^{\mu\nu}\}}_{\text{LO spin kick}} \frac{\partial}{\partial S_i^{\mu\nu}}$$



NLO scattering waveform

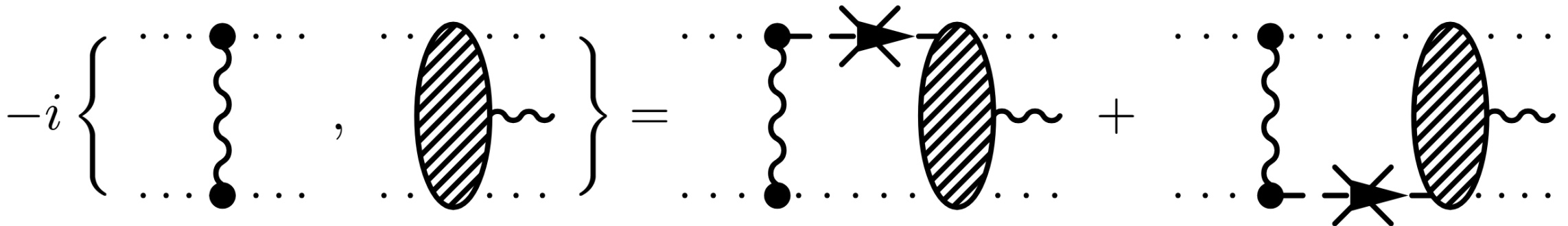
[JWK]

- Nested brackets (= cut contributions) compute rotated LO WF

$$S_{\mu\nu}^{\text{LO}} = \{\chi_{(1.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\}$$

$$S_{\mu\nu}^{\text{NLO}} = \{\chi_{(2.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\} + \frac{1}{2} \{\chi_{(1)}, \{\chi_{(1.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\}\} = \{\chi_{(2.5)}^{H^1}, H_{\mu\nu}^{\text{in}}\} + \frac{1}{2} \{\chi_{(1)}, S_{\mu\nu}^{\text{LO}}\}$$

$$\{\chi_{(1)}, \bullet\} = \underbrace{\{\chi_{(1)}, b^\mu\}}_{\text{IP rotation by LO SA}} \frac{\partial}{\partial b^\mu} + \sum_{i=1,2} \underbrace{\{\chi_{(1)}, p_i^\mu\}}_{\text{LO impulse}} \frac{\partial}{\partial p_i^\mu} + \sum_{i=1,2} \underbrace{\{\chi_{(1)}, S_i^{\mu\nu}\}}_{\text{LO spin kick}} \frac{\partial}{\partial S_i^{\mu\nu}}$$



Take-home messages

- *Spin resummation* will become important for next-gen GW detectors
- Eikonal = scattering generator [JH Kim, JWK, Lee; Gonzo, Shi]
 - Computed by the *Magnus series*
$$\mathcal{O}_f = e^{\{\chi, \bullet\}}[\mathcal{O}] = \mathcal{O} + \{\chi, \mathcal{O}\} + \frac{1}{2!}\{\chi, \{\chi, \mathcal{O}\}\} + \frac{1}{3!}\{\chi, \{\chi, \{\chi, \mathcal{O}\}\}\} + \dots$$
- New understanding of brackets
 - Impact parameter space is “noncommutative”: $\{b^\mu, b^\nu\} \neq 0$
 - Poisson brackets = *causality cuts* (retarded-minus-advanced Green’s fn)
- Radiation eikonal: radiative effects including *dissipation*
 - Adding graviton DOFs to phase space / Poisson brackets

Take-home messages

- *Spin resummation* will become important for next-gen GW detectors

- Eikonal = scattering generator [JH Kim, JWK, Lee; Gonzo, Shi]

- Computed by the *Magnus series*

$$\mathcal{O}_f = e^{\{\chi, \bullet\}}[\mathcal{O}] = \mathcal{O} + \{\chi, \mathcal{O}\} + \frac{1}{2!}\{\chi, \{\chi, \mathcal{O}\}\} + \frac{1}{3!}\{\chi, \{\chi, \{\chi, \mathcal{O}\}\}\} + \dots$$

- New understanding of brackets

- Impact parameter space is “noncommutative”: $\{b^\mu, b^\nu\} \neq 0$
- Poisson brackets = *causality cuts* (retarded-minus-advanced Green’s fn)

- Radiation eikonal: radiative effects including *dissipation*

- Adding graviton DOFs to phase space / Poisson brackets

Take-home messages

- *Spin resummation* will become important for next-gen GW detectors
- Eikonal = scattering generator [JH Kim, JWK, Lee; Gonzo, Shi]
 - Computed by the *Magnus series*
$$\mathcal{O}_f = e^{\{\chi, \bullet\}}[\mathcal{O}] = \mathcal{O} + \{\chi, \mathcal{O}\} + \frac{1}{2!}\{\chi, \{\chi, \mathcal{O}\}\} + \frac{1}{3!}\{\chi, \{\chi, \{\chi, \mathcal{O}\}\}\} + \dots$$
- New understanding of brackets
 - Impact parameter space is “noncommutative”: $\{b^\mu, b^\nu\} \neq 0$
 - Poisson brackets = *causality cuts* (retarded-minus-advanced Green’s fn)
- Radiation eikonal: radiative effects including *dissipation*
 - Adding graviton DOFs to phase space / Poisson brackets

Take-home messages

- *Spin resummation* will become important for next-gen GW detectors
- Eikonal = scattering generator [JH Kim, JWK, Lee; Gonzo, Shi]
 - Computed by the *Magnus series*
$$\mathcal{O}_f = e^{\{\chi, \bullet\}}[\mathcal{O}] = \mathcal{O} + \{\chi, \mathcal{O}\} + \frac{1}{2!}\{\chi, \{\chi, \mathcal{O}\}\} + \frac{1}{3!}\{\chi, \{\chi, \{\chi, \mathcal{O}\}\}\} + \dots$$
- New understanding of brackets
 - Impact parameter space is “noncommutative”: $\{b^\mu, b^\nu\} \neq 0$
 - Poisson brackets = *causality cuts* (retarded-minus-advanced Green’s fn)
- Radiation eikonal: radiative effects including *dissipation*
 - Adding graviton DOFs to phase space / Poisson brackets

Some future directions

- B2B continuation for *nonseparable* HJ systems [Cho, Kälin, Porto; Gonzo, Shi]
 - Eikonal-based WF models for bound binaries? [JWK, Patil, Scheopner, Steinhoff (WIP)]
- Local-in-time / hereditary separation? [Dlapa, Kälin, Liu, Porto; Buonanno, Mogull, Patil, Pompili]
 - (hereditary contributions) $\stackrel{?}{=}$ (radiation eikonal contributions)
- Nonlinear radiation eikonal (e.g. $\chi_{(2)}^{H^2}$: quadratic in $H_{\mu\nu}$)
 - Contributes to NNLO WF / 5PM impulse [Driesse, Jakobsen, Mogull, Plefka, Sauer, Usovitsch; Bern, Herrmann, Roiban, Ruf, Smirnov, Smirnov, Zeng]
 - Is 6pt relevant for classical physics? [Cristofoli, Gonzo, Moynihan, O'Connell, Ross, Sergola, White]
- Spin resummation @ 3PL / 3PM (1SF corrections: relevant for EOB)
 - Obstruction: Integrals (Tensor Integral Generating Functions / Twisted FI)
[Feng; Brunello, Crisanti, Giroux, Mastrolia, Smith; Brunello, De Angelis; Chen, JWK, Wang; JH Kim, JWK, Lim (WIP)]
 - Application to tensor reduction of high-rank Feynman integrals?

Some future directions

- B2B continuation for *nonseparable* HJ systems [Cho, Kälin, Porto; Gonzo, Shi]
 - Eikonal-based WF models for bound binaries? [JWK, Patil, Scheopner, Steinhoff (WIP)]
- Local-in-time / hereditary separation? [Dlapa, Kälin, Liu, Porto; Buonanno, Mogull, Patil, Pompili]
 - (hereditary contributions) $\stackrel{?}{=}$ (radiation eikonal contributions)
- Nonlinear radiation eikonal (e.g. $\chi_{(2)}^{H^2}$: quadratic in $H_{\mu\nu}$)
 - Contributes to NNLO WF / 5PM impulse [Driesse, Jakobsen, Mogull, Plefka, Sauer, Usovitsch; Bern, Herrmann, Roiban, Ruf, Smirnov, Smirnov, Zeng]
 - Is 6pt relevant for classical physics? [Cristofoli, Gonzo, Moynihan, O'Connell, Ross, Sergola, White]
- Spin resummation @ 3PL / 3PM (1SF corrections: relevant for EOB)
 - Obstruction: Integrals (Tensor Integral Generating Functions / Twisted FI)
[Feng; Brunello, Crisanti, Giroux, Mastrolia, Smith; Brunello, De Angelis; Chen, JWK, Wang; JH Kim, JWK, Lim (WIP)]
 - Application to tensor reduction of high-rank Feynman integrals?

Some future directions

- B2B continuation for *nonseparable* HJ systems [Cho, Kälin, Porto; Gonzo, Shi]
 - Eikonal-based WF models for bound binaries? [JWK, Patil, Scheopner, Steinhoff (WIP)]
- Local-in-time / hereditary separation? [Dlapa, Kälin, Liu, Porto; Buonanno, Mogull, Patil, Pompili]
 - (hereditary contributions) $\stackrel{?}{=}$ (radiation eikonal contributions)
- Nonlinear radiation eikonal (e.g. $\chi_{(2)}^{H^2}$: quadratic in $H_{\mu\nu}$)
 - Contributes to NNLO WF / 5PM impulse [Driesse, Jakobsen, Mogull, Plefka, Sauer, Usovitsch; Bern, Herrmann, Roiban, Ruf, Smirnov, Smirnov, Zeng]
 - Is 6pt relevant for classical physics? [Cristofoli, Gonzo, Moynihan, O'Connell, Ross, Sergola, White]
- Spin resummation @ 3PL / 3PM (1SF corrections: relevant for EOB)
 - Obstruction: Integrals (Tensor Integral Generating Functions / Twisted FI)
[Feng; Brunello, Crisanti, Giroux, Mastrolia, Smith; Brunello, De Angelis; Chen, JWK, Wang; JH Kim, JWK, Lim (WIP)]
 - Application to tensor reduction of high-rank Feynman integrals?

Some future directions

- B2B continuation for *nonseparable* HJ systems [Cho, Kälin, Porto; Gonzo, Shi]
 - Eikonal-based WF models for bound binaries? [JWK, Patil, Scheopner, Steinhoff (WIP)]
- Local-in-time / hereditary separation? [Dlapa, Kälin, Liu, Porto; Buonanno, Mogull, Patil, Pompili]
 - (hereditary contributions) $\stackrel{?}{=}$ (radiation eikonal contributions)
- Nonlinear radiation eikonal (e.g. $\chi_{(2)}^{H^2}$: quadratic in $H_{\mu\nu}$)
 - Contributes to NNLO WF / 5PM impulse [Driesse, Jakobsen, Mogull, Plefka, Sauer, Usovitsch; Bern, Herrmann, Roiban, Ruf, Smirnov, Smirnov, Zeng]
 - Is 6pt relevant for classical physics? [Cristofoli, Gonzo, Moynihan, O'Connell, Ross, Sergola, White]
- Spin resummation @ 3PL / 3PM (1SF corrections: relevant for EOB)
 - Obstruction: Integrals (Tensor Integral Generating Functions / Twisted FI)
[Feng; Brunello, Crisanti, Giroux, Mastrolia, Smith; Brunello, De Angelis; Chen, JWK, Wang; JH Kim, JWK, Lim (WIP)]
 - Application to tensor reduction of high-rank Feynman integrals?