#### **Non-Oscillation Searches at DUNE ND**

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Path to Dark Sector Discoveries at Neutrino Experiments Colorado State University 5<sup>th</sup> June 2023



## Outline

- Introduction
- Deep Underground Neutrino Experiment (DUNE)
- DUNE Near Detector Complex
- DUNE ND BSM Physics Topics
  - Light Dark Matter
  - Axion-like Particles
  - Neutrino Trident
  - Heavy Neutral Lepton
- Summary

Quick introduction of ND subdetectors, and will point out what makes it good for dark sector or Non-oscillation studies

Not going to address all the details but will focus on:

- 1) Signatures of Signals/Backgrounds
- 2) Tools and simulation workflow
- 3) Currently expected sensitivity



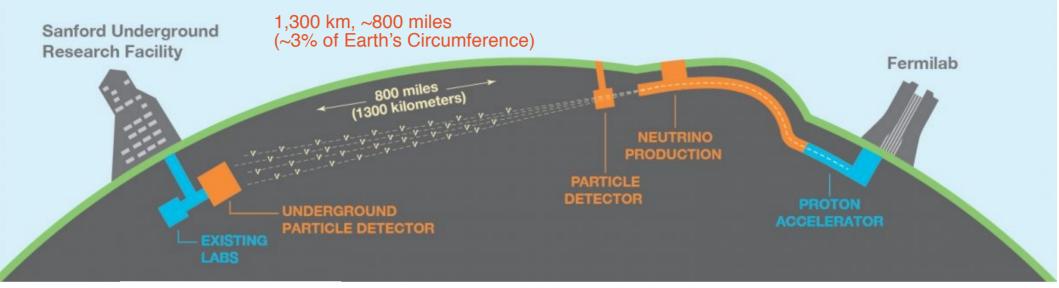
## Introduction

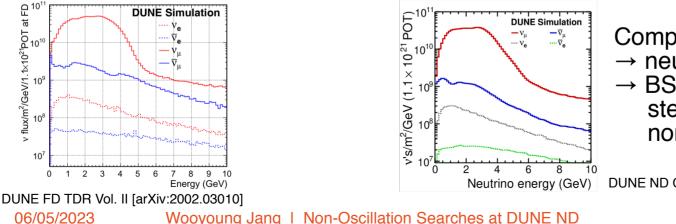
- 1998, neutrino oscillation phenomenon discovered:
  - Neutrino oscillation revealed that  $m_{\nu} \neq 0$ .
  - This is apparently contrary to the assumption of the Standard Model (SM).
  - Therefore, this means that now neutrino physics entered the into the realm of Beyond Standard Model (BSM) physics.
- This demands the precision measurement of <u>neutrino oscillation parameters</u>.
- In experimental particle physics, the word 'precision' often related to high statistics.
  - Large mass, high precision detectors and high intensity neutrino beam with longbaseline.
  - This makes cutting-edge neutrino experiments as a general purpose particle physics machine.





#### **Deep Underground Neutrino Experiment**





Comparing two measurements

- $\rightarrow$  neutrino oscillations
- $\rightarrow$  BSM physics? sterile neutrino, non-standard interaction

DUNE ND CDR [arXiv:2002.03010]



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#### **The LBNF Beam**

- The new PIP-II accelerator will provide 120 GeV proton beam to DUNE through the Main Injector.
- The Phase-I operating beam power is 1.2 MW and it can be upgraded up to 2.4 MW.

Parameter	Protons per cycle	Cycle Time (sec)	Beam Power (MW)
≤ 1.2 MW Operation - Current	Maximum Value	for LBNF	
Proton Beam Energy (GeV):	8		
60	7.5E+13	0.7	1.03
80	7.5E+13	0.9	1.07
120	7.5E+13	1.2	1.20

Day-1 configuration

Proton Beam Energy (GeV):			
60	1.5E+14	0.7	2.06
80	1.5E+14	0.9	2.14
120	1.5E+14	1.2	2.40

#### Fermilab Accelerator Complex



[Heidi Schellman, ICHEP 2018]

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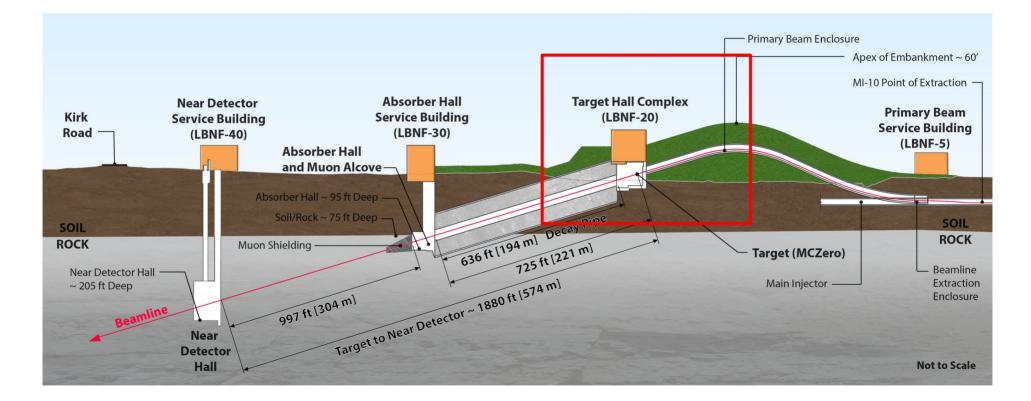
Phase-II

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# **Target / Hadron Absorber / ND**





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# **Target Hall Complex**

- Neutrino production.
- Collimate charged mesons using 3 steps of focusing horns system to focus neutrino beams and get wide bandwidth of neutrino energy.
- ~52% beam power will be used in here.
- Produce dark secto particles?

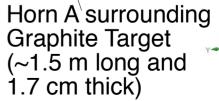
[G4LBNE] is the simulation tool for this study.

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#### Decay Pipe

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Horn C

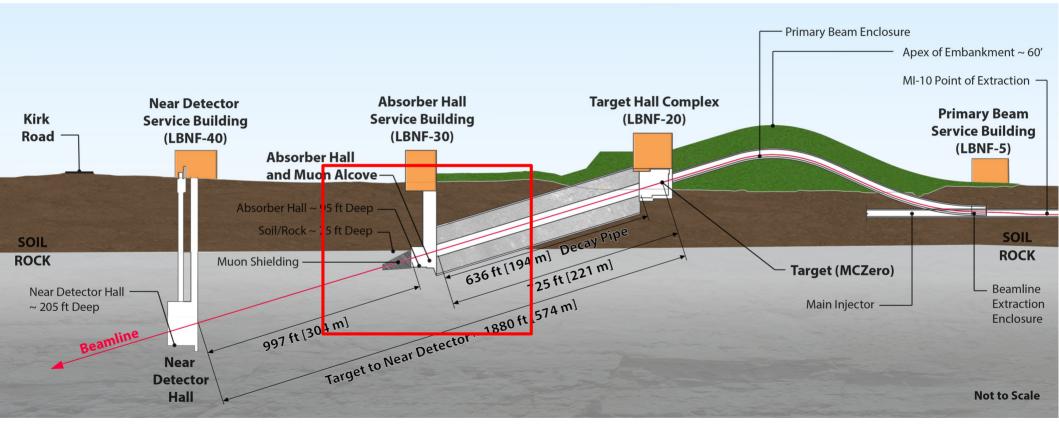


rimary bea



Horn B

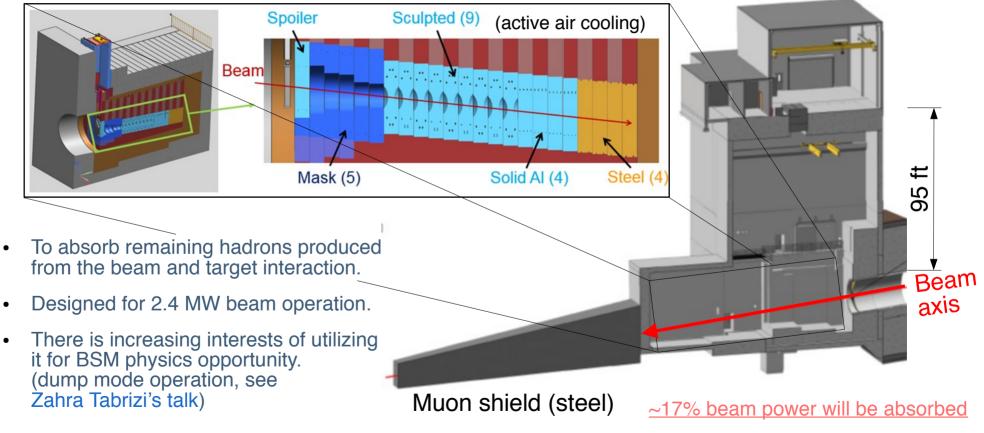
# **Target / Hadron Absorber / ND**





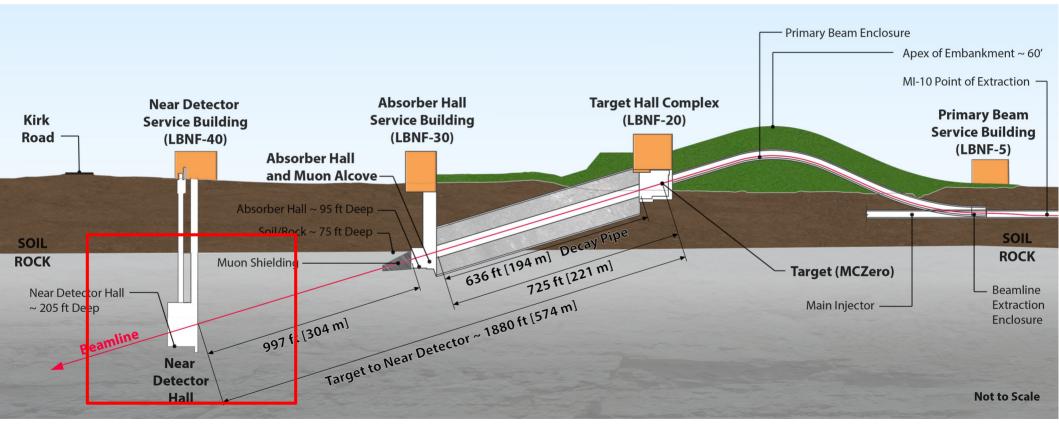
## **Hadron Absorber**







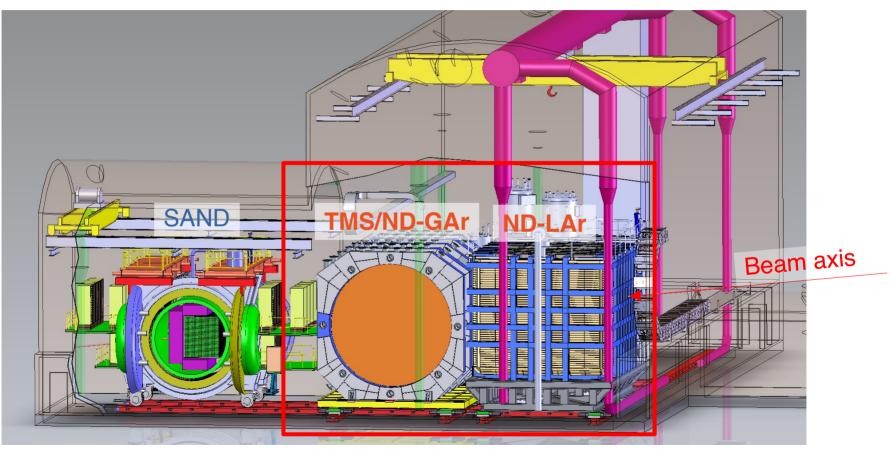
# **Target / Hadron Absorber / ND**







#### **DUNE Near Detector Complex**



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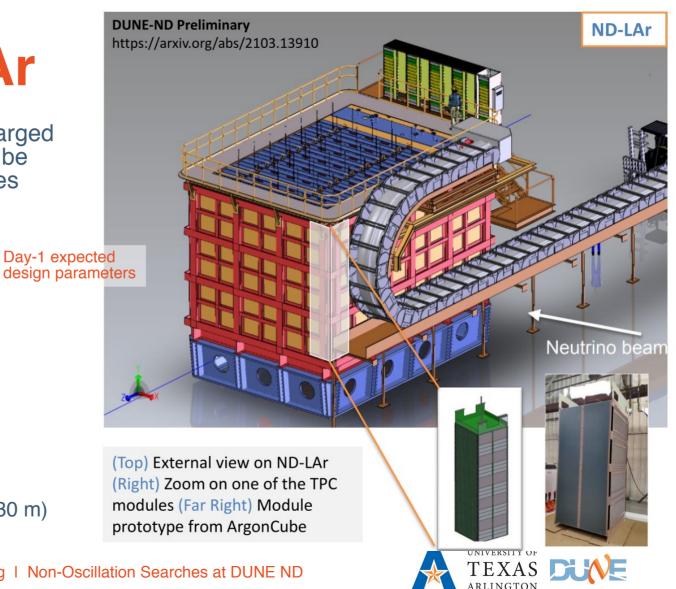
# **DUNE ND-LAr**

- LArTPC provides excellent charged particle measurement and will be used to measure neutrino fluxes precisely.
  - Spatial resolution ~ 1 mm<sup>3</sup>
  - Angular resolution ~ 5 mrad
  - Energy resolution  $\sim 5\%$
- ~150 t active material mass
  - Modular design
- $4 \text{ m} \times 3 \text{ m} \times 5 \text{ m}$  Active Vol.  $(3 \text{ m} \times 2 \text{ m} \times 3 \text{ m} \text{ Fiducial V.})$
- Detector movable Off-Axis (max ~30 m)

#### [Federico Battisti, ICHEP 2022]

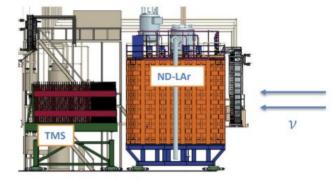


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# **DUNE ND-GAr / ND-TMS**

- Allows ND provides excellent momentum resolution for charged particle tracks ( $\mu$ , p) that were produced in ND-LAr and through interplay with ND-LAr it will provide precise neutrino flux measurement at ND site.
- ND-TMS will likely to be the DUNE day-1 configuration and will be upgraded to ND-GAr in DUNE Phase-II.

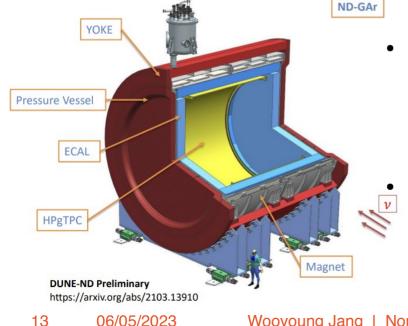


Forty 15-mm

thick steel layers

TMS

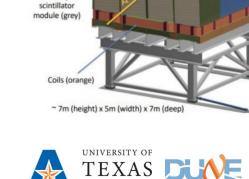
48-strip



 In the viewpoint of dark sector study, we expect opportunities to catch the decays of longlived unstable dark sector particles.

Detector movable Off-Axis as well as ND-LAr

[Federico Battisti, ICHEP 2022]



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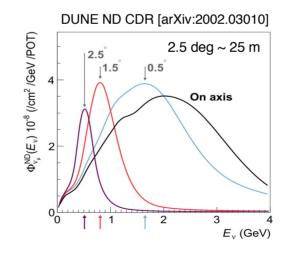
Sixty 40-mm

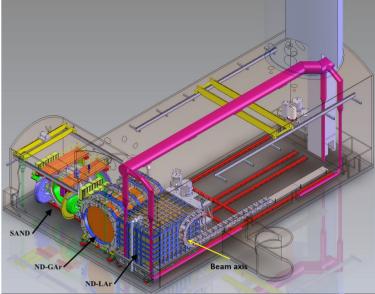
thick steel lavers

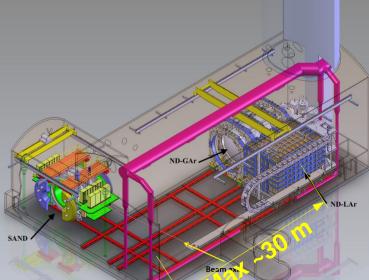
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### **DUNE ND PRISM**

- The main purpose of the PRISM system is to control systematic uncertainties of neutrino fluxes originated from materials along the neutrino pathways.
- By locating ND subdetectors to the Off-Axis position, this decreases neutrino flux and it plays a role in <u>reducing backgrounds</u> in **BSM** studies.
- Movable up to ~30 m.





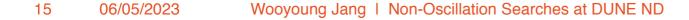


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#### What Makes DUNE ND Well-Suited for BSM Study?

- Direct Observation Signature from the Beams
  - Require high beam flux  $\rightarrow P_{beam} = 1.2 2.4 \text{ MW}$
  - Large mass, high density for scattering signatures  $\rightarrow$  ND LAr w/M<sub>A</sub> = 150 t, fine segmentation
  - Large volume, low density for decay signatures → ND-GAr ND w/ECAL and magnetized precision tracking
  - Capable near detector complex → Combinations of ND-LAr + ND-GAr on PRISM + SAND for fine tracking & beam monitoring
  - Low threshold energy → Both ND and FD TPC threshold ALA few MeV
- What do we need to know?
  - Signal flux and realistic behaviors in the detector
  - Neutrino flux and their interactions in the detector as bck  $\rightarrow$  ND Sub-detectors with PRISM

[Jae Yu, Snowmass CSS 2022]





### Selected BSM Topics at DUNE ND

- High beam power, large detector mass + highly capable, precision near and far detectors with low E threshold make DUNE a BSM machine
  - Recall the signal to background ratio grows by the sqrt of the beam power
  - Near Detector Searches è Take advantage of high beam power-
    - **Axion-like Particles (ALP)** ٠
    - Low mass Dark Matter (LDM) ٠
    - Heavy Neutral Leptons (HNL) ٠
    - **Neutrino Trident** ٠
    - Dark Photon ٠

٠

Milli-charge Particles (mCP) ٠

#### Please check arXiv:2103.13910 and EPJ C.81, 322 (2021) for more physics. And many many more ..

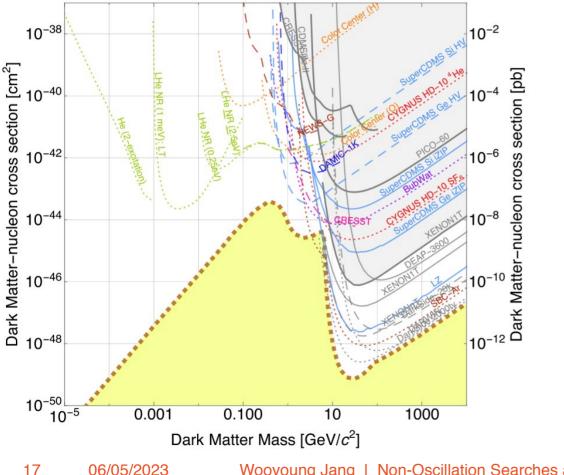
- Far Detector Searches take advantage of ND, large VA FD & long baseline
  - Sterile neutrino searches .
  - Non-standard Interactions, Non-Unitarity, CPT violation ٠
  - Large Extra Dimensions (LED) ٠
  - Boosted Dark Matter (BDM) & Inelastic Boosted Dark Matter (iBDM)
  - And many many more... ٠
- Strong collaboration of theorists and experimentalists essential
- Some of these topics covered in EPJ C.81, 322 (2021)
  - 16 Wooyoung Jang 1 Non-Oscillation Searches at DUNE ND 06/05/2023

[Jae Yu, Snowmass CSS 2022]



#### Topics I will cover today.

# **Light Dark Matter**

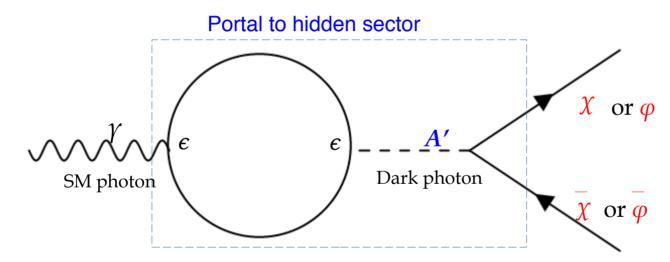


- Phase space for WIMP searches.
- ~1 GeV, a threshold to capture DM-nucleon scattering event in LXe/LAr/Ge/GHe detectors.
- Sub-GeV territory is remain unexplored.
- Many new ideas other than WIMP
  - WIMP mass lower than 2 GeV can not explain dark matter relic abundance.
  - Hidden sector or portal interaction scenario.



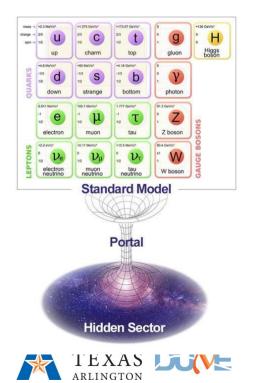
### **Light Dark Matter – Introduction**

- In our light dark matter model, we assume that standard model photon is kinetically mixed with 'dark photon'.
- Dark matter particles can be produced by decay of dark photon through the 'portal interaction'.



**Beam intensity** ~ Photon flux ~ **Dark matter flux** 

**DUNE**, equipped with **high-intensity proton beam** provides a great opportunity to test this type of dark matter scenario.

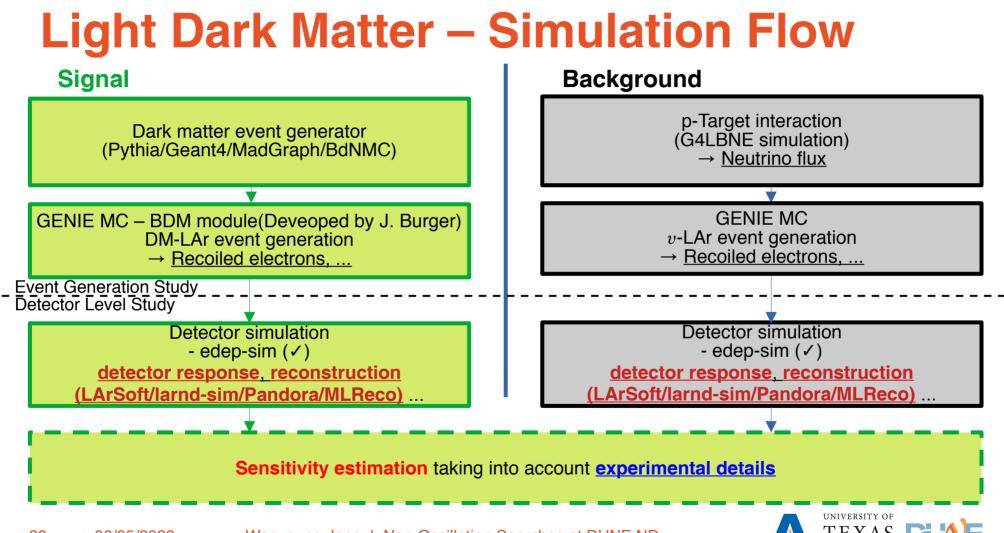


# Light Dark Matter – Event Generator

- PYTHIA: Neutral meson flux → Dark matter flux [Valentina De Romeri et. al. Phys. Rev. D 100 095010]
  - Contribution from secondary interactions?
- Signal Event Generator DMG4 [M. Bondi et. al. CPC 269 (2021) 108129]
  - Good
    - Versatility and detailed simulation powered by Geant4
  - Limitation
    - Programmed for electron beam dump experiment (NA64)
    - Supports DM production only through <u>bremsstrahlung</u> and e<sup>-</sup>+ e<sup>+</sup> <u>annihilation</u>.
       π<sup>0</sup> decay is not supported.
       [Matthew J. Dolan et al Phys. Rev. Lett. **121** 101801 (2018)]
- MadGraph
- BdNMC

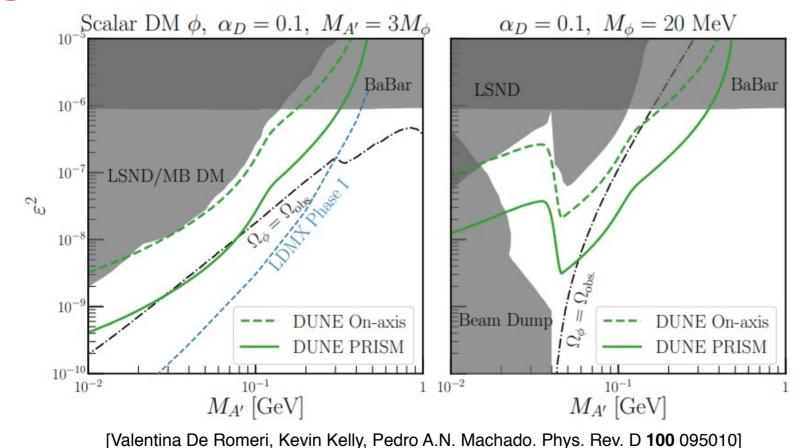
#### We didn't had chance to simulate nucleon recoil yet...

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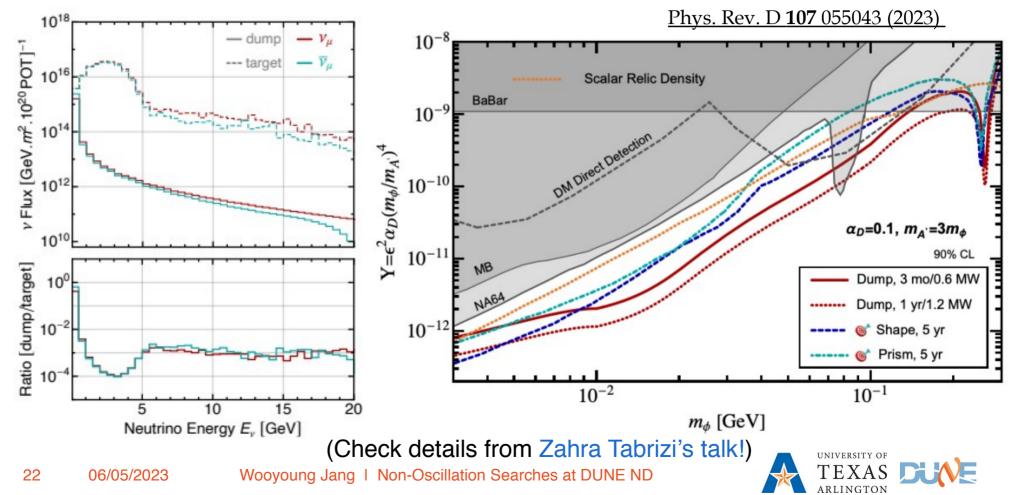
#### Light Dark Matter – Sensitivity



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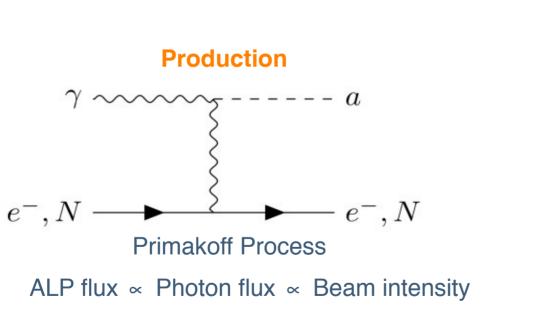
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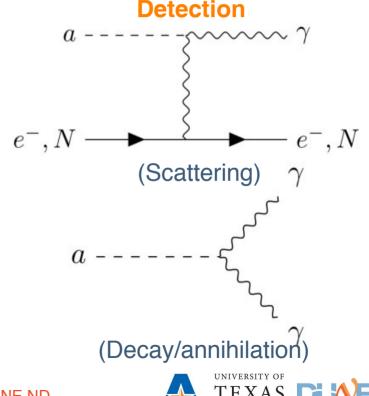
#### Light Dark Matter – Sensitivity (cont'd)



## **Axion-like Particles**

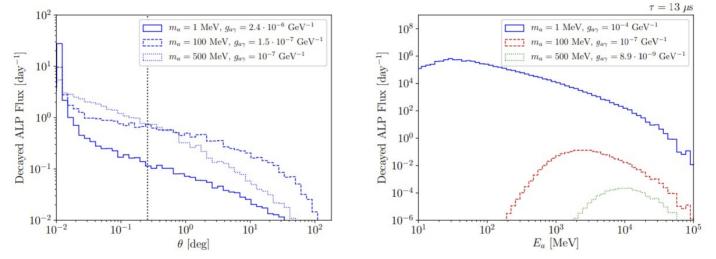
• ALPs are general extension of QCD axion to solve strong CP problem and at the same time an excellent dark matter candidate.





### **Axion-like Particles – Tools**

• Standalone Geant4 simulation to obtain photon flux → Convert it to ALP



[Brdar et. al. Phys. Rev. Lett. 126, 201801 (2021)]

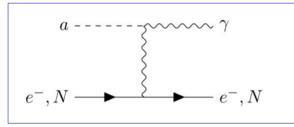
DMG4 supports ALP production

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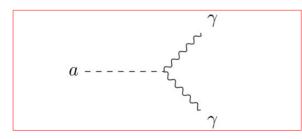


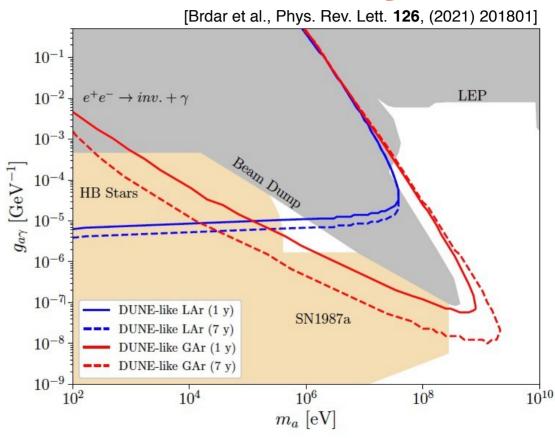
## **Axion-like Particles – Sensitivity**

- Detection through scattering
  - ND-LAr



- Detection through decay
  - ND-GAr (Phase-II)



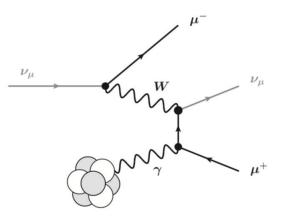


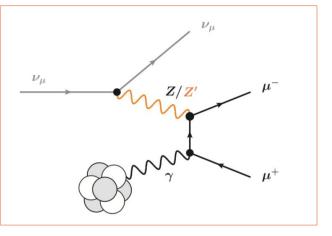


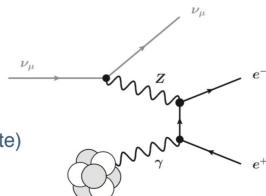
# **Neutrino Trident**

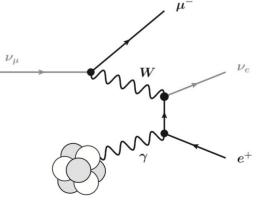
- Neutrino tridents rare SM weak processes
- Signature: a pair of charged leptons
- Δ suggests unknown gauge boson (Z') couplings.

 $\Delta$ =(SM expected rate – Observed event rate)











#### Neutrino Trident – Backgrounds Backgrounds

- Muon tridents main background:  $v_{\mu} CC 1\pi 0p$ 
  - Difficult to distinguish pions and muons using calorimetry information, range etc.
  - Need to be identified from other event properties.
- Electron tridents main background: NC π<sup>0</sup>
  - Photon showers from resulting pi-zero decay could mimic the expected two electron showers.
- Have started studying the muon trident backgrounds using initial ν<sub>μ</sub> CC 1π0p events (some plots follow). Order of magnitude more stats on its way!

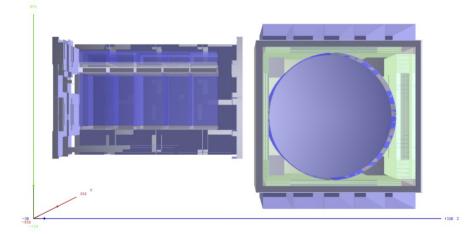
#### [M. Wallbank, DUNE BSM WG Meeting]



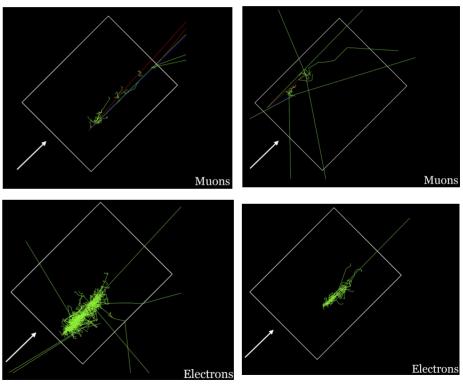
#### M Wallbank (Cincinnati)

## **Neutrino Trident – Tools**

- Standalone Event Generator Interfaced to Geant4
- Background simulation: GENIE



ND-LAr geometry (GDML based implement)



[Figures from J. Martín-Albo, DUNE CM Jan. 31. 2018] Wooyoung Jang I Non-Oscillation Searches at DUNE ND

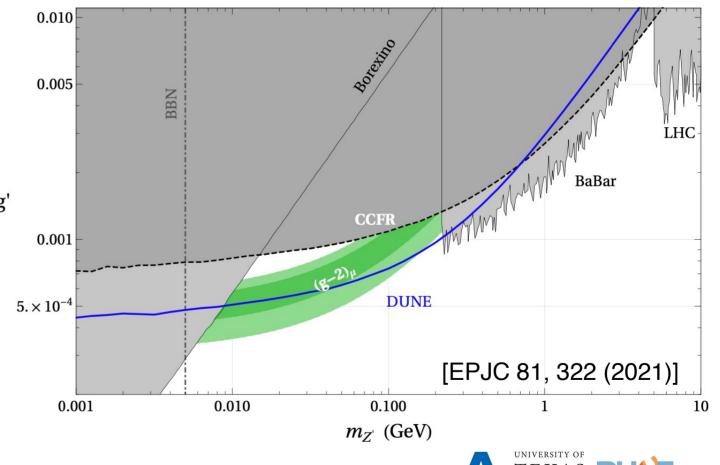


## **Neutrino Trident – Sensitivity**

DUNE is sensitive in the region where  $(g-2)_{\mu}$  anomaly can be explained at  $1\sigma$  and  $2\sigma$ 

Shaded in gray: CMS( $pp \rightarrow \mu^+\mu^- Z \rightarrow \mu^+\mu^-\mu^+\mu^-$ ) BaBar( $e^+e^- \rightarrow \mu^+\mu^- Z \rightarrow \mu^+\mu^-\mu^+\mu^-$ ) Borexino(solar *v*-e<sup>-</sup> scattering) CCFR(Tevatron trident meas.) g'

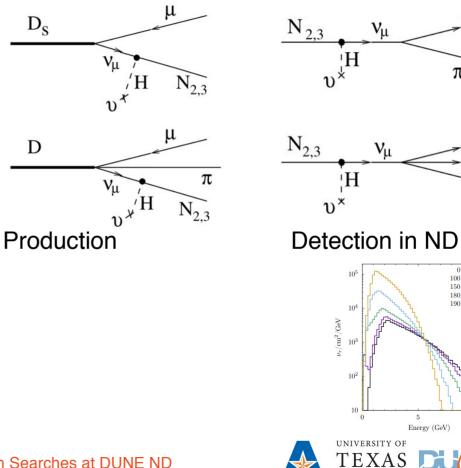
Theoretical constraints: BBN



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## **Heavy Neutral Leptons**

- Energetic collision of proton beam to the target often create heavy mesons such as D.
- Such heavy mesons could be a source of Heavy Neutral Leptons.
- HNLs are assumed to be stable enough to be able to stay alive until it fly ~500 m from the target to ND and then they decay-inflight in the ND.
- HNL Signature:
  - Charged leptons + lighter mesons



D

π

μ

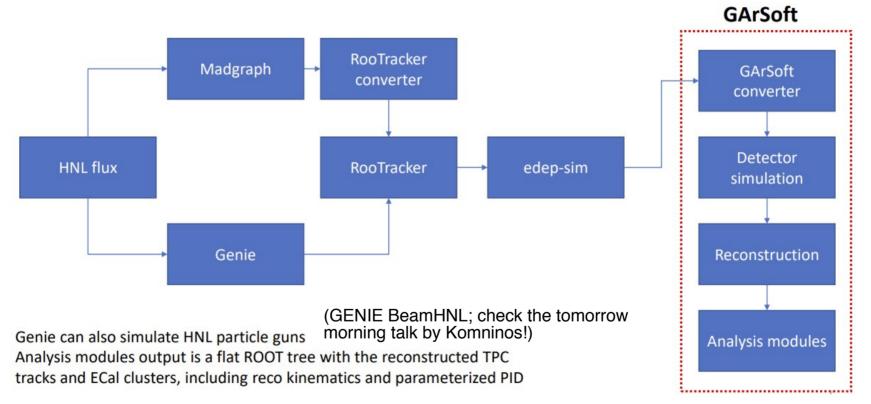
 $v_e$ 

0 MeV -

100 MeV -150 MeV -180 MeV 190 MeV

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# **Heavy Neutral Leptons (2)**



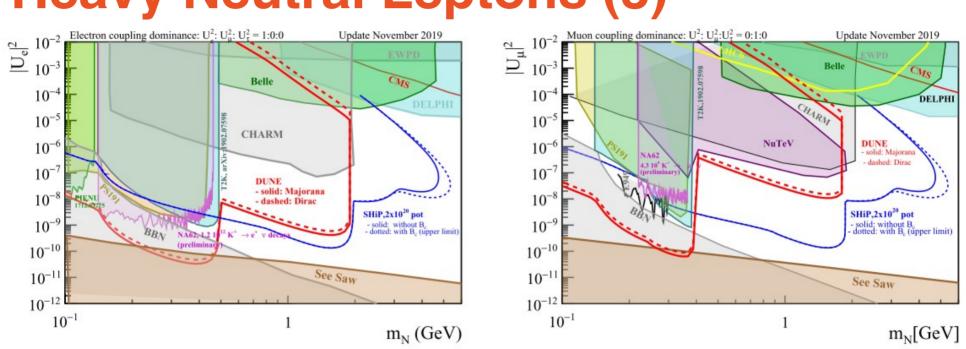
#### [Haifa Sfar, DUNE BSM WG Meeting]



### **Heavy Neutral Leptons (3)**

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90% CL sensitivity regions for dominant mixings  $|U_{eN}|^2$  and  $|U_{\mu N}|^2$  are presented for DUNE ND (red). solid: Majorana neutrino / dashed: Dirac neutrino





#### **Event Signatures and List of Tools**

	Signal	Background	Evt. Gen.	Det. Sim.
LDM	$\chi e^- \rightarrow \chi e^-$	$v_e e^- \rightarrow v_e e^-$	Pythia/Geant4/	Edep-Sim
	$\chi N \to \chi N$	$vN \rightarrow vN$	MadGraph/BdNMC/ DMG4	Larnd-sim
ALP	(S) γ <b>e</b> , γ <b>N</b>	$v$ coherent, NC w/ $\pi^0$ , $v_e$ CC w/ $\pi^0$ , etc	Pythia/Geant4/DMG4	Edep-Sim Larnd-sim
	(D) <b>үү</b>			
Trident	$\nu \rightarrow \nu e^- e^+$	$v_{\mu}N \rightarrow v_{\mu}\pi N' (v CC)$ NC $\pi^{0}$	Geant4 (Standalone) GENIE	Edep-Sim Larnd-sim
	$\nu \rightarrow \nu \mu^{-} \mu^{+}$			
	$\nu \rightarrow \nu e \mu$			
HNL	$N \rightarrow \nu e^- e^+$	<i>v</i> CC + mis-ID p, <i>v</i> e CC w/ π <sup>0</sup>	Standalone HNL flux GENIE MadGraph5	Edep-Sim GArSoft
	$N \rightarrow \nu \mu^{-} \mu^{+}$			
	$N \rightarrow \nu \gamma$			
	$N \rightarrow \nu e \mu$			
	$N \rightarrow \nu \pi^0$			
	$N  ightarrow e \pi$			
	$N  ightarrow  u\pi$			



## Conclusion

- DUNE is a powerful BSM machine as well as it is excellent for neutrino physics.
- High-intensity proton beam and precision detectors of DUNE provides great opportunity to explore the dark sector or BSM physics.
- We have discussed the capabilities of DUNE ND subdetectors and a variety of Non-oscillatory physics topics that can be unveiled by utilizing them.
- Collaboration between theorist and experimentalist is essential to accomplish this.

