



Contribution ID: 16

Type: not specified

## Searching for gravitational waves from inspiraling planetary-mass primordial black holes in LIGO O3 data

*Wednesday 21 May 2025 14:50 (20 minutes)*

Gravitational waves from sub-solar mass inspiraling compact objects would provide almost smoking-gun evidence for primordial black holes (PBHs). We perform the first search for inspiraling planetary-mass compact objects in equal-mass and highly asymmetric mass-ratio binaries using data from the first half of the LIGO-Virgo-KAGRA third observing run. Though we do not find any significant candidates, we determine the maximum luminosity distance reachable with our search to be of  $\mathcal{O}(0.1 - 100)$  kpc, and corresponding model-independent upper limits on the merger rate densities to be  $\mathcal{O}(10^3 - 10^{-7}) \text{ kpc}^{-3} \text{ yr}^{-1}$  for systems with chirp masses of  $\mathcal{O}(10^{-4} - 10^{-2}) M_{\odot}$ , respectively. Furthermore, we interpret these rate densities as arising from PBH binaries and constrain the fraction of dark matter that such objects could comprise. {For equal-mass PBH binaries, we find that these objects would compose less than 4-100% of DM for PBH masses of  $10^{-2} M_{\odot}$  to  $2 \times 10^{-3} M_{\odot}$ , respectively. For asymmetric binaries, assuming one black hole mass corresponds to a peak in the mass function at  $2.5 M_{\odot}$ , a PBH dark-matter fraction of 10% and a second, much lighter PBH, we constrain the mass function of the second PBH to be less than 1 for masses between  $1.5 \times 10^{-5} M_{\odot}$  and  $2 \times 10^{-4} M_{\odot}$ . Our constraints, released on Zenodo, are robust enough to be applied to any PBH or exotic compact object binary formation models, and complement existing microlensing results.

**Author:** MILLER, Andrew

**Presenter:** MILLER, Andrew

**Session Classification:** PBHs and GWs