

CALCULATION OF PRIMORDIAL n_G FROM USR INFLATION

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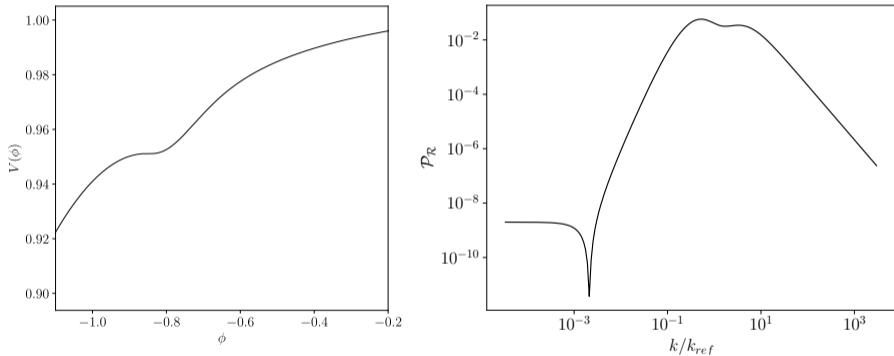
OUTLINE

1. Setup: USR inflation
2. The δN formalism: non-Gaussianity from non-linearity
3. Going beyond: non-Gaussian $\delta\phi$

Based on:

- 2406.02417 (G. Ballesteros, T. Konstandin, **APR**, M. Pierre, J. Rey)
- 2412.14106 (G. Ballesteros, J. Gambín Egea, T. Konstandin, **APR**, M. Pierre, J. Rey)

1. THE USR SETUP | Inflection point



Inflection point in $V(\phi) \rightarrow$ Violation of slow-roll \rightarrow
Peak in $\mathcal{P}_{\mathcal{R}}$ and non-Gaussianity

2. THE δN FORMALISM

Fully non-linear *classical* relation between ζ and $\delta\phi$, $\delta\pi = \delta\phi'$ to first order in *gradients*

$$\zeta = \delta N \equiv N \left[\{\bar{\phi}(N_i) + \delta\phi, \bar{\pi}(N_i) + \delta\pi\} \rightarrow \bar{\phi}(N_f) \right] - N \left[\{\bar{\phi}(N_i), \bar{\pi}(N_i)\} \rightarrow \bar{\phi}(N_f) \right] ,$$

- No dependence on final time (ζ conserved) [Lyth, Malik, Sasaki '05]
- Number of e-folds N computed solving background EOMs (no \mathbf{x} dependence)
 \leftrightarrow *Separate universe approach* [Sugiyama, Komatsu, Futamase '13]
- Reference N_i chosen once relevant Fourier modes \mathcal{R}_k have frozen
 $\rightarrow \delta\pi = (\pi'/\pi)|_N \delta\phi$, cf. [Jackson et al. '24] (gradient contamination)
- N.B. Gradient neglecting enters at two points in the formula

2. THE δN FORMALISM | Gaussian input

Main idea

Assuming Gaussian $\delta\phi$, non-linear relation with ζ yields non-Gaussian ζ :

$$P[\zeta] = P[\delta\phi] \left| \frac{d\delta\phi}{d\zeta} \right|$$

- Assumption of $\delta\phi$ gaussianity motivated by smallness of self-interactions

[Biagetti et al. '21], [Hooshangi, Namjoo, Noorbala '23]

- Gaussian PDF fully determined by its variance $\sigma_{\delta\phi}^2 = \int_{k_{IR}}^{k(N_i)} d \log k \mathcal{P}_{\delta\phi}(k)$
- This approach provides only *local non-Gaussianity* of ζ

2. THE δN FORMALISM | CR attractor

Constant-roll $\leftrightarrow \epsilon_2 = \frac{2\phi''}{\phi'} = \text{cst} \leftrightarrow$ linear EOM for $\phi(N) \leftrightarrow \phi(N) \propto e^{\frac{\epsilon_2}{2}N}$.

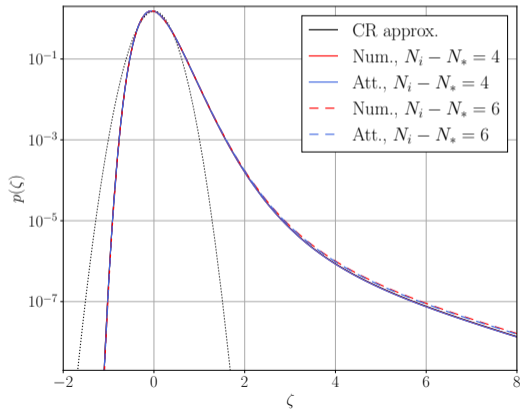
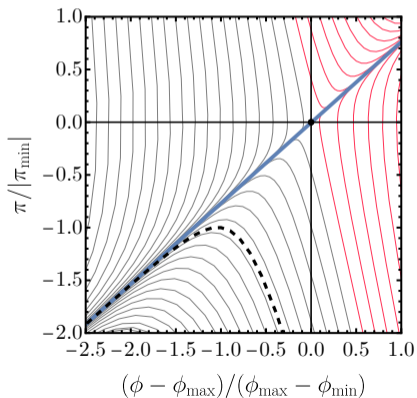
$$\boxed{\delta\phi = \bar{\phi}'(N_i) \left(e^{-\frac{\epsilon_2}{2}\delta N} - 1 \right)} \quad \text{Recall: } \zeta = \delta N$$

Change of variables: $P(\delta N) = P(\delta\phi) \left| \frac{d\delta\phi}{d\delta N} \right|$ ($\delta\phi$: Gaussian with $\sigma_{\delta\phi}^2$)

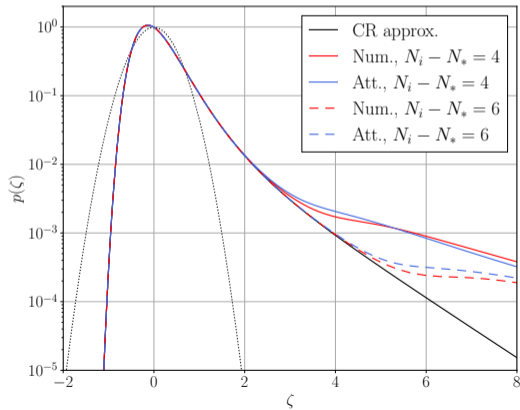
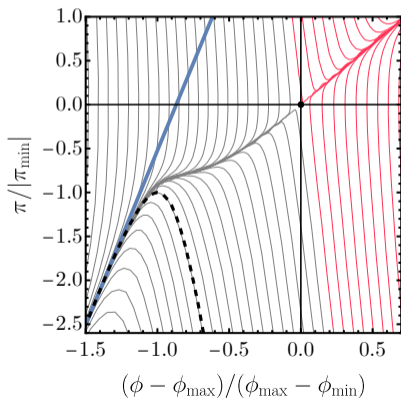
$$P(\zeta) = \frac{1}{\sqrt{2\pi} \sigma_{\zeta}^2} \exp \left[-\frac{1}{2\sigma_{\zeta}^2} \left(\frac{2}{\epsilon_2} \right)^2 \left(1 - e^{-\frac{\epsilon_2}{2}\zeta} \right)^2 - \frac{\epsilon_2}{2} \zeta \right].$$

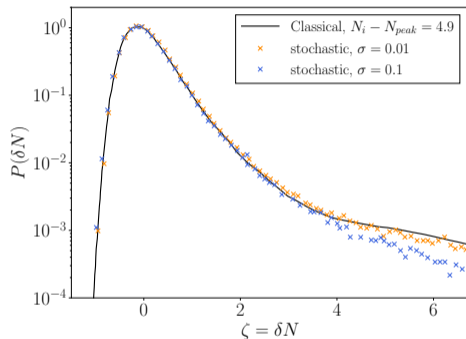
- Small $\epsilon_2\delta N$: $P(\zeta) \sim e^{-\frac{\zeta^2}{2\sigma_{\mathcal{R}}^2}}$. Gaussian with variance $\sigma_{\mathcal{R}}^2$ (agrees with LPT)
- Large $\epsilon_2\delta N$: $\boxed{P(\zeta) \sim e^{-\frac{\epsilon_2}{2}\zeta}}$. For $\epsilon_2 > 0$, exponential tail

N.B.: $\epsilon_2 \propto d\pi/d\phi$. Straight line in phase space \leftrightarrow CR trajectory ☞ [2406.02417]



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Models studied: Same result as classical δN provided $k_{\text{peak}} \simeq \sigma a(N_i)H(N_i)$ [2406.02417] (irrelevant for CR attractor [Tomberg '23])

$$N_i - N_{\text{peak}} \approx -\log(0.01) \approx 4.6 \neq -\log(0.1) \approx 2.3$$

N.B. Stochastic and classical δN do not predict same result for attractor with $\epsilon_2 = 0$ (Exponential vs. Gaussian) [Jackson et al. '24], cf. *first-passage time analysis* [Pattison, Vennin, Assadullahi, Wands '17], [Ezquiaga, García-Bellido, Vennin '20]

3. NON-GAUSSIAN $\delta\phi$ | Main idea

Recent update: 2412.14106

[Ballesteros, Gambín Egea, Konstadin, APR, Pierre, Rey '24 (*arXiv preprint*)]

Question:

- δN formalism: assumes *as input* Gaussian $\delta\phi$
- Computes local NG due to *non-linear* relation between $\delta\phi$ and ζ
- Is $\delta\phi$ at N_i really Gaussian?

Approach to answer:

- Use in-in formalism to compute higher-order (≥ 3) correlators of $\delta\phi$
- Construct non-Gaussian PDF of $\delta\phi$ (*intrinsic* non-Gaussianity)
- Use non-linear δN transformation, check how intrinsic NG affects the PDF of ζ

3. NON-GAUSSIAN $\delta\phi$ | A first estimate

Neglecting gradients,

$$\zeta = \varphi + \frac{\epsilon_2}{4}\varphi^2 + \mathcal{O}(\varphi^3), \quad \varphi = -\delta\phi/\phi'$$

Truncation of δN formula to first non-linear order.

Two first contributions to the bispectrum of ζ :

$$\langle \zeta_{\mathbf{p}} \zeta_{\mathbf{k}} \zeta_{\mathbf{q}} \rangle = \langle \varphi_{\mathbf{p}} \varphi_{\mathbf{k}} \varphi_{\mathbf{q}} \rangle + \frac{\epsilon_2}{4} \langle (\varphi^2)_{\mathbf{p}} \varphi_{\mathbf{k}} \varphi_{\mathbf{q}} \rangle + \mathcal{O}(\varphi^5) \equiv [\mathcal{I}(p, q, k) + \mathcal{N}(p, q, k)] (2\pi)^3 \delta(\mathbf{p} + \mathbf{k} + \mathbf{q})$$

3. NON-GAUSSIAN $\delta\phi$ | A first estimate

$$\langle \zeta_{\mathbf{p}} \zeta_{\mathbf{k}} \zeta_{\mathbf{q}} \rangle = \langle \varphi_{\mathbf{p}} \varphi_{\mathbf{k}} \varphi_{\mathbf{q}} \rangle + \frac{\epsilon_2}{4} \langle (\varphi^2)_{\mathbf{p}} \varphi_{\mathbf{k}} \varphi_{\mathbf{q}} \rangle + \mathcal{O}(\varphi^5) \equiv [\mathcal{I}(p, q, k) + \mathcal{N}(p, q, k)] (2\pi)^3 \delta(\mathbf{p} + \mathbf{k} + \mathbf{q})$$

where

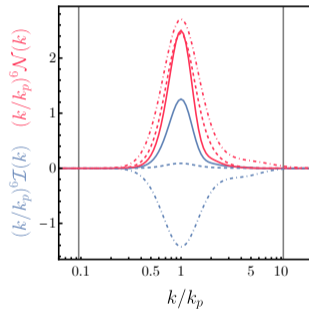
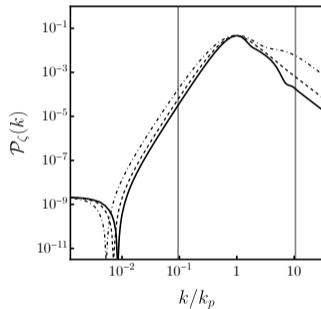
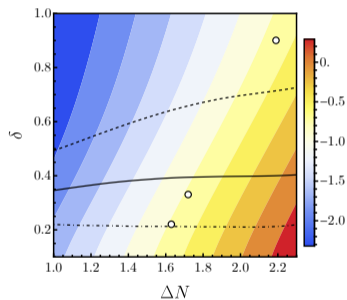
$$\langle \varphi_{\mathbf{p}} \varphi_{\mathbf{k}} \varphi_{\mathbf{q}} \rangle = 2\text{Im} \left[\int dt' \langle 0 | \varphi_{\mathbf{p}} \varphi_{\mathbf{k}} \varphi_{\mathbf{q}} H_I(t') | 0 \rangle \right] \quad (1)$$

$$\frac{\epsilon_2}{4} \langle (\varphi^2)_{\mathbf{p}} \varphi_{\mathbf{k}} \varphi_{\mathbf{q}} \rangle = \frac{\epsilon_2}{2} (|\varphi_{\mathbf{p}}|^2 |\varphi_{\mathbf{q}}|^2 + |\varphi_{\mathbf{p}}|^2 |\varphi_{\mathbf{k}}|^2 + |\varphi_{\mathbf{k}}|^2 |\varphi_{\mathbf{q}}|^2) \quad (2)$$

- (1) First correction to **Gaussianity of $\delta\phi$** . H_I : cubic interaction Hamiltonian in the $\delta\phi$ -gauge. *Intrinsic non-Gaussianity*
- (2) **Local non-Gaussianity** from change of gauge. First term of the non-linear result obtained through δN formalism. *Non-linear non-Gaussianity*

3. NON-GAUSSIAN $\delta\phi$ | A first estimate

Relative size of \mathcal{I} and \mathcal{N} depending on smoothness of SR-USR transition



See details in [\[2406.02417\]](#)

3. NON-GAUSSIAN $\delta\phi$ | PDF of ζ

Next step: impact of first order correction to Gaussianity of $\delta\phi$ in all orders of ζ

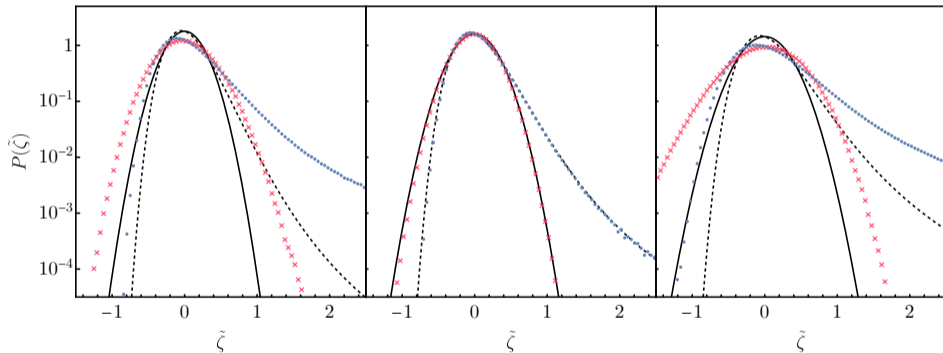
Procedure: (details ↗ 2412.14106)

1. Compute $\mathcal{P}_{\delta\phi}(k)$, $\mathcal{I}(p, k, q)$ (spectrum and bispectrum of $\delta\phi$)
2. Generate a (random) lattice of non-Gaussian $\delta\phi(\mathbf{x})$ with the right spectrum and bispectrum:
 - 2.1 Fourier transform Gaussian $\delta\phi(\mathbf{x}) \rightarrow \delta\phi_{\mathbf{k}}$
 - 2.2 Multiply by $\mathcal{P}_{\delta\phi}(k)$, convolve with $\mathcal{I}(p, k, q)$
 - 2.3 Inverse-Fourier transform non-Gaussian $\delta\phi_{\mathbf{k}} \rightarrow \delta\phi(\mathbf{x})$
3. From the lattice, sample the PDF of $\delta\phi$
4. Use δN formula to compute non-linear ζ from non-Gaussian $\delta\phi$

Remark: Convergence? Relation to perturbativity breakdown in the in-in formalism (large loop corrections)? See 2412.14106

3. NON-GAUSSIAN $\delta\phi$ | PDF of ζ

See details in [\[2406.02417\]](#)



- Gaussian $\delta\phi$
- × non-Gaussian $\delta\phi$
- - - non-linear ζ from Gaussian $\delta\phi$
- non-linear ζ from non-Gaussian $\delta\phi$

CONCLUSIONS

- δN formula: non-linear relation between inflaton and curvature perturbations.
Gaussian $\delta\phi \rightarrow$ non-Gaussian ζ
- USR phase: Exponential tail in PDF of ζ due to non-linear change of variables along attractor in phase space
- Refinement: account for inflaton non-Gaussianity
 - Two sources of non-Gaussianity: *intrinsic non-Gaussianity* and non-linearities
 - A first approach: sample non-Gaussian $\delta\phi$ (using QFT input), apply non-linear relation to ζ .