

Electroweak monopoles and electroweak baryogenesis ¹²

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¹SA & A. Kobakhidze, 1702.04068 Eur. Phys. J. C (2017) 77: 444

²SA, Daniel Collison and Archil Kobakhidze, [arXiv:1810.10696]

Outline

- 1 Motivation
- 2 The nature of the monopoles
 - Topological Stability
 - Monopole Mass
- 3 Monopoles and cosmology
 - The Electroweak Phase transition
 - Monopole production
 - Sphaleron Processes
 - Big Bang Nucleosynthesis
- 4 Baryogenesis

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Motivation

- There is an asymmetry between the matter and antimatter abundances in the universe
- Sakharov conditions must be satisfied
- Sphaleron washout processes must be suppressed in the broken phase
- We propose a mechanism for generating the asymmetry through the production of electroweak monopoles in a Born-Infeld extension to the standard model.

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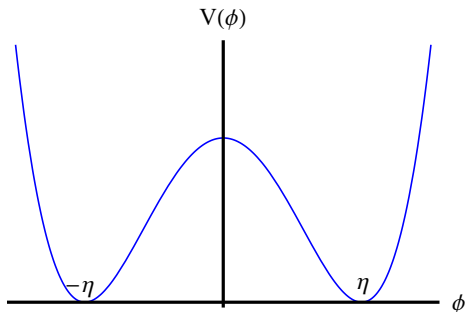
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1-D topological defects

- Consider a 1D potential:

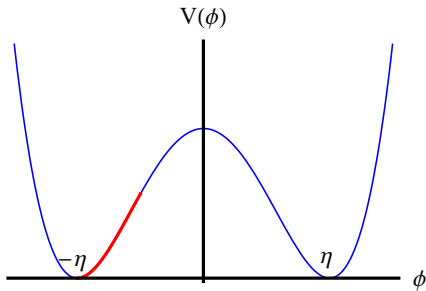
$$V(\phi) = \frac{\lambda}{4} (\phi^2 - \eta^2)^2$$



- For $\int_{-\infty}^{\infty} V(\phi) dx < \infty$, $\phi(\pm\infty) \rightarrow \pm\eta$

1-D topological defects

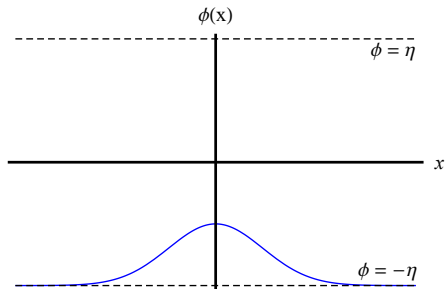
- Suppose $\phi(\infty) = \phi(-\infty) = -\eta$



- Decays to the constant solution

1-D topological defects

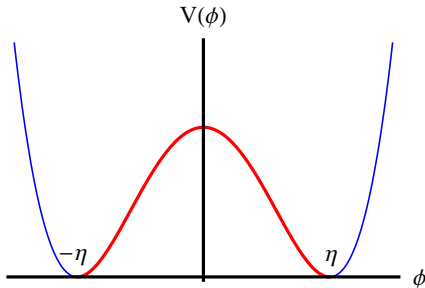
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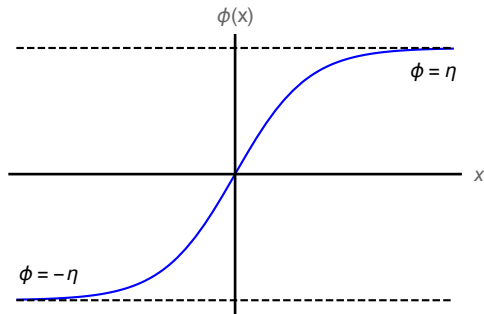
- Suppose $\phi(\infty) = -\phi(-\infty)$



- Heuristically requires an infinite amount of energy to transition to constant solution.
- Topological stability from disconnected vacuum manifold
- $\pi_0(M_{vac}) \neq 0$.

1-D topological defects

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Monopoles

- Monopoles are an extension of this idea to 3 spatial dimensions
- Spatial infinity is described by a 2-sphere
- Finite energy requires $\phi : S_{\infty}^2 \rightarrow M_{vac}$.
- Topologically non-trivial solutions exist when $\pi_2(M_{vac}) \neq 0$
- For the standard model, $M_{vac} = (SU(2)_L \times U(1)_Y)/U(1)_{EM}$
- $\pi_2(M_{vac}) = \pi_2(S^3) = 0$
- No electroweak monopoles?

The Ansatz

- Cho and Maison (1997) found electroweak monopoles through the ansatz:

$$\phi = \frac{1}{\sqrt{2}}\rho\xi$$

$$\rho = \rho(r)$$

$$\xi = i \begin{pmatrix} \sin(\theta/2)e^{-i\varphi} \\ -\cos(\theta/2) \end{pmatrix}$$

$$A_\mu = \frac{1}{g}A(r)\partial_\mu t \hat{r} + \frac{1}{g}(f(r) - 1)\hat{r} \times \partial_\mu \hat{r}$$

$$B_\mu = -\frac{1}{g'}B(r)\partial_\mu t - \frac{1}{g'}(1 - \cos\theta)\partial_\mu\varphi$$

Why is this stable?

$$\xi = i \begin{pmatrix} \sin(\theta/2) e^{-i\varphi} \\ -\cos(\theta/2) \end{pmatrix}$$
$$B_\mu = -\frac{1}{g'}(1 - \cos\theta)\partial_\mu\varphi$$

- Gauge invariance under $U(1)_Y$ implies that the vacuum manifold is defined up to a phase.
- String singularities in both fields at $\theta = \pi$
- Can be removed using a Wu-Yang construction
- Each hemisphere maps onto \mathbb{C}^1
- By definition, this corresponds to the Riemann sphere, $\mathbb{C}P^1$
- $\pi_2(M_{\text{vac}}) = \mathbb{Z}$

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The energy

$$E = E_0 + E_1$$

$$E_0 = 4\pi \int_0^\infty \frac{dr}{2r^2} \left\{ \frac{1}{g'^2} + \frac{1}{g^2} (f^2 - 1)^2 \right\}$$

$$E_1 = 4\pi \int_0^\infty dr \left\{ \frac{1}{2} (r\dot{\rho})^2 + \frac{1}{g^2} \left(\dot{f}^2 + \frac{1}{2} (r\dot{A})^2 + f^2 A^2 \right) \right. \\ \left. + \frac{1}{2g'^2} (r\dot{B})^2 + \frac{\lambda r^2}{8} (\rho^2 - \rho_0^2)^2 \right. \\ \left. + \frac{1}{4} f^2 \rho^2 + \frac{r^2}{8} (B - A)^2 \rho^2 \right\}$$

- The first term of E_0 is divergent at the origin.

Regularisation

- Cho, Kim and Yoon(2015) proposed a regularisation of the form:

$$g' \rightarrow \frac{g'}{\sqrt{\epsilon}}$$

$$\epsilon = \left(\frac{\phi}{\phi_0} \right)^n$$

- However, g' becomes non-perturbative as $\phi \rightarrow 0$.
- This is undesirable in an EFT framework.
- We instead propose a Born-Infeld modification for the $U(1)_Y$ kinetic term.

Born-Infeld modification

- We regularise the $U(1)_Y$ kinetic term by replacing it with:

$$\begin{aligned} & \beta^2 \left[1 - \sqrt{-\det \left(\eta_{\mu\nu} + \frac{1}{\beta} B_{\mu\nu} \right)} \right] \\ &= \beta^2 \left[1 - \sqrt{1 + \frac{1}{2\beta^2} B_{\mu\nu} B^{\mu\nu} - \frac{1}{16\beta^4} (B_{\mu\nu} \tilde{B}^{\mu\nu})^2} \right] \end{aligned}$$

- As $\beta \rightarrow \infty$, the SM is recovered.
- The corresponding energy is

$$\begin{aligned} & \int_0^\infty dr \beta^2 \left[\sqrt{(4\pi r^2)^2 + \left(\frac{4\pi}{g'\beta} \right)^2} - 4\pi r^2 \right] \\ &= \frac{4\pi^{5/2}}{3\Gamma\left(\frac{3}{4}\right)^2} \sqrt{\frac{\beta}{g'^3}} \approx 72.8 \sqrt{\beta} \end{aligned}$$

- Hence, β acts as a mass parameter for the monopoles.

- Extend the $SU(2)$ sector as well with an independent Born-Infeld term:

$$\beta_1^2 \left[1 - \sqrt{-\det \left(\eta_{\mu\nu} + \frac{1}{\beta_1} \mathbf{B}_{\mu\nu} \right)} \right] + \beta_2^2 \left[1 - \sqrt{-\det \left(\eta_{\mu\nu} + \frac{1}{\beta_2} \mathbf{F}_{\mu\nu} \right)} \right]$$

- Constrained by light by light scattering results (Ellis et al. 2017):

$$\sqrt{\beta_{EM}} = \frac{\sqrt{\beta_2}}{\sqrt[4]{\sin^4 \theta_W + \cos^4 \theta_W \left(\frac{\beta_2}{\beta_1}\right)^2}} \gtrsim 100\text{GeV}$$

- For $\beta_2 \gg \beta_1$ (perturbative unitarity) gives a lower bound for monopole mass of $\sim 9 - 11\text{TeV}$

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The Electroweak Phase transition

- SM high temperature effective potential:

$$V(\phi, T) = D(T^2 - T_0^2)\phi^2 - ET\phi^3 - \frac{1}{4}\lambda_T\phi^4$$

- curvature at the origin changes at $T = T_0$
- the nature of the transition depends on the values of the SM parameters.

First order phase transition

- The minima become degenerate before T_0
- Bubbles of the broken phase form
- collisions lead to gravitational waves, baryogenesis etc.

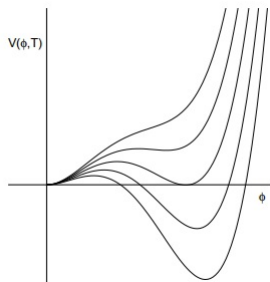


Figure: First order phase transition (Petropoulos, 2003)

Second order phase transition

- The minima never become degenerate
- the universe rolls homogeneously into the broken phase
- predicted by SM parameters

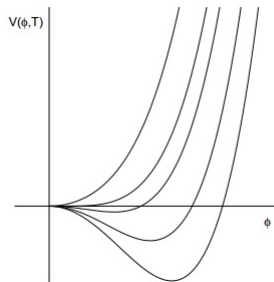


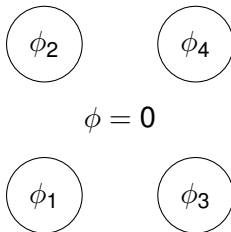
Figure: Second order phase transition (Petropoulos, 2003)

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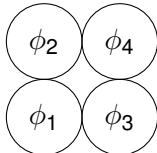
The Kibble Mechanism (Kibble, 1976)

- At $T = T_c$, domains of the broken phase will appear
- The higgs field in each domain takes independent directions on the vacuum manifold



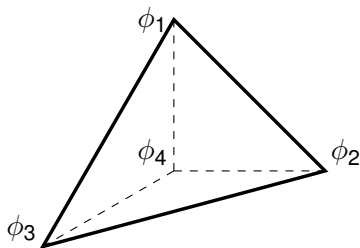
The Kibble mechanism (Kibble, 1976)

- As the Higgs field is continuous, it must be interpolated at the intersections.
- Consider an intersection of four of these domains:



The Kibble mechanism (Kibble, 1976)

- In field space, these points form the vertices of a tetrahedron.
- This tetrahedron should be shrunk to a point at the intersection.
- If these cannot be shrunk to a point continuously, a topological defect in the form of a monopole which continuously joins the two minima.
- The tetrahedron is homotopically equivalent to S^2 .
- Therefore, $\pi_2(\mathbb{C}P^1) = \mathbb{Z}$ implies the existence of monopoles



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Sphaleron Processes

- Sphaleron mediated scattering processes occur in the unbroken phase
- They violate $B + L$ in units of $\Delta B = \Delta L = 3$
- If unsuppressed, they washout any pre-existing baryon number.
- Suppression in the broken phase requires a 1st order EWPT with $\frac{\phi_c}{T_c} \gtrsim 1$.

The electroweak phase transition

- The Gibbs free energy:

$$G_u = V(0)$$

$$G_b = V(\phi_c(T)) + E_{\text{monopoles}}$$

- At the critical temperature:

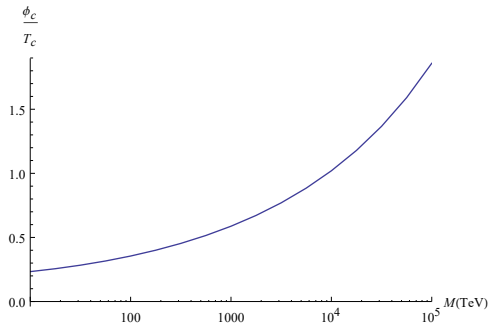
$$V(0) = V(\phi_c(T_c)) + E_{\text{monopoles}}$$

- Assuming $T \ll M$, the monopoles are decoupled and $E_{\text{monopoles}} = M \times n_M$

The initial density

- $n_M \approx \frac{1}{d^3}$ where d is the separation of two uncorrelated monopoles.
- This is chosen to be the Coulomb capture distance.
- Hence, $n_M \approx \left(\frac{4\pi}{h^2}\right)^3 T^3$

Results



- Sphaleron processes are suppressed for $M > 0.9 \cdot 10^4$ TeV.

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The constraint

- The monopole density should not dominate the universe at the time of helium synthesis. This implies:

- $\frac{n}{T^3} \Big|_{T=1\text{MeV}} < \frac{1\text{MeV}}{M}$

- Hence, the evolution of the number density over time must be considered:

$$\frac{dn_M}{dt} = -Dn_M^2 - 3Hn_M$$

- This constrains the mass of the monopole to $M \lesssim 2.3 \cdot 10^4 \text{ TeV}$.

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Sakharov conditions

- In 1967, Andrei Sakharov proposed three conditions for baryogenesis to occur:
 - ① Baryon number violation
 - ② C and CP violation
 - ③ Departure from thermal equilibrium- Heavy monopoles.

C and CP -violation

- Consider the θ - terms:

$$\mathcal{L}_\theta = \theta_2 F_{\mu\nu}^a \tilde{F}^{a\mu\nu} + \theta_1 B_{\mu\nu} \tilde{B}^{\mu\nu}$$

- In the usual case:
 - hypercharge sector is topologically trivial, and hence, θ_1 is unphysical
 - θ_2 can be rotated away by a $B + L$ -rotation of quarks and leptons.
 - no CP -violation

C and CP-violation

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- In the usual case:
 - hypercharge sector is topologically trivial, and hence, θ_1 is unphysical
 - θ_2 can be rotated away by a $B + L$ -rotation of quarks and leptons.
 - no CP-violation
- With electroweak monopoles:
 - Monopoles gain an electric charge through the Witten effect
 - Supports θ_1
 - Only one can be rotated away
 - a new source of CP violation

$B + L$ -violation

$$\mathcal{L}_\theta = \theta_{ew} F_{\mu\nu}^a \tilde{F}^{a\mu\nu} ,$$

- Topologically inequivalent vacuum configurations related by large gauge transformations $g \in SU(2)_L$ give rise to the θ_{ew} -vacuum structure.

$$|M, \theta_{ew}\rangle = \sum_{n=-\infty}^{n=+\infty} e^{in\theta_{ew}} (U[g])^n |M, 0\rangle .$$

- monopole-antimonopole pair that carries $\Delta n = 1$ topological charge, would annihilate into 9 quarks and 3 leptons, giving rise to $\Delta B = \Delta L = 3$.
- not suppressed even at zero temperature (Callan, 1982) (Rubakov, 1981)

Baryon asymmetry of the universe

$$\frac{d\bar{n}_B}{dt} = -\kappa\theta \frac{dn_M}{dt}$$

- \bar{n}_B is the difference in the number densities of matter and antimatter
- κ describes the asymmetry generated in each collision
- for monopoles, $n_{M0} \gg n_{Mf}$.
- Hence,

$$\bar{n}_B \approx \kappa\theta n_0 = \kappa\theta\alpha_{EM}^3 T_c^3$$

Baryon asymmetry of the universe

- The asymmetry parameter, η_B , can now be evaluated:

$$\eta_B = \frac{\bar{n}_B}{s} = \kappa\theta \frac{45\alpha_{\text{EM}}^3 T_c^3}{2\pi^2 g_* T_f^3}$$

- $1.6 \times 10^{-8} \kappa\theta \leq \eta_B \leq 2.5 \times 10^{-7} \kappa\theta$.
- Empirical values for the asymmetry parameter $\eta_B \approx 10^{-10}$ can be accommodated for with $\kappa\theta_{ew} \sim 10^{-3} - 10^{-2}$.

Conclusion

- Finite energy monopoles exist in the Standard model with a Born-Infeld extension.
- The mass is related to the Born-Infeld parameters
- Sphaleron mediated processes can be made ineffective in the broken phase while remaining under the nucleosynthesis constraints.
- This occurs for monopoles with a mass of $(0.9 - 2.3) \cdot 10^4 \text{TeV}$.
- Baryon asymmetry of the universe can be accounted for through this mechanism