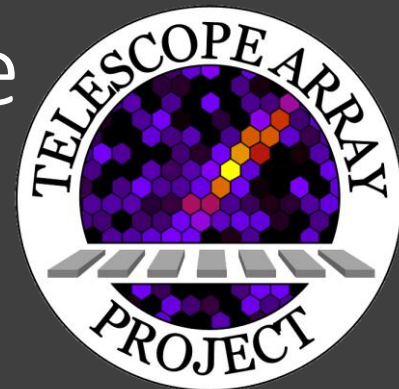




Recent results and status of the Telescope Array Experiment

Eiji Kido for the Telescope Array Collaboration

Riken



Outline

- The Telescope Array (TA) Experiment
 - Detectors
 - Energy Spectrum
 - Anisotropy
 - Composition
- Extension of the TA Experiment
 - The TAx4 Experiment
- Summary

Telescope Array collaboration

S. Ogio,
ICRC2019

147 collaborators from 36 institutes
in 6 countries



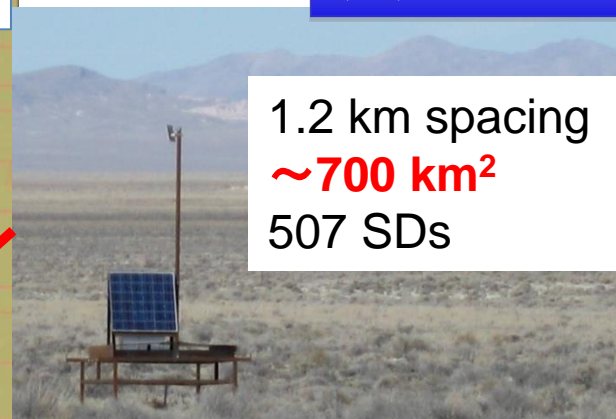
R.U. Abbasi (1), M. Abe (2), T. Abu-Zayyad (1), M. Allen (1), R. Azuma (3), E. Barcikowski (1), J.W. Belz (1), D.R. Bergman (1), S.A. Blake (1), R. Cady (1), B.G. Cheon (4), J. Chiba (5), M. Chikawa (6), A. di Matteo (7), T. Fujii (8), K. Fujita (9), R. Fujiwara (9), M. Fukushima (10,11), G. Furlich (1), W. Hanlon (1), M. Hayashi (12), Y. Hayashi (9), N. Hayashida (13), K. Hibino (13), K. Honda (14), D. Ikeda (15), T. Inadomi (16), N. Inoue (2), T. Ishii (14), R. Ishimori (3), H. Ito (17), D. Ivanov (1), H. Iwakura (16), H.M. Jeong (18), S. Jeong (18), C.C.H. Jui (1), K. Kadota (19), F. Kakimoto (3), O. Kalashev (20), K. Kasahara (21), S. Kasami (22), H. Kawai (23), S. Kawakami (9), S. Kawana (2), K. Kawata (10), E. Kido (10), H.B. Kim (4), J.H. Kim (1), J.H. Kim (24), S. Kishigami (9), V. Kuzmin (20), M. Kuznetsov (7,20), Y.J. Kwon (25), K.H. Lee (18), B. Lubsandorzhiev (20), J.P. Lundquist (1), K. Machida (14), K. Martens (11), H. Matsumiya (9), T. Matsuyama (9), J.N. Matthews (1), R. Mayta (9), M. Minamino (9), K. Mukai (14), I. Myers (1), S. Nagataki (17), K. Nakai (9), R. Nakamura (16), T. Nakamura (26), Y. Nakamura (16), T. Nonaka (10), H. Oda (9), S. Ogio (9,27), M. Ohnishi (10), H. Ohoka (10), Y. Oku (22), T. Okuda (28), Y. Omura (9), M. Ono (17), R. Onogi (9), A. Oshima (9), S. Ozawa (21), I.H. Park (18), M.S. Pshirkov (20,29), J. Remington (1), D.C. Rodriguez (1), G. Rubtsov (20), D. Ryu (24), H. Sagawa (10), R. Sahara (9), K. Saito (10), Y. Saito (16), N. Sakaki (10), T. Sako (10), N. Sakurai (9), K. Sano (16), L.M. Scott (30), T. Seki (16), K. Sekino (10), P.D. Shah (1), F. Shibata (14), T. Shibata (10), H. Shimodaira (10), B.K. Shin (9), H.S. Shin (10), J.D. Smith (1), P. Sokolsky (1), N. Sone (16), B.T. Stokes (1), S.R. Stratton (1,30), T.A. Stroman (1), T. Suzawa (2), Y. Takagi (9), Y. Takahashi (9), M. Takamura (5), M. Takeda (10), R. Takeishi (18), A. Taketa (15), M. Takita (10), Y. Tameda (22), H. Tanaka (9), K. Tanaka (31), M. Tanaka (32), Y. Tanoue (9), S.B. Thomas (1), G.B. Thomson (1), P. Tinyakov (7,20), I. Tkachev (20), H. Tokuno (3), T. Tomida (16), S. Troitsky (20), Y. Tsunesada (9,27), Y. Uchihori (33), S. Udo (13), T. Uehama (16), F. Urban (34), T. Wong (1), II. M. Yamamoto (16), H. Yamaoka (32), K. Yamazaki (13), J. Yang (35), K. Yashiro (5), M. Yosei (22), H. Yoshii (36), Y. Nakamura (16), Y. Zhezher (20), III. Z. Zundel (1)

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TA Detectors

Surface Detector (SD)

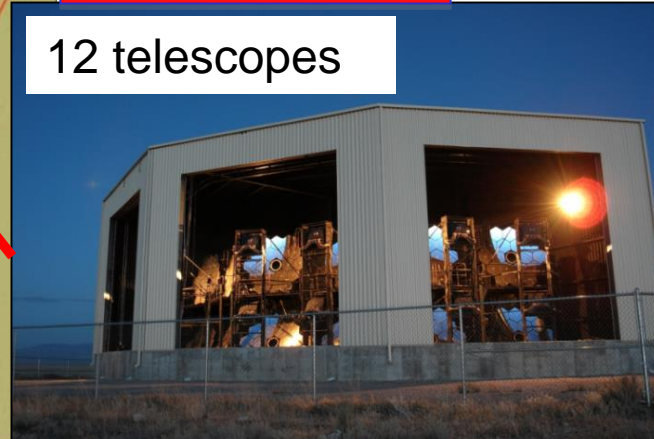
1.2 km spacing
~700 km²
507 SDs



- **Largest cosmic-ray observatory in the northern hemisphere**
- **Hybrid observation by SD and FD**

Fluorescence Detector (FD) station

12 telescopes



FD station

MDFD

TALE SD array

SD array

Latitude 39.30° N
Longitude 112.91° W
Height: 1382 m

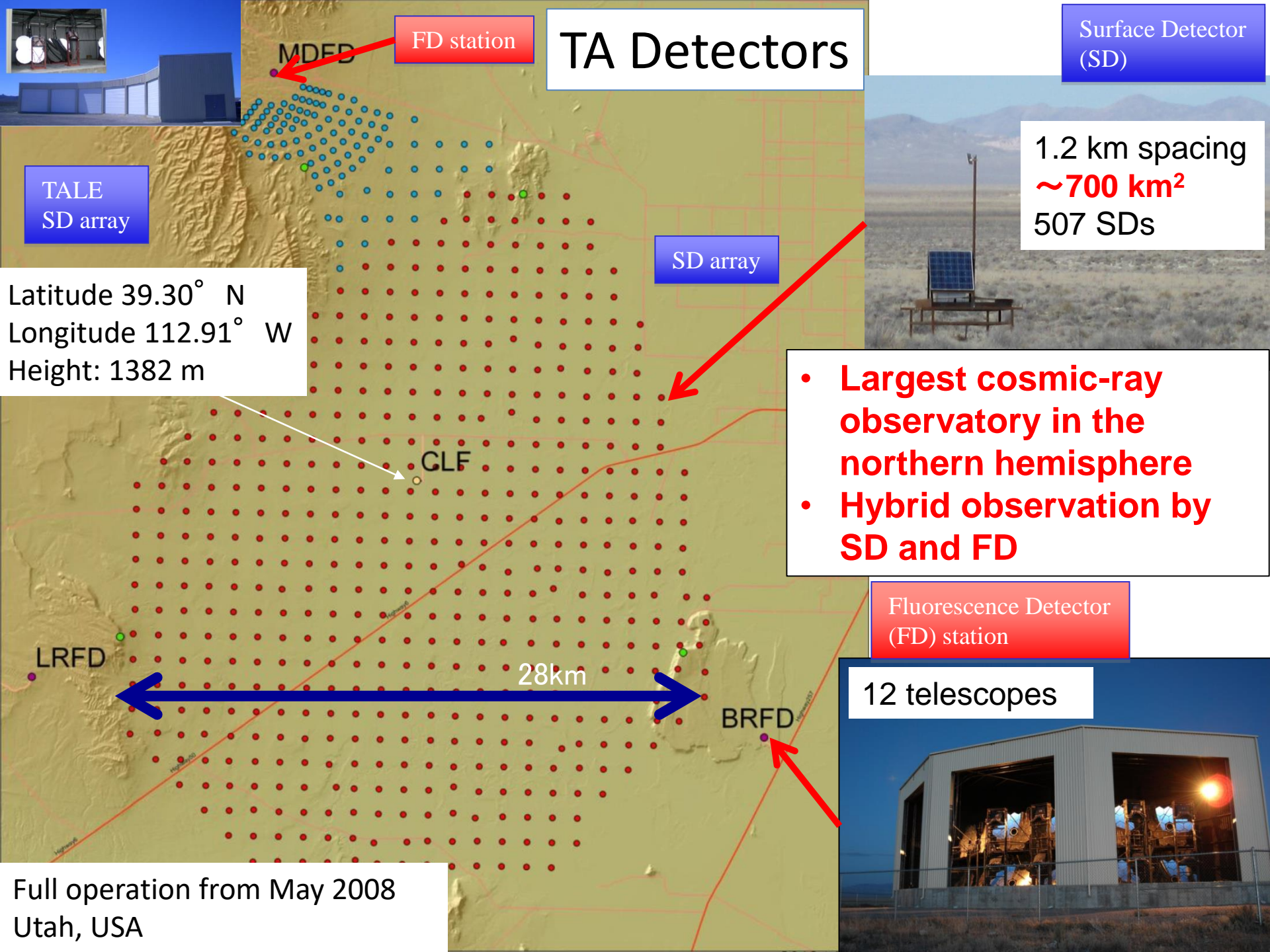
GLF

28km

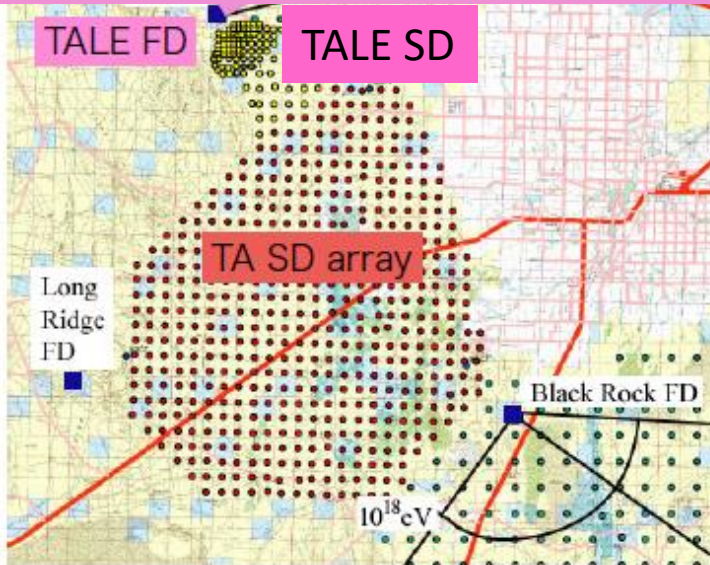
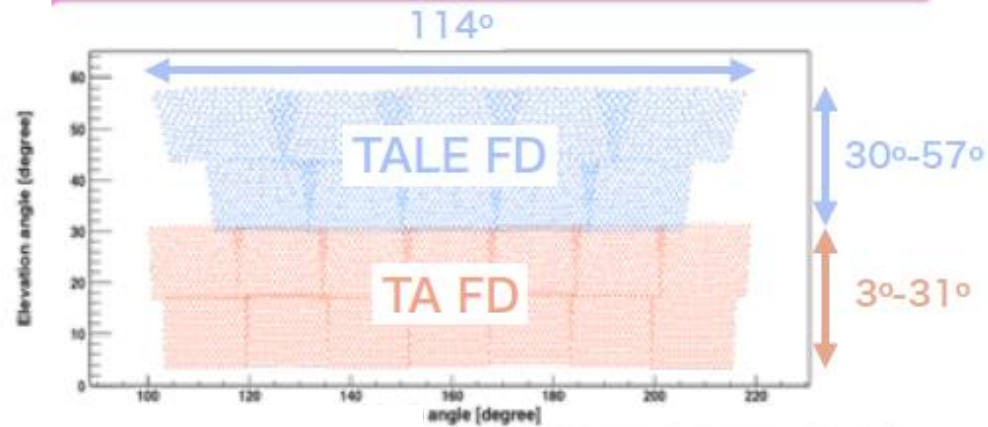
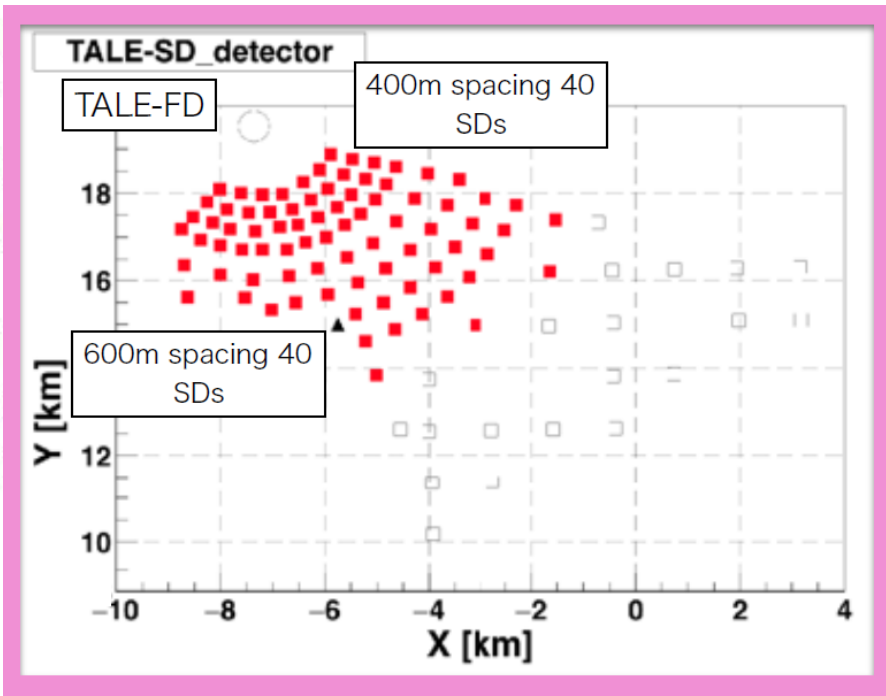
LRFD

BRFD

Full operation from May 2008
Utah, USA



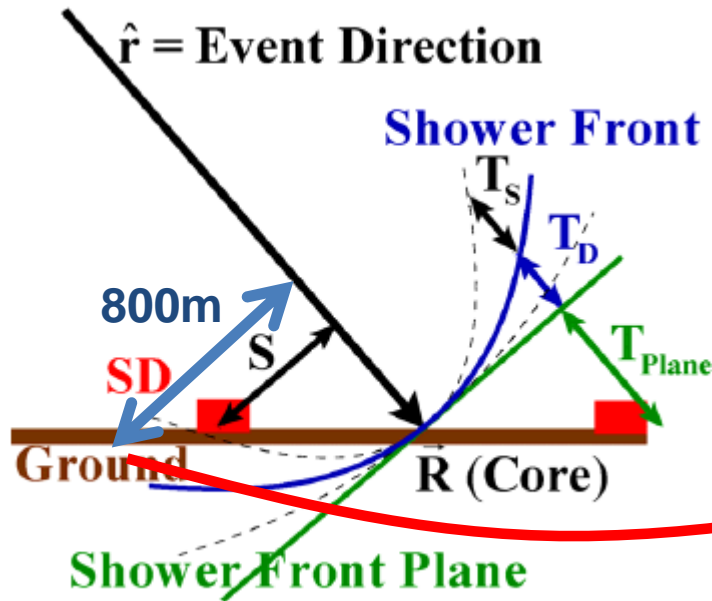
TALE (TA Low-energy Extension) Detectors



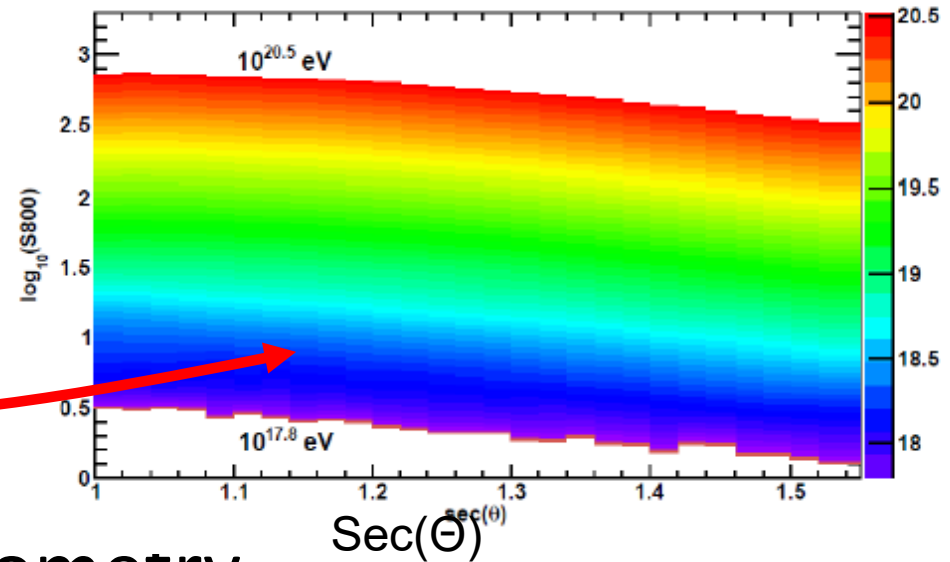
TALE FD was installed in Nov. 2012
Operation since Sep. 2013
Hybrid trigger: Sep. 2018



Event Reconstructions with SDs



$\text{Log}_{10}(S(800))$



- Timing fit \rightarrow Shower Geometry
- Lateral distribution fit $\rightarrow S(800) \rightarrow$ Energy from MC
 \rightarrow rescale to calorimetrically measured E_{FD} using

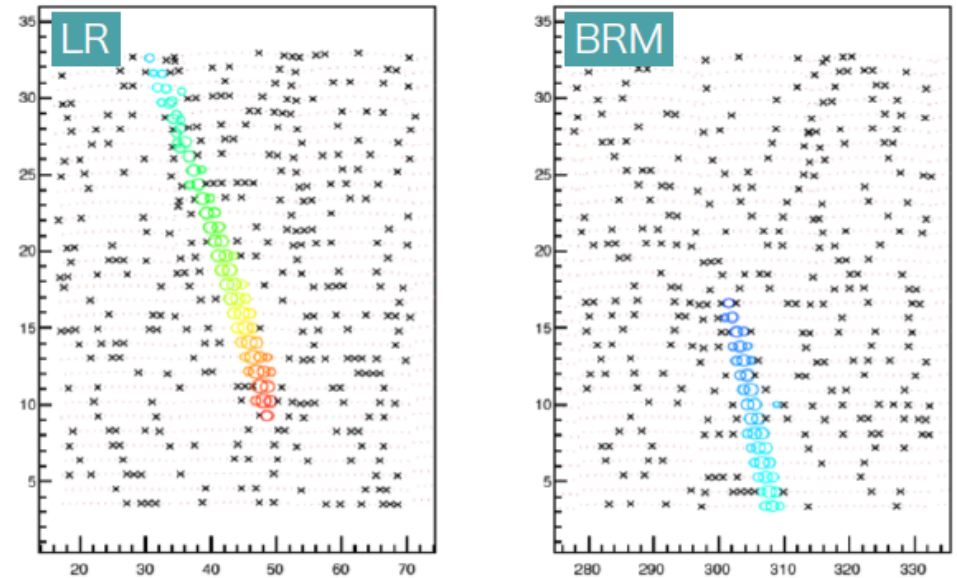
SD and FD hybrid events

$S(800)$: energy depositions which are converted in VEM unit.

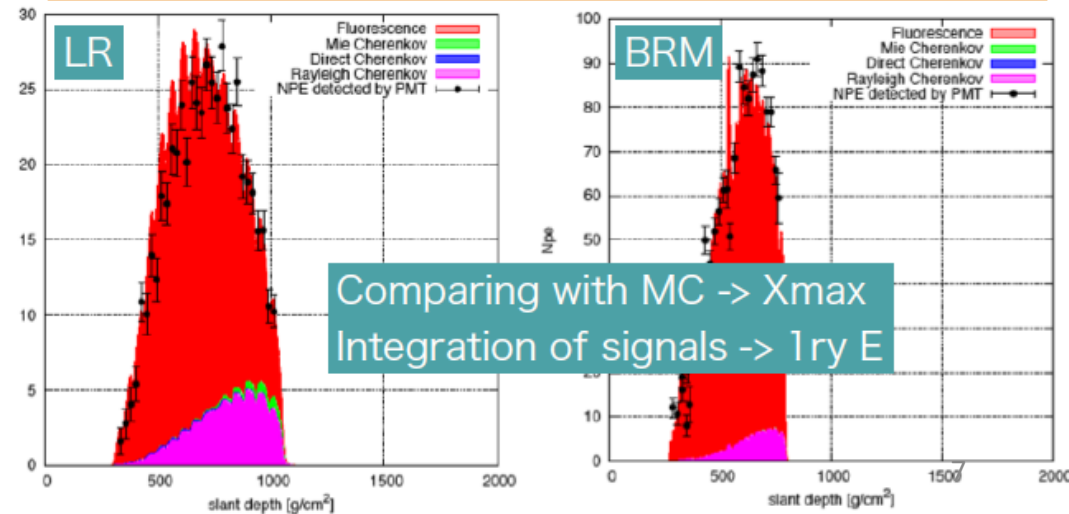
Event reconstruction

S. Ogio,
ICRC2019

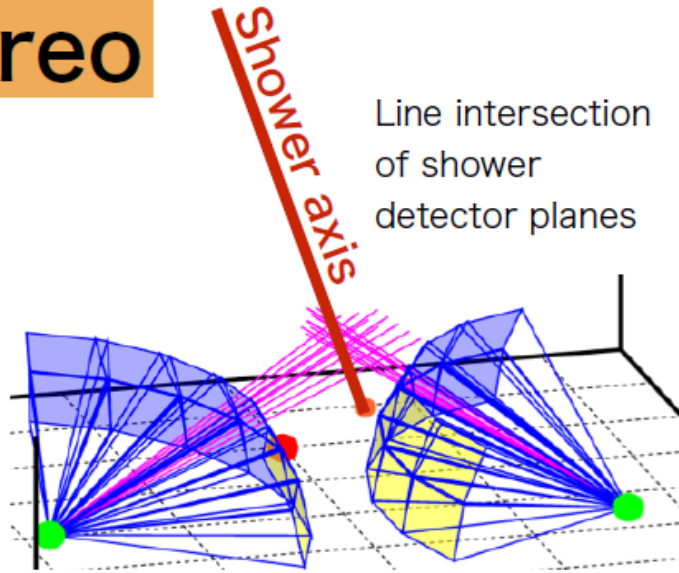
observed images



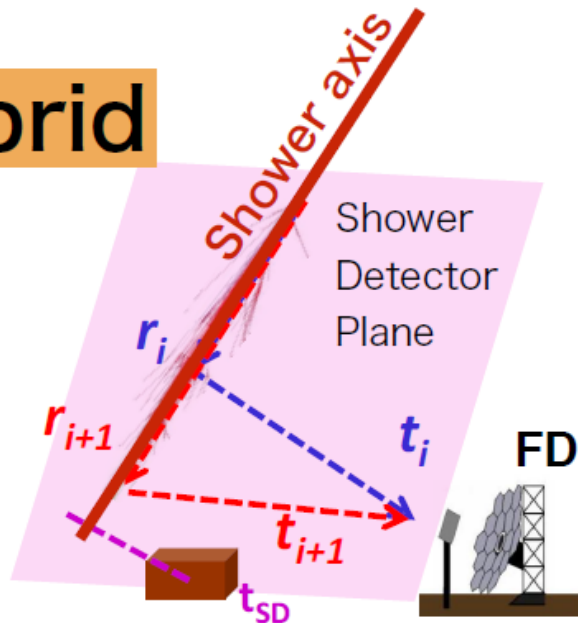
reconstructed shower profiles



Stereo

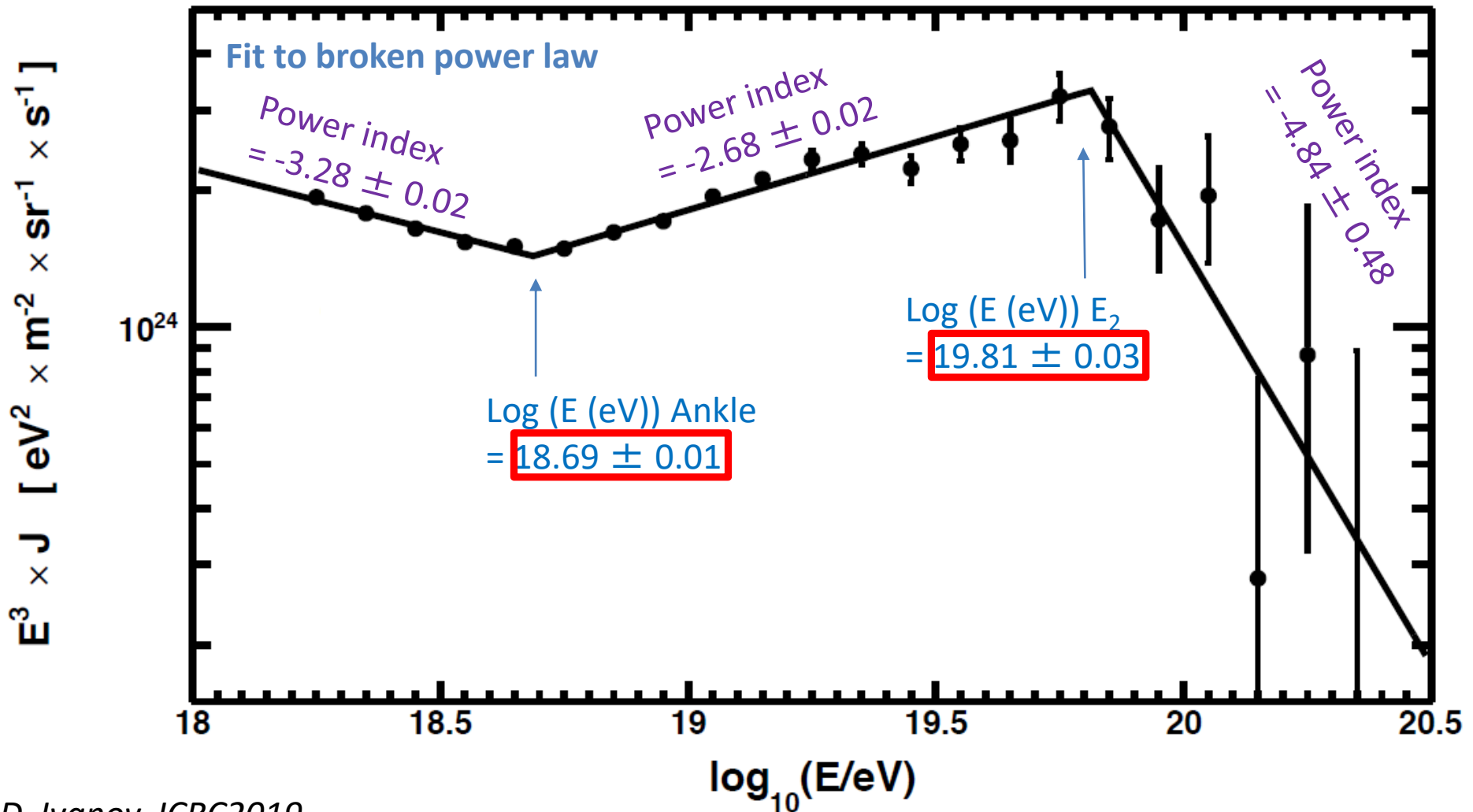


Hybrid

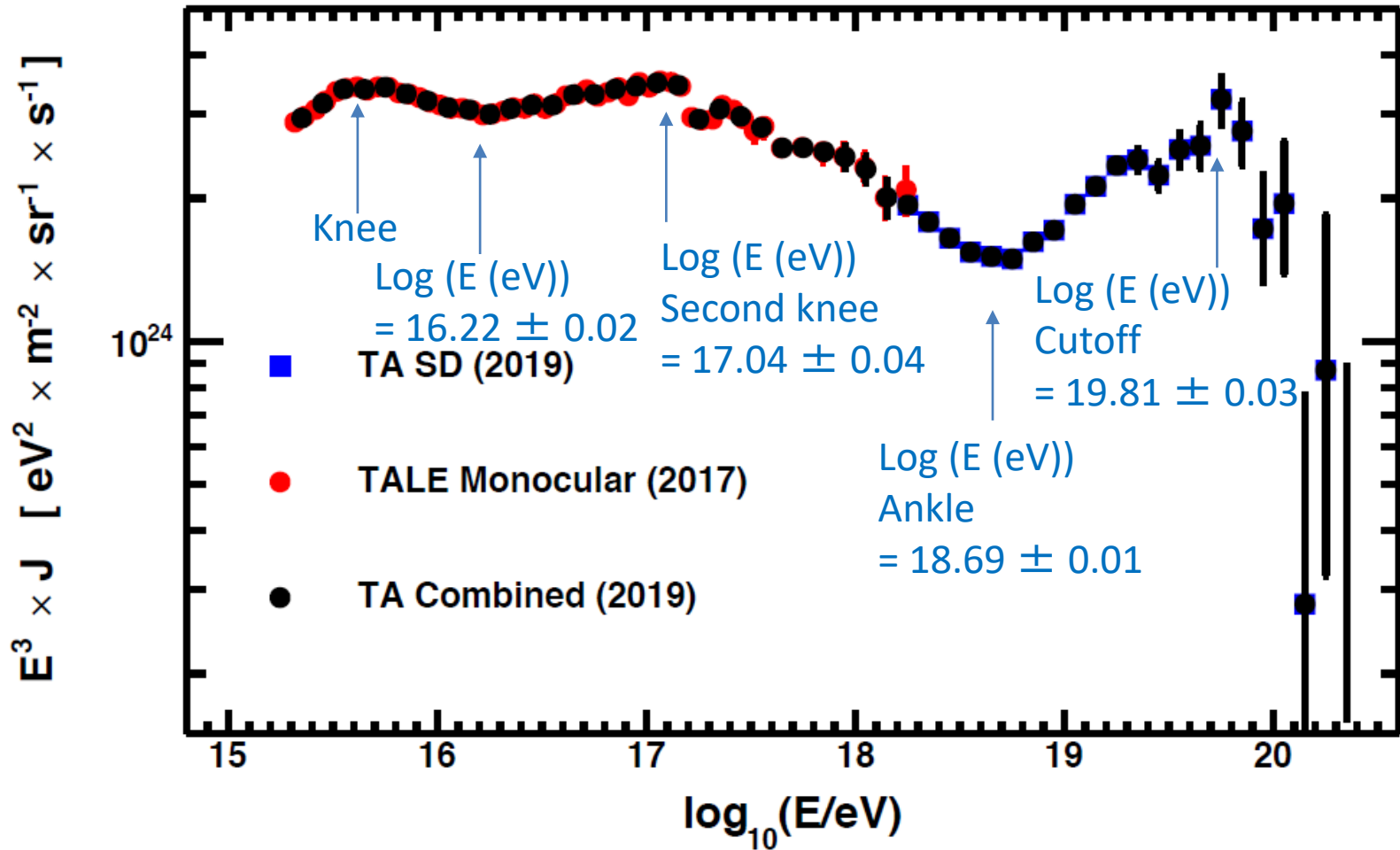


Energy Spectrum with TA SD

TA SD 11 years data



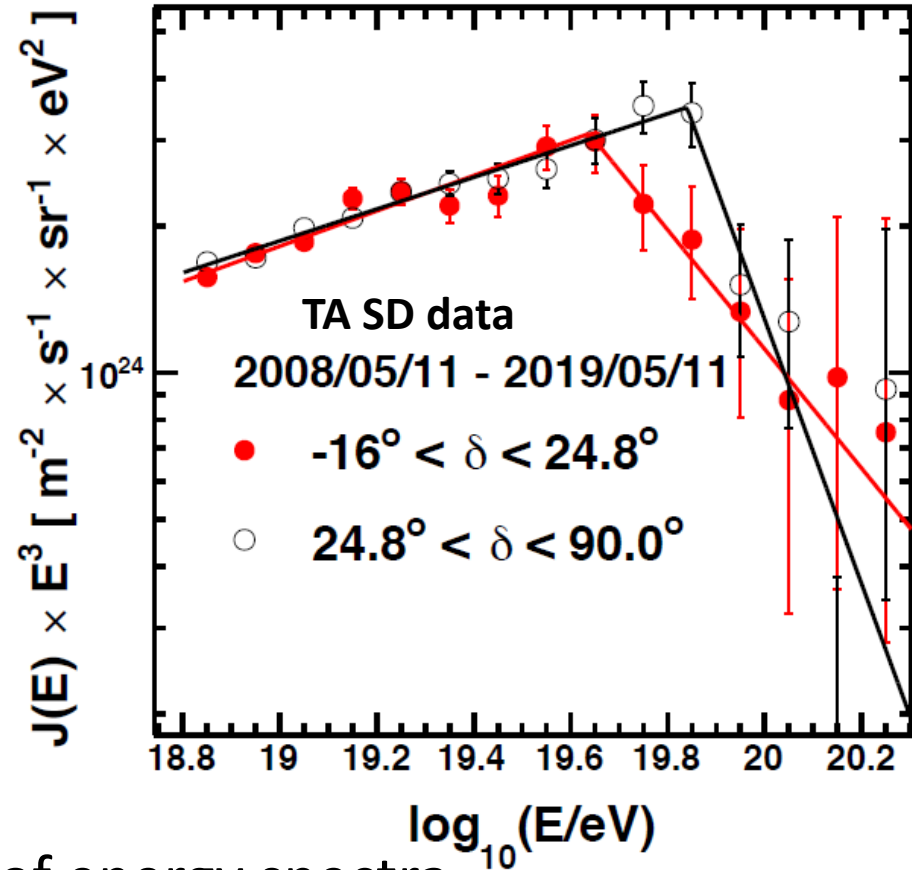
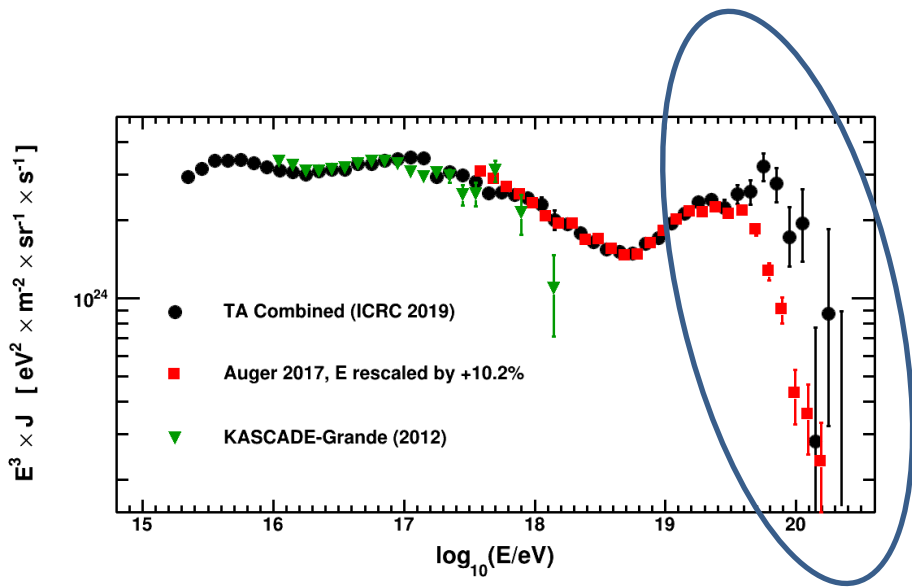
Combined Energy Spectrum with TALE FD Mono



Combined TA spectrum using
22 months TALE FD monocular data +
11 years TA SD data

Declination Dependence of Energy Spectrum

D. Ivanov, ICRC2019



- Difference of the cutoff energies of energy spectra
 - $\log(E/\text{eV}) = 19.64 \pm 0.04$ for lower dec. band ($-16^\circ - 24.8^\circ$)
 - $\log(E/\text{eV}) = 19.84 \pm 0.02$ for higher dec. band ($24.8^\circ - 90^\circ$)
- The global significance of the difference was estimated to be 4.3σ

Hotspot: $E > 57$ EeV

Original hotspot reported in 2014,
from 5 years of data

Ap. J., 790, L21(2014)

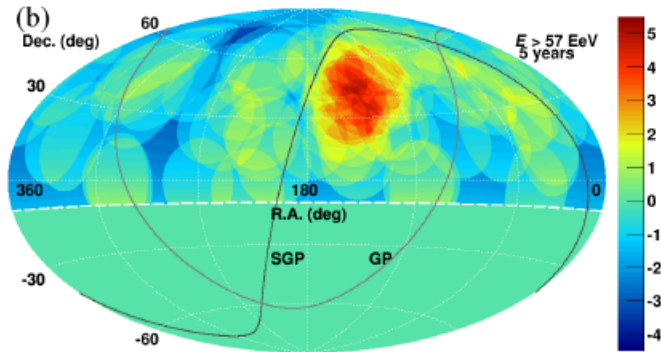
$E > 57$ EeV (Observed 72 events)

20° over-sampling circle

19 events fall in "Hotspot" centered at $(146.7^\circ, 43.2^\circ)$

(Expected = 4.5 events)

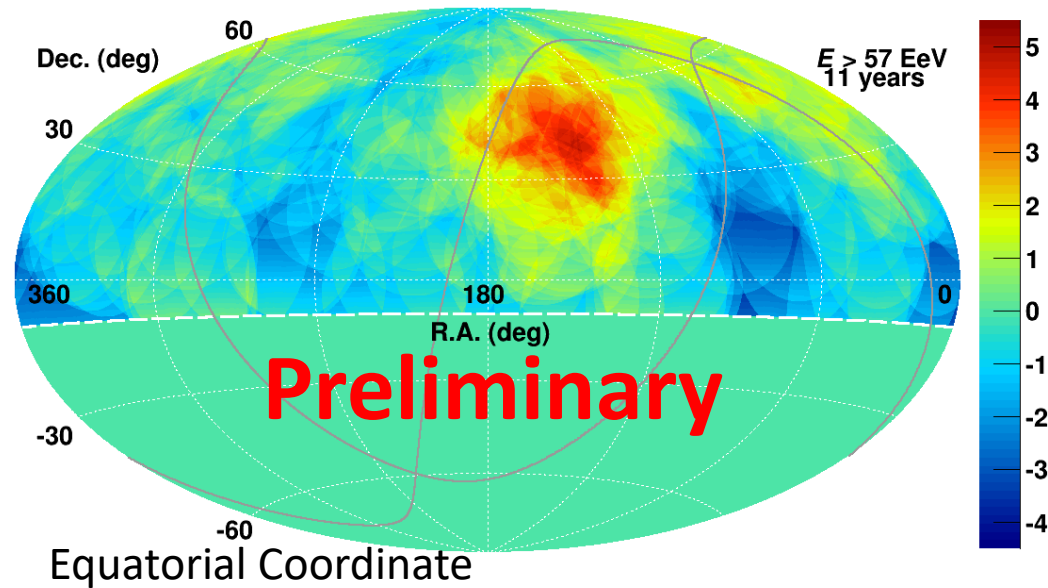
local significance 5.1σ , post trial significance 3.4σ



K. Kawata, ICRC2019

TA SD 11 years data

Significance map from isotropy expectation



$E > 57$ EeV, in total 168 events

38 events fall in Hotspot ($\alpha=144.3^\circ$, $\delta=40.3^\circ$, 25° radius, 22° from SGP), expected=14.2 events

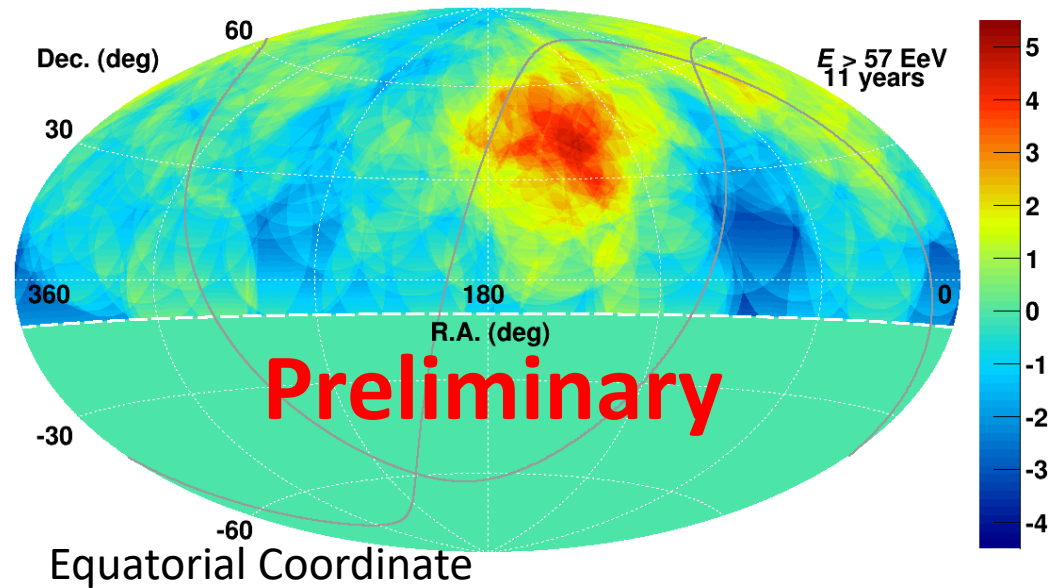
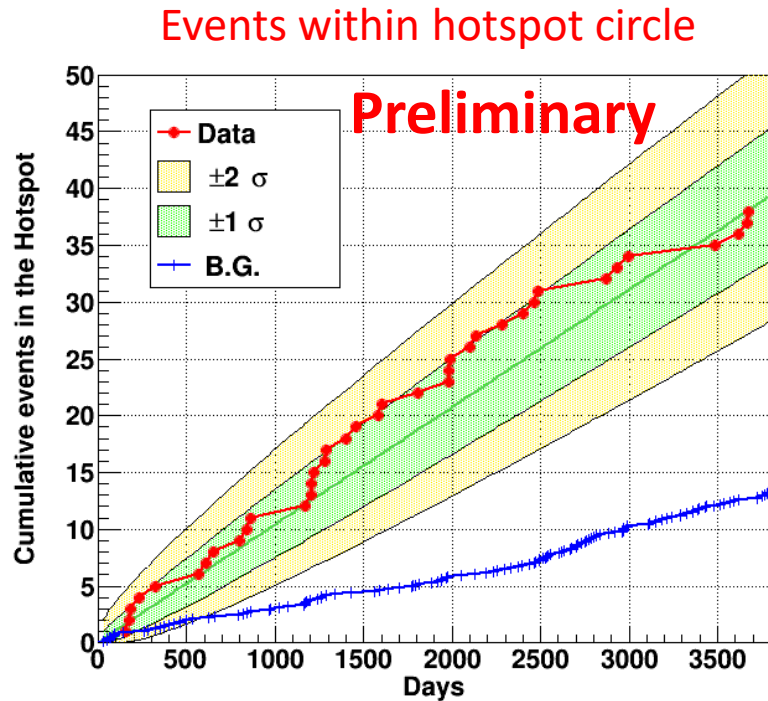
local significance = 5.1σ , chance probability $\rightarrow 2.9\sigma$

25° over-sampling radius shows the highest local significance (scanned 15° to 35° with 5° step)

Hotspot : $E > 57 \text{ EeV}$

K. Kawata, ICRC2019

Significance map from isotropy expectation



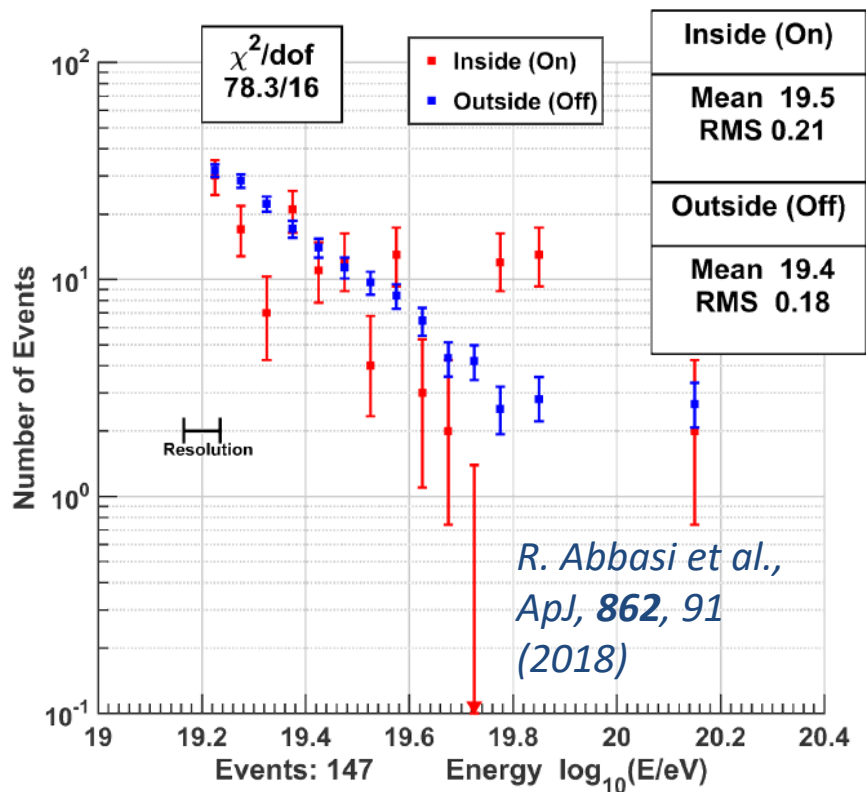
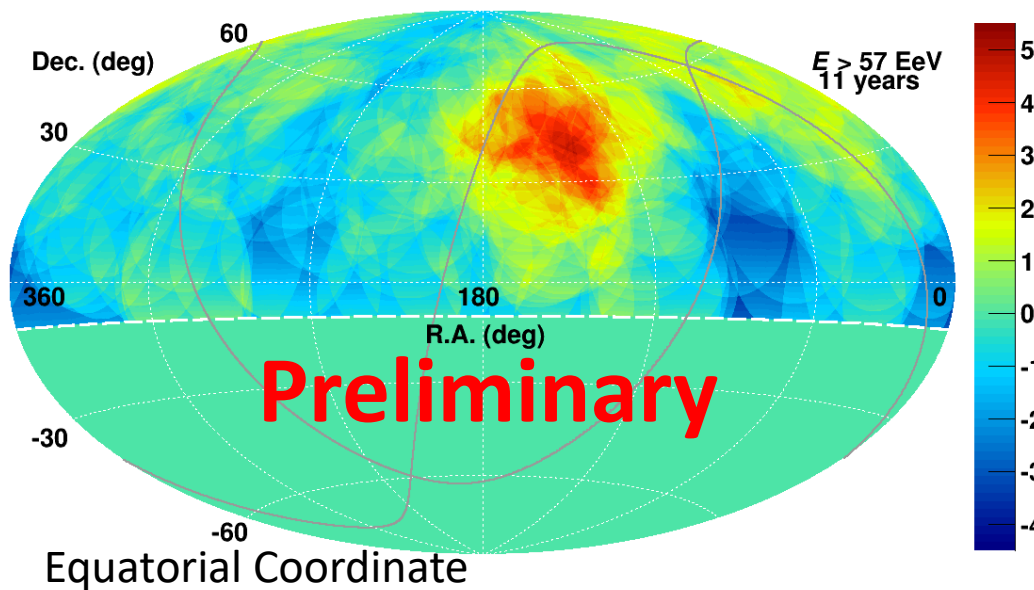
The cumulative events inside the hotspot circle (25° -radius circle) defined by the 11-year.
The increase rate of the events inside the hotspot circle:

Consistent with the linear increase within $\sim 1\sigma$

Energy Spectrum at around the Hotspot

K. Kawata, ICRC2019

Significance map from isotropy expectation



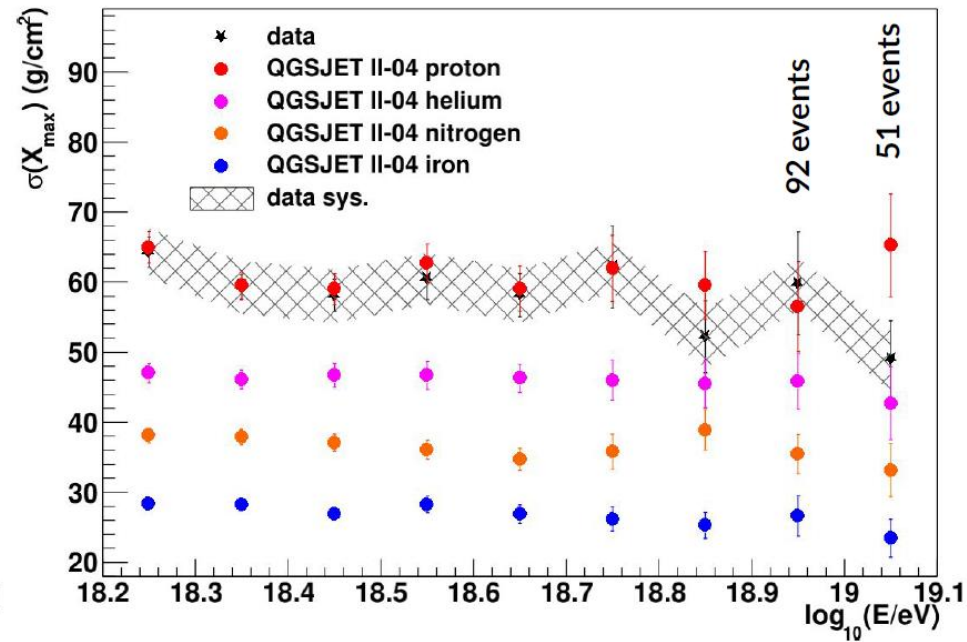
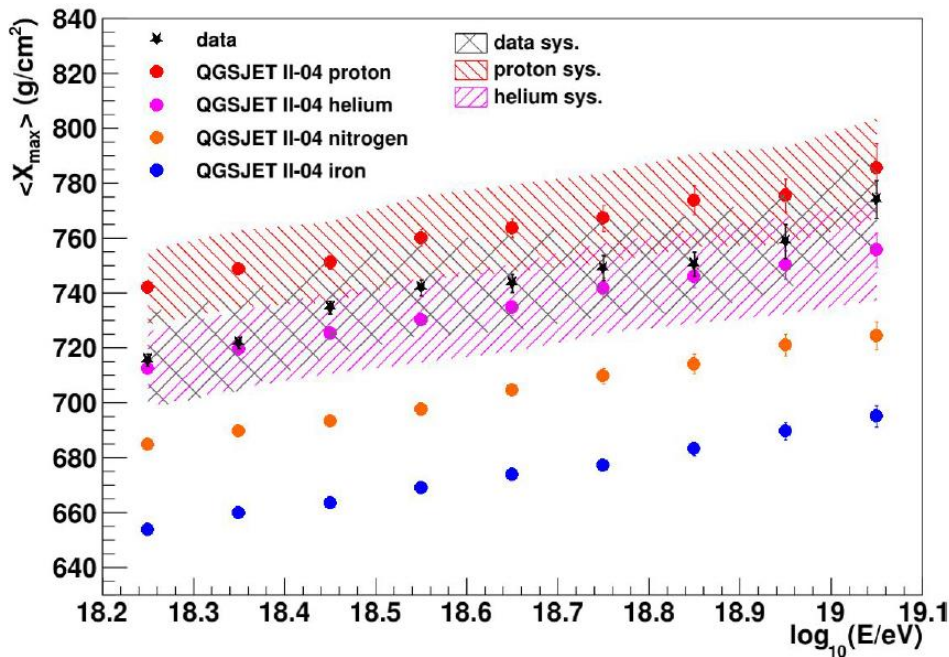
Energy spectrum anisotropy $E > 10^{19.2} \text{ eV}$
@7° from hotspot, ~30 deg. circle
post trial significance: 3.7σ

Composition Analysis with TA SDFD Hybrid Xmax

W. Hanlon, ICRC2019

10 years SD and FD hybrid data
 $\sigma(X_{\max})$

Mean X_{\max}

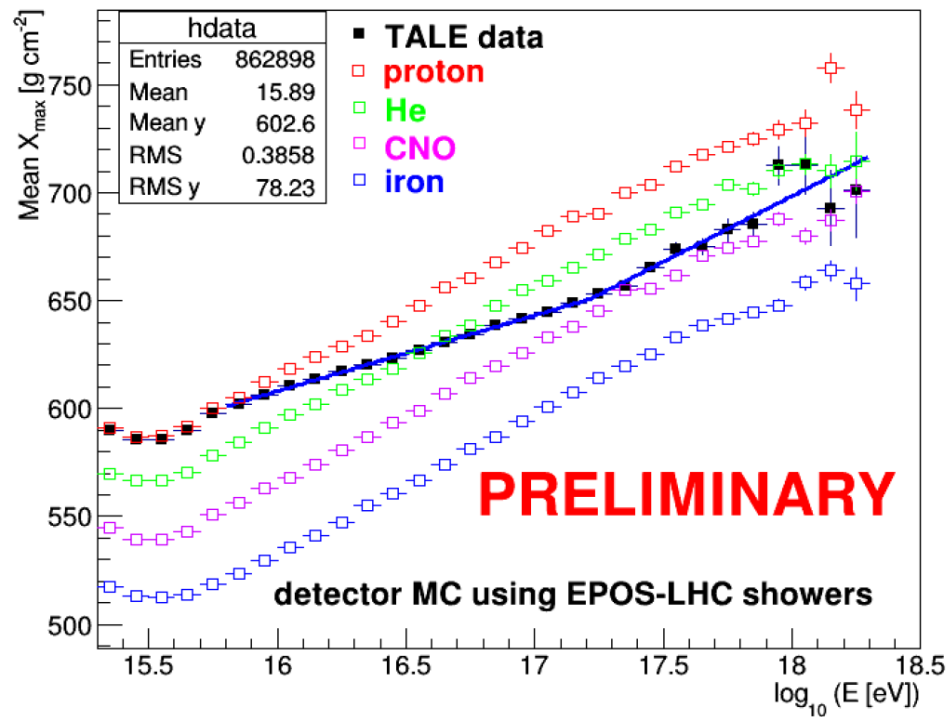


- Energy Range: $10^{18.2} \text{ eV} - 10^{19.1} \text{ eV}$
- 3560 events after the quality cuts
- Systematic uncertainty of $\langle X_{\max} \rangle$: $\pm 17 \text{ g/cm}^2$
- QGSjetII-04 interaction model was compared with the data
→ agreement with light composition
- More events are needed to study highest energies

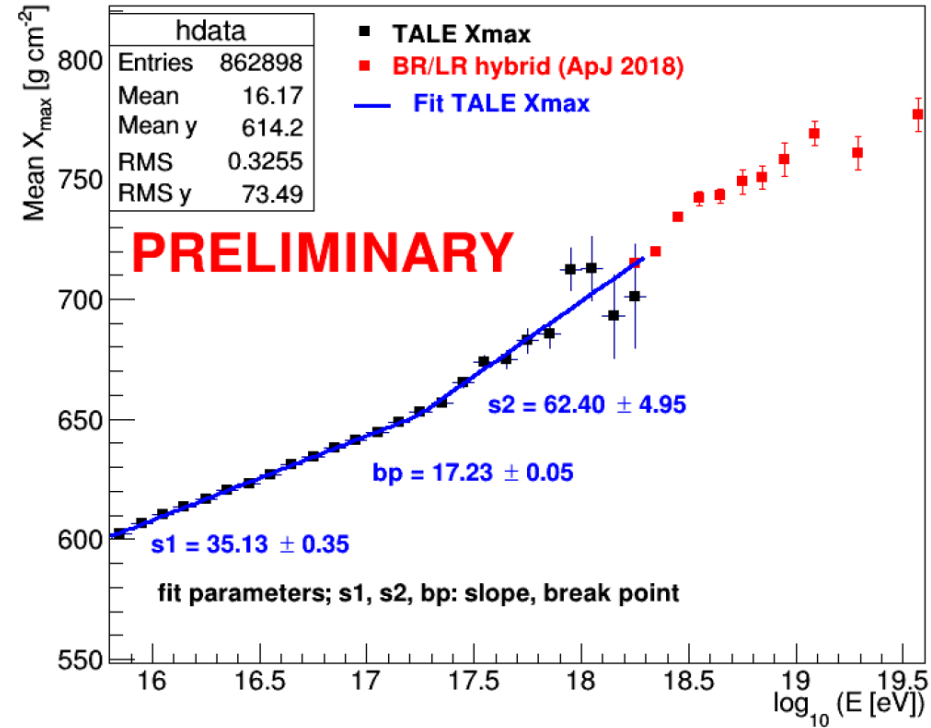
Composition Analysis with TALE FD Mono Xmax

T. Abu-Zayyad, ICRC2019

TALE Reconstructed Shower X_{\max} vs Shower Energy



TALE Mean X_{\max} vs energy

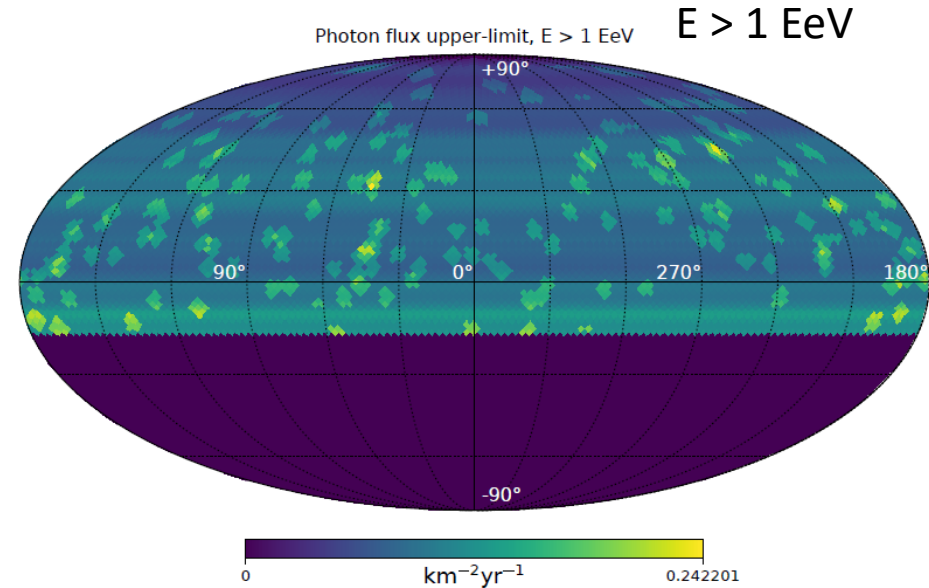
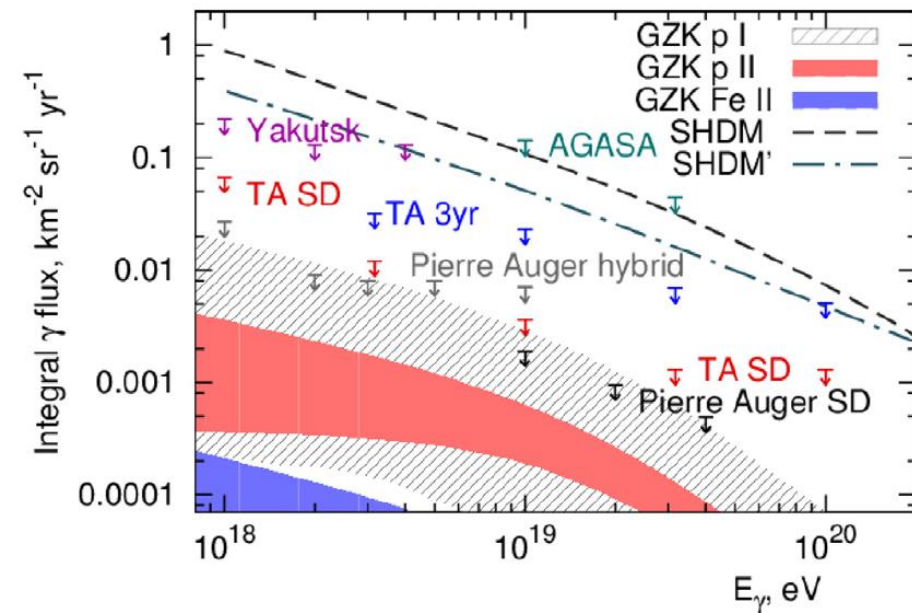


- Jun. 2014 – Nov. 2018 TALE FD mono data
- Energy Range: $10^{15.3} \text{ eV} - 10^{18.3} \text{ eV}$
- Break point $\log (E/\text{eV}) = 17.23 \pm 0.05$

Photon Search

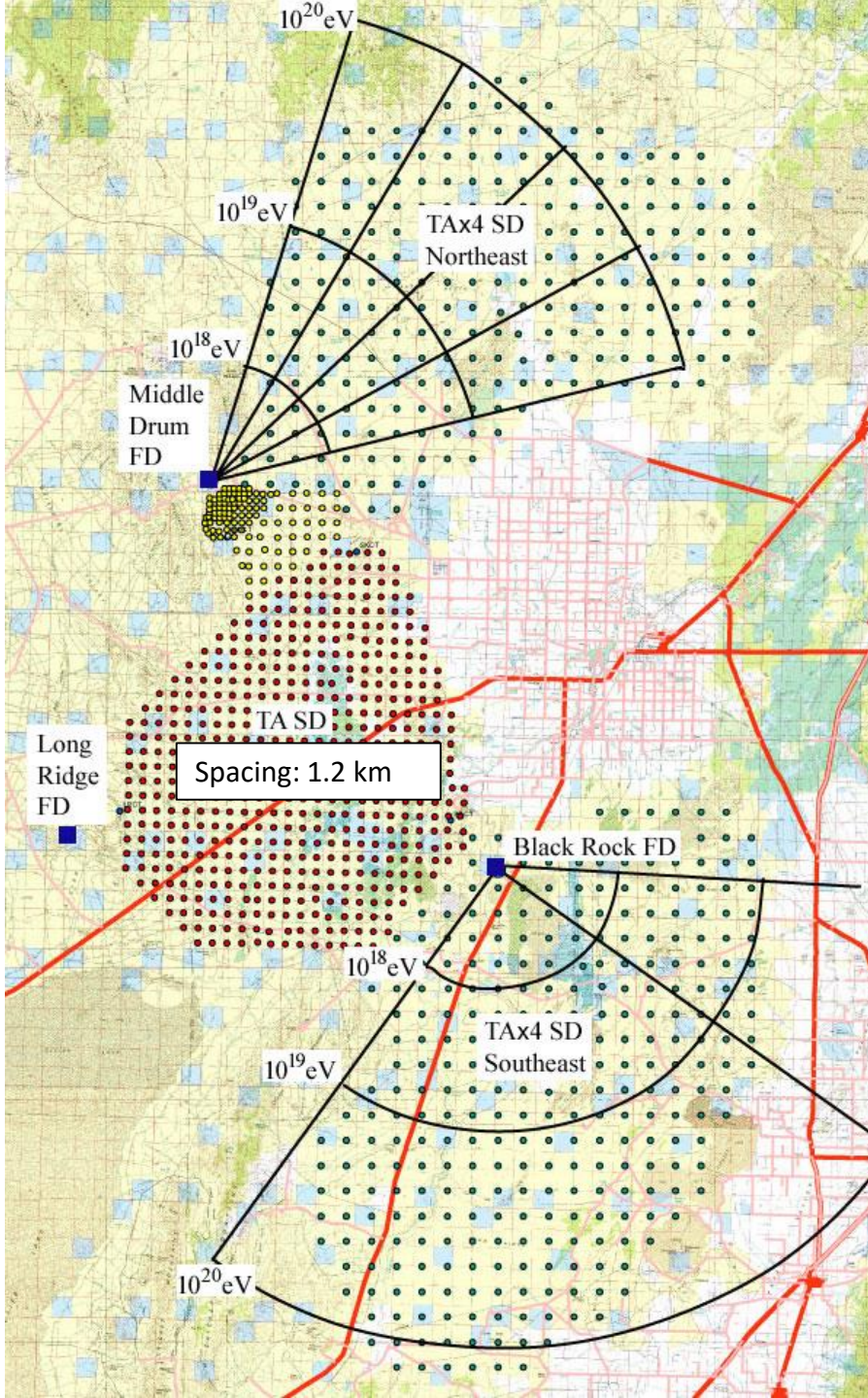
Astropart. Phys. **110**, 8 (2019)

M. Kuznetsov, ICRC2019



- Left: updated upper limit on GZK photons with 9 years TA SD data
 - Multivariable classifier is built on 16 reconstructed parameters
- Right: upper limit for directions in the field of view

The TAx4 experiment



To study more about the highest energies and examine the implications obtained by TA

500 new SDs with 2.08 km spacing

$E > 57 \text{ EeV}$:

- Reconstruction efficiency $> 95\%$
- Angular resolution: 2.2°
- Energy resolution: $\sim 25\%$

and TA SDs cover

$4 \times \text{TA SD}$ detection area ($\sim 3000 \text{ km}^2$)

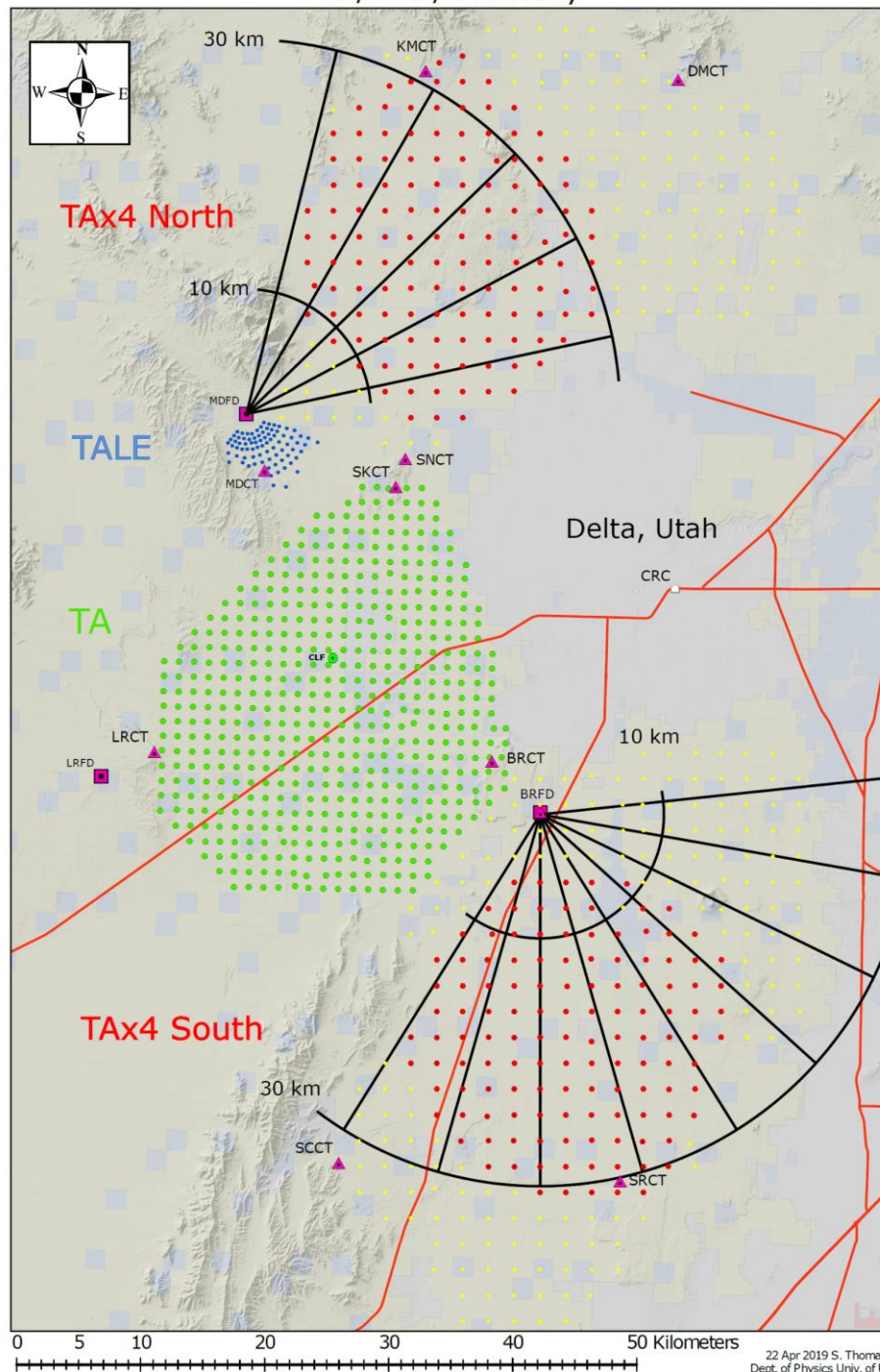
2 new Fluorescence Detector (FD) stations (4+8 HiRes Telescopes)

Deployment of Assembled SDs

<https://www.flickr.com/photos/142880279@N06/albums/72157689940402503>

Helicopter for the transportation of SDs

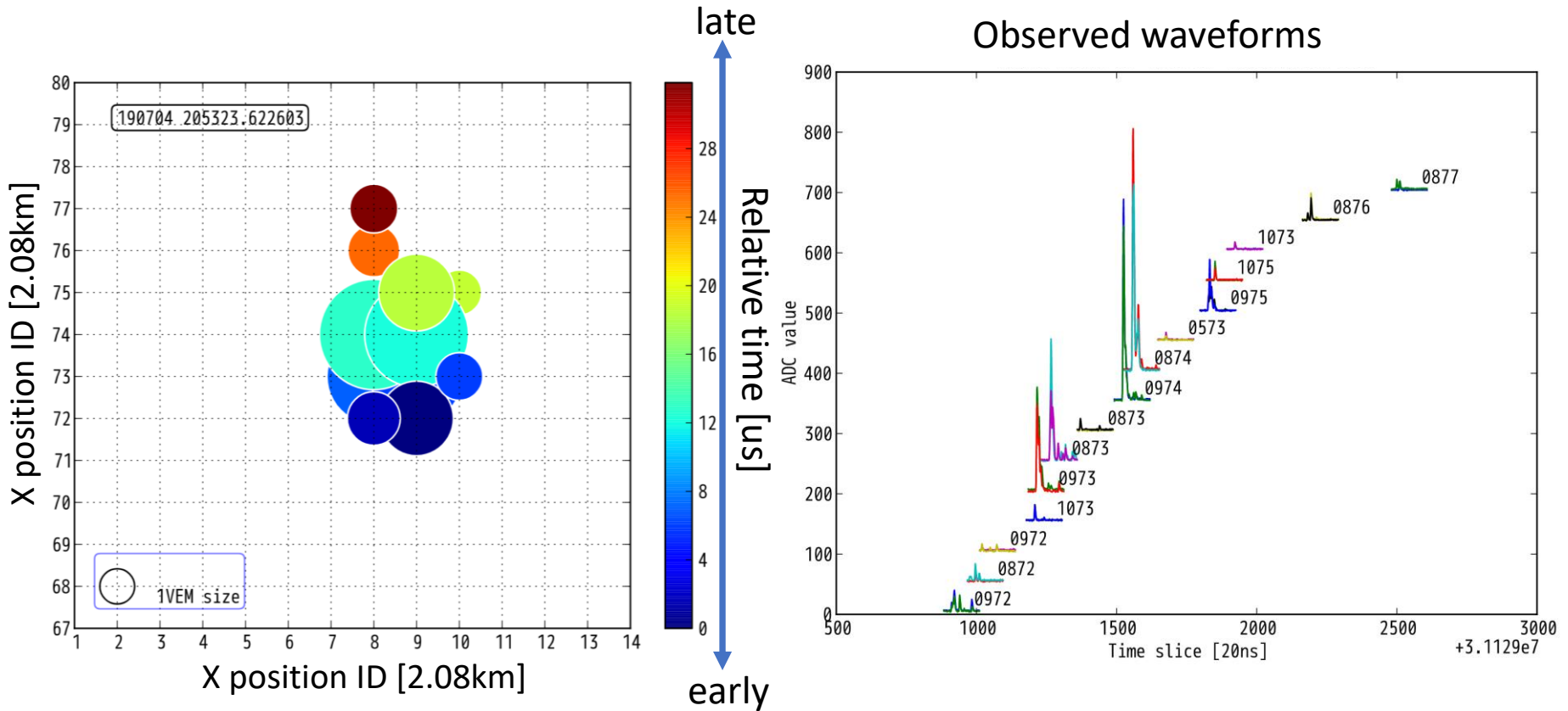




Deployed SDs and Communication Towers

- **More than half of SDs (257 SDs)** were deployed on 19 Feb. – 12 Mar. 2019.
- Locations of SDs were decided to optimize hybrid events above 10 EeV and consider practical conditions of wireless communications

Cosmic Ray Event



- Largest number of SDs were fired in this event
- **DAQ of each sub-array was started** from the end of **Apr. 2019** when the sub-array is ready.
- **Cosmic ray events** are being collected.
- $\sim 4 \times$ TA SD equivalent cosmic ray events are collected when the full operation is started.

Construction of FD Station



16th Feb. 2018 **First light** at the north FD station

22nd Oct. 2019 **First light** at the south FD station

Cosmic Ray Event

TA4XMDFD: 180616 074411.875382

Time, [μs]

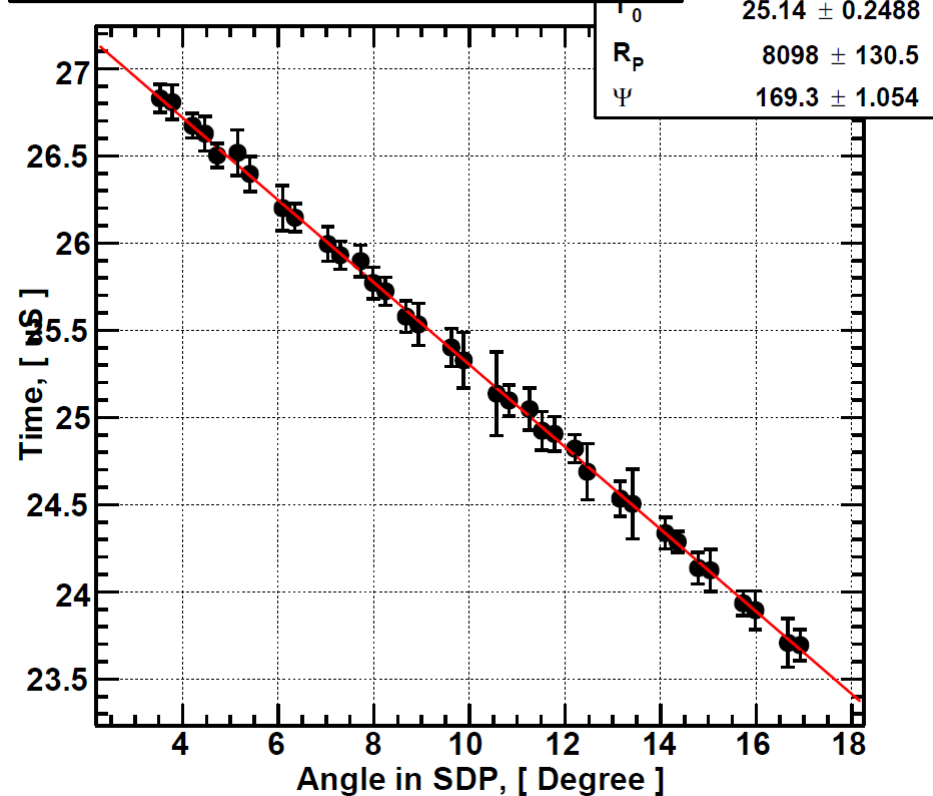
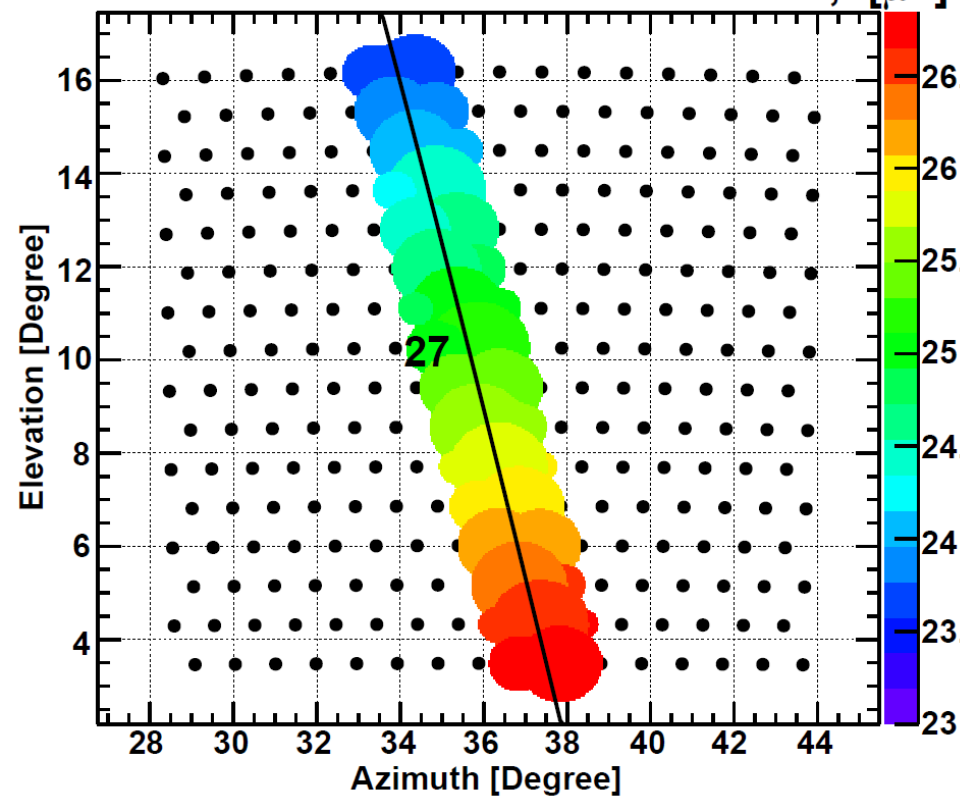
Time vs Angle, TA4XMDFD-Mono

df 2.59 / 32

25.14 ± 0.2488

R_p 8098 ± 130.5

Ψ 169.3 ± 1.054



- Stable operation of north FD station was started from 8th June 2018.
- Data analysis is ongoing.
- X_{max} : $\sim 3 \times$ TA SDFD equivalent events are collected at the highest energies when the full operation is started.

Summary

- Energy Spectrum
 - Combined spectrum of TA SD with TALE FD mono
 - Energy spectrum with $\log(E/\text{eV}) > 15.3$ was obtained.
 - Declination dependence of TA SD energy spectrum was implicated in 4.3σ
- Anisotropy
 - Hotspot ($> 57 \text{ EeV}$): $\sim 3\sigma$ global significance was obtained from 11 years TA SD data
- Composition:
 - TA SD and FD hybrid: consistent with light composition with $\log(E/\text{eV}) > 18.2$ and $\log(E/\text{eV}) < 19.1$
 - TALE FD mono: preliminary Xmax results were obtained with $\log(E/\text{eV}) > 15.3$
- Photon limit was updated with 9 years TA SD data and the dependence of the arrival directions was also obtained.
- Implications on anisotropy were obtained by the TA experiment.
- TAx4 experiment is **in operation**
- TAx4 detectors:
 - **500 new** SDs with **2.08 km** spacing + TA SDs → Coverage of **$4 \times \text{TA SDs}$** $\sim 3000 \text{ km}^2$
 - **2 new** Fluorescence Detector (FD) stations (4+8 Telescopes)
- **More than half of TAx4 SDs** were deployed, and **2 TAx4 FD** stations were constructed.
- **Data acquisition was started.** SD: from **Apr. 2019**, FD: from **Jun. 2018**. Cosmic ray events are being collected.
- Prospects
 - $\sim 4 \times \text{TA SD}$ equivalent cosmic ray events with $E > 57 \text{ EeV}$ will be collected when the full operation is started.
 - $\sim 3 \times \text{TA SDFD}$ equivalent hybrid events will be collected especially for Xmax at the highest energies when the full operation is started.

Backup