

Electron Injection in Merger Shocks of Galaxy Clusters

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Introduction

- radio and X-ray observations of so-called radio relics indicate electron acceleration at merger shocks in galaxy clusters
- these large-scale shocks are also candidate sites for ultra-high-energy cosmic ray production
- most energetic merger shocks have low Mach numbers ($M_s < 4$, $M_A < 10$) and propagate in high beta ($\beta >> 1$) plasmas.
- Diffusive Shock Acceleration (DSA) assumed to operate at merger shocks but mechanism of electron injection is poorly known for galaxy cluster conditions

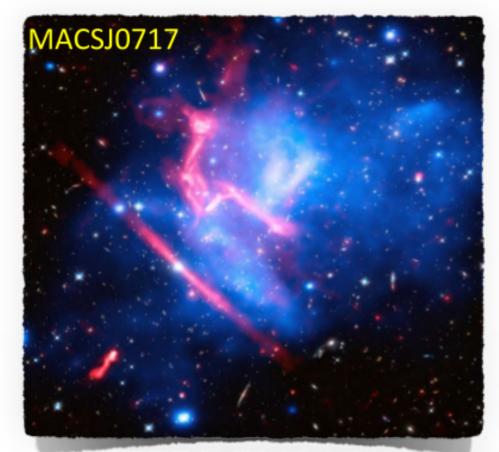
Alfvenic Mach number: $M_{
m A}=rac{v_{
m sh}}{v_{
m A}}$

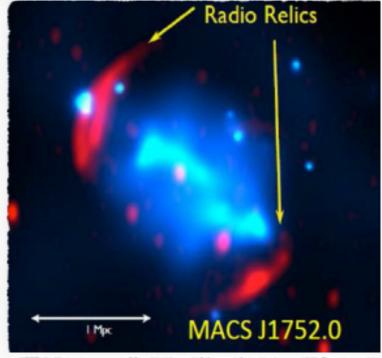
Sonic Mach number: $M_{
m s} = rac{v_{
m sh}}{c_{
m s}}$

Plasma beta: $\beta = p_{\rm th}/p_{\rm mag}$

$$v_{\rm A} = \frac{B_0}{\sqrt{\mu_0(N_e m_e + N_i m_i)}}$$

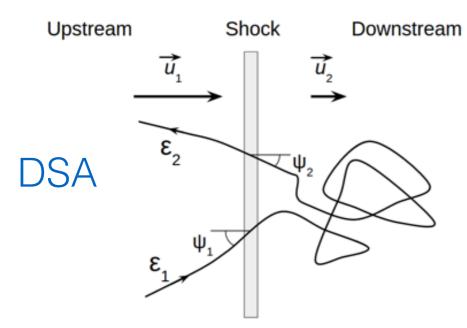
$$c_s = \sqrt{2\Gamma k_B T_i / m_i}$$





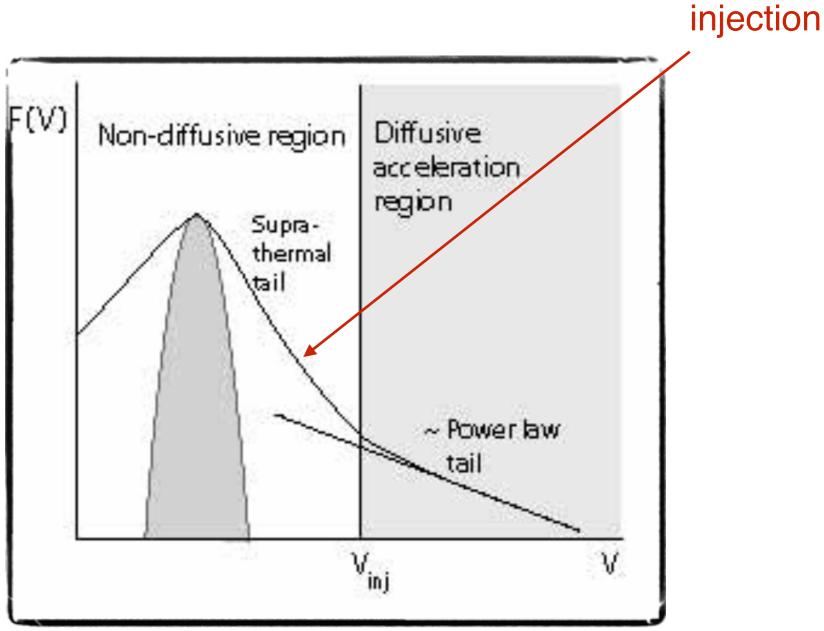
White – optical (Hubble)
Blue – X-ray (Chandra)
Red – radio (VLA)

Particle injection to DSA

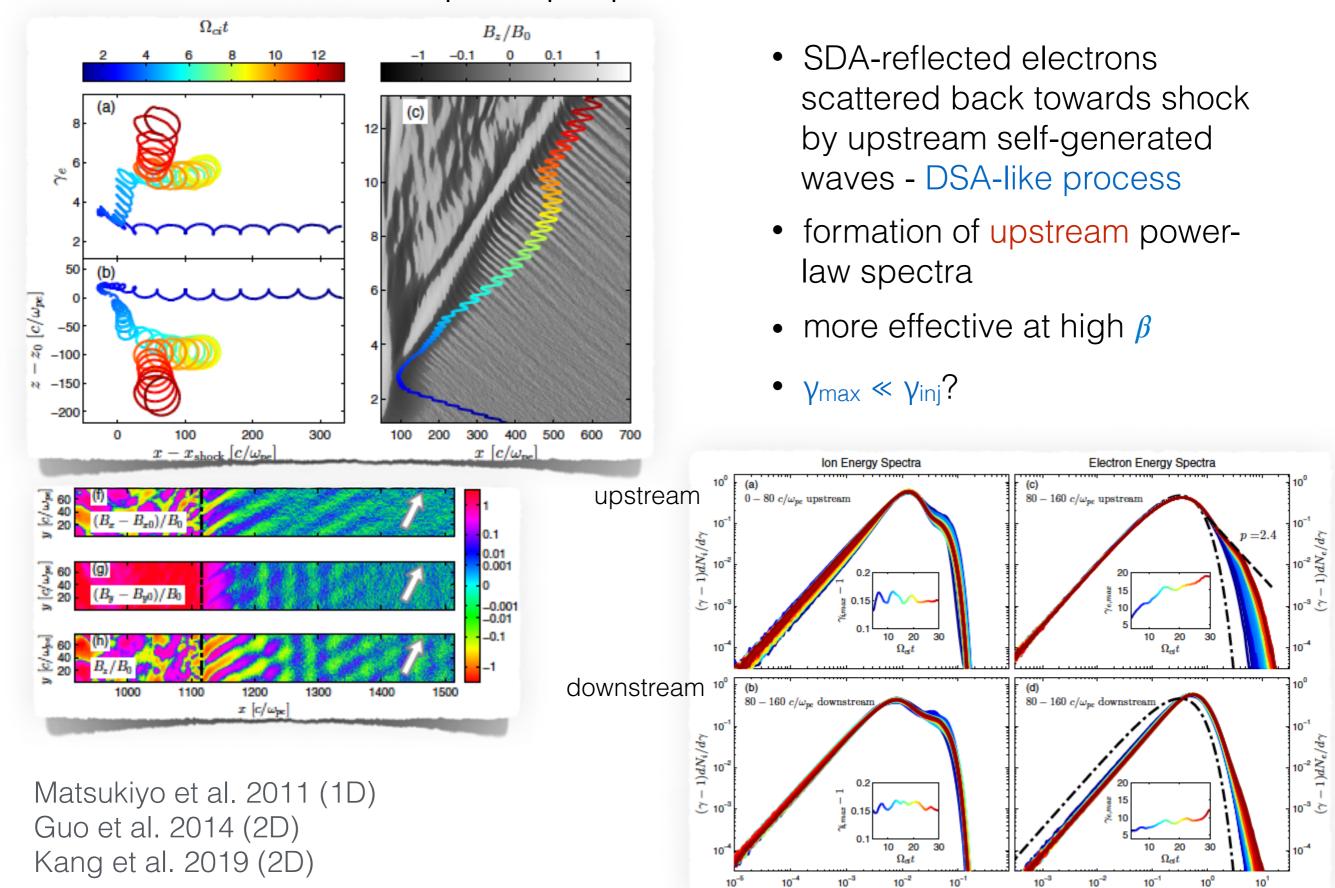


$$d_{sh} \sim (1\text{-}100) \; \lambda_{gi}$$

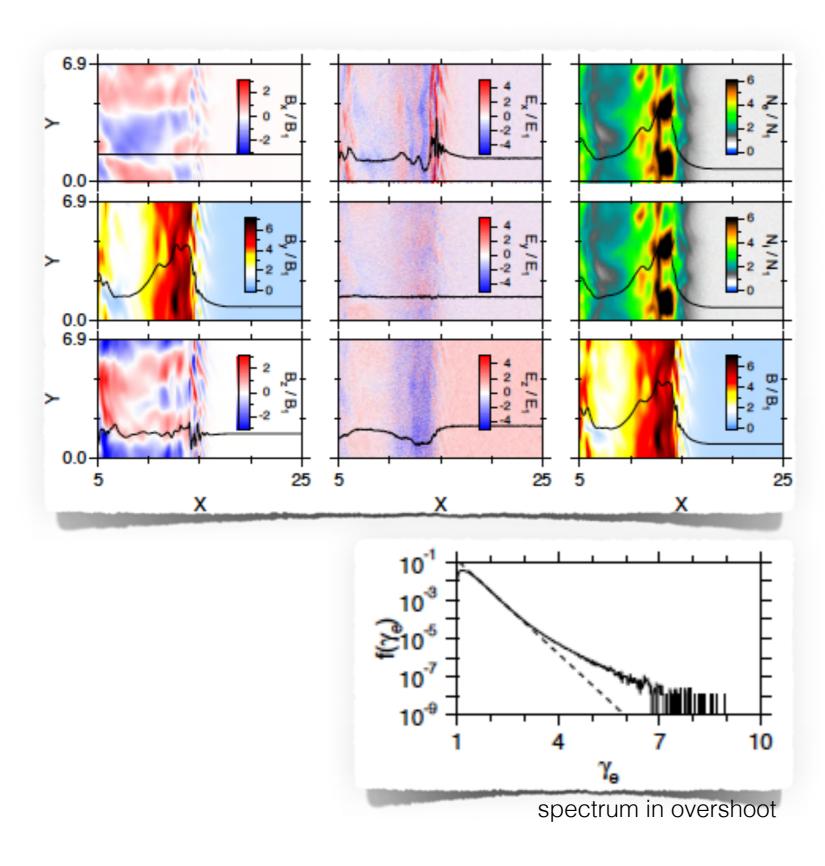
$$r_g(\epsilon_{inj}) > d_{sh}$$



Previous work: multiple Shock Drift Acceleration cycles at quasi-perpendicular shocks



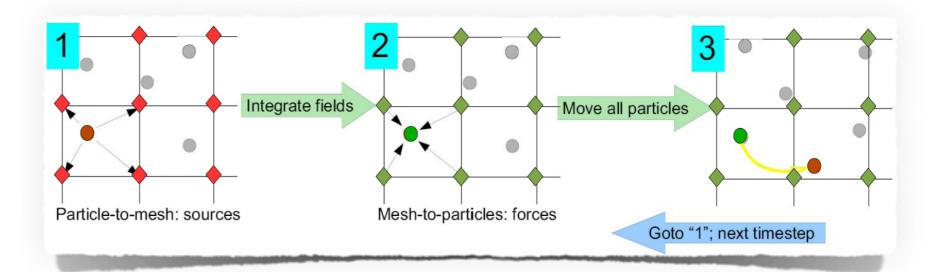
Effects of the shock rippling

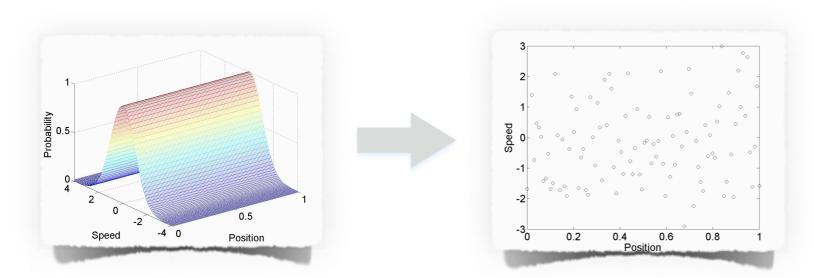


- no reflected electrons because of ion-scale shock ripples
- SDA does not work
- acceleration by scattering on the waves in the shock transition instead
- if the same shock rippling mechanism operates for conditions assumed in Guo et al. 2014 and Kang et al. 2019, then their simulations might have not been able to resolve it

Method of Particle-In-Cell Simulations

- Fully self-consistent description of collisionless plasma:
 - Vlasov equation (kinetic theory; time evolution of particle distribution function $f(\mathbf{x}, \mathbf{v}, t)$ in phase-space) + Maxwell's equations
- Particle-In-Cell modeling an ab-initio method of Vlasov equation solution through:
 - integration of Maxwell's equations on a numerical grid
 - integration of relativistic particle equations of motion in collective self-consistent EM field



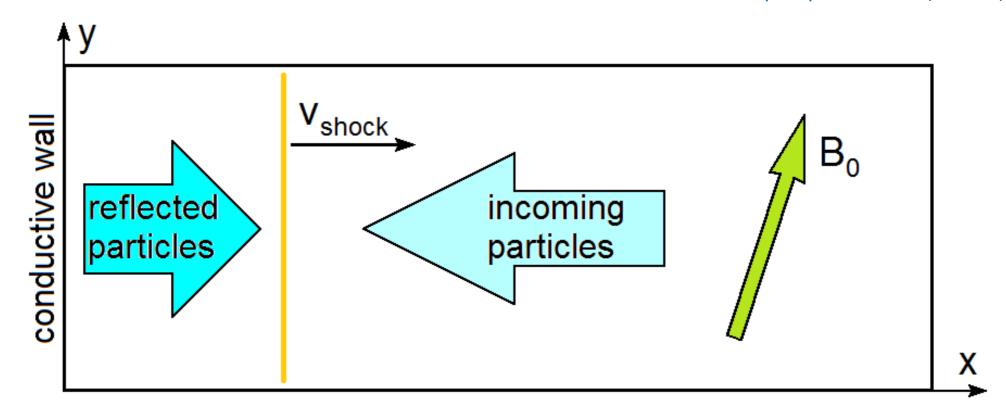


Particle distribution function represented by macroparticles on a numerical grid.

(Macroparticles represent a small volume of particle phase-space; equations of motion as for realistic particles)

Our current work: investigate multi-scale electron acceleration physics with large-scale 2D Particle-In-Cell simulations

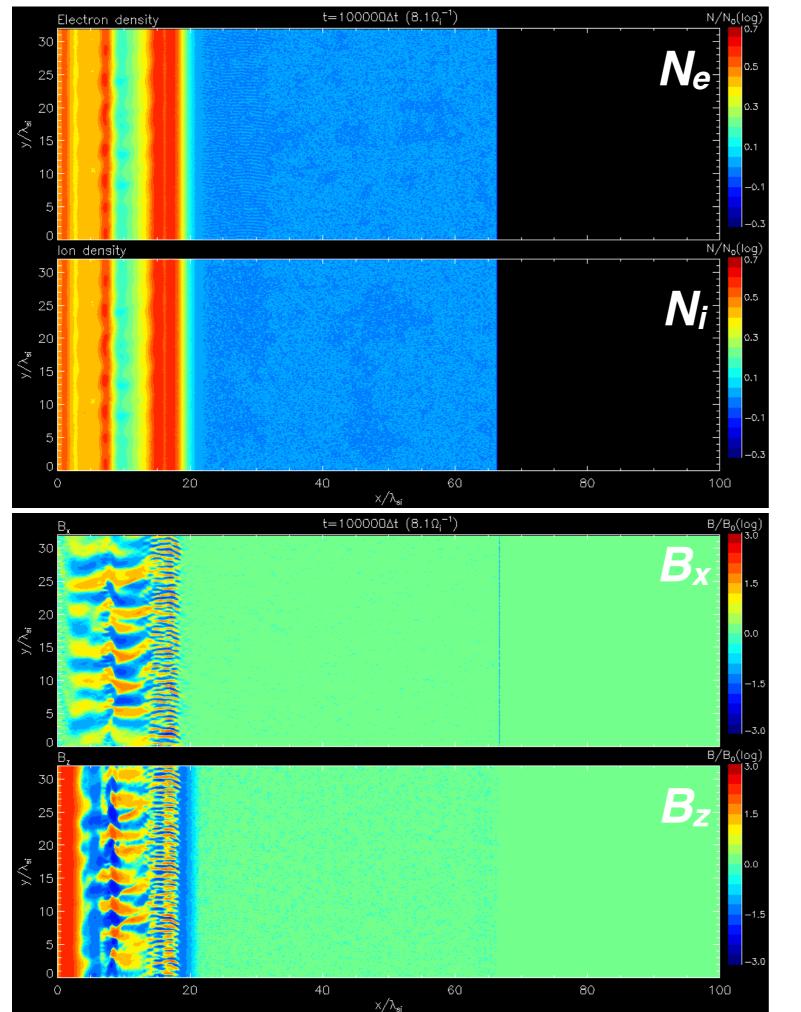
Kobzar et al., in preparation (2020)



2D3V (Lx x Ly = 333
$$\lambda_{si}$$
 x 32 λ_{si}), M_A=6.1, M_s=3, m_i/m_e=100, v₀=0.1c, β =5 (plasma temperature $k_BT \approx 40$ keV)

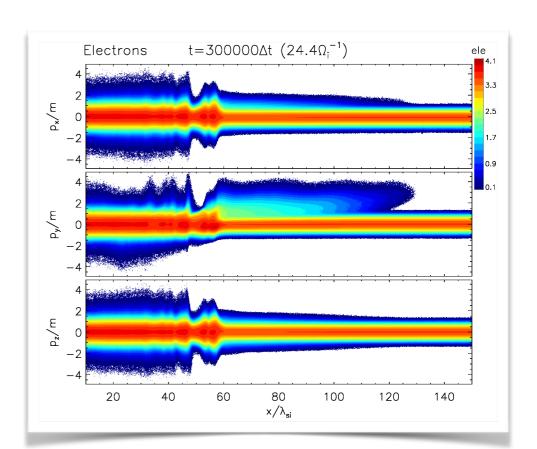
subluminal shock: $\theta = 75^{\circ} (\theta_{cr} \approx 81^{\circ})$

$$v_t \simeq 1.5 v_{\rm th,e} > v_{\rm th,e} \quad (v_t = v_{\rm sh}^{\rm up}/\cos\vartheta)$$

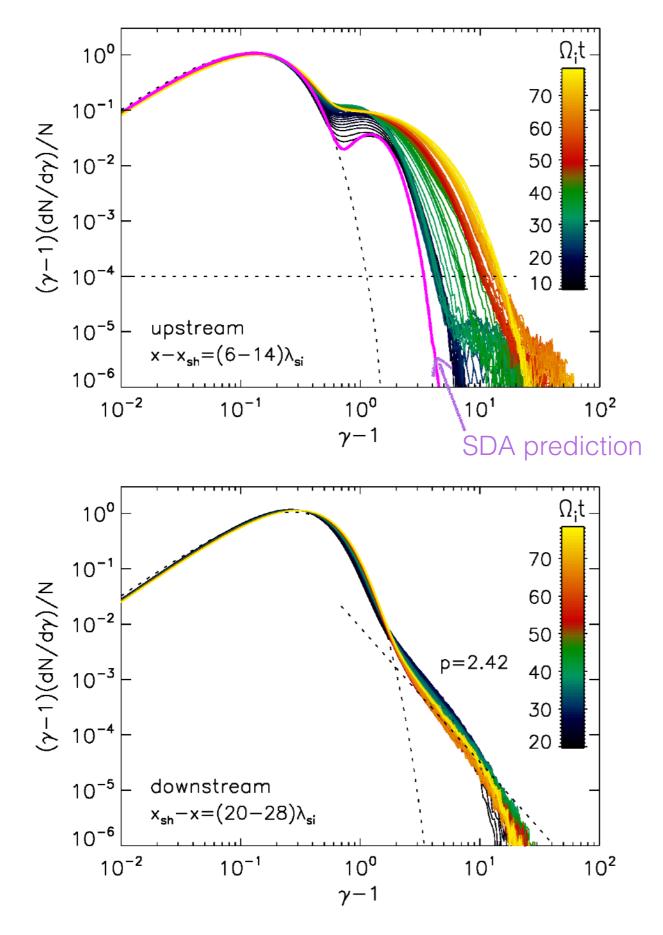


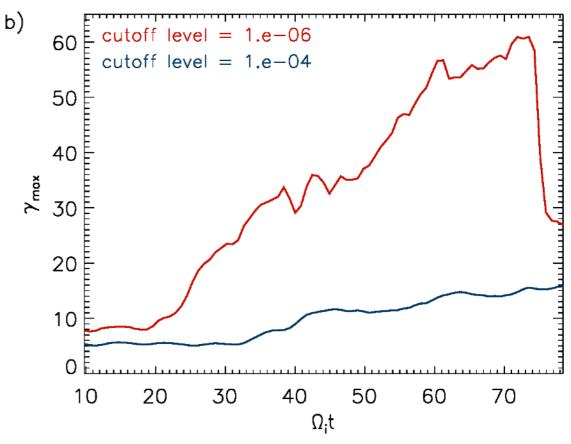
Global shock structure: multi-scale turbulence

- rippling in the shock transition on different scales (overshootundershoot-2nd overshoot)
 - AIC and mirror modes
- short-scale whistler waves in the overshoot
- oblique and perpendicular modes of the electron firehose instability in the upstream, enhanced and modulated by the ripples



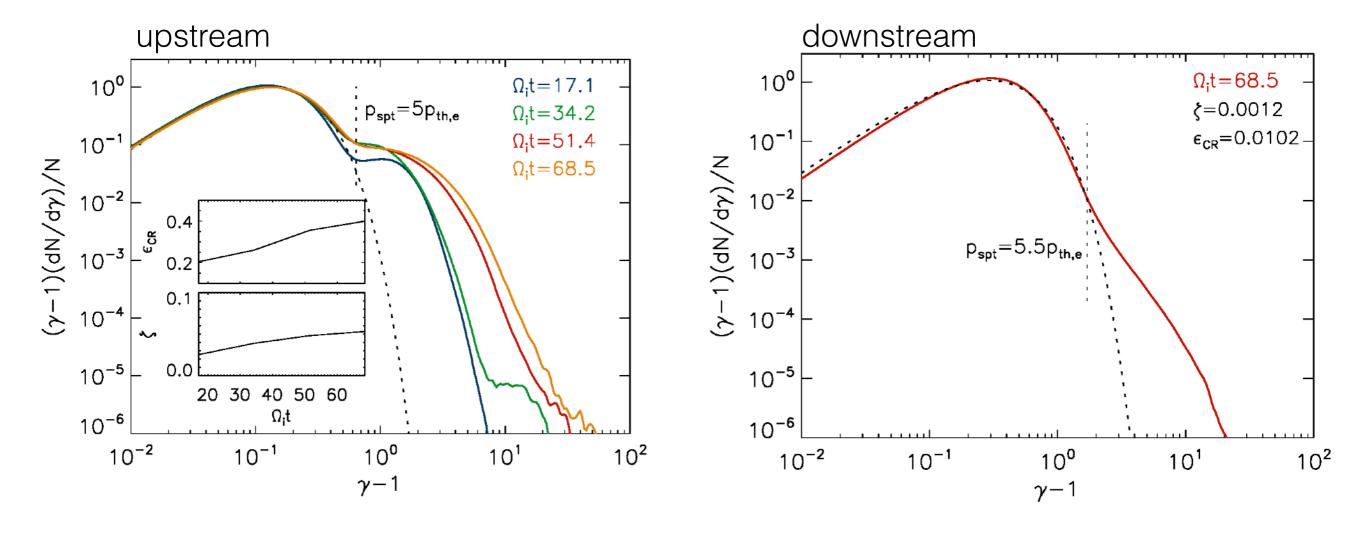
Electron spectra – time evolution





- substantial increase in non-thermal tail production efficiency coincident with the onset of the shock rippling at $\Omega_{ci}t\approx 20$
- power-law spectra downstream in agreement with observations

Electron spectra – injection efficiency



$$\zeta = \frac{4\pi}{N_2} \int_{p_{\rm spt}}^{p_{\rm max}} \langle f(p) \rangle \, p^2 dp \quad \text{- fraction of supra-thermal electrons}$$

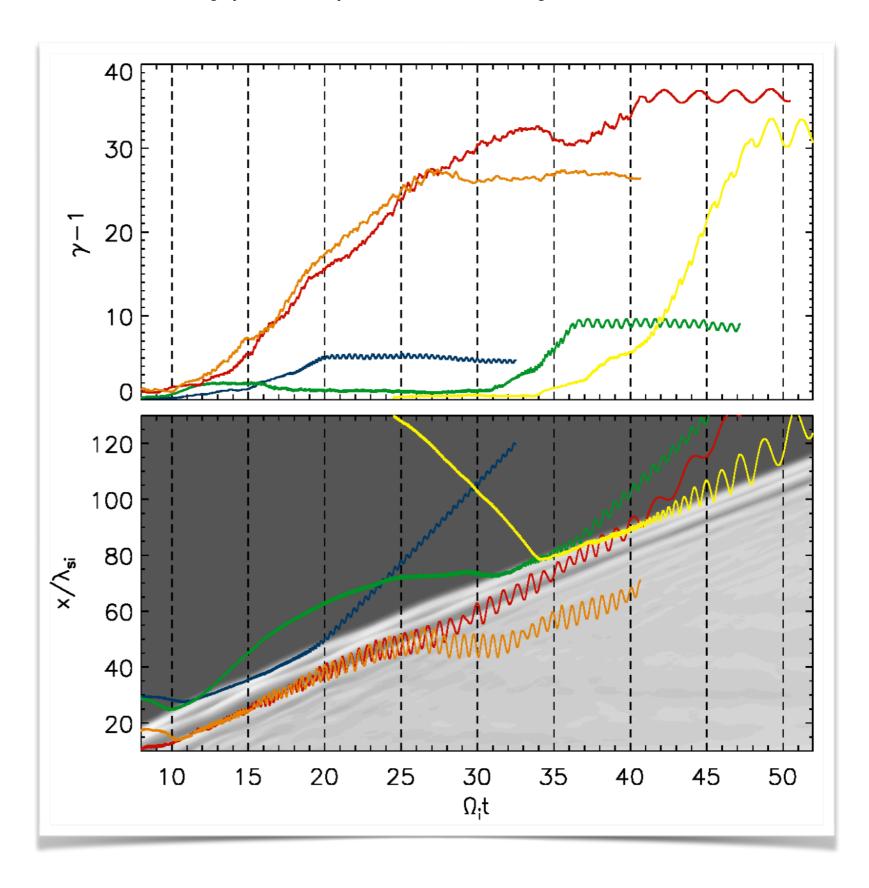
 ϵ_{CR} - corresponding energy density fraction

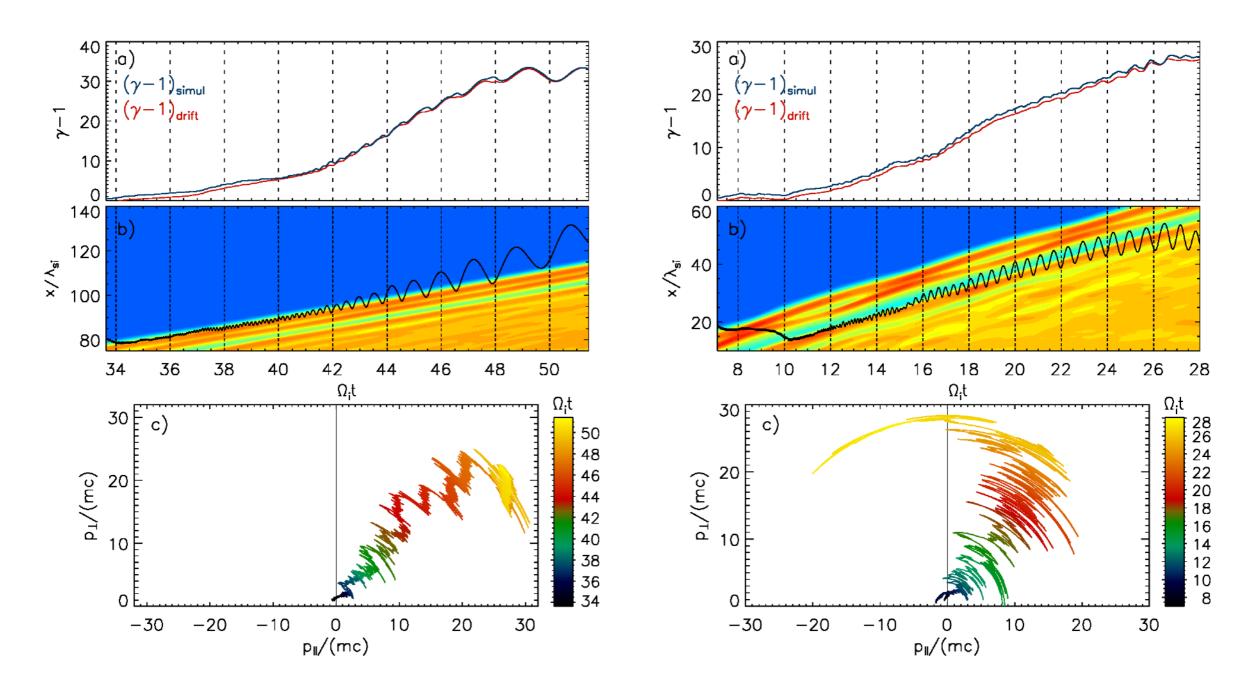
$$\zeta_{\rm max,up} \simeq 5\%, \ \epsilon_{\rm CRmax,up} \simeq 40\%$$

$$\gamma_{\rm max,up} \approx 40 - 60 \quad \leftarrow \gamma_{\rm inj} \approx 25 \ (p_{\rm inj} \sim 3 \, p_{\rm th,i})$$

Acceleration processes - typical particle trajectories

- most particles gain their energies in a single interaction with the shock
- acceleration time much longer than predicted by SDA ($\sim 1/\Omega_i$)
- acceleration takes place also deep in the shock transition





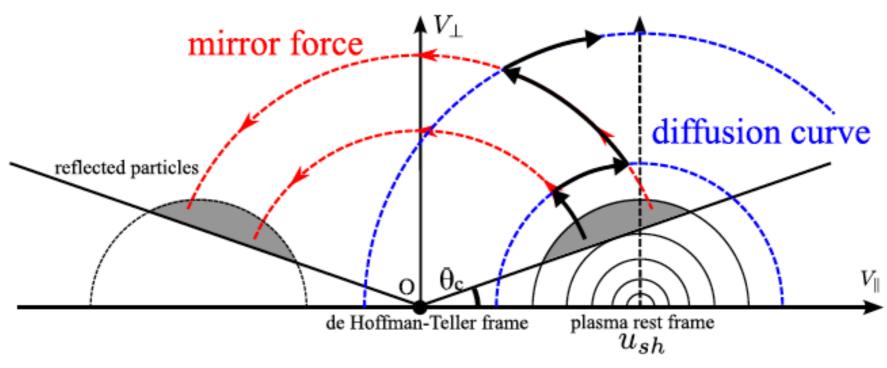
- most accelerations associated with an increase in p_
- strong pitch-angle scattering (arcs in p_{II}-p_⊥ momentum space)
- energy gain mostly through the drift along motional electric field:

$$\Delta \gamma_{\rm drift} = (-e/m_{\rm e}c^2) \int E_z \, dz$$

→ Stochastic Shock-Drift Acceleration (SSDA)

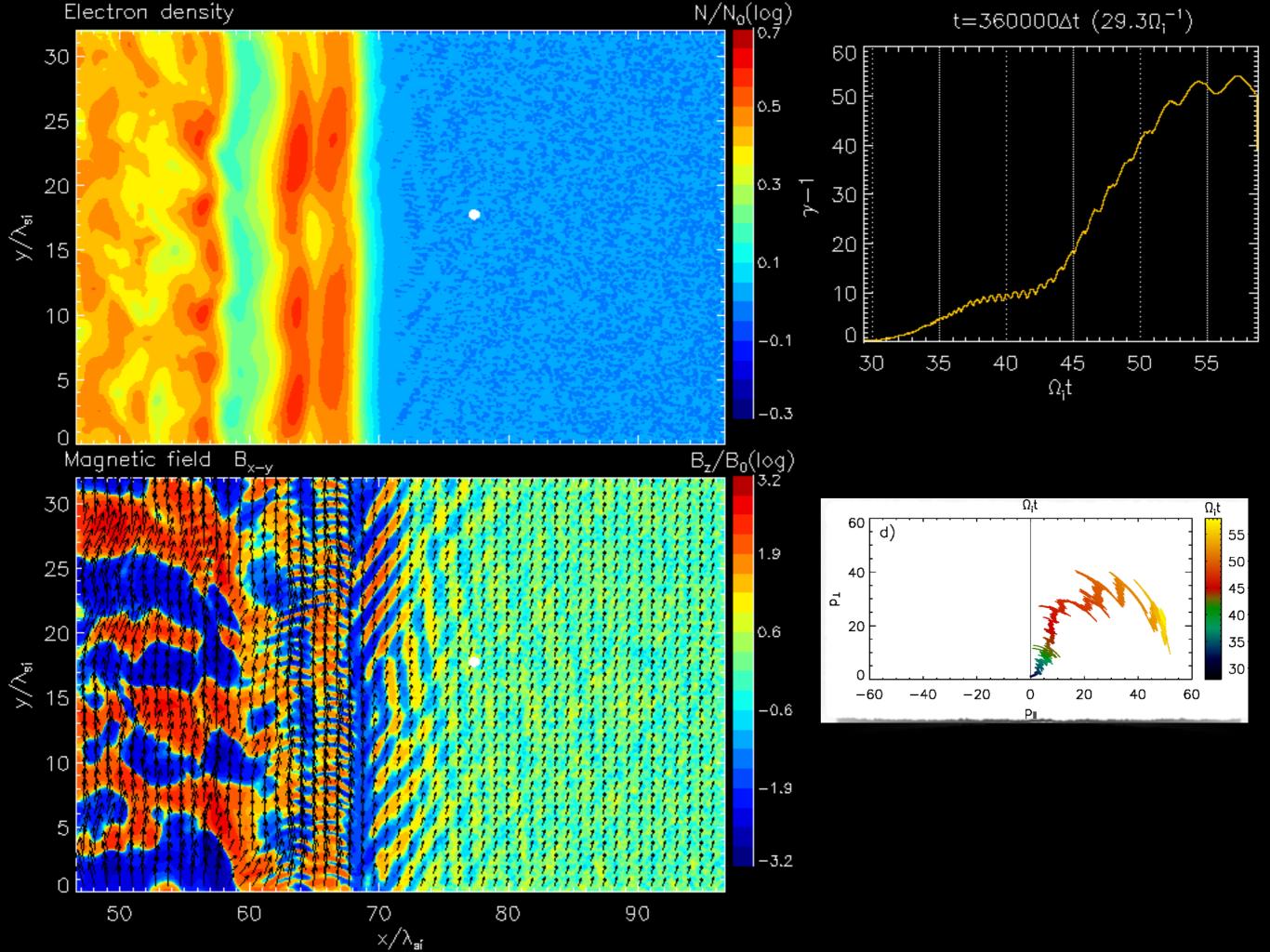
Stochastic Shock Drift Acceleration (SSDA)

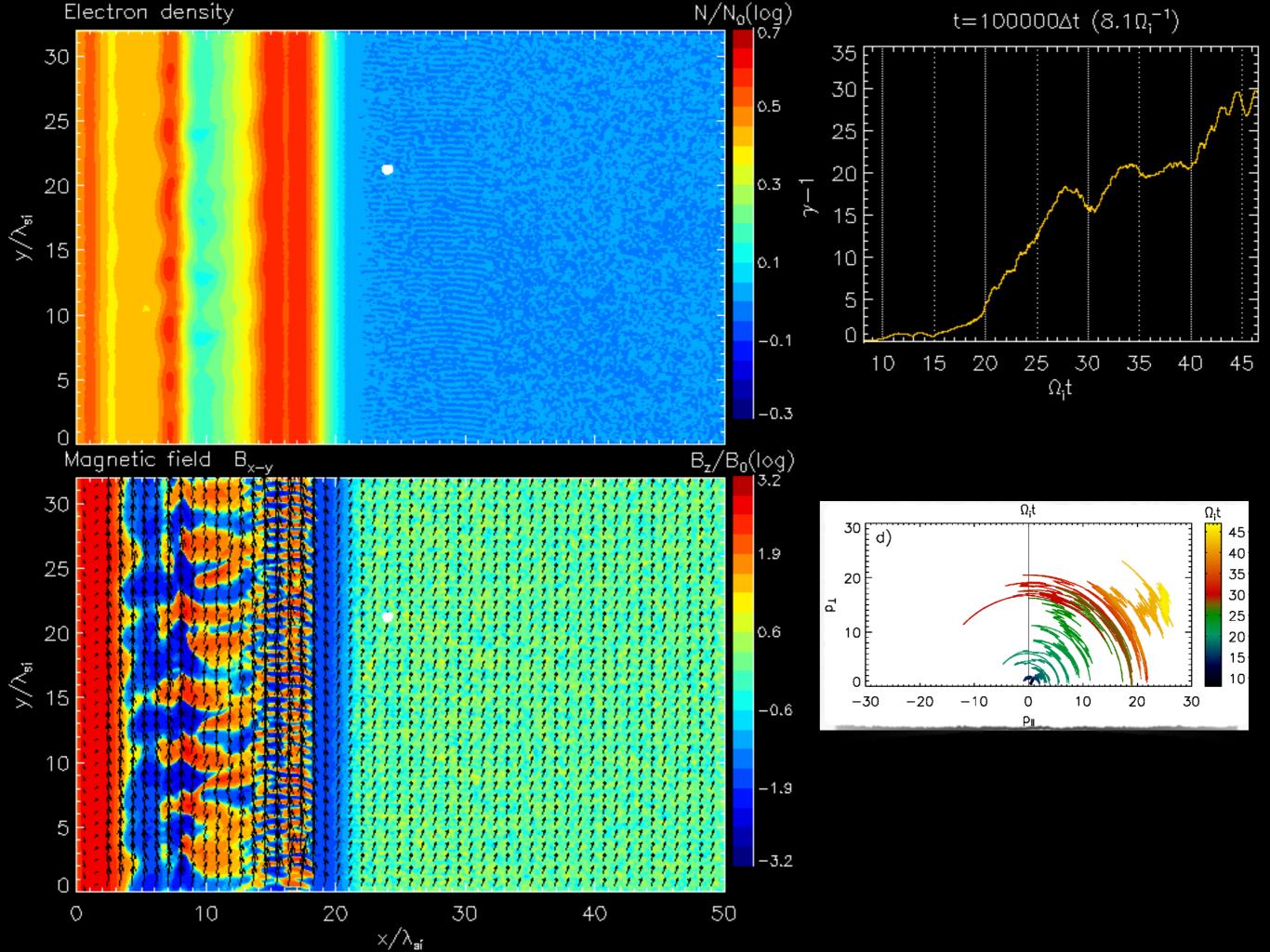
Katou & Amano (2019)



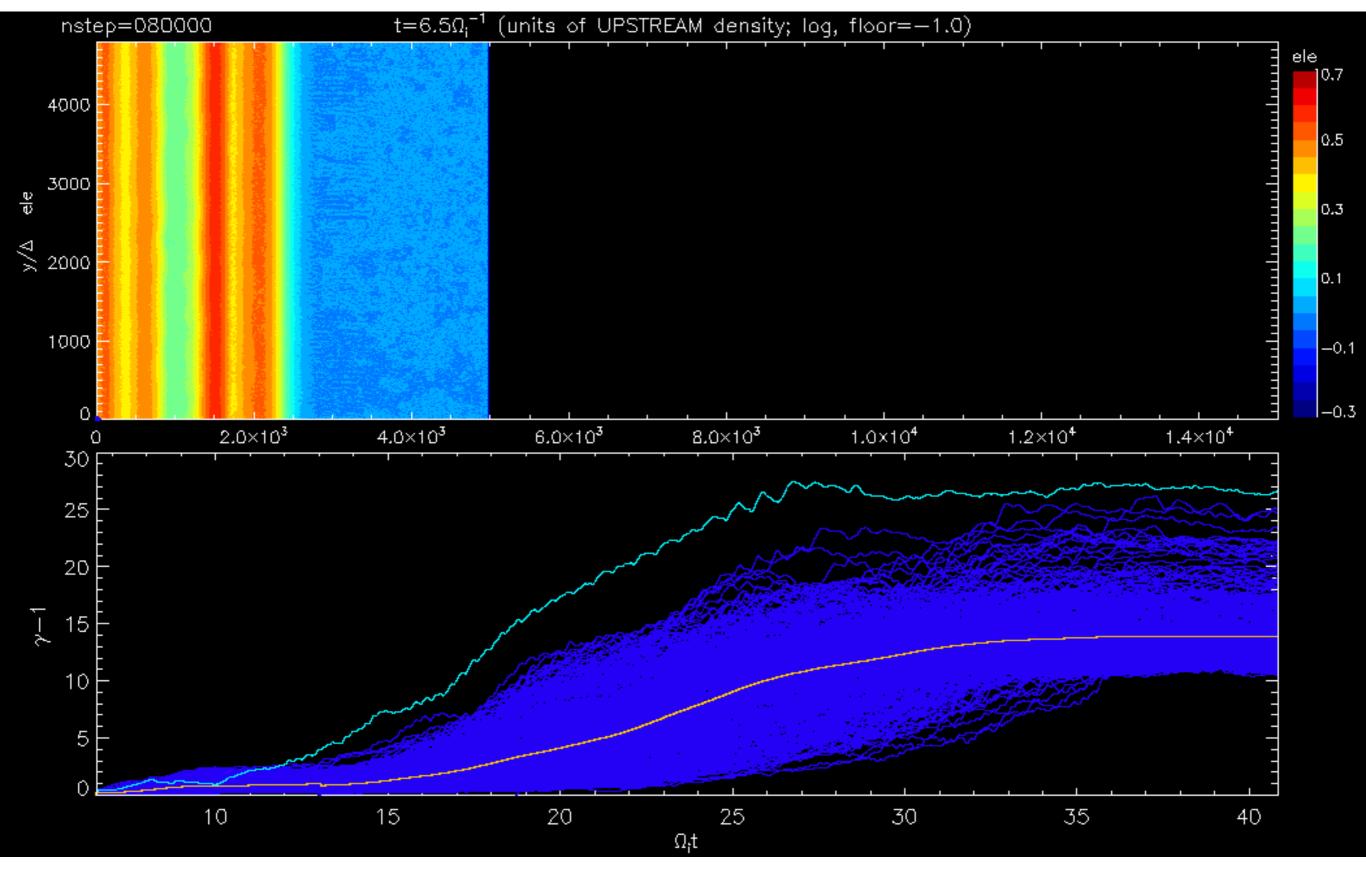
- adiabatic mirror reflection in the HTF
- elastic scattering (diffusion) in the plasma rest frame

- electrons are confined in the shock transition region by stochastic pitch-angle scattering off magnetic turbulence and gain energy through SDA (non-adiabatic acceleration)
- longer particle confinement increases energy gains and enables more efficient acceleration





Downstream spectrum formation



- downstream particles accelerated in the shock through SSDA
- advection of upstream-reflected particles plays minor role

Summary and conclusions

- kinetic modeling of particle acceleration at low Mach number shocks in highbeta plasmas requires multi-dimensional and large-scale effects to be taken into account
- for parameters analysed in this work we find the presence of multi-scale turbulence, including ion-scale shock rippling modes, to be critical for efficient electron acceleration
- electron injection proceeds mainly through the stochastic SDA process, effects of multi-SDA cycles are also observed
- acceleration to very high energies occurs that should lead to electron injection to DSA in the presence of long-wave (MHD) upstream turbulence
- work to be continued