

Detecting Gauged $L_\mu - L_\tau$ using Neutron Star Binaries

Tuesday 12 November 2019 12:25 (5 minutes)

We show that gravitational wave emission from neutron star binaries can be used to discover ultra-light $U(1)_{L_\mu - L_\tau}$ vectors by making use of the large inevitable abundance of muons inside neutron stars. In pulsar binaries the $U(1)_{L_\mu - L_\tau}$ vectors induce an anomalously fast decay of the orbital period through the emission of dipole radiation. We study a range of different pulsar binaries, finding the most powerful constraints for vector masses below $\mathcal{O}(10^{-18} \text{ eV})$. For merging binaries the presence of muons in neutron stars can result in dipole radiation as well as a modification of the chirp mass during the inspiral phase. We make projections for a prospective search using the GW170817 event and find that current data can discover light vectors with masses below $\mathcal{O}(10^{-18} \text{ eV})$. In both cases, the limits attainable with neutron stars reach gauge coupling $g' < \sim 10^{-20}$, which are many orders of magnitude stronger than previous constraints. We also show projections for next generation experiments, such as Einstein Telescope.

Author: Dr LAHA, Ranjan (CERN)

Co-authors: OPFERKUCH, Toby Oliver (CERN); Dr DROR, Jeff (Lawrence Berkeley National Laboratory)

Presenter: Dr LAHA, Ranjan (CERN)

Session Classification: Lightning talks