# Mechanical quantum sensing in the search for dark matter

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#### Macroscopic quantum coherence





New, open frontier: quantum coherence, entanglement with many particles, large volumes, distances, ...

Sensing at the level of vacuum fluctuations of macroscopic objects becoming routine

True quantum control of larger systems possible

Technology  $\leftarrow \rightarrow$  theory

### **Quantum-limited detection**





The Sensitivity of the Advanced LIGO Detectors at the Beginning of Gravitational Wave Astronomy LIGO Collaboration 1604.00439



Teufel et al, Nature 2011



Matsumoto et al, PRA 2015



Aspelmeyer ICTP slides 2013

![](_page_3_Picture_5.jpeg)

#### PHYSICAL REVIEW LETTERS

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Featured in Physics

Demonstration of Displacement Sensing of a mg-Scale Pendulum for mm- and mg-Scale Gravity Measurements

Nobuyuki Matsumoto, Seth B. Cataño-Lopez, Masakazu Sugawara, Seiya Suzuki, Naofumi Abe, Kentaro Komori, Yuta Michimura, Yoichi Aso, and Keiichi Edamatsu Phys. Rev. Lett. **122**, 071101 – Published 19 February 2019

Physics See Synopsis: Gravity of the Ultralight

 $\rm F_{grav}$  =  $\rm G_{N}~m^{2}/d^{2} \sim 10^{-17}~N$  for two masses m = mg separated by d = mm

cf.  $10^{-21}$  N/ $\sqrt{Hz}$  (and better) sensitivities achieved optomechanically

Dark Matter Mass  $\log[m/\text{GeV}]$ 

![](_page_5_Figure_1.jpeg)

Part 1: ultralight DM detection

Part 2: "heavy" DM detection

### **Ultralight DM detection**

Ex: DM coupled to *neutrons* (B-L charge),  $m\phi \le 1 \text{ meV}$  ( $\lambda \ge 10^{-3} \text{ m}$ ).

Coherent, persistent, oscillating force on mechanical sensor  $\rightarrow$  acceleration signal

$$\mathcal{L}_{int} = g_{B-L} \mathcal{A} \overline{n} n \quad \longrightarrow \quad F = g_{B-L} N_n F_0 \sin(\omega_s t)$$

Different couplings to different neutron/proton ratios ("EP-violating")  $\rightarrow$  use two sensors, material types to eliminates common mode backgrounds

*Ultralight dark matter detection with mechanical quantum sensors* **Carney**, Hook, Liu, Taylor, Zhao 1908.04797 (NJP)

*Dark matter direct detection with accelerometers* Graham et al 1512.06165 (PRD)

![](_page_6_Picture_7.jpeg)

### Some upcoming experiments

![](_page_7_Figure_1.jpeg)

![](_page_7_Figure_2.jpeg)

#### Particle DM detection

![](_page_8_Figure_1.jpeg)

#### Impulse detection

![](_page_9_Picture_1.jpeg)

![](_page_9_Figure_2.jpeg)

Sharp, rapid impulse signal (eg. particle colliding with a sensor)

Highly broadband in frequency domain--what are the quantum limits?

(NB on terminology: impulse =  $\int Fdt$  = momentum transfer)

# Quantum limits in impulse sensing

Standard quantum limit for momentum transfer:  

$$\Delta p_{SQL} = \sqrt{\hbar m_s \omega}$$
600 keV (m = 1 ng,  $\omega$  = 1 kHz)  
Again this is just a benchmark. "Simple" and natural ways to go  
below this level:

- Squeezing
- Non-demolition/backaction-evasion

*Back-action evading impulse measurements with mechanical quantum sensors* Ghosh, **Carney**, Shawhan, Taylor 1910.11892 (PRA)

See Clerk PRB 2004 for review

#### Yale experiment

Search for new Interactions in a Microsphere Precision Levitation Experiment (SIMPLE) @ D. Moore group

![](_page_11_Picture_2.jpeg)

~ 1 ng dielectric spheres, optically levitated, stable for days

~ 75 MeV momentum transfer resolution (~ 100 x SQL), currently technical-noise limited

Continuous monitoring of two or three spatial axes  $\rightarrow$  directional sensitivity

Search for composite DM ("dark nucleons"), coupled to SM via some long-range force (for example, B-L boson). Novel constraints with ~1 day of data, impulse bump search:

![](_page_12_Figure_2.jpeg)

Some model realizations: Lin, Yu, Zurek 1111.0293 Krnjaic, Sigurdson 1406.1171

![](_page_12_Figure_4.jpeg)

Monteiro, Afek, Carney, Krnjaic, Wang, Moore 2007.12067 (PRL)

![](_page_13_Picture_0.jpeg)

### Gravitational direct detection of dark matter

The only coupling dark matter is guaranteed to have is through gravity.

Roughly one proton mass per cubic cm of dark matter--hopeless to try to detect it gravitationally in a local lab. Right?

Extremely hard, but maybe possible...

![](_page_14_Picture_4.jpeg)

#### The issues

$$F = \frac{G_N m_s m_\chi}{r^2}$$

Signal ~  $G_N m_s / b^2$  $\rightarrow$  want heavy DM, heavy device, small impact parameter

$$n_{DM} \approx \frac{0.3}{\mathrm{cm}^3} \left( \frac{1 \mathrm{~GeV}}{m_{\chi}} \right)$$

Observable flux ~ A/m $\chi$  $\rightarrow$  want large area

![](_page_15_Picture_5.jpeg)

# The solution: array

Signal = correlated track of macroscopic motion

Complete directional info

Exquisite background rejection

NB: here drawing pendulums w/o readout. In practice, probably don't want to use optics. Microwave circuit readout more likely, many other possibilities.

![](_page_16_Figure_5.jpeg)

### The situation

~10 million sensors, mg-scale, mm-cm spacing ~thermally limited detection (substantially sub-SQL)

 $\rightarrow$  DM of mass ~ m\_{\_{Planck}} detectable @ 1-10 events/yr

This is primarily a proof-of-principle that this could be possible with concrete setup

Just the beginning of the story--many improvements possible! Better optimized architecture, quantum track reconstruction, frequency multiplexing, ...

*Gravitational Direct Detection of Dark Matter* **Carney**, Ghosh, Krnjaic, Taylor 1903.00492 (PRD)

![](_page_17_Figure_6.jpeg)

# Windchime collaboration

![](_page_18_Picture_1.jpeg)

Rafael Lang

2020-21 goals:

Build array of ~100 sensors

70 mg masses (two types: Si and Ge  $\rightarrow$  EP violation sensitive), chip-etched, mass-producible by S. Bhave (Purdue mech eng), optical readout

Demonstrate track sensing and ultralight DM search capacity

Collaboration members: Purdue, Fermilab, Berkeley Lab, Oak Ridge Lab, Minnesota, Maryland Funded by US National Quantum Initiative

# Thanks to collaborators!

![](_page_19_Picture_1.jpeg)

![](_page_19_Picture_2.jpeg)

![](_page_19_Picture_3.jpeg)

![](_page_19_Picture_4.jpeg)

- C. Regal

S. Bhave

N. Matsumoto

![](_page_19_Picture_9.jpeg)

![](_page_19_Picture_10.jpeg)

D. Moore (EXO)

![](_page_19_Picture_12.jpeg)

P. Shawhan

(LIGO)

![](_page_19_Picture_13.jpeg)

R. Lang (XENON)

quant

![](_page_19_Picture_16.jpeg)

J. Taylor

![](_page_19_Picture_18.jpeg)

S. Ghosh

![](_page_19_Picture_20.jpeg)

P. Stamp

![](_page_19_Picture_22.jpeg)

ex

B. Unruh

![](_page_19_Picture_24.jpeg)

G. Krnjaic

![](_page_19_Picture_26.jpeg)

A. Hook

![](_page_19_Picture_28.jpeg)

![](_page_19_Picture_29.jpeg)

hep/gr

![](_page_19_Picture_31.jpeg)

Y. Zhao

![](_page_19_Picture_33.jpeg)

G. Semenoff

![](_page_19_Picture_35.jpeg)

# **Conclusions & future**

- Quantum mechanics of measurement imposes fundamental noise floor
- Meso-to-macroscopic mechanical devices offer range of exciting possibilities in DM detection
- Many experiments online now or planned. All first-gen pathfinders. Next step: scale up
- Improvements to sensitivity: squeezing/backaction evasion to get below SQL, QEC-assisted sensing, quantum coherent readout/track reconstruction, ...

• What other physics targets can we aim for??

Gravitational direct detection of dark matter: 1903.00492 **DC**, Ghosh, Krnjaic, Taylor

Ultralight dark matter detection with mechanical quantum sensors: 1908.04797 **DC**, Hook, Liu, Taylor, Zhao

First experimental result with optomechanical detection (not Windchime): 2007.12067 Monteiro, Afek, **DC**, Krnjaic, Wang, Moore

Review paper: 2008.06074 **DC**, Krnjaic, Regal, Moore + 35 others