Gomel state university of F. Scorina

Quasipotential equations solutions for three dimensional harmonic oscillator in the relativistic configuration representation

Grishechkin Yu.A. Kapshai V.N.

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Plan

- Quasipotential equations
- Transformation of integral equations to the Sturm-Liouville problem
- The methods for approximate solution of the Logunov-Tavkhelidze equation
- Numerical solution of quasipotential integral equations in the RCR
- Analysis of results

Quasipotential equations

• Quasipotential equations in the momentum representation (MR):

Quasipotential equations
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\n
$$
\psi_{(j)}(E_q, p) = -\frac{2m}{\pi} G_{(j)}(E_q, p) \int_0^\infty \frac{dk}{E_k} V(p, k) \psi_{(j)}(E_q, k) \qquad E_k = \sqrt{k^2 + m^2}
$$
\n
$$
j=1 (j=3) - is the Logunov-Tavkhelidze equation (modified)
$$
\n
$$
j=2 (j=4) - is the Kadyshevsky equation (modified)
$$
\n
$$
2E_q \geq 2m - is the energy of two-particle system
$$
\n
$$
m - is the mass of each particle
$$
\nThe Green functions:

j=1 (j=3) – is the Logunov-Tavkhelidze equation (modified) j=2 (j=4) – is the Kadyshevsky equation (modified)

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\nThe Green functions:
\n
$$
G_{(1)}(E_q, p) = \frac{1}{E_p^2 - E_q^2 - i0} \qquad G_{(2)}(E_q, p) = \frac{1}{2E_p(E_p - E_q - i0)}
$$
\n
$$
G_{(3)}(E_q, p) = \frac{1}{(E_p^2 - E_q^2 - i0)} \frac{E_p}{m} \qquad G_{(4)}(E_q, p) = \frac{1}{2(E_p - E_q - i0)} \frac{1}{m}
$$
\n
$$
E_p = \sqrt{p^2 + m^2}
$$
\n
$$
V(p, k) - potential
$$
\n
$$
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$$

 $V(p, k)$ - potential

Quasipotential equations in the relativistic configurational representation (RCR): The relativistic configurational representation (RCR):
 $\int_{0}^{\infty} dr' G_{(j)}(\chi_q, r, r') V(r') \psi_{(j)}(r')$

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\n
$$
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$$
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\nmeter related to energy as $2E_q = 2mch \chi_q$
\nn the RCR:
\n
$$
G_{(j)}(\chi_q, r, r') = G_{(j)}(\chi_q, r - r') - G_{(j)}(\chi_q, r + r')
$$

 r - is the module of radius-vector in the RCR

 $\chi_{q} \geq 0$ - is the parameter related to energy as $-2E_{q} = 2m\mathrm{ch}\,\chi_{q}$

The Green functions in the RCR:

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\n
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G_{(1)}(\chi_q, r) = \frac{-i}{msh 2\chi_q} \frac{\sin[(\pi/2 + i\chi_q)mr]}{\sin[\pi mr/2]} \frac{G_{(2)}(\chi_q, r) = \frac{(4mch \chi_q)^{-1}}{ch[\pi mr/2]} - \frac{i}{msh 2\chi_q} \frac{\sin[(\pi + i\chi_q)mr]}{\sin[\pi mr]}
$$
\n
$$
G_{(3)}(\chi_q, r) = \frac{-i}{2msh \chi_q} \frac{\cosh[(\pi/2 + i\chi_q)mr]}{\cosh[\pi mr/2]} \frac{G_{(4)}(\chi_q, r) = \frac{-i}{2msh \chi_q} \frac{\sin[(\pi + i\chi_q)mr]}{\sin[\pi mr]}
$$
\n
$$
V(r) -
$$
potential in the RCR

The relationship between the values in the MR and in the RCR:

Ship between the values in the MR and in the RCR:

and in the MR are interrelated by means of the Shapiro

which in the spherically symmetric case is the Fourier
 $\frac{1}{(j)}(r) = \frac{2m}{\pi} \int_{0}^{s} dz \sin(zmr) \psi_{(j)}(m \text{ch } \chi_{q}, m \text{$ (hip between the values in the MR and in the

md in the MR are interrelated by means of

which in the spherically symmetric case
 $\lim_{(j)}(r) = \frac{2m}{\pi} \int_{0}^{\infty} d\chi \sin(\chi mr) \psi_{(j)}(m \text{ch } \chi_q, m \text{sh } \chi)$

related to momentum as **butter in the values in the MR and in the RCR:**
 r and in the MR are interrelated by means of the Shapiro
 r, which in the spherically symmetric case is the Fourier
 $\psi_{(j)}(r) = \frac{2m}{\pi} \int_{0}^{\infty} dz \sin(\chi m r) \psi_{(j)}(m \text$ The values in the RCR and in the MR are interrelated by means of the Shapiro integral transformation, which in the spherically symmetric case is the Fourier transformation lationship between the values in the

RCR and in the MR are interrelation, which in the spherically sy
 $\psi_{(j)}(r) = \frac{2m}{\pi} \int_0^{\infty} d\chi \sin(\chi mr) \psi_{(j)}(mc) \chi_{(j)}(mc)$

idity related to momentum as $p =$

:
 $\psi_{(j)}(\chi_q, r, r') = \frac{-2$ ationship between the values in the MR and in the RCR:

RCR and in the MR are interrelated by means of the S

tion, which in the spherically symmetric case is the F
 $\Psi_{(j)}(r) = \frac{2m}{\pi} \int_0^{\infty} dz \sin(\chi m r) \Psi_{(j)}(m \text{ch } \chi_q, m$ **Example 11**
 Example 12
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 EXER and in the spherically symmetric case is the Fourier ip between the values in the MR and in the RCR:

1 in the MR are interrelated by means of the Shapiro

hich in the spherically symmetric case is the Fourier
 $(x) = \frac{2m}{\pi} \int_0^{\pi} dz \sin(\chi m r) \psi_{(j)}(m \text{ch } \chi_q, m \text{sh } \chi)$

lated

Wave functions:

$$
\psi_{(j)}(r) = \frac{2m}{\pi} \int_{0}^{\infty} d\chi \sin(\chi mr)\psi_{(j)}(m \operatorname{ch} \chi_q, m \operatorname{sh} \chi)
$$

 $\chi \geq 0$ - is the rapidity related to momentum as $p = m \sin \chi$

Green functions:

$$
G_{(j)}(\chi_q, r, r') = \frac{-2m}{\pi} \int_0^\infty d\chi \sin(\chi mr) G_{(j)}(m \operatorname{ch} \chi_q, m \operatorname{sh} \chi) \sin(\chi mr')
$$

Potential:

$$
V(p,k) = \int_{0}^{\infty} dr \sin(\chi mr) V(r) \sin(\chi'mr)
$$

The non-relativistic limit of the above formulas and equations leads to the corresponding formulas and equations of the non-relativistic quantum theory.

Transformation of integral equations to the Sturm-Liouville problem ormation of inte

he Sturm-Liouvi

scillator type potential
 $V(r) = \omega^2 r^2$

i the momentum repress
 $\frac{\pi \omega^2}{2m^3} \left(\sqrt{m^2 + p^2} \frac{d}{dp} \right)^2 \sqrt{m^2 + k}$

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il equation (DE):
 $\frac{2}{\sqrt{2}} \psi_{(j)}(\chi_q, \chi$ or primation of integral equations to

ine Sturm-Liouville problem

scillator type potential in the RCR:
 $V(r) = \omega^2 r^2$

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 $\frac{\pi \omega^2}{km^3} \left(\sqrt{m^2 + p^2} \frac{d}{dp} \right)^2 \sqrt{m^2 + k^2} \delta(p - k) = - \frac{\pi \omega^2}{2m^$ Formation of integral equations

the Sturm-Liouville problem

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The harmonic oscillator type potential in the RCR:

$$
V(r) = \omega^2 r^2
$$

The potential in the momentum representation:
\n
$$
V(p,k) = -\frac{\pi \omega^2}{2m^3} \left(\sqrt{m^2 + p^2} \frac{d}{dp} \right)^2 \sqrt{m^2 + k^2} \delta(p-k) = -\frac{\pi \omega^2}{2m^3} \frac{d^2}{d\chi^2} \delta(\chi - \chi')
$$

The substitution of the potential in the integral equation in the MR leads to the differential equation (DE):

$$
\frac{d^2}{dx^2}\psi_{(j)}(\chi_q,\chi) = \frac{m^2}{\omega^2}G_{(j)}^{-1}(m\operatorname{ch}\chi_q, m\operatorname{ch}\chi)\psi_{(j)}(\chi_q,\chi)
$$

The boundary conditions:

$$
\boxed{\psi(\chi_q,0)=0} \quad \left|\psi(\chi_q,\chi)\right|_{\chi\to\infty} \equiv 0
$$

The Sturm-Liouville problem

The methods for approximate solving of the Logunov-Tavkhelidze equation

The Sturm-Liouville problem for the Logunov-Tavkhelidze equation:
\n
$$
\frac{d^2}{d\chi^2}\psi(\chi_q, \chi) = \frac{m^4}{\omega^2} (\text{ch}^2 \chi - \text{ch}^2 \chi_q) \psi(\chi_q, \chi) \qquad \chi \ge 0
$$
\n
$$
\psi(\chi_q, 0) = 0 \qquad \psi(\chi_q, \chi)|_{\chi \to \infty} \approx 0
$$

The DE solution can be expressed through the modified Mathieu functions. However, the study of such solutions is a cumbersome problem. Let us consider the approximate analytical solution of the Sturm-Liouville problem.

Reduction of the equation to the modified Bessel equation

Let us replace the variable $z = \omega^{-1} m^2 \exp(\chi)/2$
 $\left[\left(\frac{d}{2} \right)^2 - z^2 + \frac{m^4}{2} \exp(2\chi)/2 \right]$

$$
\left[\left(z \frac{d}{dz} \right)^2 - z^2 + \frac{m^4}{2\omega^2} \text{ch} \, 2 \chi_q \right] \psi(\chi_q, z) = \frac{m^8 \omega^4}{\sqrt{6z^4}} \psi(\chi_q, z) \qquad z \ge \omega^{-1} m^2 / 2
$$

$$
\psi(\chi_q, \omega^{-1} m^2 / 2) = 0 \qquad \psi(\chi_q, z) \Big|_{z \to \infty} \equiv 0
$$

In the presented DE, we omit the right-hand side. The modified Bessel functions satisfy the equation obtained in this way.

The second of the boundary conditions holds for the Macdonald function

$$
K_{i\nu}(z) \text{, } \text{rge } \nu = m^2/\omega \sqrt{(1/2)\text{ch } 2\chi_q}
$$

Taking into account the first boundary condition leads to a transcendental equation for the quantity *ν*

$$
K_{i\nu}\left(\omega^{-1}m^2/2\right) = 0
$$

which is the condition for quantization of energy.

The approximate solution of the Logunov-Tavkhelidze equation in the MR:

$$
\psi(\chi_q^{(n)}, \chi) = C_n K_{i\nu_n} \left(\omega^{-1} m^2 \exp(\chi)/2 \right)
$$

n=1,2,3,… - is the state number of the relativistic harmonic oscillator

- C_n is the normalization constant
- v_n is the root of the transcendental equation, connected with the energy of the relativistic harmonic oscillator by the formula

$$
2E_q^{(n)} = \sqrt{2m^2 + (2v_n \omega/m)^2}
$$

The approximate wave function in the RCR:

$$
\psi_{n}(r) = \frac{C_{n}}{4i} \left\{ \frac{1}{2} \sum_{s=1}^{n} s \left(\frac{4\omega}{m^{2}} \right)^{i s m r} \Gamma \left(\frac{i s m r - i v_{n}}{2} \right) \Gamma \left(\frac{i s m r + i v_{n}}{2} \right) + \right.
$$

+
$$
\left(\frac{4\omega}{m^{2}} \right)^{i v_{n}} \Gamma(i v_{n}) \sum_{s=\pm 1}^{n} \frac{1}{i s v_{n} - i m r} {}_{1}F_{2} \left(\frac{i s m r - i v_{n}}{2}; 1 - i v_{n}, 1 + \frac{i s m r - i v_{n}}{2}; \frac{m^{4}}{16\omega^{2}} \right) - \left. - \left(\frac{4\omega}{m^{2}} \right)^{-i v_{n}} \Gamma(-i v_{n}) \sum_{s=\pm 1}^{n} \frac{1}{i s v_{n} + i m r} {}_{1}F_{2} \left(\frac{i s m r + i v_{n}}{2}; 1 + i v_{n}, 1 + \frac{i s m r + i v_{n}}{2}; \frac{m^{4}}{16\omega^{2}} \right) \right\}
$$

 $\Gamma({\rm z})$ - is the gamma function, ${}_1\rm{F}_2({\rm a};$ ${\rm b},$ ${\rm c};$ ${\rm z})$ - is the generalized hypergeometric series

To find the constants C_n we use the normalization conditions for the wave functions :

in the MR

\n
$$
\lim_{n \to \infty} \frac{\sin \theta}{2\pi} \frac{d\chi \psi^2(\chi_q^{(n)}, \chi) = 1}{\int_0^\infty \frac{d\chi \psi_n^2(r)}{r}} = 1
$$

The advantages of the solving method:

• the possibility of finding the analytical solution: wave functions in the MR and in the RCR

Disadvantages of the solving method:

• the absence of a non-relativistic limit of the results obtained

Solution by the Galerkin method

Performing a change of variable $p = msh \chi$, we represent the Sturm-Liouville problem in the form: on by the Galerkin method
 $p = msh \chi$, we represent the Sturm-Liouville problem in the form:
 $\sqrt{2} \frac{d}{d\chi} \int_{W(\chi_1, p)}^2 = (p^2 - m^2sh^2 \chi_1)w(\chi_2, p)}$

Solution by the Galerkin method
\n
$$
\begin{array}{c}\n\text{Solution by the Galerkin method} \\
\text{Reforming a change of variable } p = msh \chi, \text{ we represent the Sturm-Liouville problem in the form:} \\
\frac{\omega^2}{m^2} \left(\sqrt{m^2 + p^2} \frac{d}{dp} \right)^2 \psi(\chi_q, p) = \left(p^2 - m^2 \sin^2 \chi_q \right) \psi(\chi_q, p) \\
\psi(\chi_q, 0) = 0 \qquad \psi(\chi_q, p) \Big|_{p \to \infty} \equiv 0\n\end{array}
$$
\nWe represent the unknown wave function as the sum:
\n
$$
\begin{array}{c}\n\psi(\chi_q, p) = \sum_{s=0}^{N} C_s \varphi_s(\chi_q, p) \\
C_s - \text{ is the unknown coefficients}\n\varphi_s(\chi_q, p) = \int p^2 - \lambda_s \varphi_s(\chi_q, p) \\
\frac{\omega^2}{dp^2} \frac{d^2}{\varphi_s(\chi_q, p)} = \left(p^2 - \lambda_s \right) \varphi_s(\chi_q, p)\n\end{array}
$$
\nThe number of terms N in the sum depends on the required accuracy of the solution obtained.
\nThe functions $\varphi_s(p)$ have the form analogous to the wave functions of a non-relativistic harmonic oscillator:
\n
$$
\varphi_s(p) = \frac{1}{\sqrt{p^2}} \exp\left(-\frac{1}{2}p^2\right) H_{2+1}(p/\sqrt{\omega})
$$

We represent the unknown wave function as the sum:

$$
\psi(\chi_q, p) = \sum_{s=0}^N C_s \varphi_s(\chi_q, p)
$$

C^s – is the unknown coefficients

$$
\omega^2 \frac{d^2}{dp^2} \varphi_s(\chi_q, p) = (p^2 - \lambda_s) \varphi_s(\chi_q, p)
$$

The number of terms N in the sum depends on the required accuracy of the solution obtained.

The functions $\varphi_s(p)$ have the form analogous to the wave functions of a non-relativistic
harmonic oscillator:
 $\varphi_s(p) = \frac{1/\sqrt{\omega}}{\left[\frac{2^{2s}(2s+1)!\sqrt{\pi}}{\sigma}\right]^{1/2}} \exp\left(-\frac{1}{2\omega}p^2\right) H_{2s+1}\left(p/\sqrt{\omega}\right)$ harmonic oscillator: λ_s) $\varphi_s(\chi_q, p)$

he required accuracy of the solution obtained.

o the wave functions of a non-relativistic
 $\exp\left(-\frac{1}{2\omega}p^2\right)H_{2s+1}(p/\sqrt{\omega})$
 $\lambda_s = \omega(3+4s)$
 $11/20$

$$
\omega^2 \frac{a}{dp^2} \varphi_s(\chi_q, p) = (p^2 - \lambda_s) \varphi_s(\chi_q, p)
$$
\nThe number of terms N in the sum depends on the required accuracy of the so
\nThe functions $\varphi_s(p)$ have the form analogous to the wave functions of a non-
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$$
\varphi_s(p) = \frac{1/\sqrt{\omega}}{\left[2^{2s}(2s+1)!\sqrt{\pi}\right]^{1/2}} \exp\left(-\frac{1}{2\omega}p^2\right) H_{2s+1}\left(p/\sqrt{\omega}\right)
$$
\n
$$
H_n(x) - \text{ is the Hermite polynomials}
$$
\nThe corresponding eigenvalues are defined as $\lambda_s = \omega(3+4s)$

 $H_n(x)$ - is the Hermite polynomials

After substituting the sum into the equation, we obtain the following equality:
\n
$$
\left(\frac{\omega}{m}\right)^2 \sum_{s=0}^N C_s \left(p \frac{d}{dp}\right)^2 \varphi_s(p) - \sum_{s=0}^N C_s \lambda_s \varphi_s(p) = -q^2 \sum_{s=0}^N C_s \varphi_s(p)
$$

Multiplying the resulting equality by the function $\varphi_n(p)$, and integrating the resulting equality from zero to infinity, we obtain the linear system of $N + 1$ equations

$$
MC = E_q^2 C
$$

C – is the vector composed of unknown coefficients

M - is the five-diagonal matrix whose elements have the form
\n
$$
M_{ns} = \left[\lambda_n + 1\right]\delta_{n,s} + \left(\frac{\omega}{m}\right)^2 \left[a_n^2 \delta_{ns} + b_s \delta_{n+1,s} - b_n \delta_{ns+1} - c_s \delta_{n+2,s} - c_n \delta_{ns+2}\right]
$$
\n
$$
\delta_{ns}
$$
- are the elements of the identity matrix

$$
a_n^2 = \frac{1}{4} \left(8n^2 + 12n + 5 \right) \qquad b_n = \frac{1}{2} \sqrt{2n(2n+1)} \qquad c_n = \frac{1}{4} \sqrt{(2n-2)(2n-1)2n(2n+1)}
$$

We used recurrence relations for the Hermite polynomials when calculating the matrix elements

$$
2xH_n(x) = H_{n+1}(x) + 2nH_{n-1}(x) \qquad \frac{d}{dx}H_n(x) = 2nH_{n-1}(x)
$$

The advantages of the solving method:

• the possibility to calculate quickly a large number of energy values simultaneously

Disadvantages of the solving method:

• the need for cumbersome preliminary analytical calculations

Numerical solving quasipotential integral equations in the RCR

The solution was found by the method that we used to study resonant states earlier on the basis covariant two-particle integral equations in the RCR

Using quadrature formulas we replace integrals in the equations by the sums. As the results we obtain homogeneously systems of linear algebraic equations

$$
M\psi = 0 \qquad \qquad M_{nm} = \delta_{nm} - W_m G_l^{(j)}(\chi_q, r_n, r_m) V(r_m)
$$

*W*_n, r_n – are the coefficients and nodes of the quadrature formula

 $\mathscr U$ − is the vector of wave functions in the nodes

The condition for existence of nontrivial system solution

 $f(\chi_q)$ = $\det M=0$ $\,$ - the energy quantization conditions

It is advisable to represent roots of equation graphically on the complex plane $\,{\mathcal X}_q$.

In the case of the potential under consideration the roots are located on the real axis.

The advantages of the solving method:

• the possibility to apply this method to solve various equations with a wide class of potentials, in the case of bound states and resonant states

Disadvantages of the solving method:

low computing speed

Energy eigenvalues of the relativistic harmonic oscillator

- energy levels are not equidistant;
- the accuracy of the solutions by the method of reduction to the modified Bessel equation is improved with increasing of coupling constant ω;
- the accuracy of the solutions by the Galerkin's method is worsens with increasing of coupling constant ω

The dependence of energy on the value of coupling constant at *m=1*

The wave functions in the momentum representation at *m=1 ω=1*

- the dependence of energy on the value of coupling constant ω in this interval is almost linear;
- the graphics of the approach wave function are indistinguishable visually from numerical ones for the indicated quantities *m* и *ω*;
- the number of wave functions zeros in the MR is equal to state number of relativistic harmonic oscillator

The wave functions in the RCR at $m=1$, $\omega=5$

- the graphics of the approach wave function are indistinguishable visually from numerical ones for the indicated quantities *m* и *ω*;
- the wave functions in the RCR have additional zeros in comparison with the wave functions in the MR and the wave functions of non-relativistic harmonic oscillator

Conclusions and results

- the solutions of the quasipotential equations for harmonic oscillator are found in the spherically symmetric case;
- the Logunov-Tavkhelidze equation in the momentum representation was transformed to the Sturm-Liouville problem. The approximate analytical and numerical solutions of this problem were found;
- the obtained wave functions in the RCR have additional zeros in comparison with corresponding wave functions in the MR and the wave functions of non-relativistic harmonic oscillator. The number of zeros for the fixed quantum state depends on value of coupling constant of relativistic harmonic oscillator.

Thank you for your attention!