

# Directional-iDBSCAN

*first look at LIME data*

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# Last presentation

- A preliminary analysis of the LIME data was done with Ambe runs.
  - [3737-3791]: Ambe runs (focus on 3790).
  - [3792-3794]: Cosmics runs soon after source off (focus on 3793).
  
- It was agreed that the ideal would be to make an analysis with  $^{55}\text{Fe}$  runs.
  - Run 4433: Cosmic run.
  - Run 4455:  $^{55}\text{Fe}$  run.
  - Both 200 ms of exposure time.

# iDBSCAN optimizations

- The DBSCAN seeding improvement discussed in the last meeting was tested and implemented.
- Isolation seeding condition was moved out of the last clusterization loop.
  - The old version was too slow with these new data.

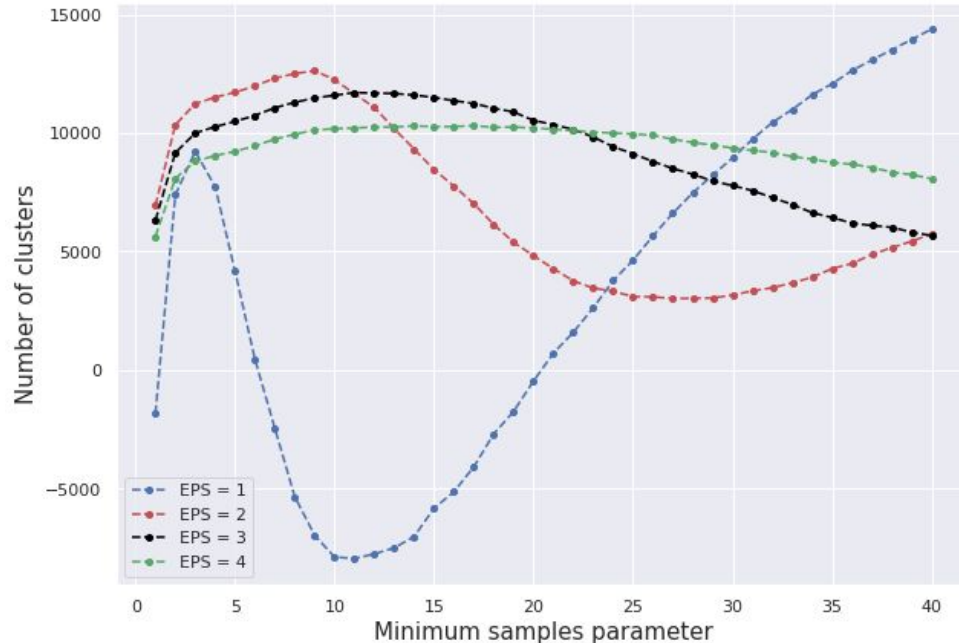
# Parameters validation

- An attempt to validate the iDBSCAN parameters was done, inspired by the iDBSCAN article.
  - Scan with the parameters aiming the maximum of:  $ER_{\text{dataset}} - NRAD_{\text{dataset}}$
  - It was done initially varying the eps and min\_samples.
- Since the currently metric distance used is the cityblock (manhattan distance), only integer values of eps will change the results.

$$\sum_i |u_i - v_i|.$$

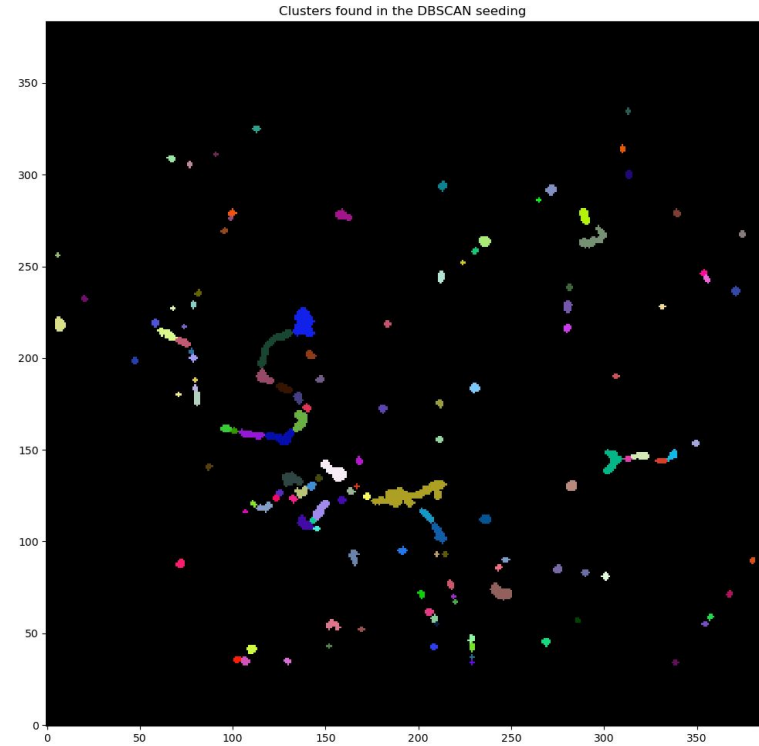
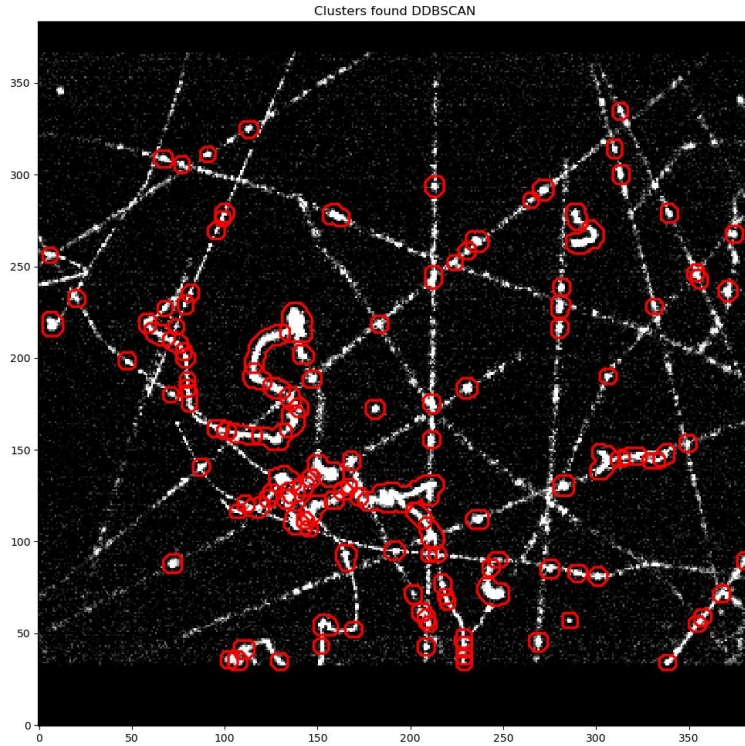
# Parameters validation

Difference between total number of clusters found in ER and NRAD



- Maximum at (eps, min\_pts) = (2, 9).
- The second peak of EPS equal to 1 was discarded due to inconsistencies. (cosmic tracks are splitted into circular clusters)

# Run 4433 - Event 50

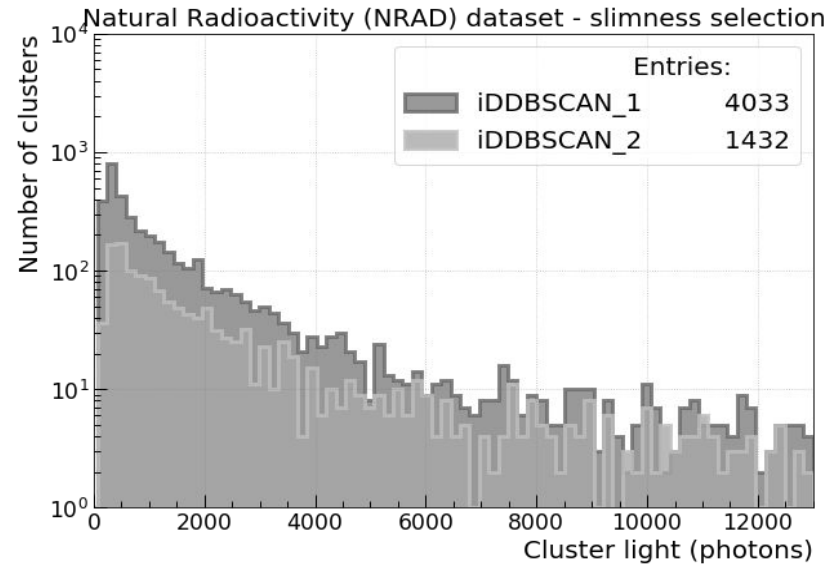
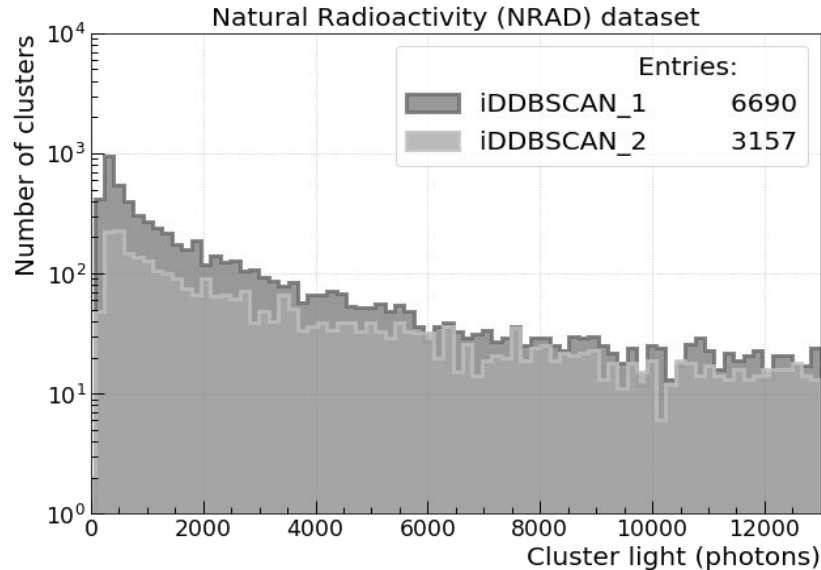


Clusterization with  $(\text{eps}, \text{min\_pts}) = (1, 40)$

# Parameters validation

- A quantitative analysis was done to compare the results of the iDDBSCAN by using the (eps, min\_pts) currently used and the one found in the previous analysis.
  - iDDBSCAN\_1: (eps, min\_pts) = (1, 5).
  - iDDBSCAN\_2: (eps, min\_pts) = (2, 9).
  - The other parameters remained the same.

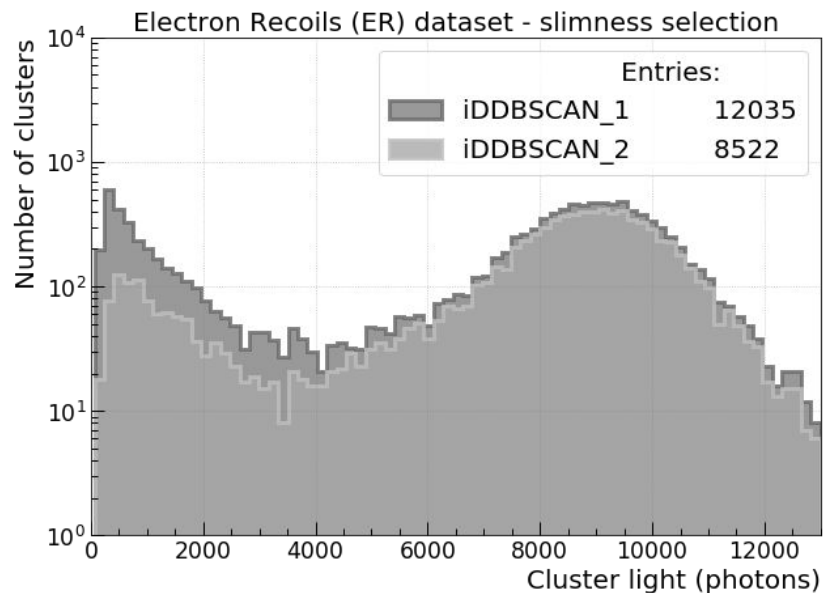
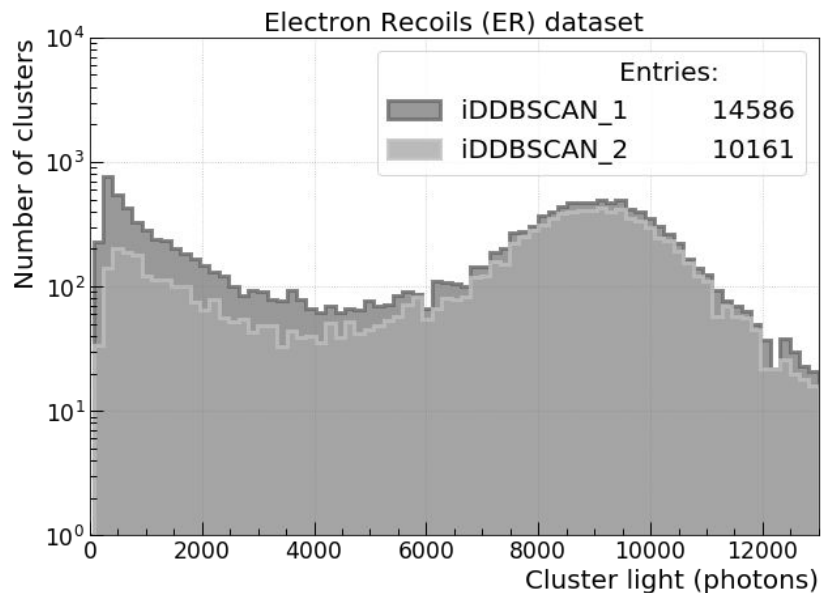
# Low energy analysis



Run 4433 - Clusters found by the iDDBSCAN with different parameters and slimness selection disabled (left) and enabled (right).

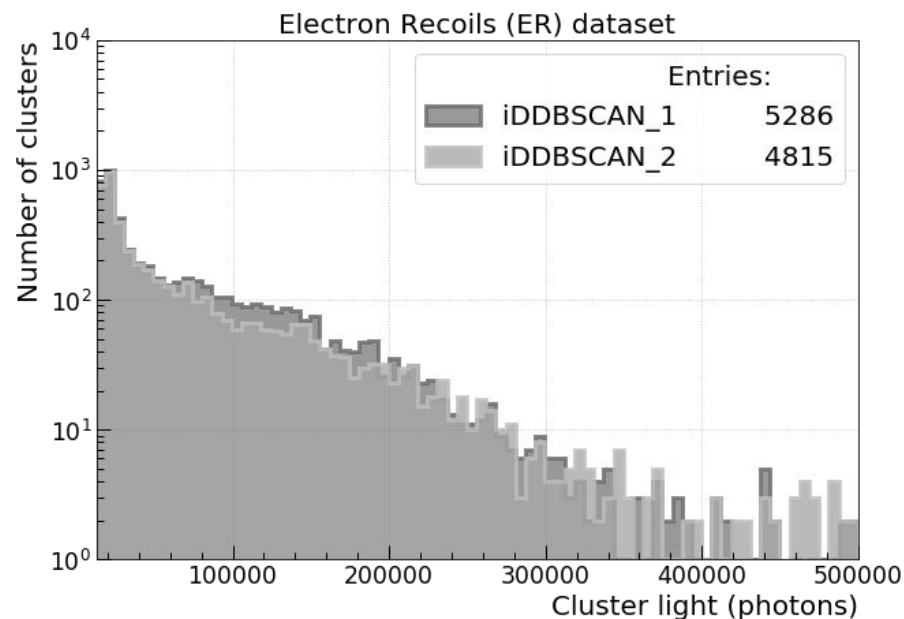
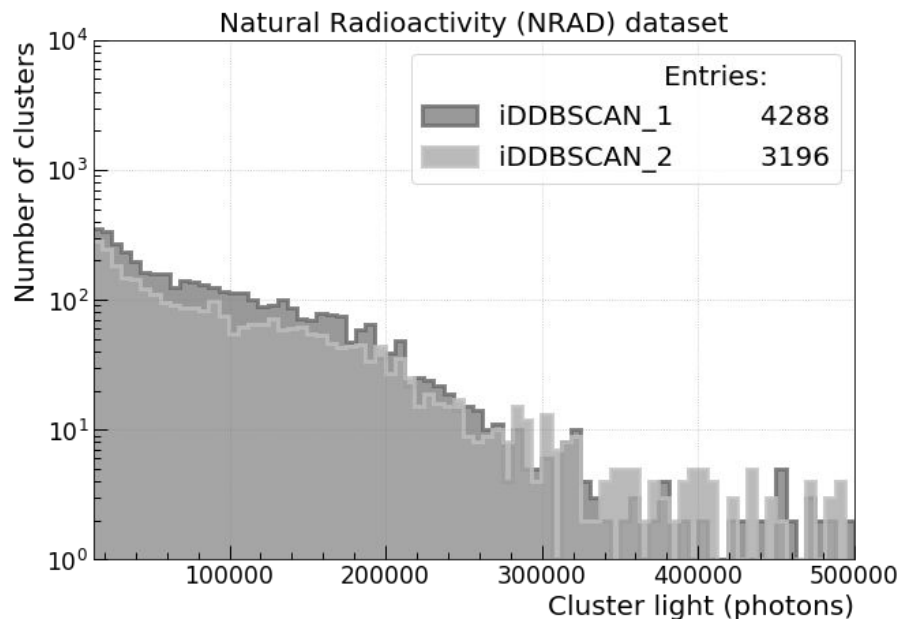


# Low energy analysis



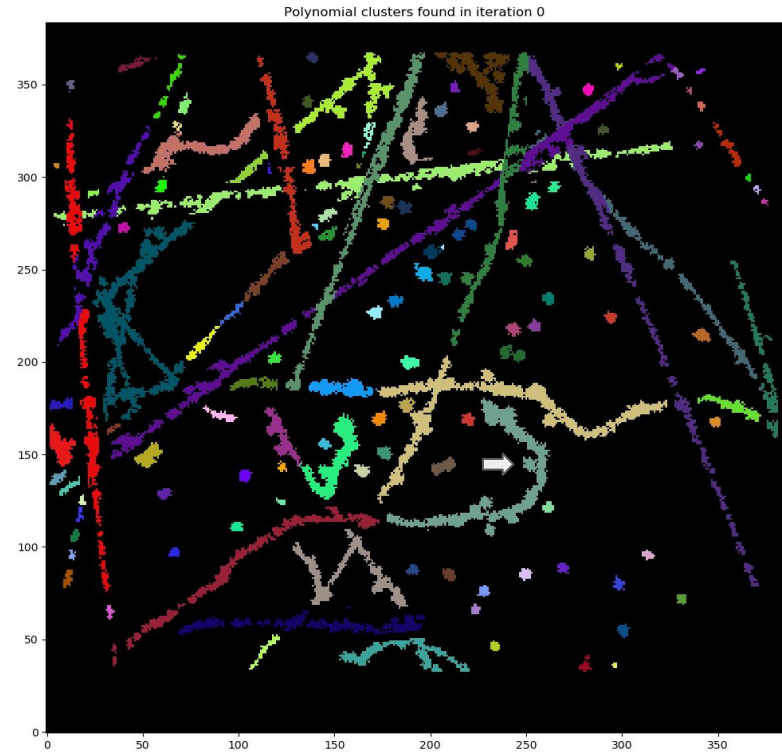
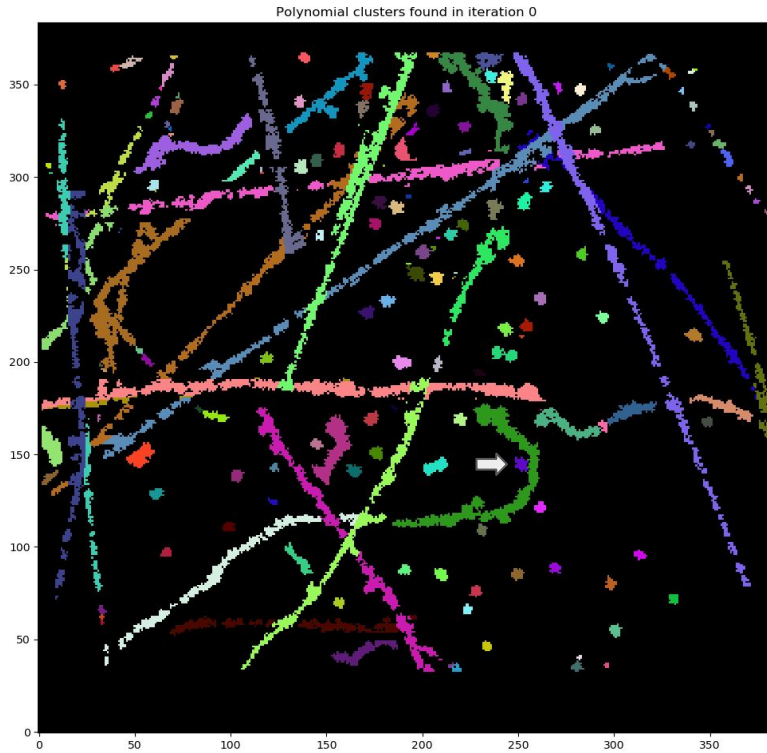
Run 4455 - Clusters found by the iDDBSCAN with different parameters and slimness selection disabled (left) and enabled (right).

# High energy analysis



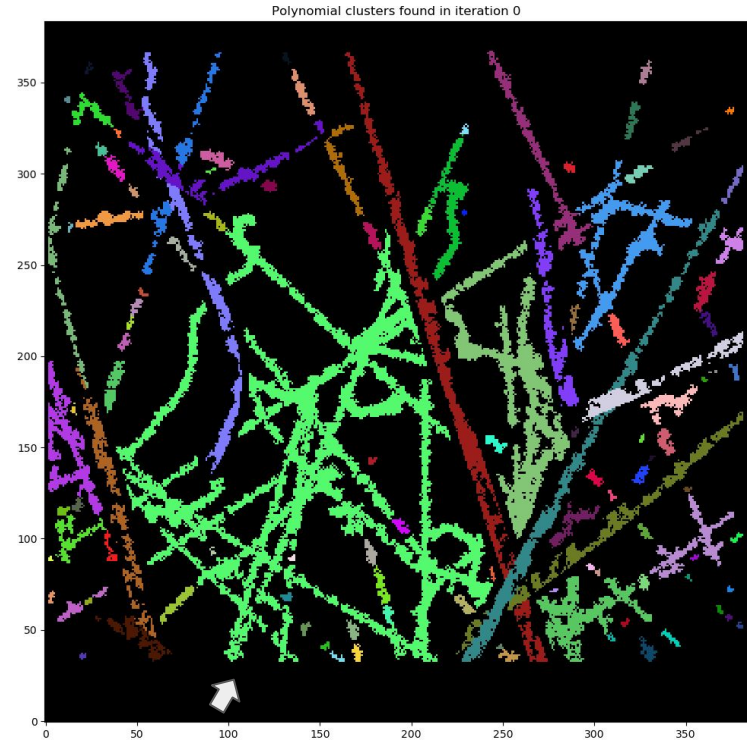
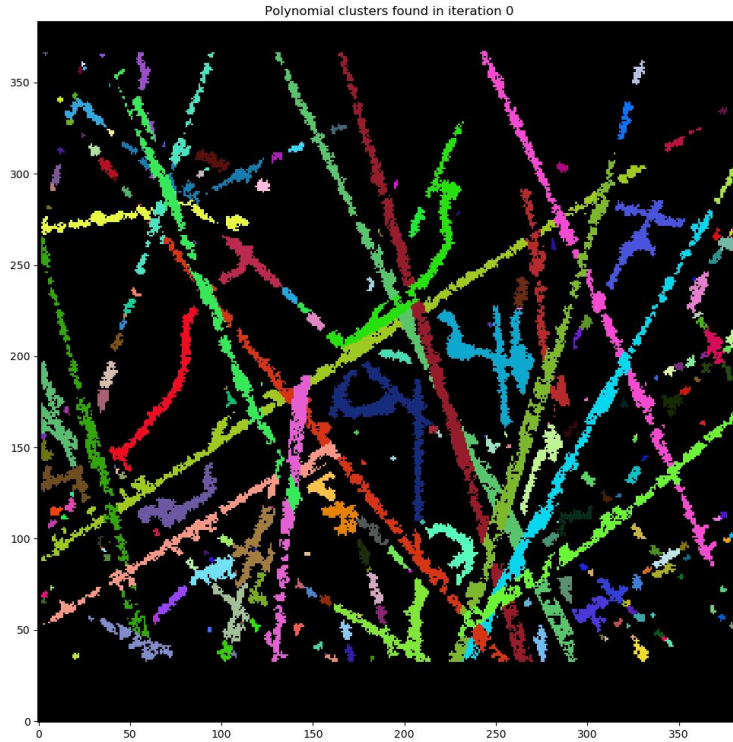
Clusters found by the iDDBSCAN with different parameters in Run 4433 (left) and 4455 (right).

# Run 4455 - Event 13



iDDBSCAN output with  $(\text{eps}, \text{min\_pts}) = (1, 5)$  - left;  $(2, 9)$  - right.

# Run 4433 - Event 13



iDDBSCAN output with  $(\text{eps}, \text{min\_pts}) = (1, 5)$  - left;  $(2, 9)$  - right.

# Conclusions

- Although the histograms show a great improvement in the low energy region, this does not mean that the cosmic tracks are being reconstructed correctly.
- Not sure if this is the best approach considering this high occupancy data.
- Any other ideas to validate these parameters?