

Testing the cosmological principle

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January 18 2022, Copernicus Webinars

Cosmological principle

The Universe is **homogeneous** and **isotropic**

Translation and Rotation invariance

Cosmological principle

The Universe is **homogeneous** and **isotropic**

$$ds^2 = -c^2 dt^2 + a^2(t)(dx^2 + dy^2 + dz^2)$$

FLRW

Cosmological principle

The Universe is **homogeneous** and **isotropic**

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FLRW

Homogeneous but anisotropic  Axis

Cosmological principle

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$$ds^2 = -c^2 dt^2 + a^2(t)(dx^2 + dy^2 + dz^2)$$

FLRW

Homogeneous but anisotropic  Axis

$$ds^2 = -dt^2 + a_x(t)^2 dx^2 + a_y(t)^2 dy^2 + a_z(t)^2 dz^2$$

Bianchi

Cosmological principle

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Bianchi

Inhomogeneous & isotropic  Centre

Cosmological principle

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Bianchi

Inhomogeneous & isotropic  Centre

$$ds^2 = -dt^2 + X^2(r, t) dr^2 + A^2(r, t) (d\theta^2 + \sin^2 \theta d\varphi^2)$$

LTB

Cosmological principle

The Universe is **homogeneous** and **isotropic**

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$$ds^2 = -dt^2 + X^2(r, t) dr^2 + A^2(r, t) (d\theta^2 + \sin^2 \theta d\phi^2)$$

LTB

Inhomogeneous & anisotropic

$$ds^2 = dt^2 - A^2 dx^2 - B^2 (dy^2 + dz^2)$$

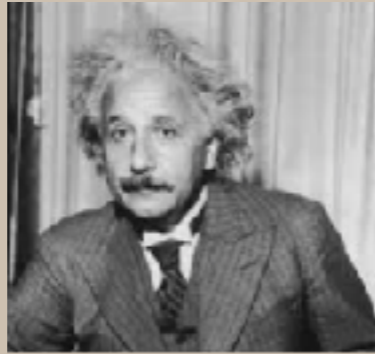
$$\begin{aligned} ds^2 = & dt^2 - (A_{\parallel}^2 \sin^2 \theta + A_{\perp}^2 \cos^2 \theta) dr^2 \\ & - (A_{\parallel}^2 \cos^2 \theta + A_{\perp}^2 \sin^2 \theta) d\theta^2 \\ & - (A_{\parallel}^2 - A_{\perp}^2) \sin \theta \cos \theta dr d\theta + -A_{\parallel}^2 \sin^2 \theta d\phi^2. \end{aligned}$$

The Cosmological principle

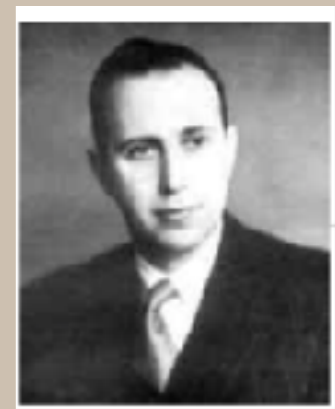
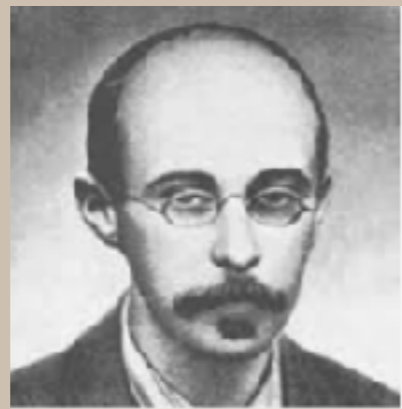
2015

$$R_{\mu\nu} - \frac{1}{2} Rg_{\mu\nu} = 8\pi GT_{\mu\nu}$$

2017



1922-1935



1932



1935

The Cosmological principle
Milne



A new basis for cosmology

BY P. A. M. DIRAC, F.R.S.

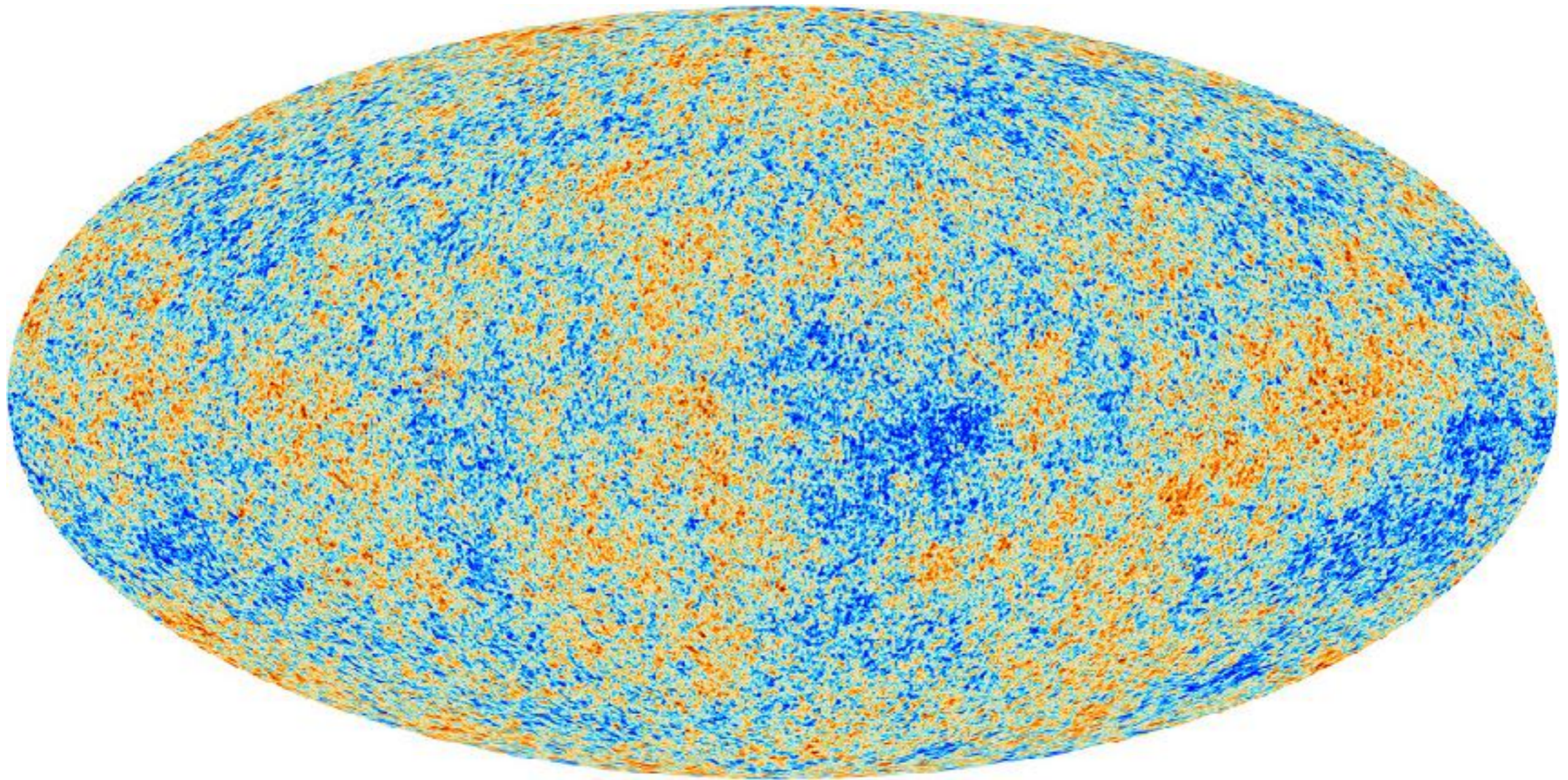
St John's College, Cambridge

(Received 29 December 1937)

We now feel the need for some new assumptions on which to build up a theory of cosmology. This need is partially satisfied by the assumptions, which Milne calls the Cosmological Principle, that, apart from local irregularities, the universe is everywhere uniform and has spherical symmetry (in three dimensions) about every point, for an observer moving with the natural velocity at that point. these assumptions are fairly plausible and have a great simplifying effect on the subject, and until there is more definite evidence of their inadequacy it does not seem worth while to try more complicated schemes.

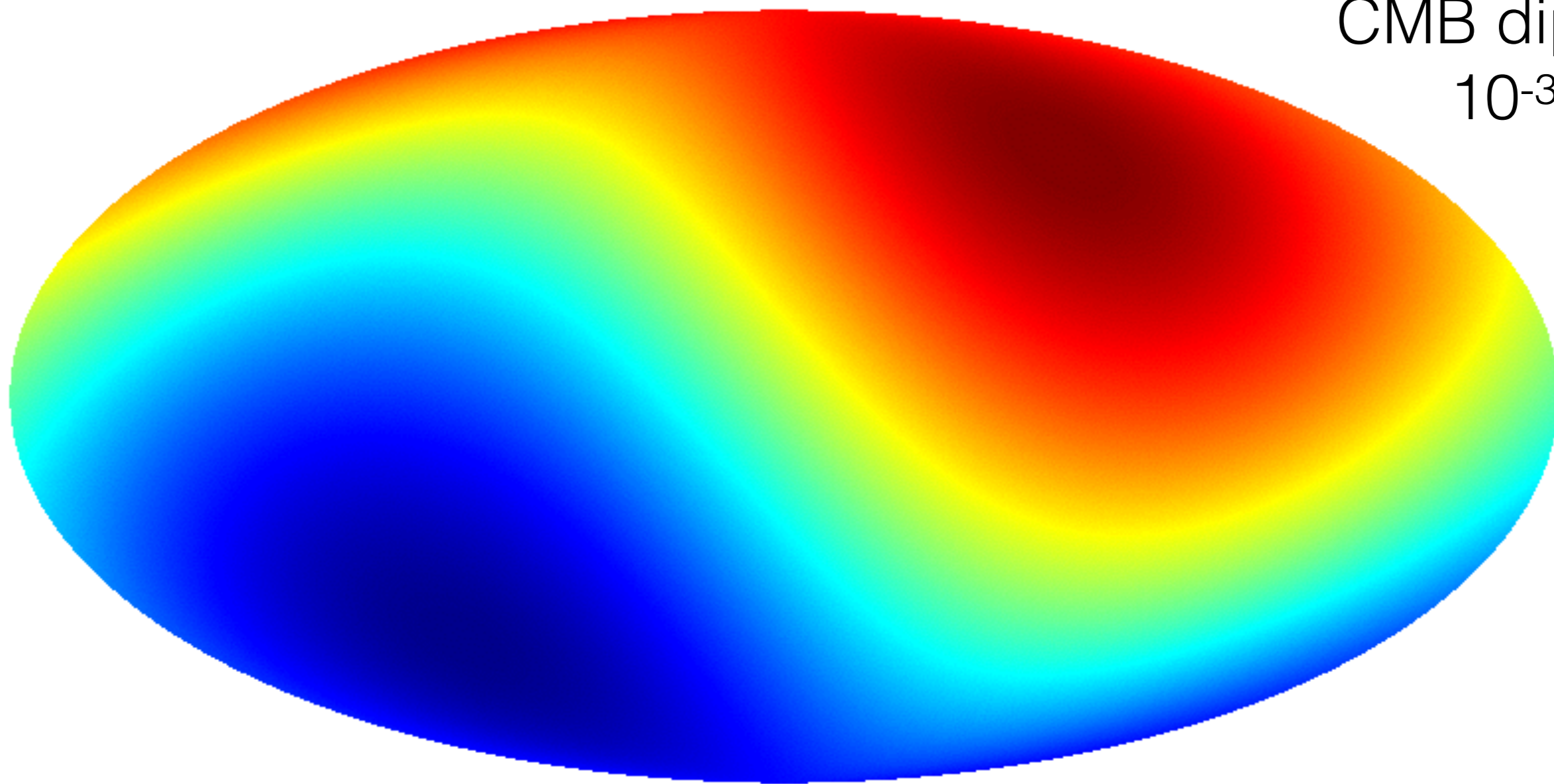
Observational evidence for the cosmological principle

Cosmic microwave background



Cosmic microwave background Dipole — Anisotropy

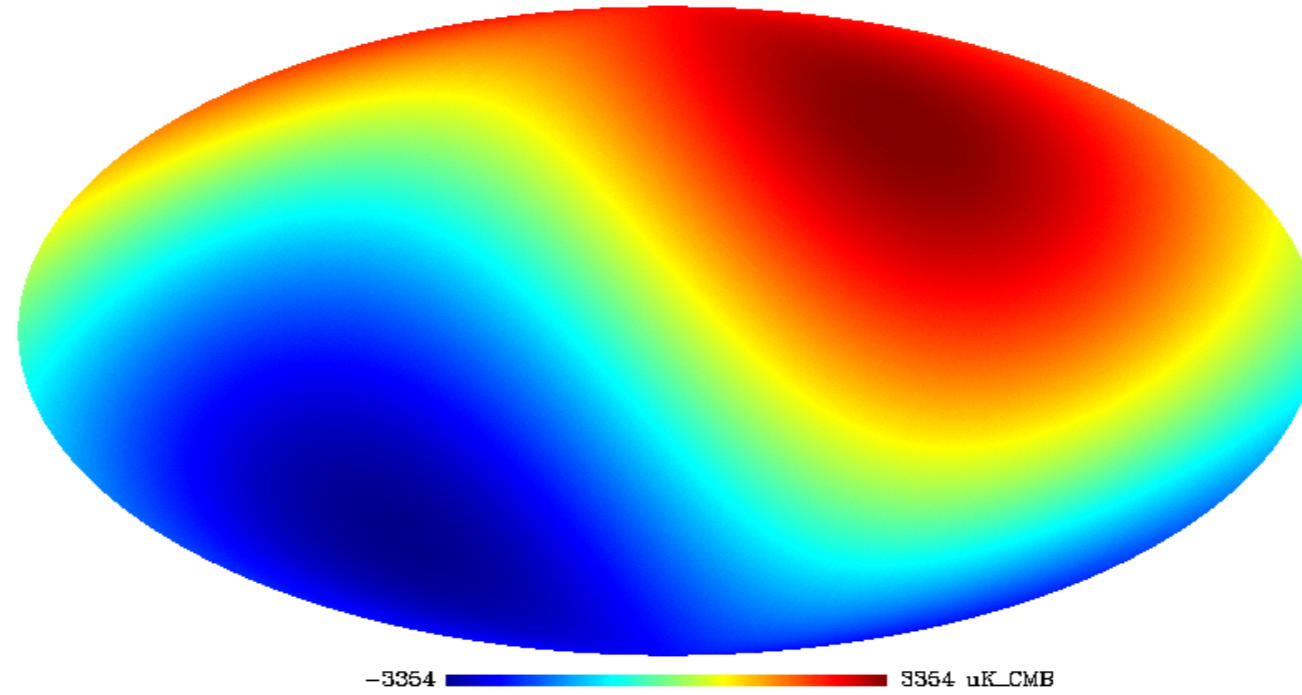
CMB dipole
 10^{-3}



-3354  3354 μK_{CMB}

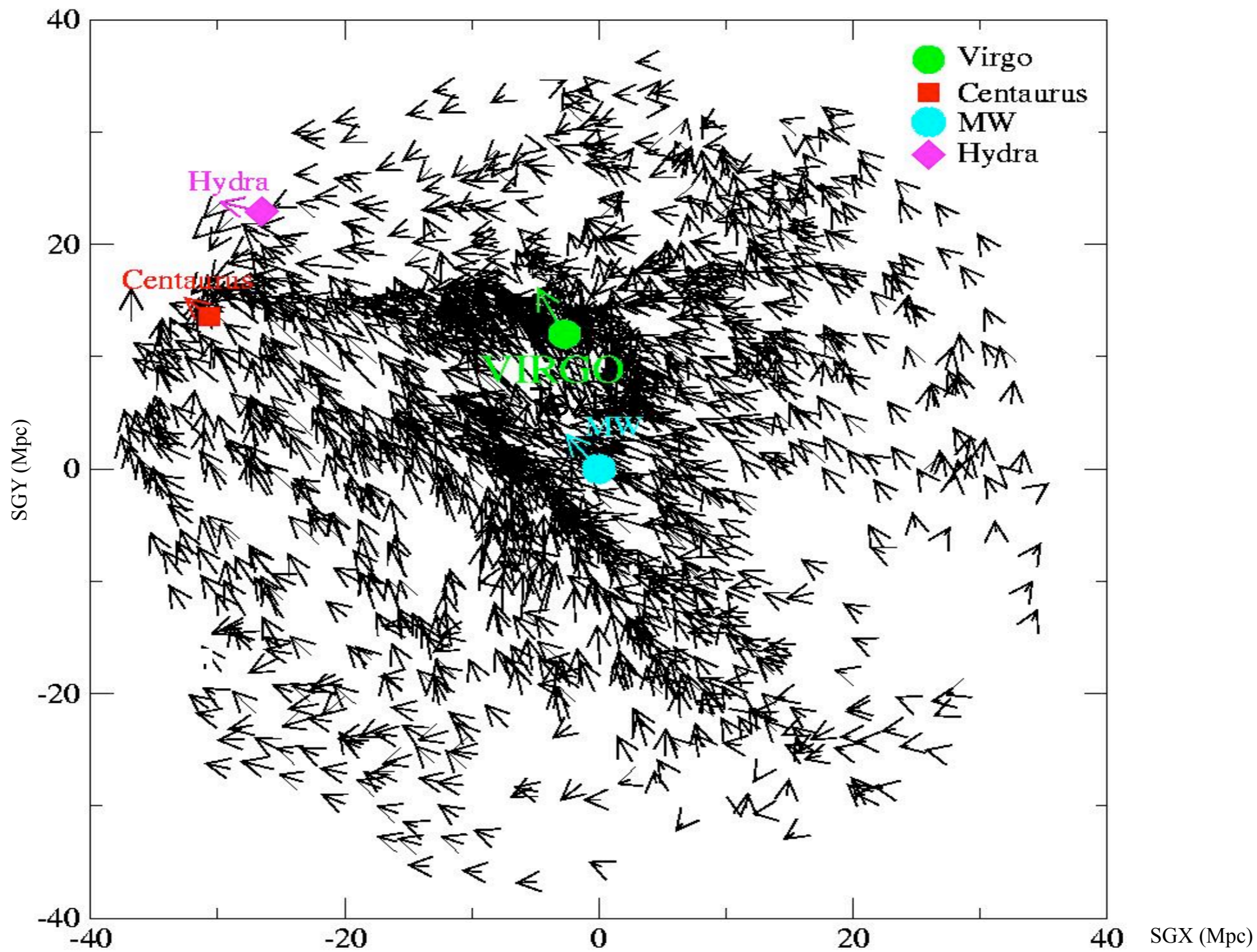
The origin of the CMB dipole ?

CMB dipole

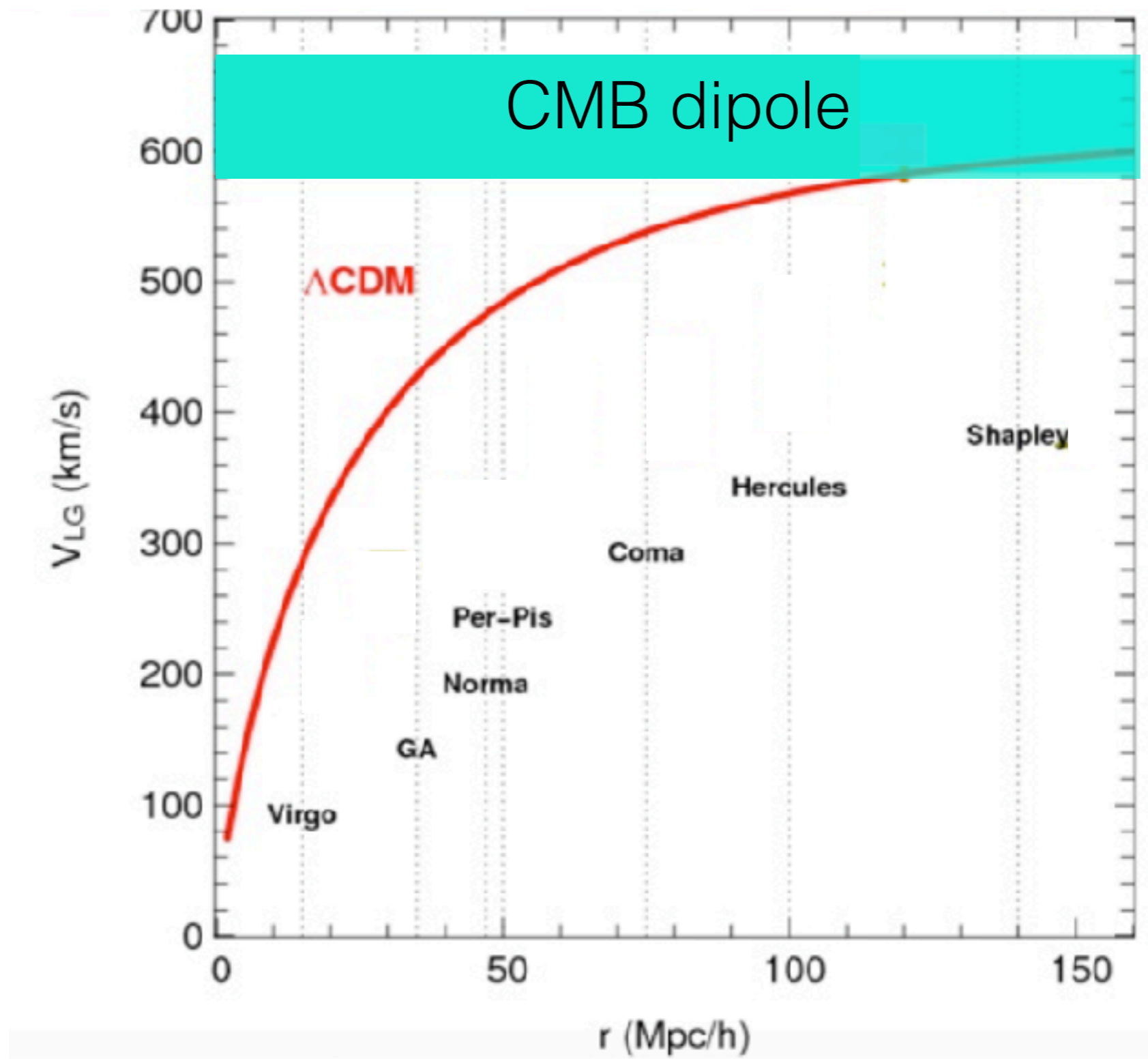


Dipole is purely Kinematic
Universe is anisotropic

Local Flow

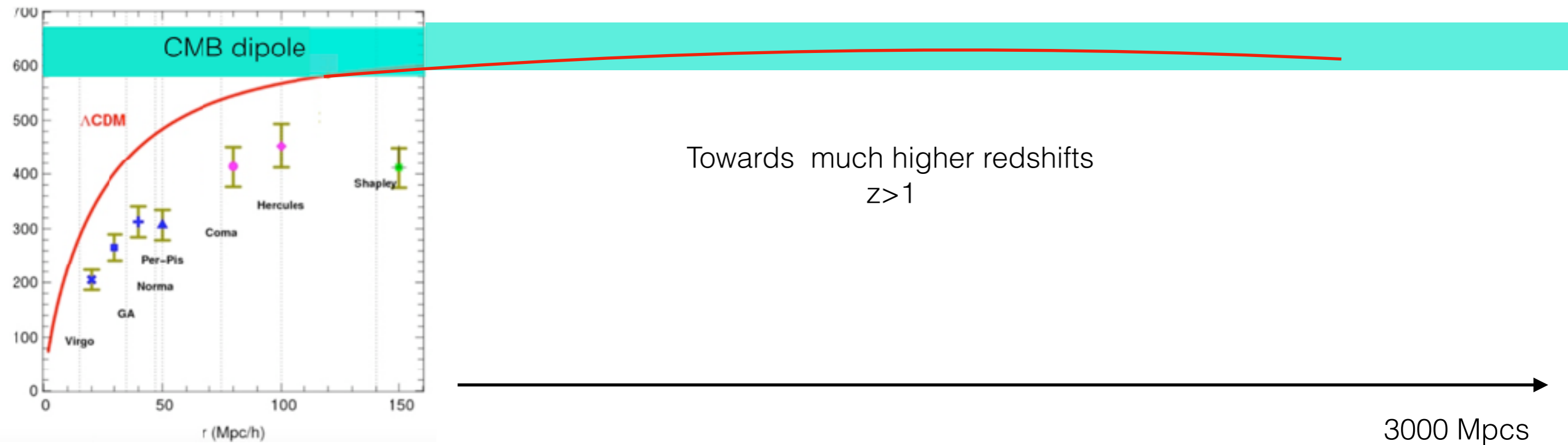


Test of cosmological principle : searching for CMB rest frame



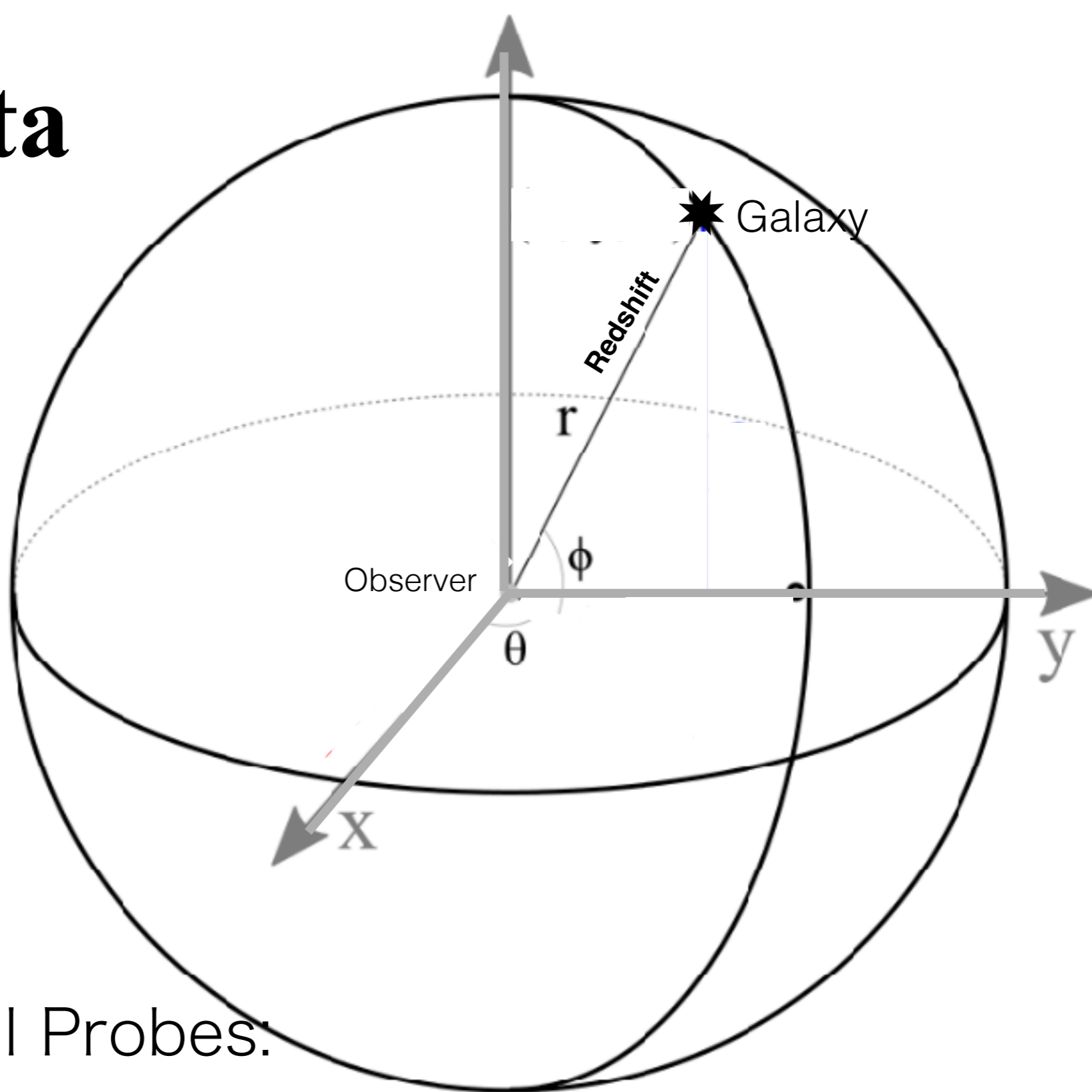
Dipole in the rest-frame of high redshift sources = Dipole in the CMB rest frame

Test of cosmological principle : searching for CMB rest frame



The rest-frame of high redshift "sources" = Rest frame of CMB

Real Data



Observational Probes:

(I) Θ , φ , z , d (distance catalogues)

(II) Θ , φ , z (redshift surveys)

(III) Θ , φ (Imaging surveys)

Real Data



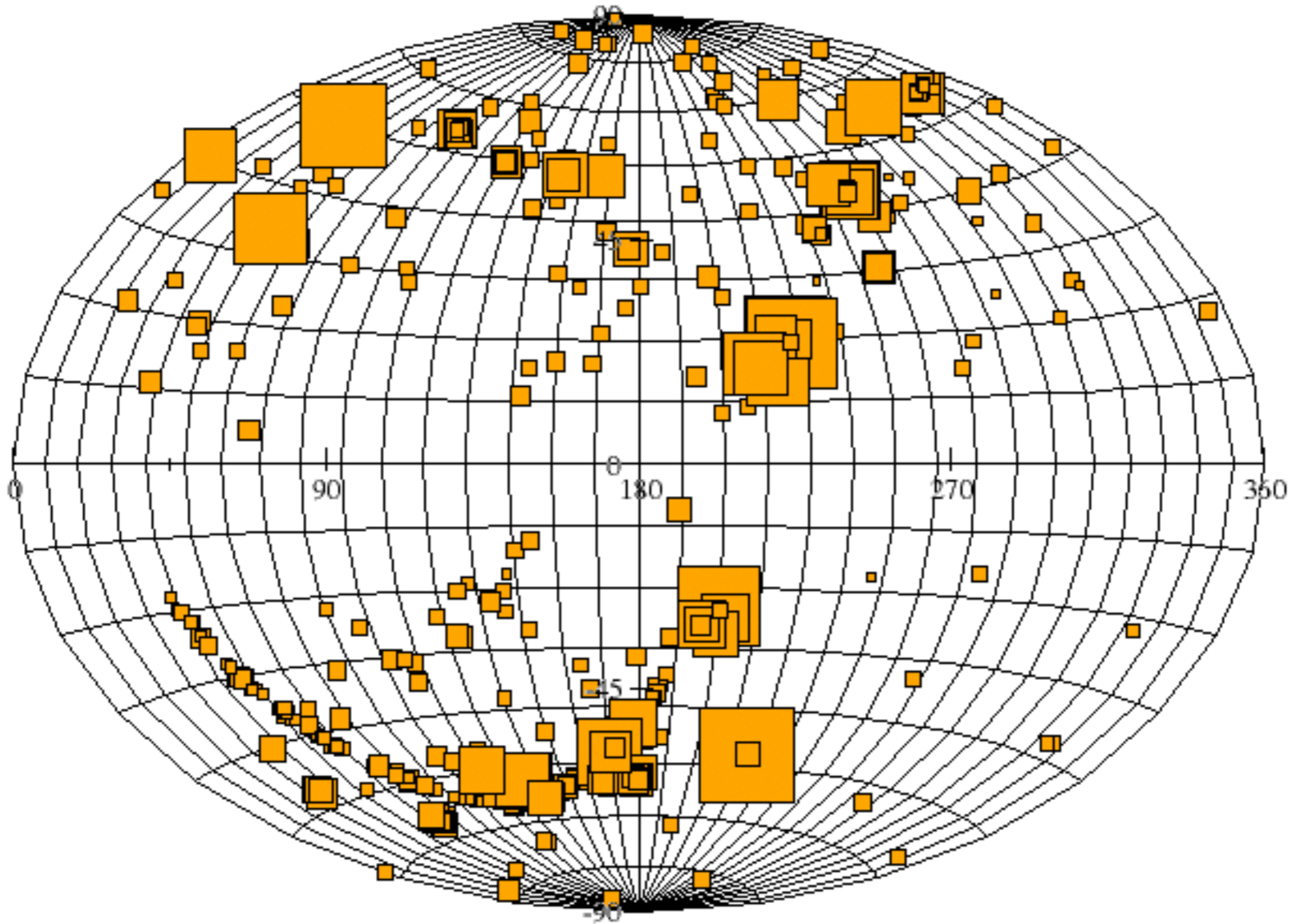
Observational Probes:

(I) Θ, φ, z, d (distance catalogues)

(II) Θ, φ, z (spectro surveys)

(III) Θ, φ (imaging surveys)

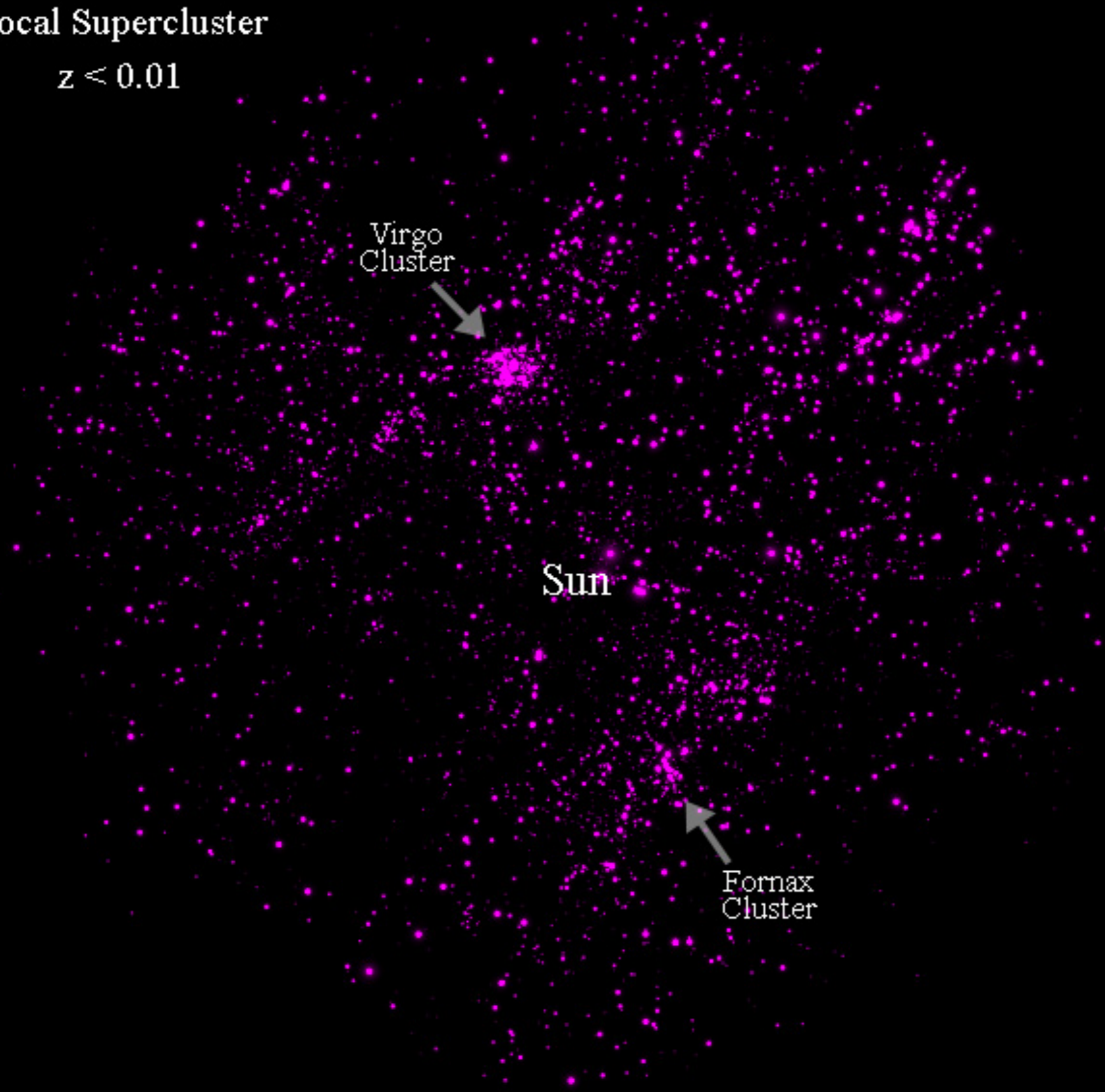
Velocities (distances) from SNe Ia Union II compilation



Bulk flow of increasingly volume: CMB rest frame ?

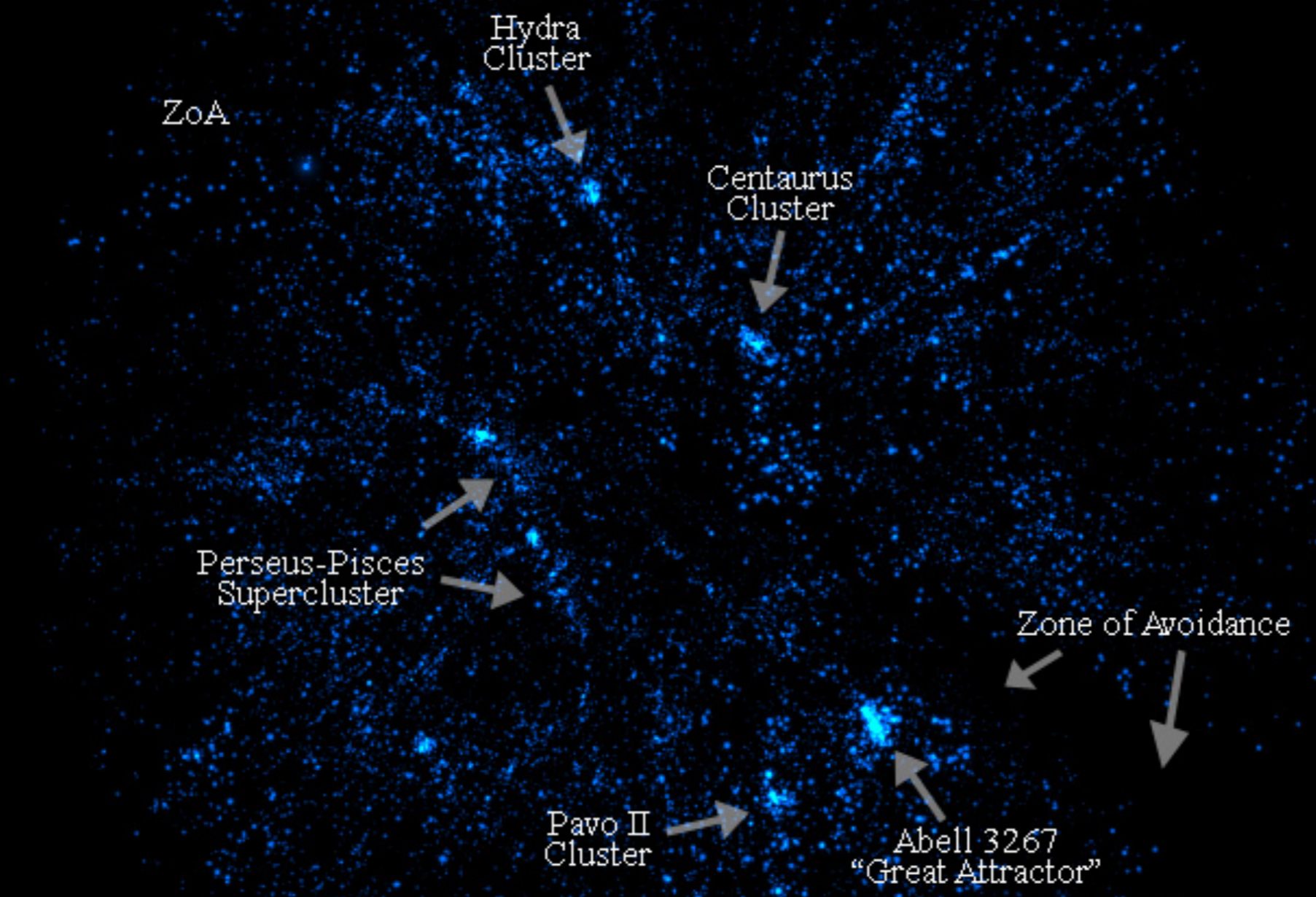
Local Supercluster

$z < 0.01$

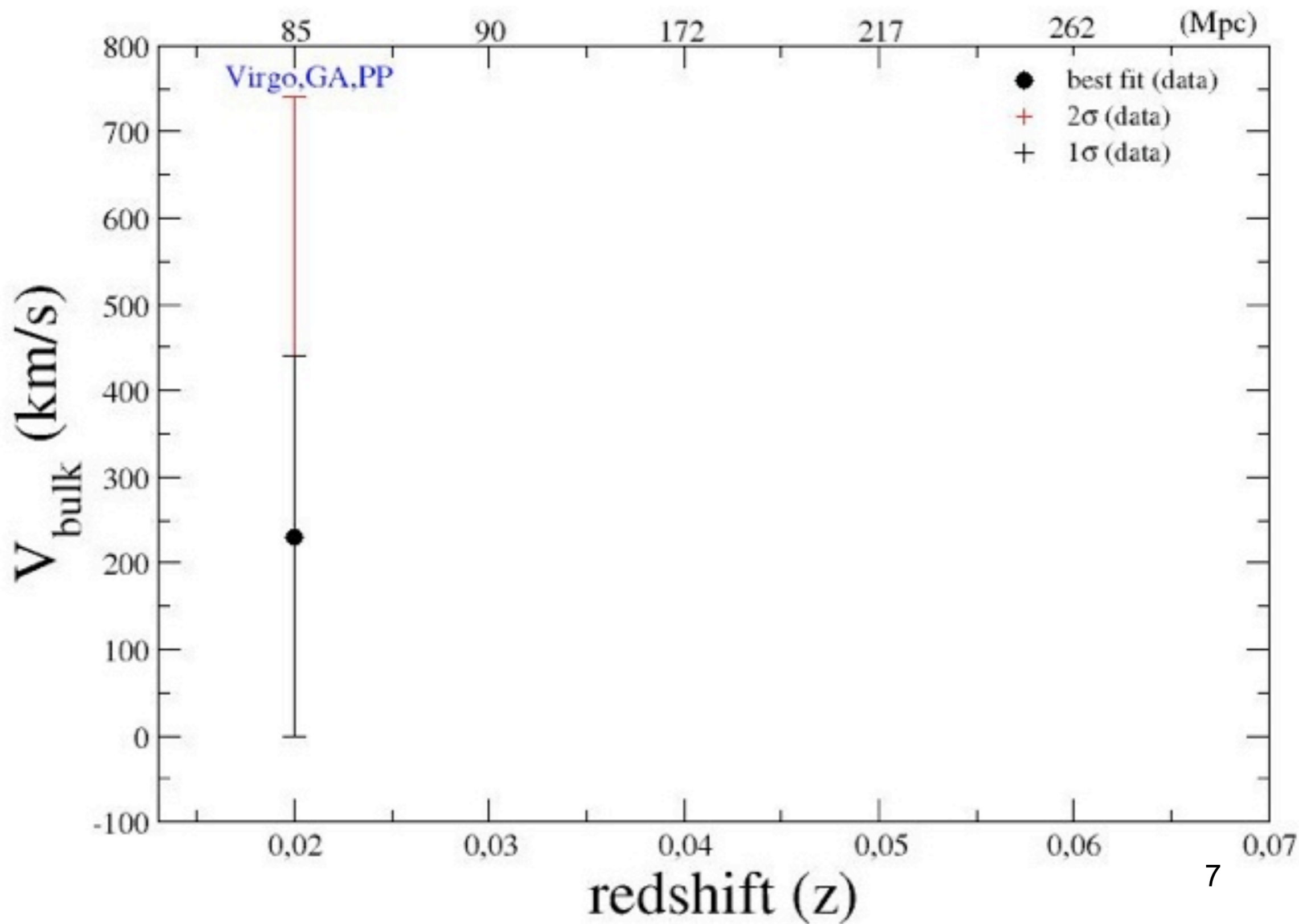


Bulk flow of increasingly larger volume: CMB rest frame ?

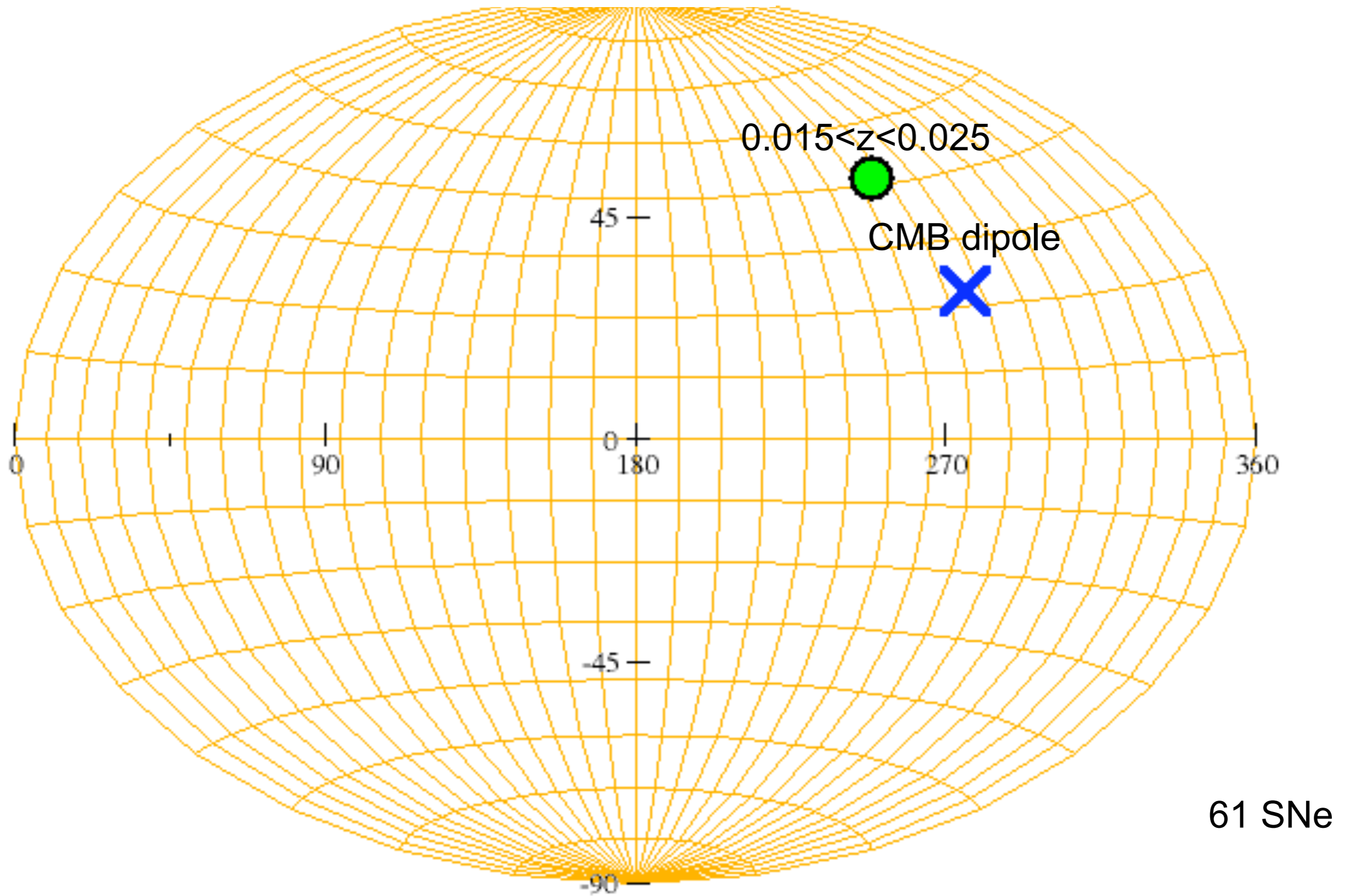
$0.01 < z < 0.02$



|Bulk flow| from SNe Ia data

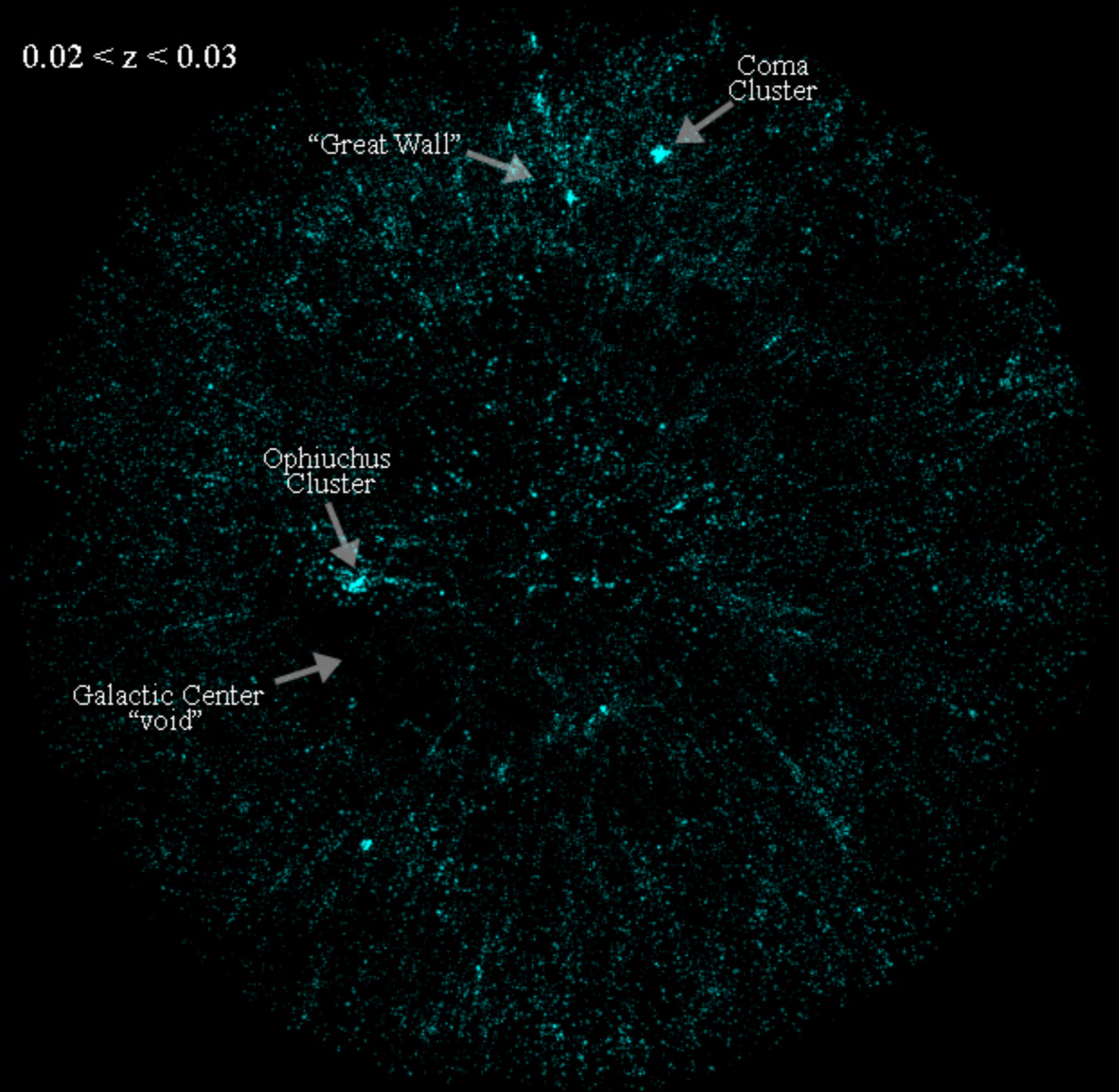


Bulk flow direction from data

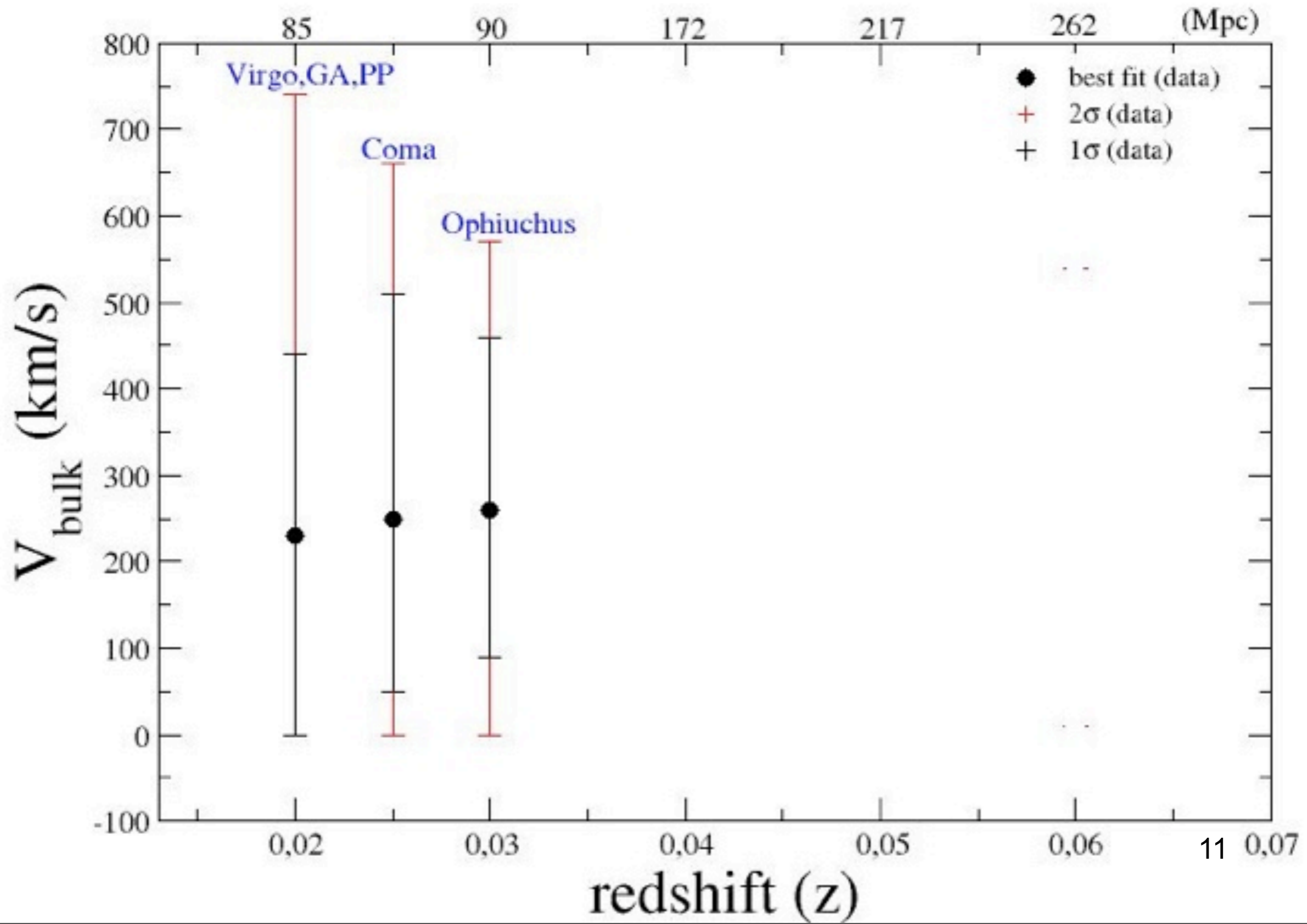


Bulk flow of increasingly larger volume: CMB rest frame ?

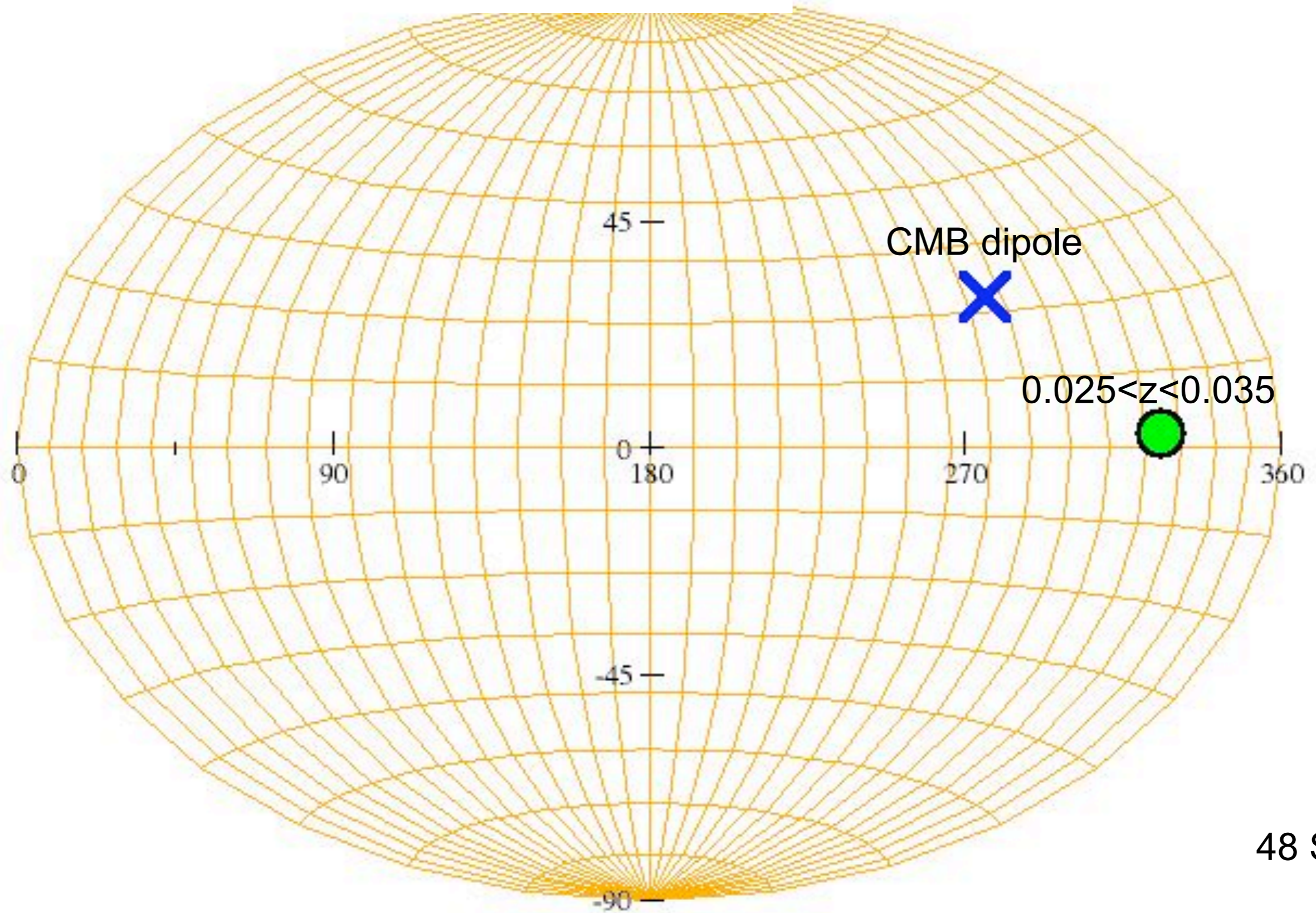
$0.02 < z < 0.03$



|Bulk flow| from SNe Ia data



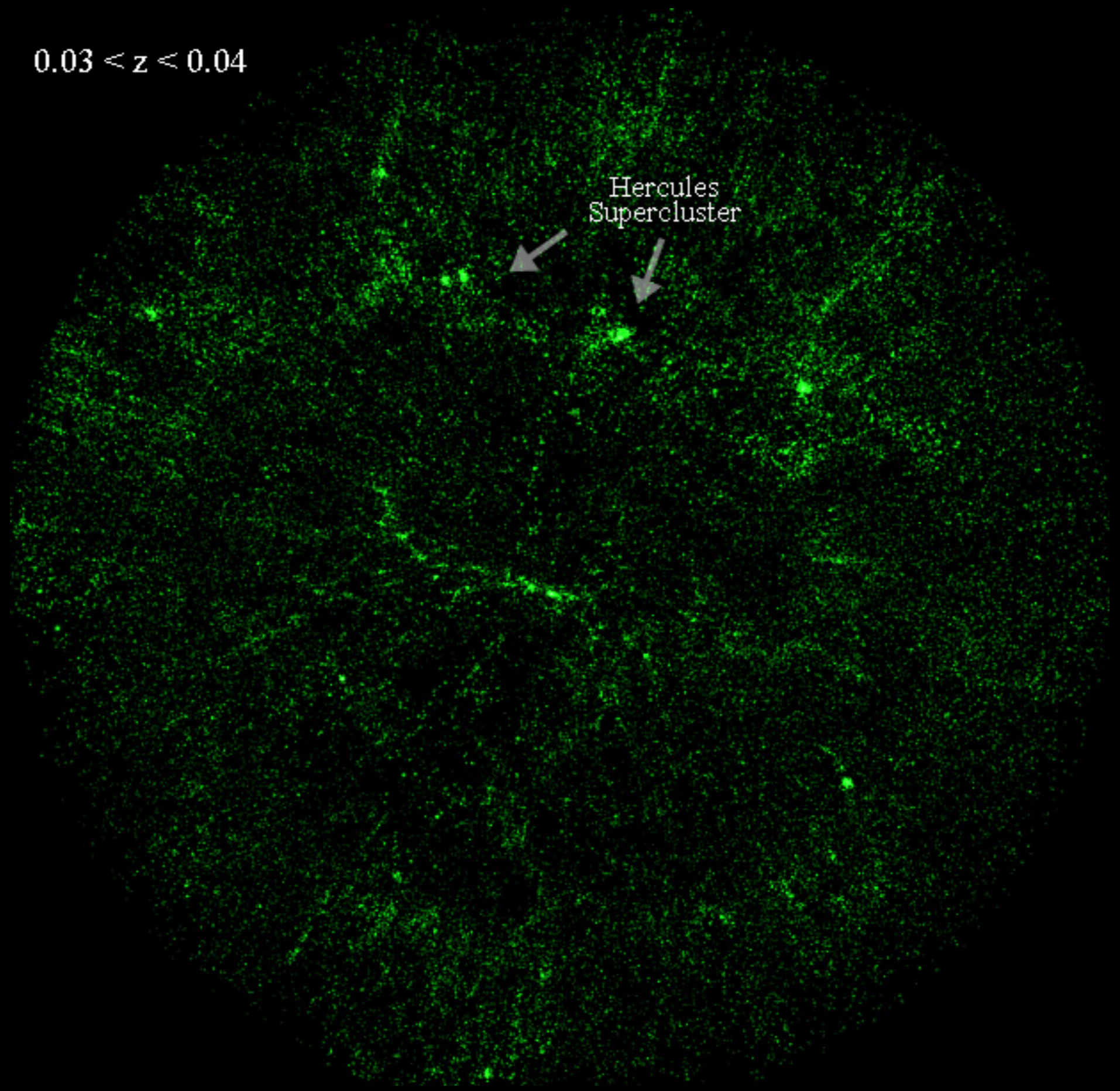
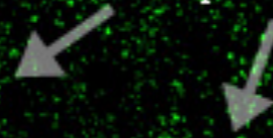
Bulk flow direction from data



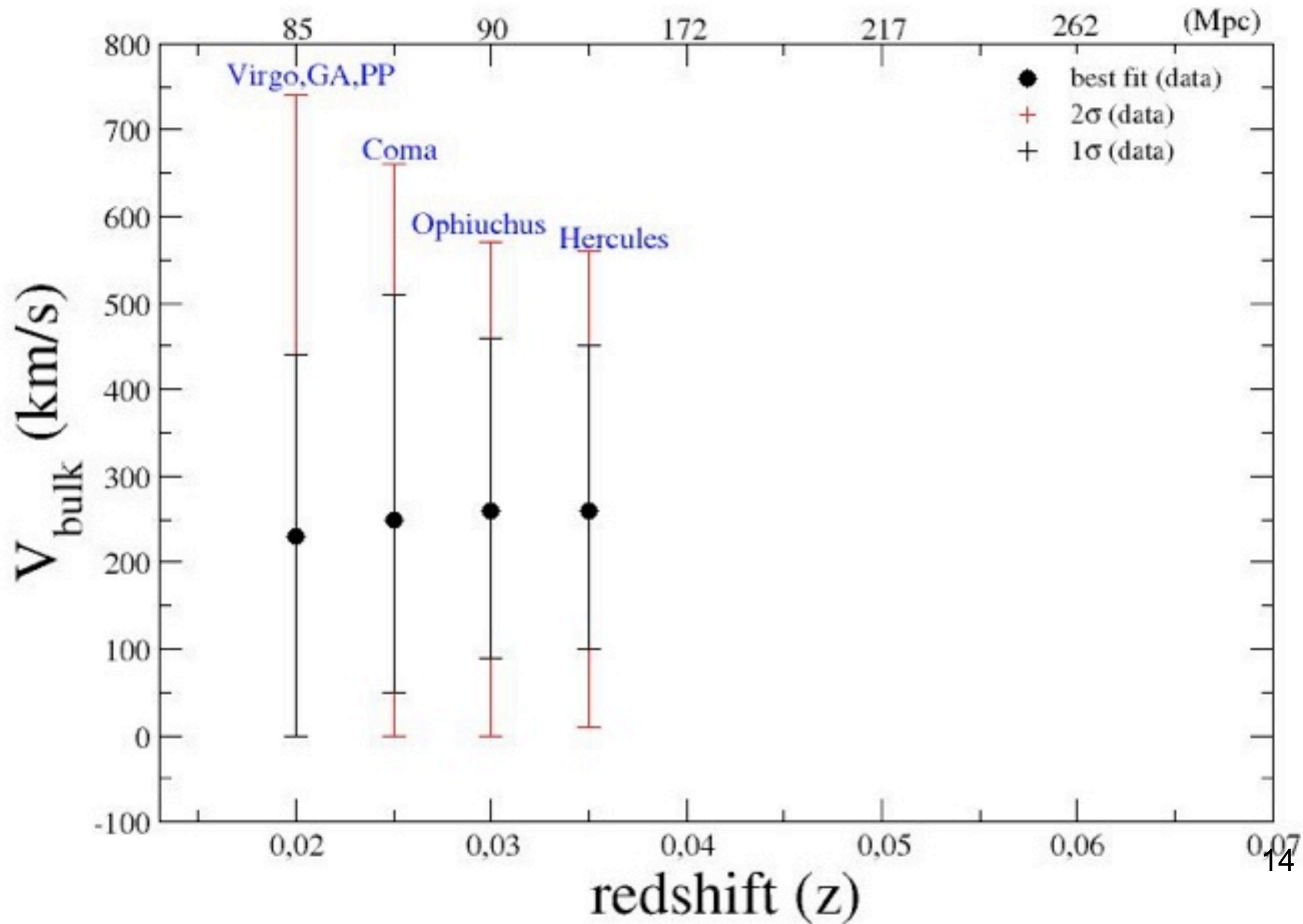
48 SNe

$0.03 < z < 0.04$

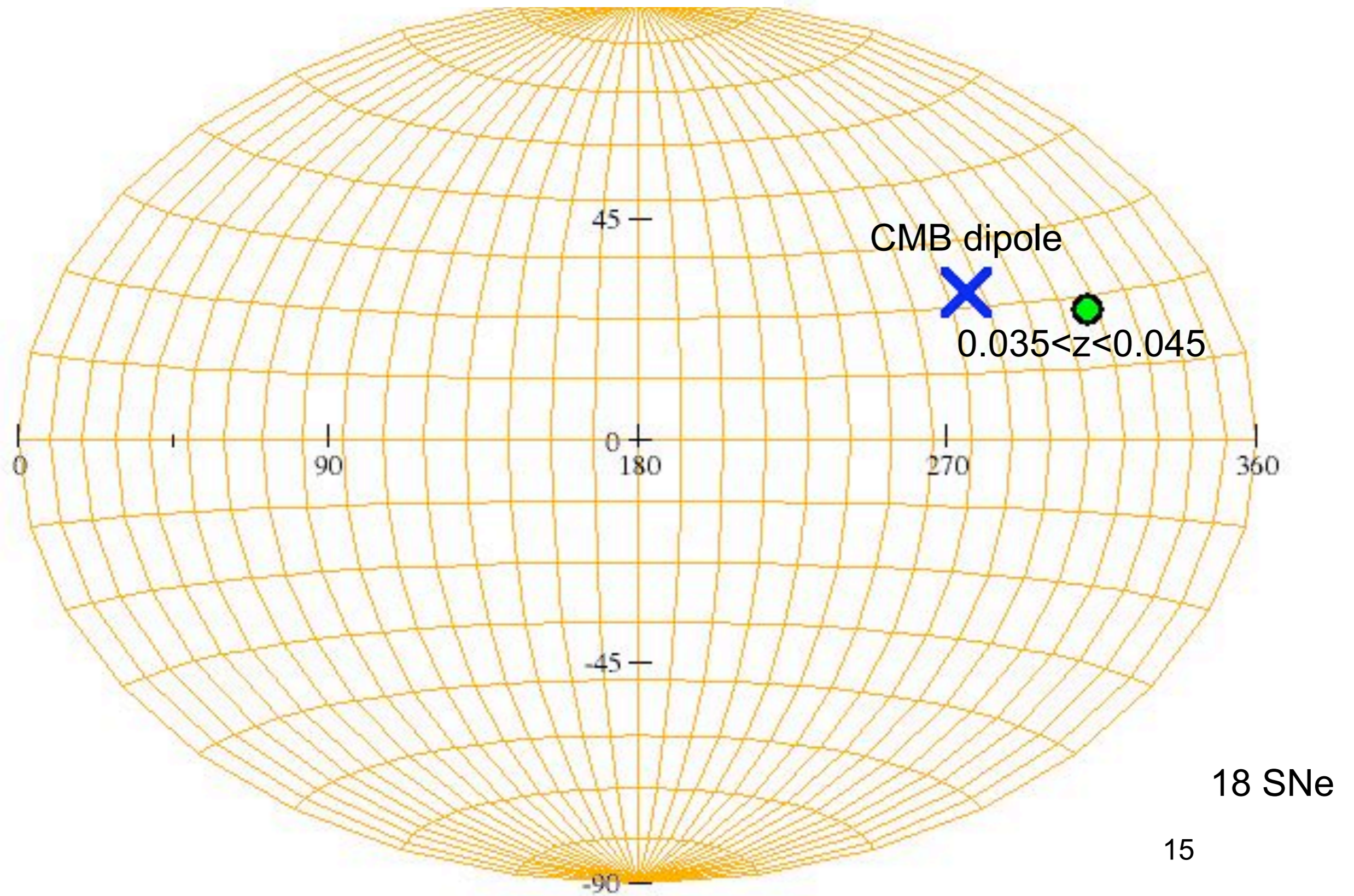
Hercules
Supercluster



|Bulk flow| from SNe Ia data

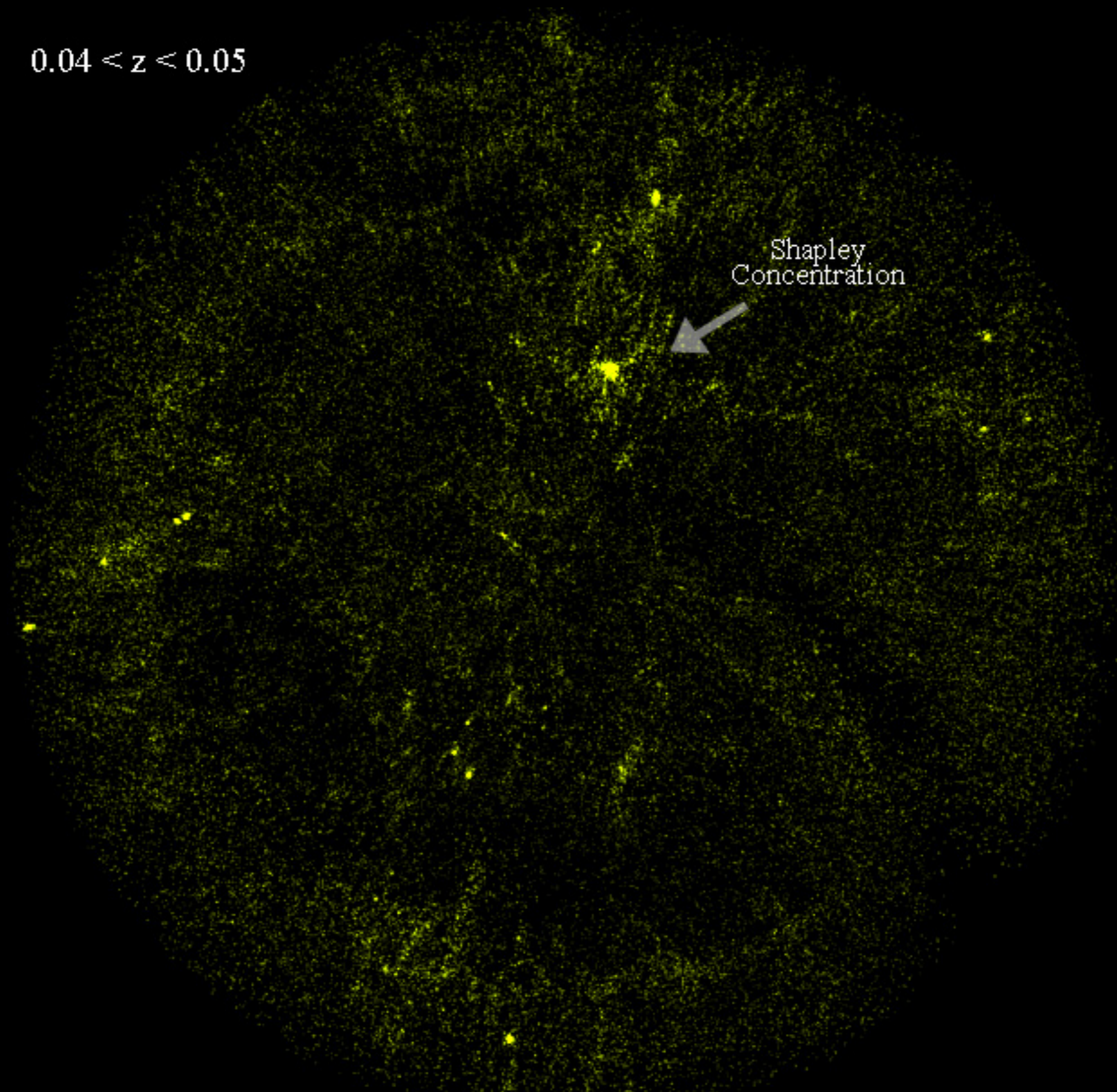


Bulk flow direction from data

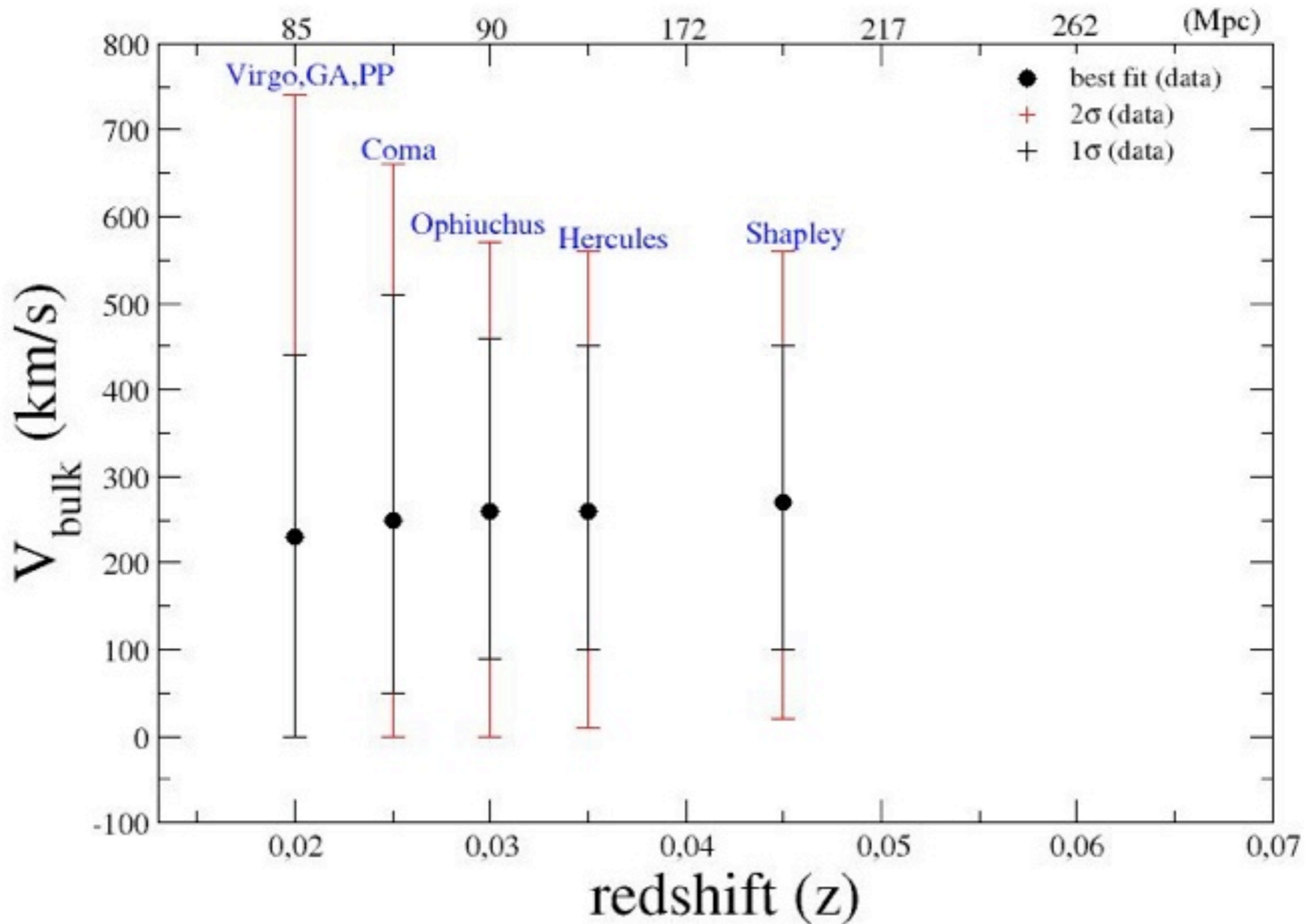


$0.04 < z < 0.05$

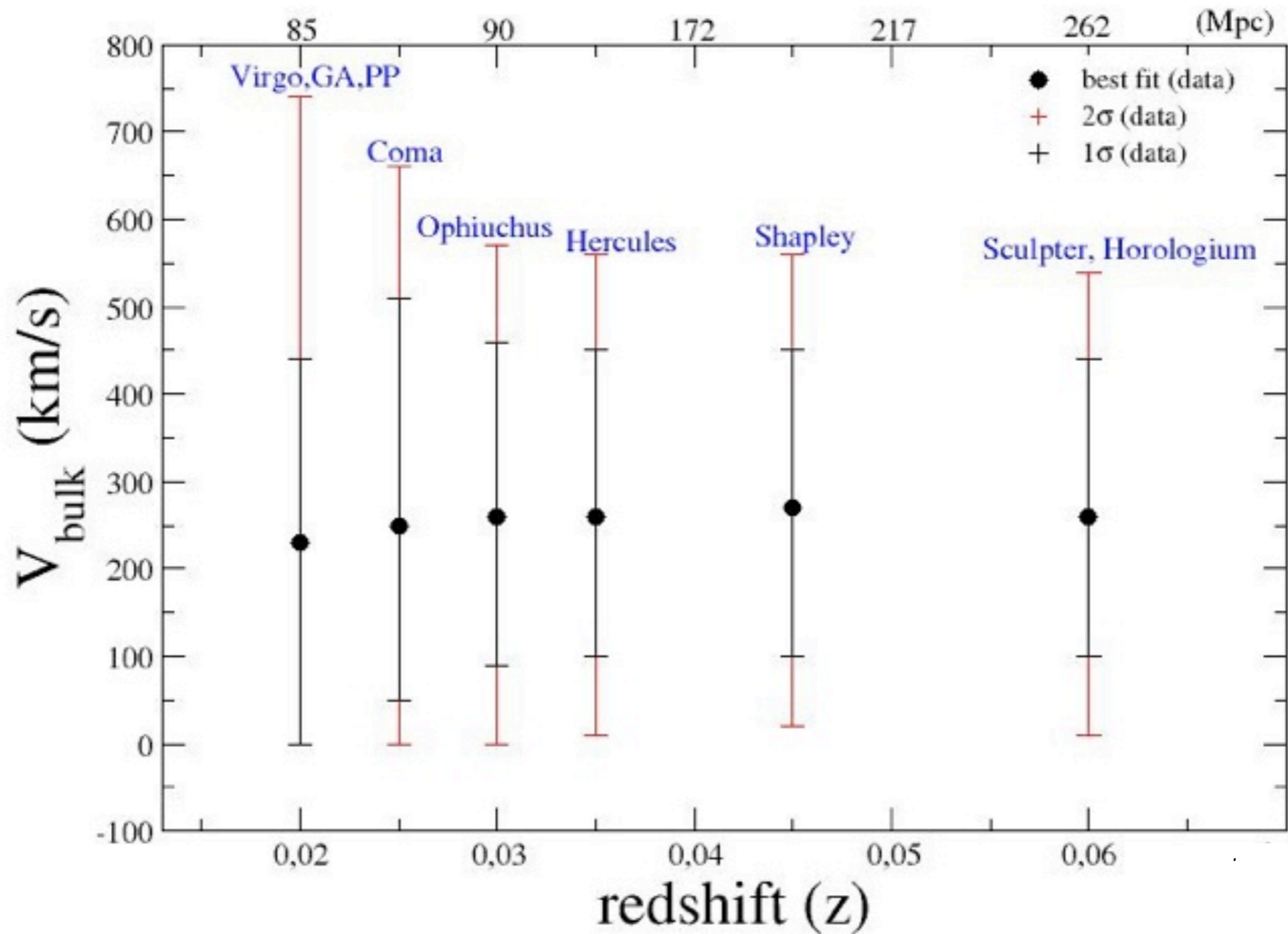
Shapley
Concentration



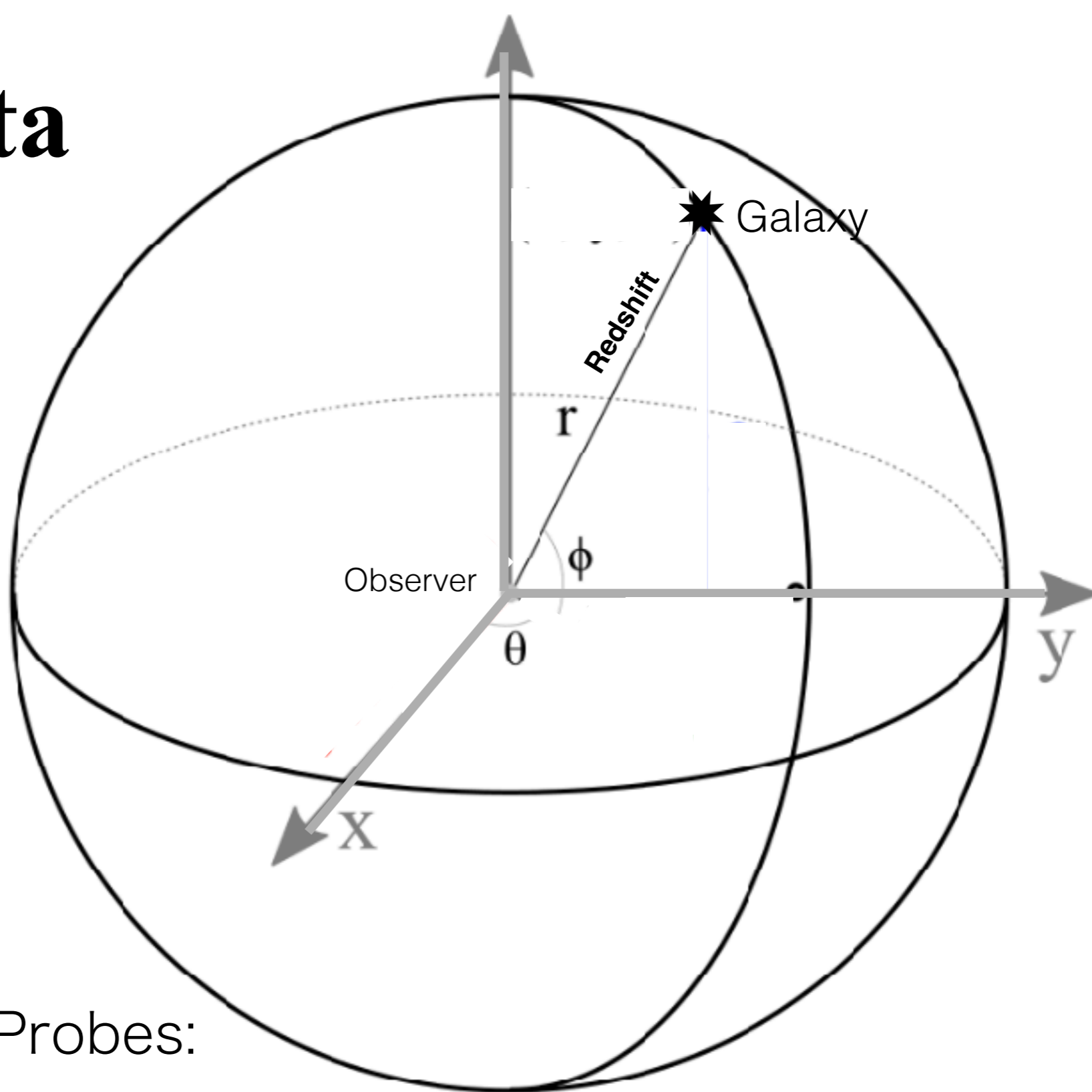
|Bulk flow| from SNe Ia data



|Bulk flow| from SNe Ia data



Real Data



Observational Probes:

(I) Θ , φ , z , d (SNe Ia ... catalogues)

(II) Θ , φ , z (redshift catalogues)

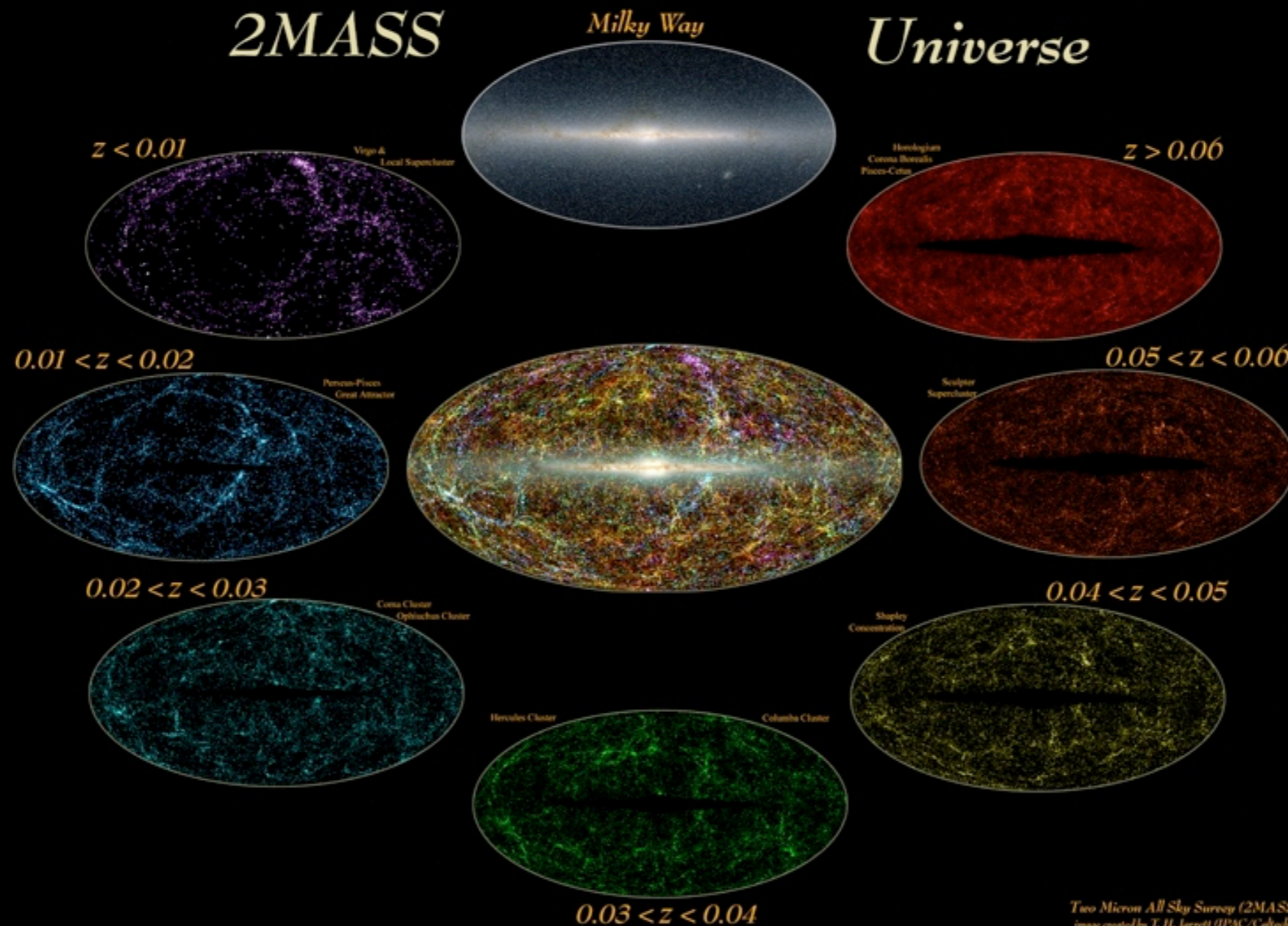
(III) Θ , φ (photometric catalogues)

2MRS redshift survey

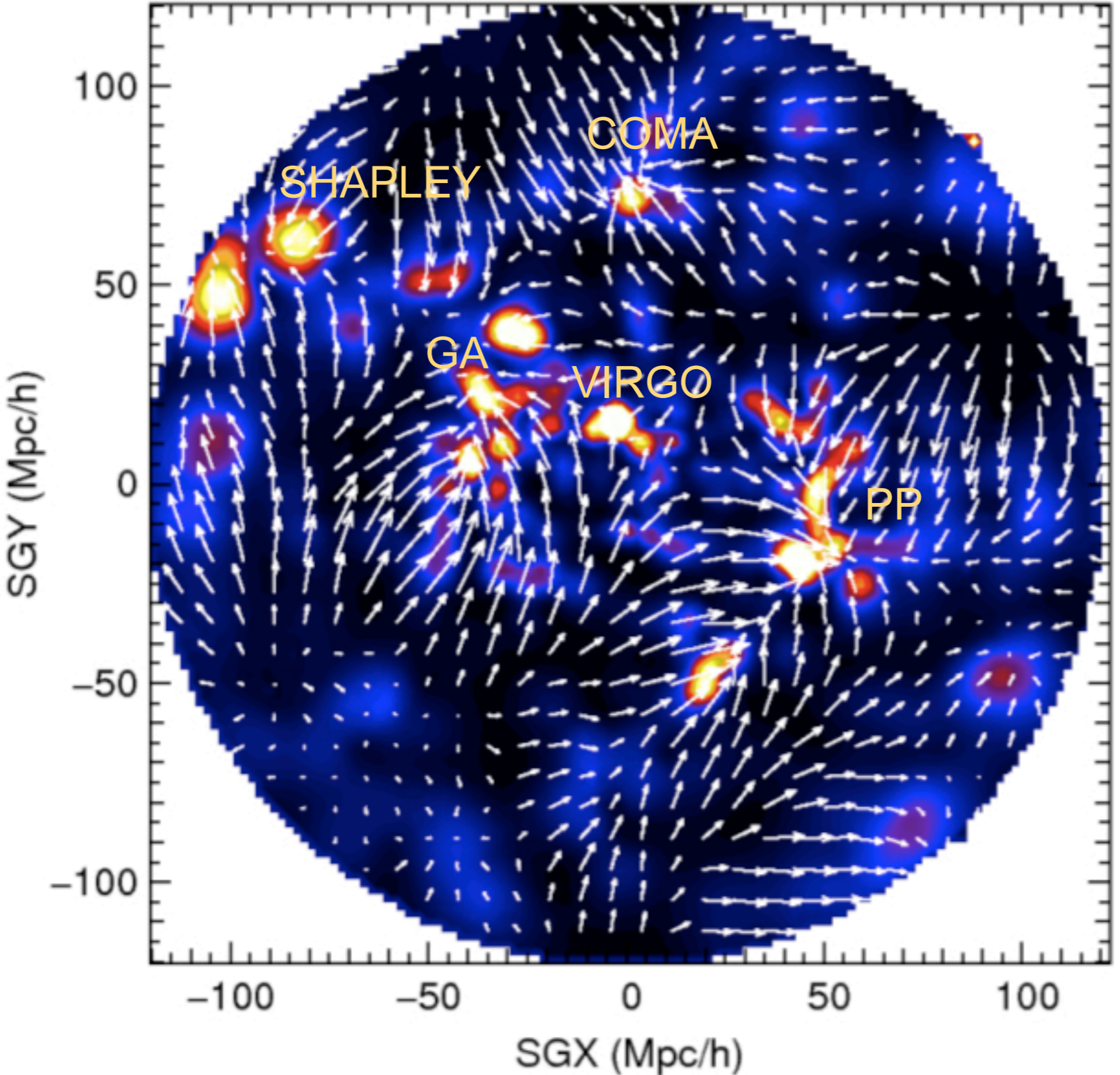
(Huchra et al 2005,...)

Based upon the 2MASS photometric galaxy catalog , Full sky
~25000 galaxies, selected with $K_s < 11.25$

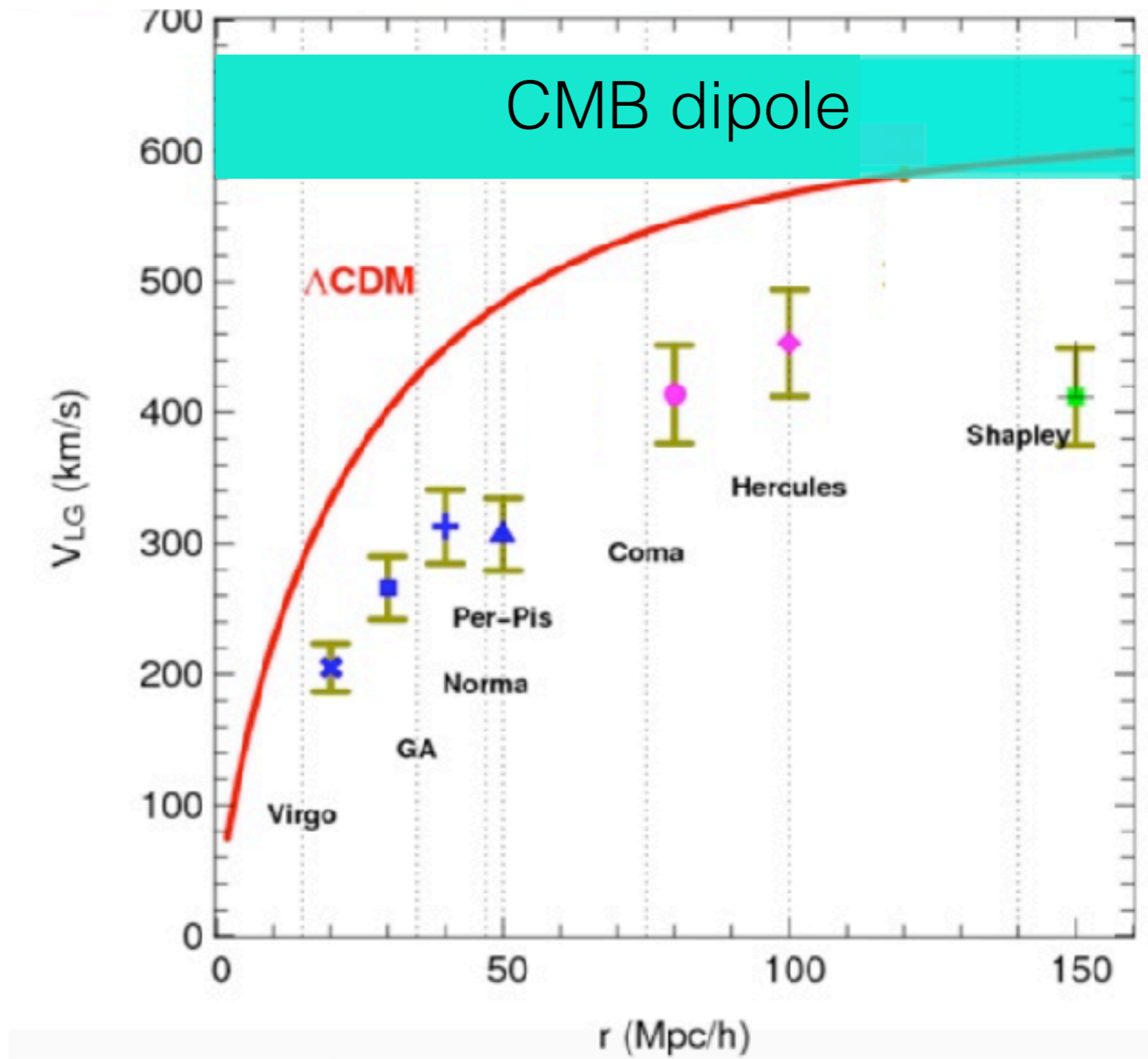
~250 Mpc/h ($z \sim 0.08$) deep , Distribution peaks at ~90 Mpc/h ($z \sim 0.03$)



Velocity field of 2MRS: from great attractor to Shapley infall

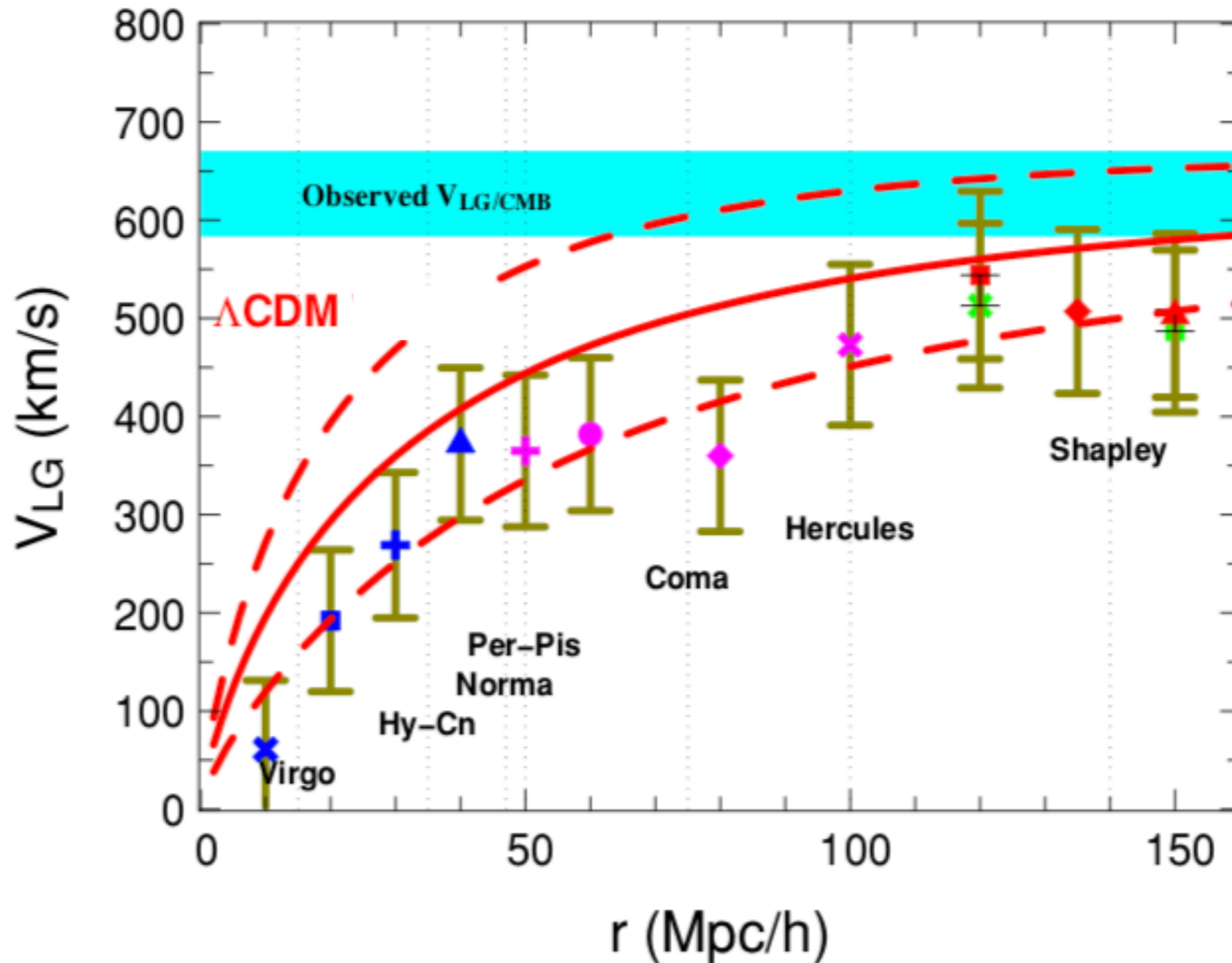


Test of cosmological principle : CMB rest frame



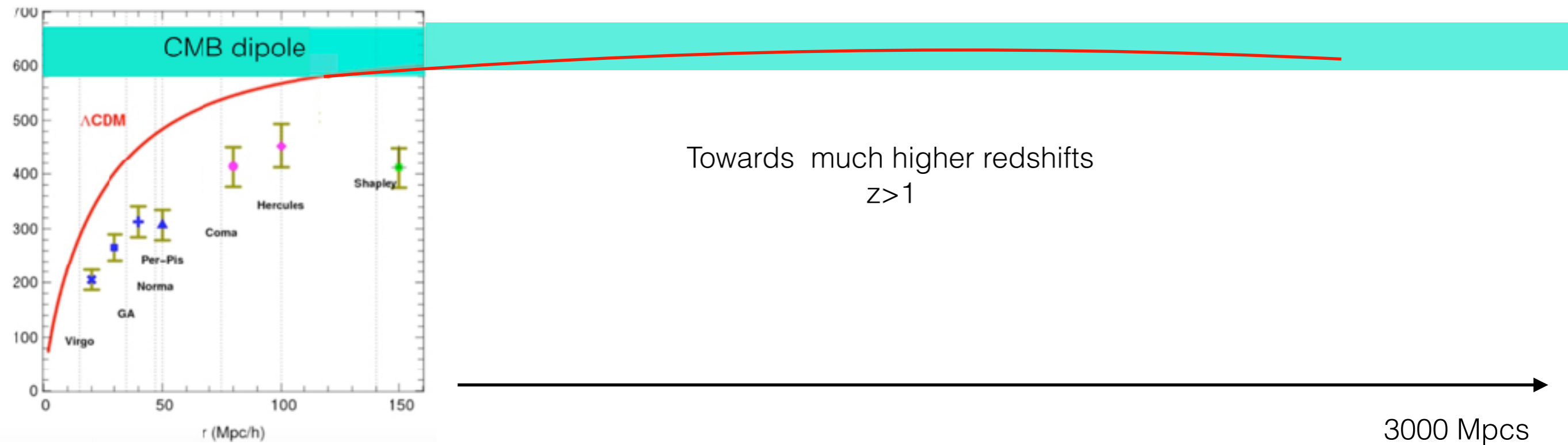
Cosmological principle:
rest-frame of high redshift "sources" = CMB rest frame

Test of cosmological principle : CMB rest frame



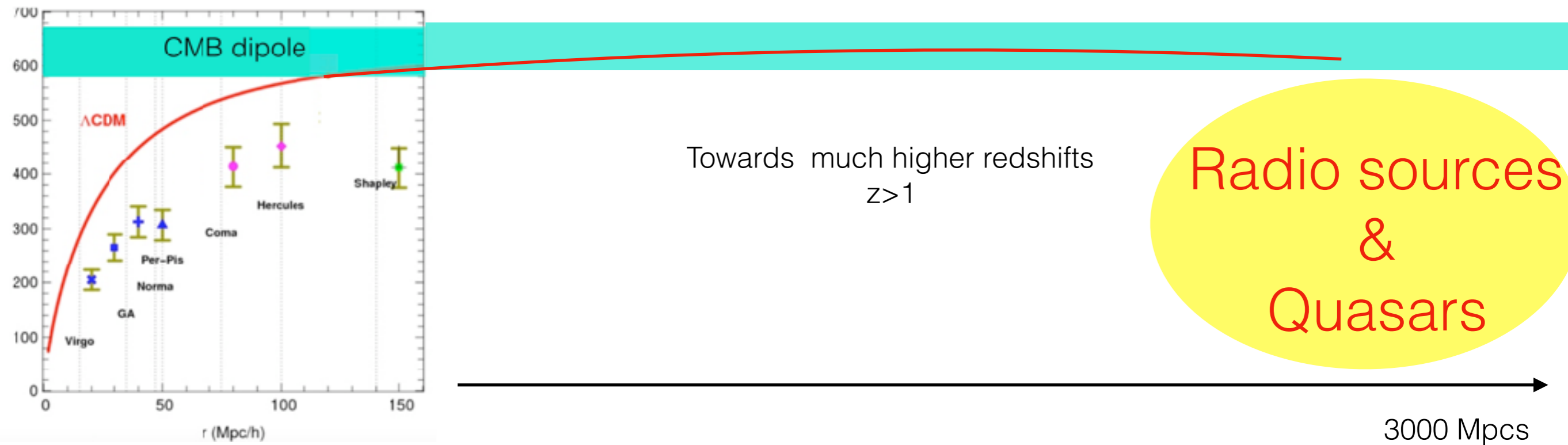
Cosmological principle:
rest-frame of high redshift "sources" = CMB rest frame

Test of cosmological principle : CMB rest frame



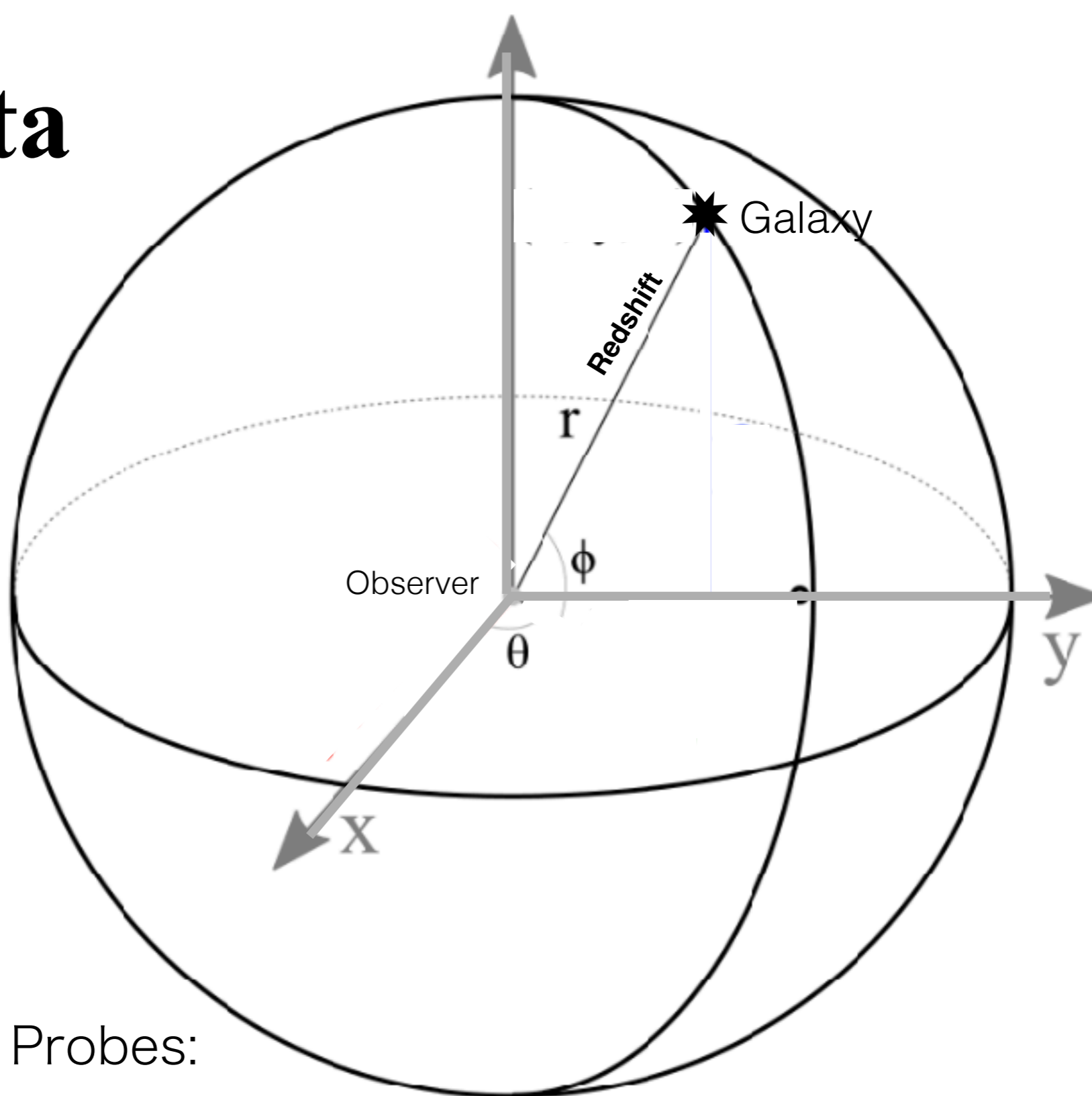
Cosmological principle:
rest-frame of high redshift "sources" = CMB rest frame

Test of cosmological principle : searching for CMB rest frame



Dipole in the rest-frame of high redshift sources = Dipole in the CMB rest frame

Real Data



Observational Probes:

(I) Θ , φ , z , d (distance catalogues)

(II) Θ , φ , z (redshift surveys)

(III) Θ , φ (Imaging surveys)

Probe 3 : photometric data

Θ , ϕ (imaging surveys)

Aberration and Doppler boosting



Colin, Mohayaee, Rameez, Sarkar, MNRAS 2017

Rameez, Mohayaee, Sarkar, Colin, MNRAS 2018

Colin, Mohayaee, Rameez, Sarkar, MNRAS 2019

Mohayaee, Rameez, Sarkar, 2020

Secret, von Hausegger, Rameez, Mohayaee, Sarkar, Colin ..., 2021

Aberration

Ellis and Baldwin 1984

Anisotropy in source distribution  observer's velocity

Aberration and Doppler boosting

$$\text{Dipole} = [2 + x(1 + \alpha)]v/c.$$

$$dN/d\Omega(>S_\nu) \propto S_\nu^{-x}$$

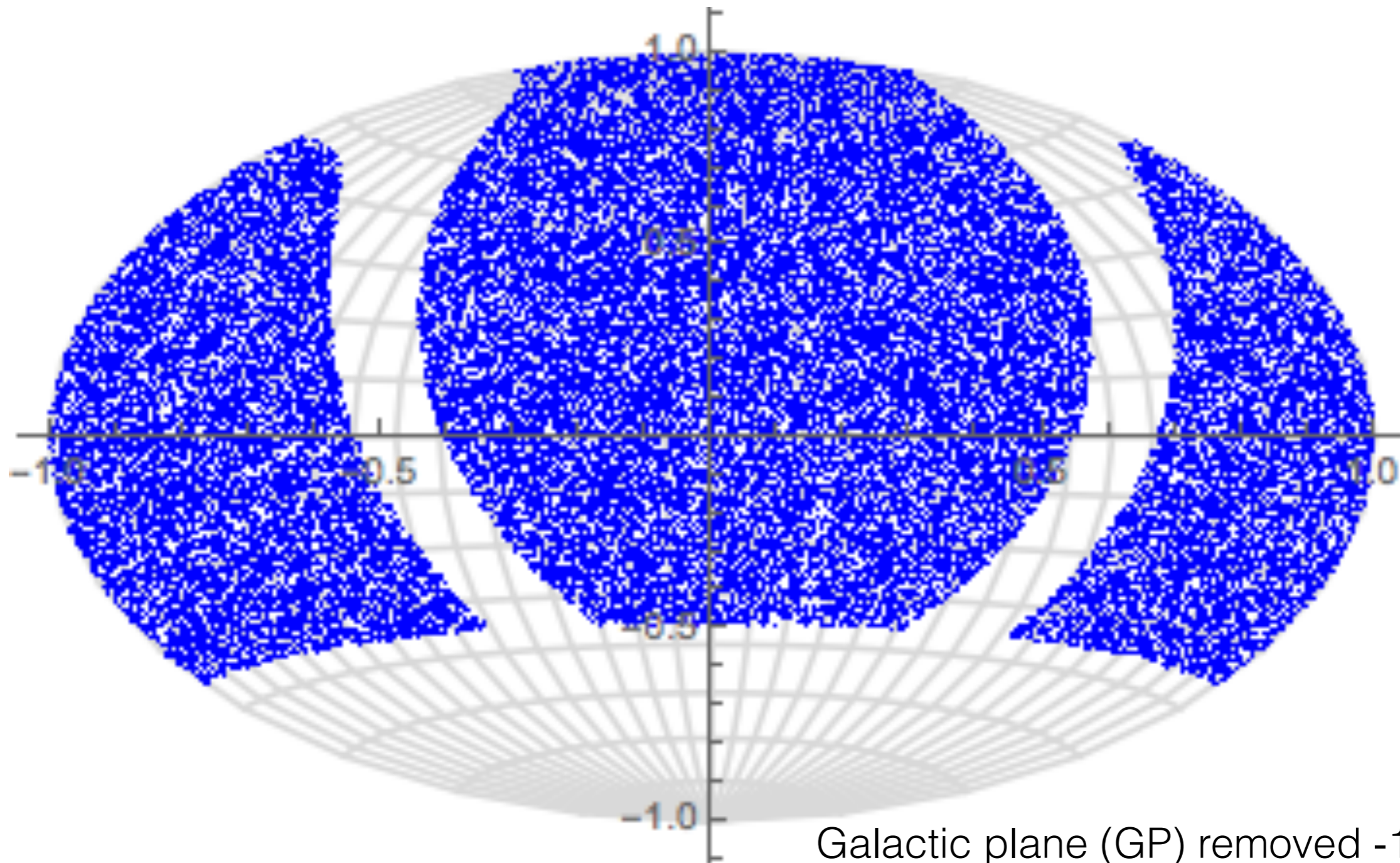
$$S_\nu \propto \nu^{-\alpha}$$

Independent of distance to the source

DATA: NRAO VLA Sky Survey Catalogue (NVSS)

1773488 Radio galaxies

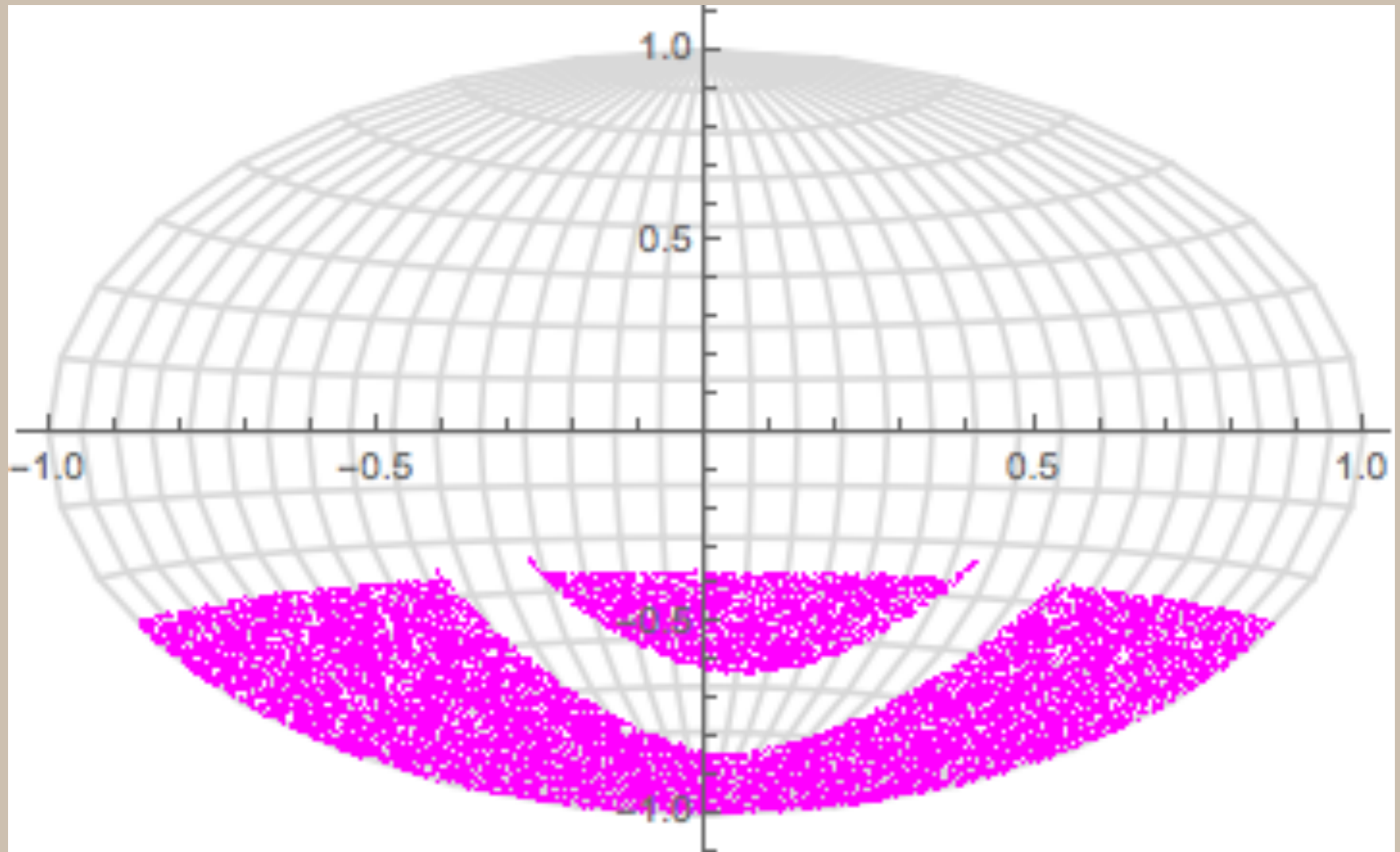
583587 Radio galaxies in $10 \text{ mJy} < \text{Flux} < 1000 \text{ mJy}$



DATA: The Sydney University Molonglo Sky Survey SUMSS

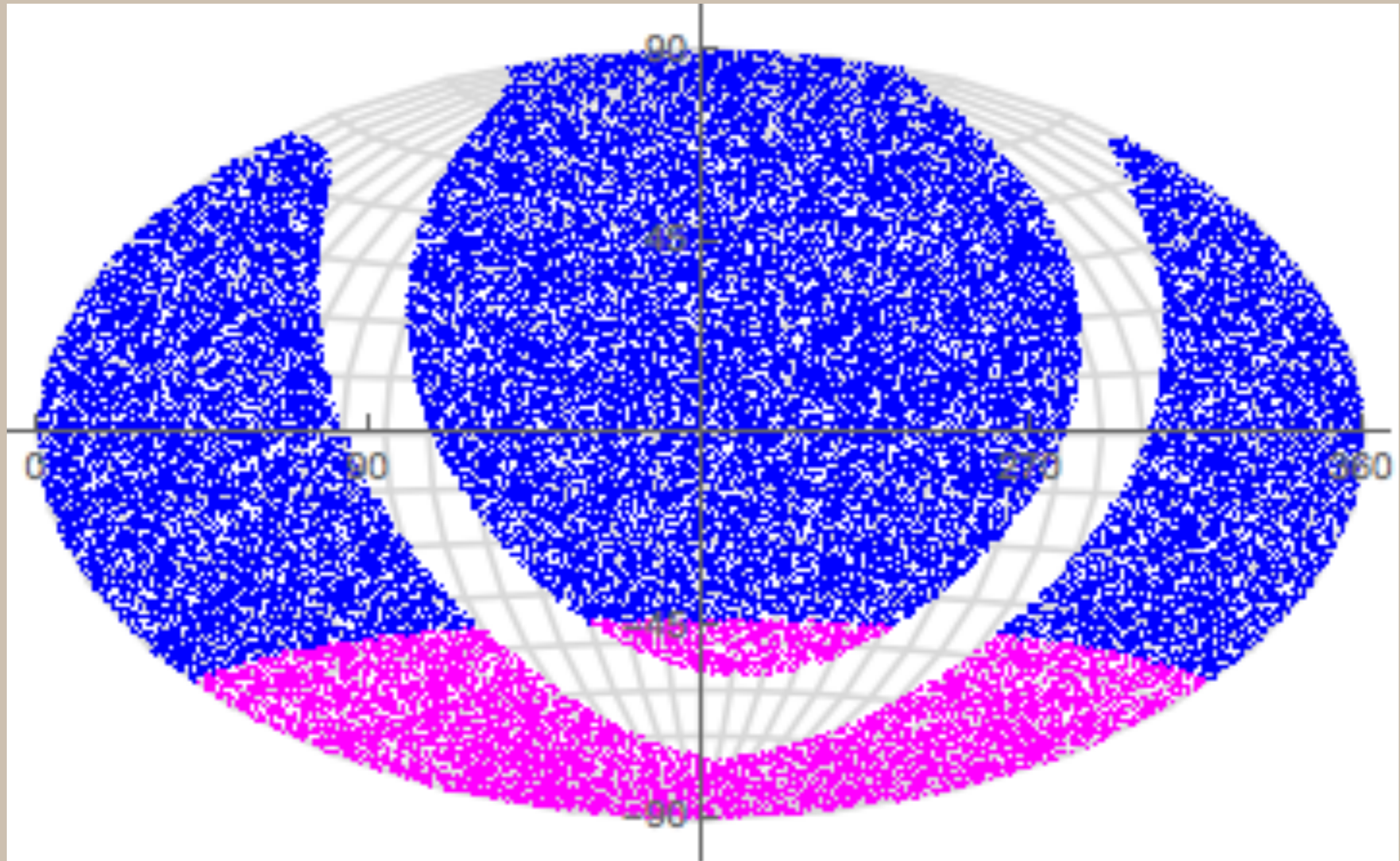
211050 Radio galaxies

183720 Radio galaxies in $10 \text{ mJy} < \text{Flux} < 1000 \text{ mJy}$



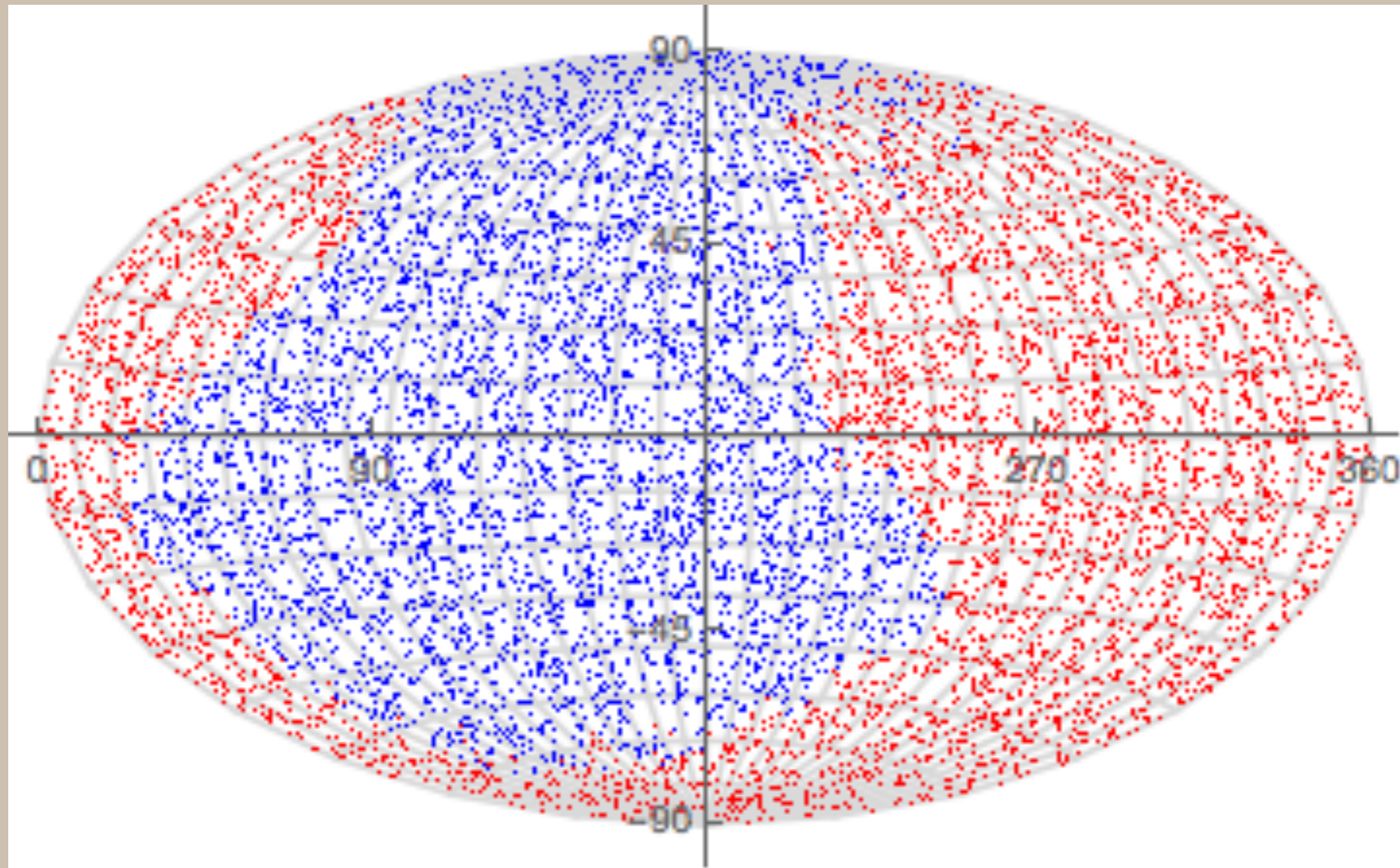
DATA: NVSS+SUMSS

576461 Radio galaxies in $10 \text{ mJy} < \text{Flux} < 1000 \text{ mJy}$



Searching for dipole

We randomly select a direction ($\theta = \{-\pi/2, \pi/2\}$ and $\varphi = \{0, 2\pi\}$) and count Number of galaxies in each hemisphere



Mean number of galaxies in each hemisphere: 345192., Max difference between two hemispheres: 5185 galaxies

Red: hemispheres containing LESS galaxies than the mean

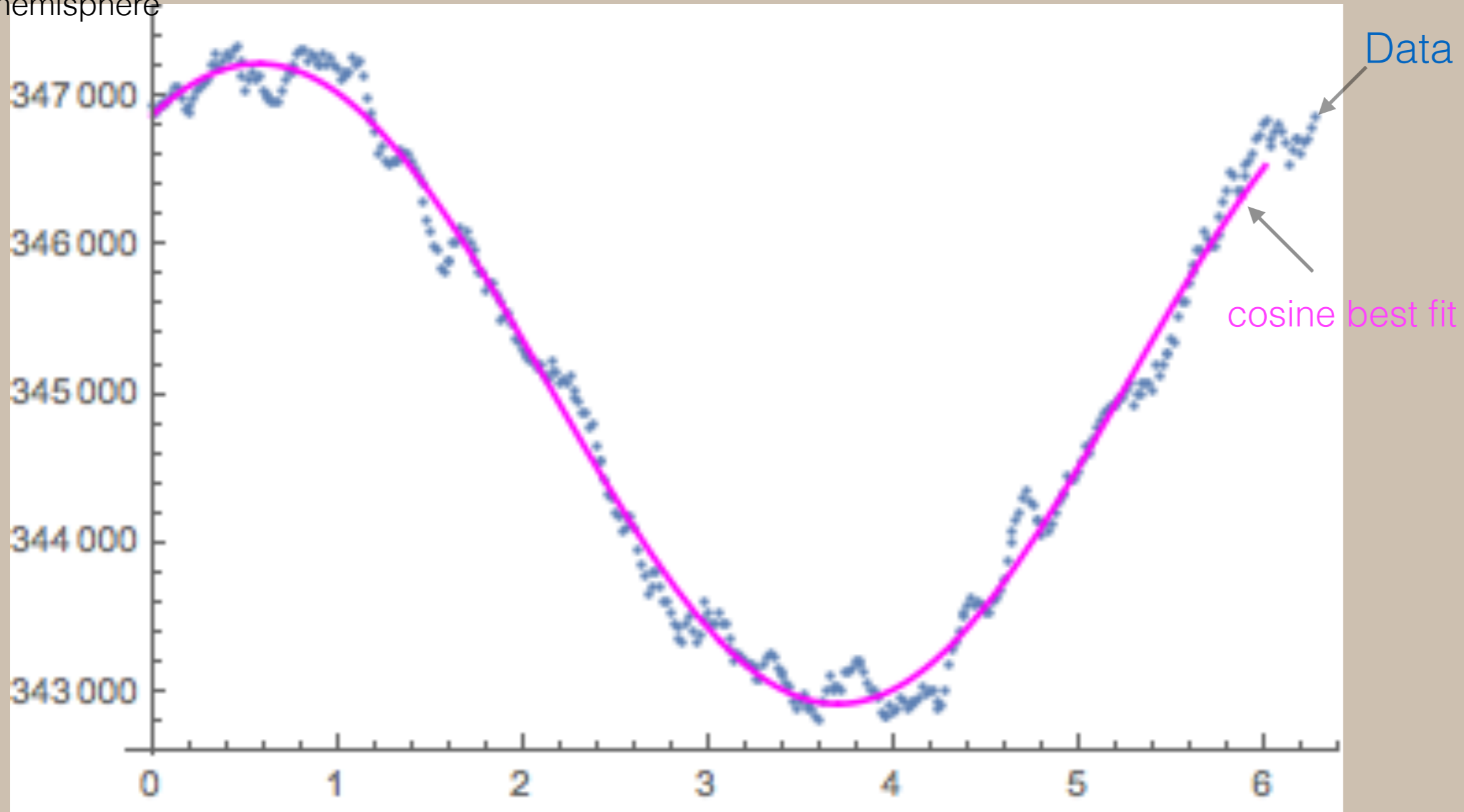
Blue: Hemispheres containing MORE galaxies than the mean

Searching for dipole

Example of hemispherical counting:

Here we fix the axis $\theta=\{0,90\}$ and turn φ every one degree

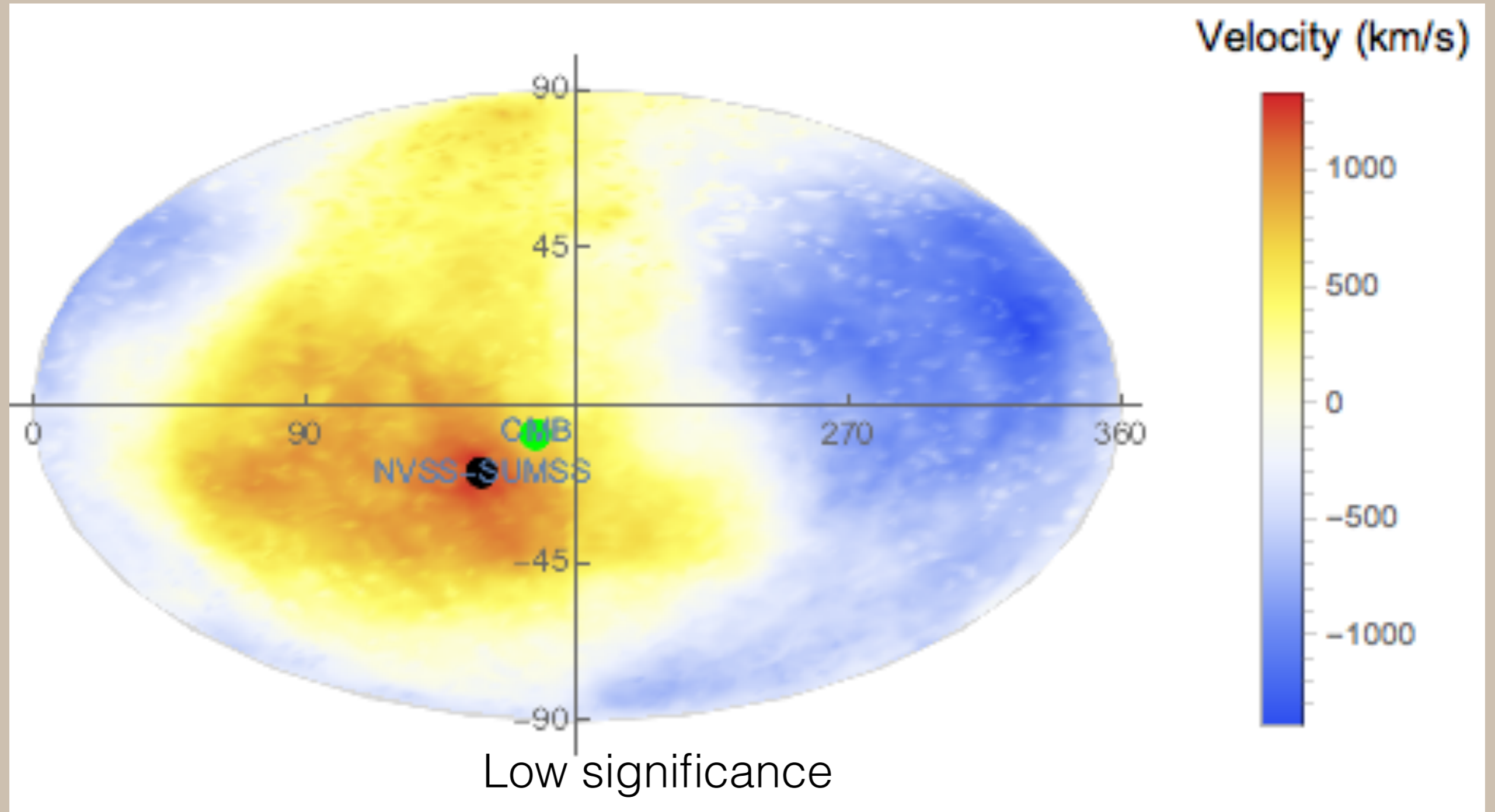
N in hemisphere



Dipole

Dipole direction: {RA=156°, DEC=-17°} compare to CMB Dipole {RA=168°, DEC=-7°}

Dipole Amplitude: velocity of barycentre of solar system w.r.t. Radio galaxies restframe = 1097 km/s
velocity of barycentre of solar system w.r.t. CMB restframe = 369 km/s



Wide-field Infrared Survey Explorer

WISE : (Wright et al. 2010) & NEOWISE (Mainzer et al. 2011)

CatWISE : Eisenhardt et al 2020

positions and the four-band photometry for 747,634,026 objects

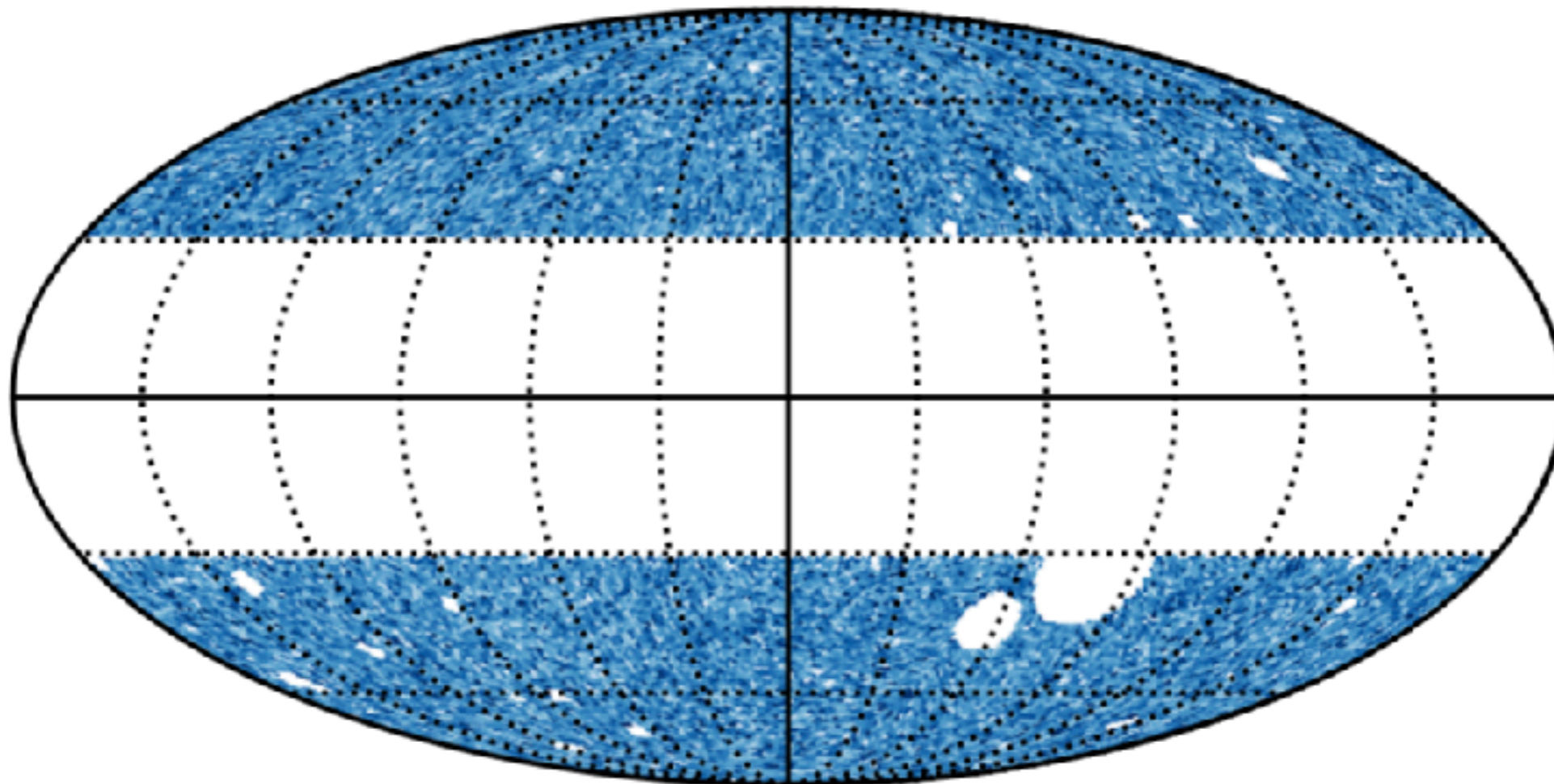
Full-sky **mid-infrared** survey in:

3.4um (W1) (2009 – present)

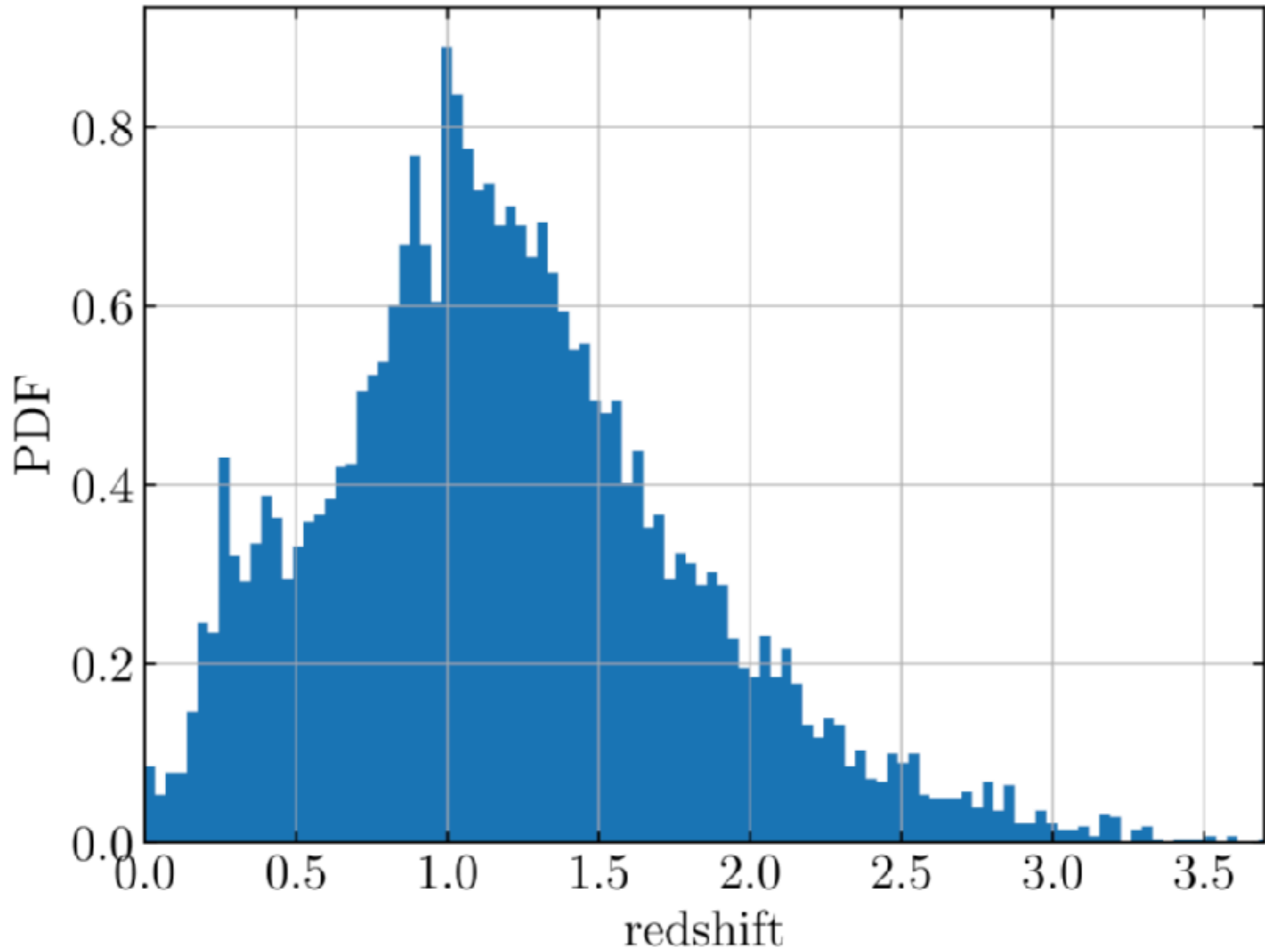
4.6um (W2) (2009 – present)

12um (W3) (2009 – 2010)

22um (W4) (2009 – 2010)



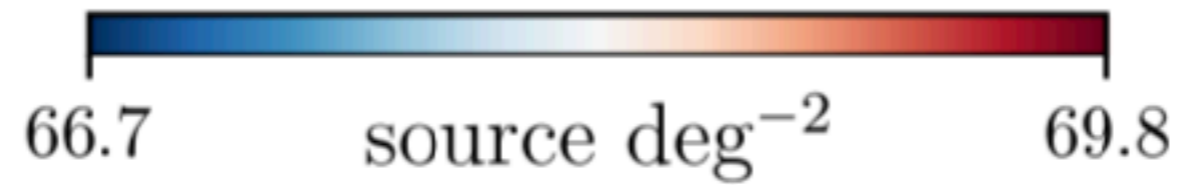
Redshift distribution



The Dipole

Quasar Dipole = 0.01554, $(l, b) = (238^\circ.2, 28^\circ.8)$.

CMB dipole. = 0.007, $(l, b) = (276^\circ, 30^\circ)$

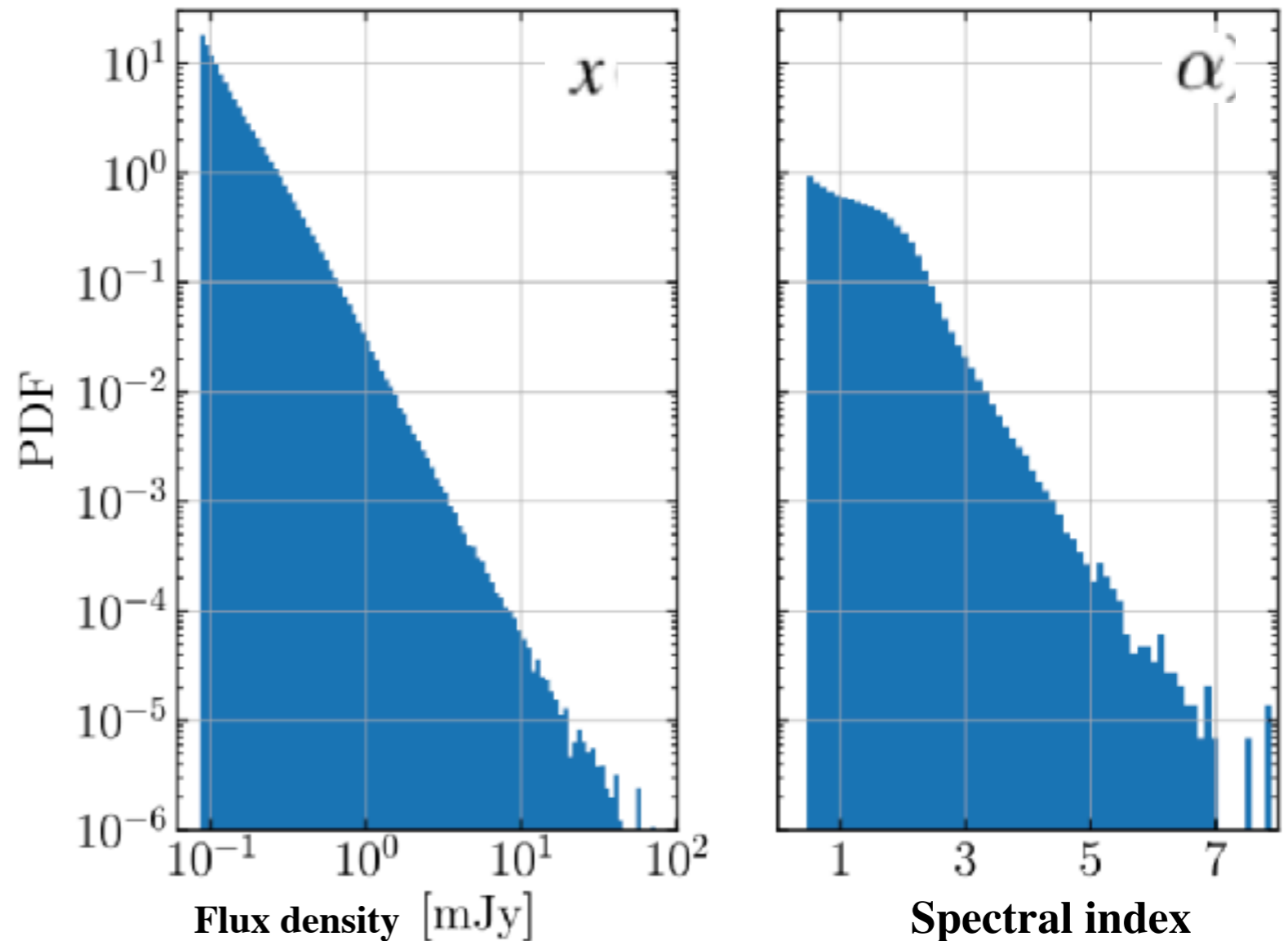


Statistical significance

10^7 random sky

mimicking CatWISE
same masks,
estimator, flux....

$$\text{Dipole} = [2 + x(1 + \alpha)]v/c.$$



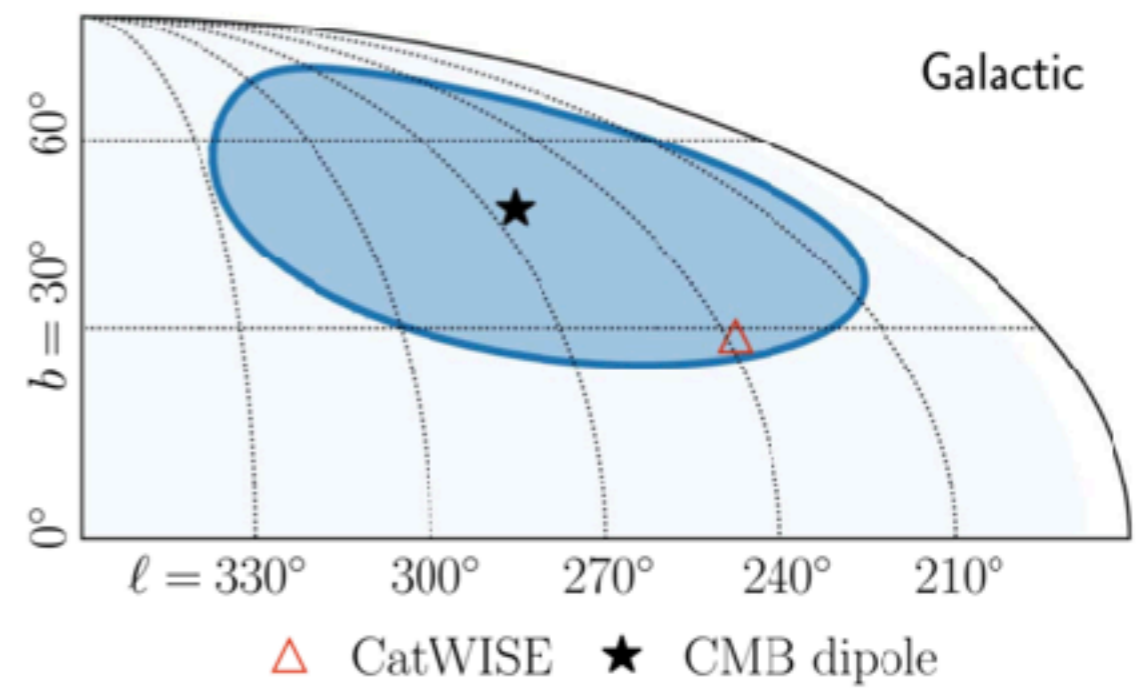
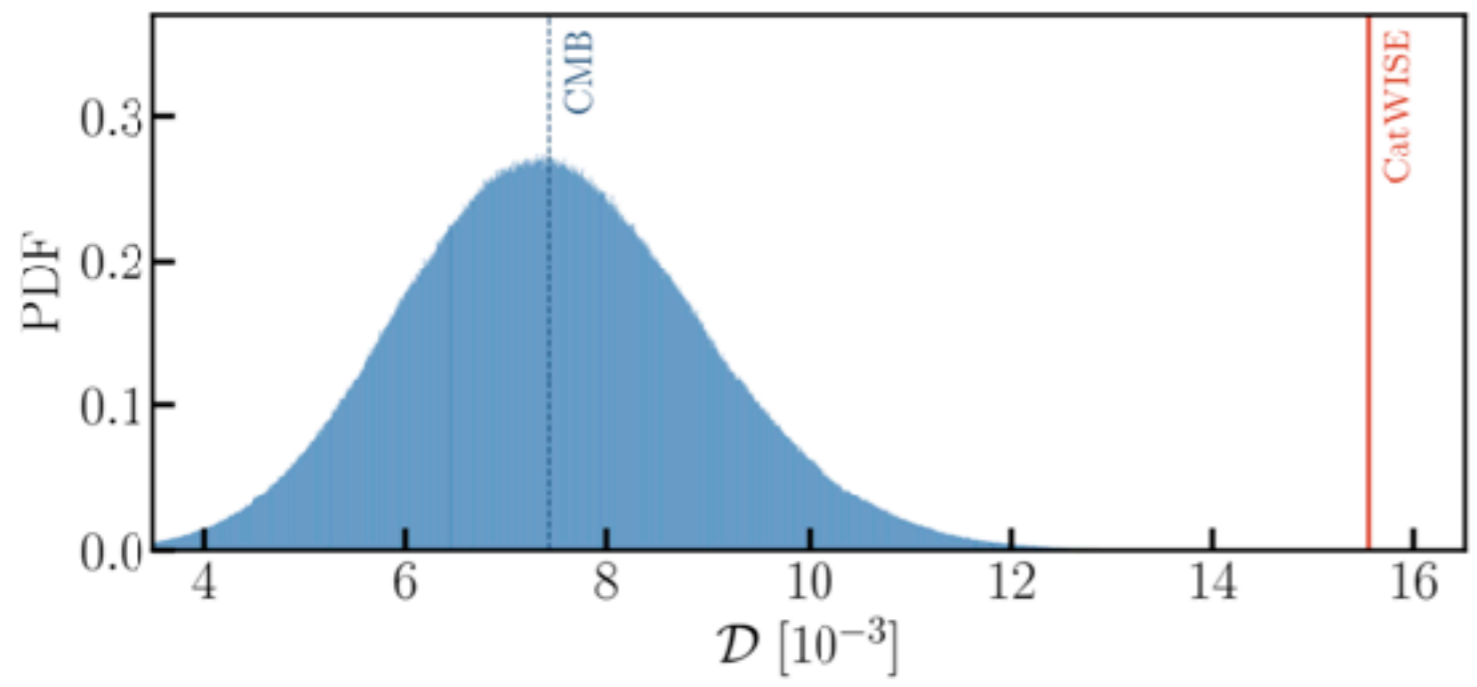
The null Hypothesis:

An observer moving with a velocity of 369.82 km/s (CMB expectation) can see a dipole twice that of CMB" !

Rejected : p value of 5×10^{-7}

Statistical significance

4.9 sigma



The cosmological principle

All data and codes made public:

Code and data → <https://doi.org/10.5281/zenodo.4431089>

Our paper (open access):

<https://ui.adsabs.harvard.edu/abs/2021ApJ...908L..51S/abstract>

Improving Physical Cosmology: An Empiricist's Assessment

P. J. E. Peebles

Princeton University

June 4, 2021

This broad variety of ways to look at the universe adds up to a compelling empirical case for the Λ CDM theory as an impressively good approximation to reality.

Maybe more tensions will be found as the constraints improve. If so then I expect the case for Λ CDM as a useful approximation will remain compelling, and there will be more clues to a still better theory.